

# **Subaru Telescope, LGSAO, & TMT**

**Masanori Iye (NAOJ)**

- 1) Subaru Telescope**
- 2) Laser Guide Star Adaptive Optics**
- 3) 30m telescope TMT**

# Subaru Telescope, its Science Achievements

Enclosure  $\Phi 40\text{m}$ , H42m, 2500ton

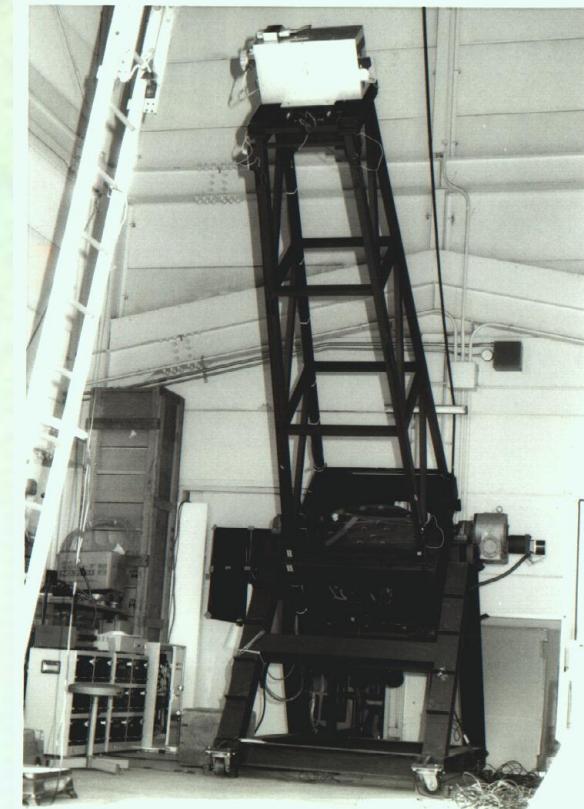
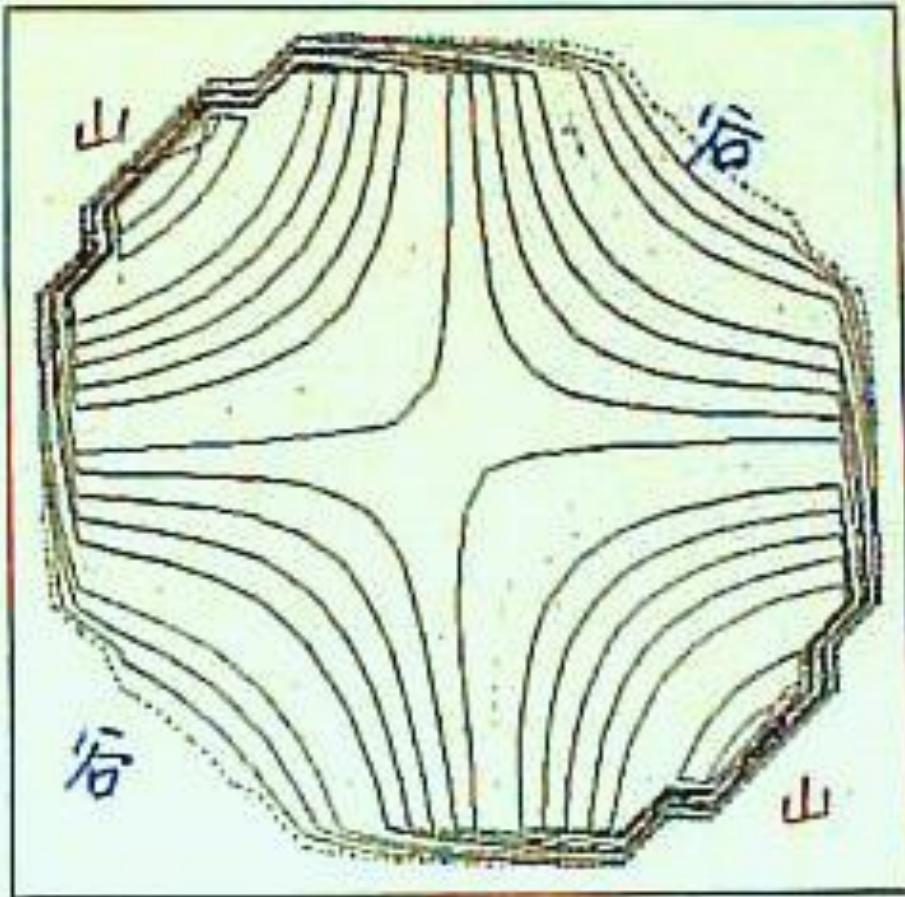


400M\$, 16yr for prep+construction, completion in 1999

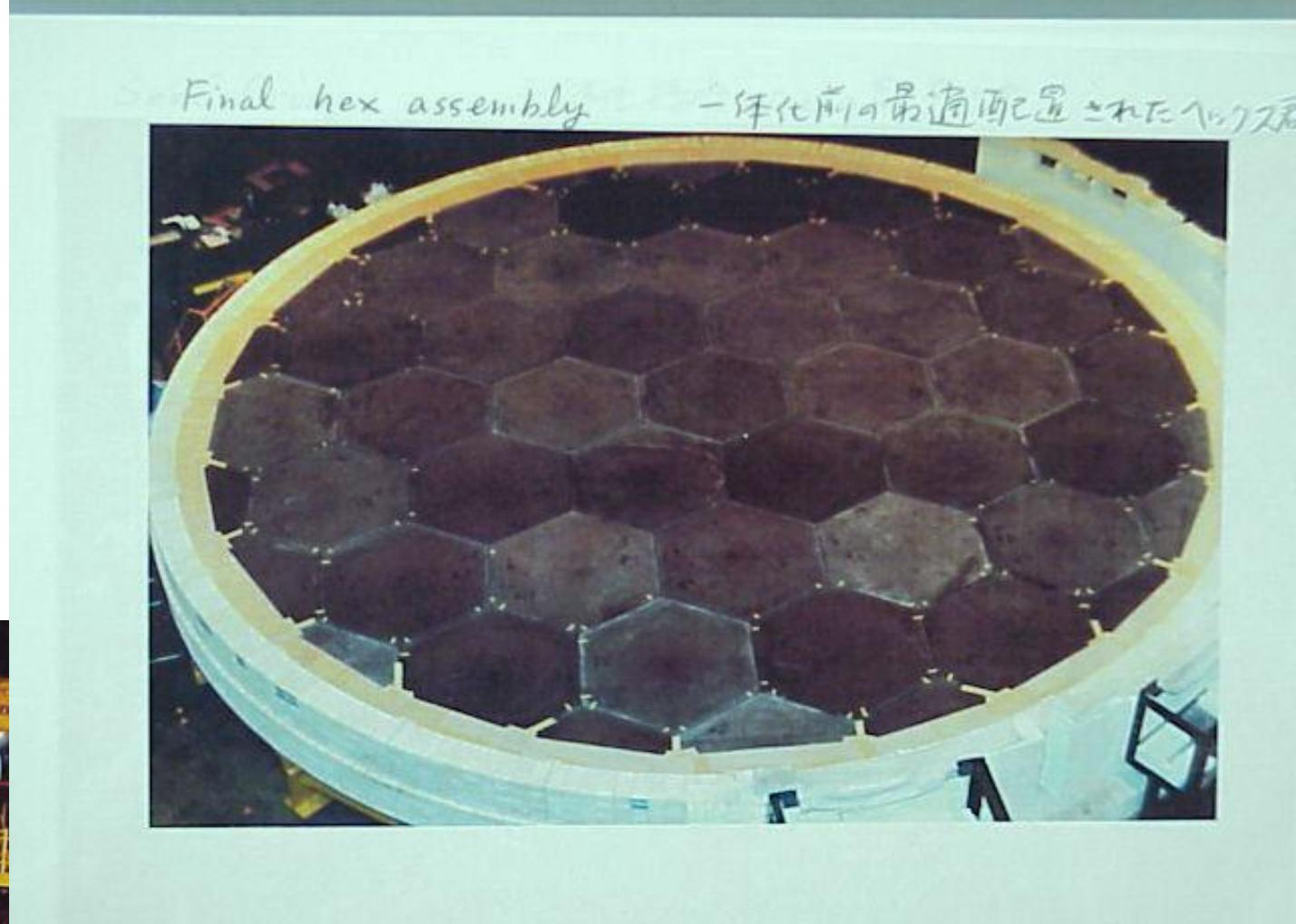
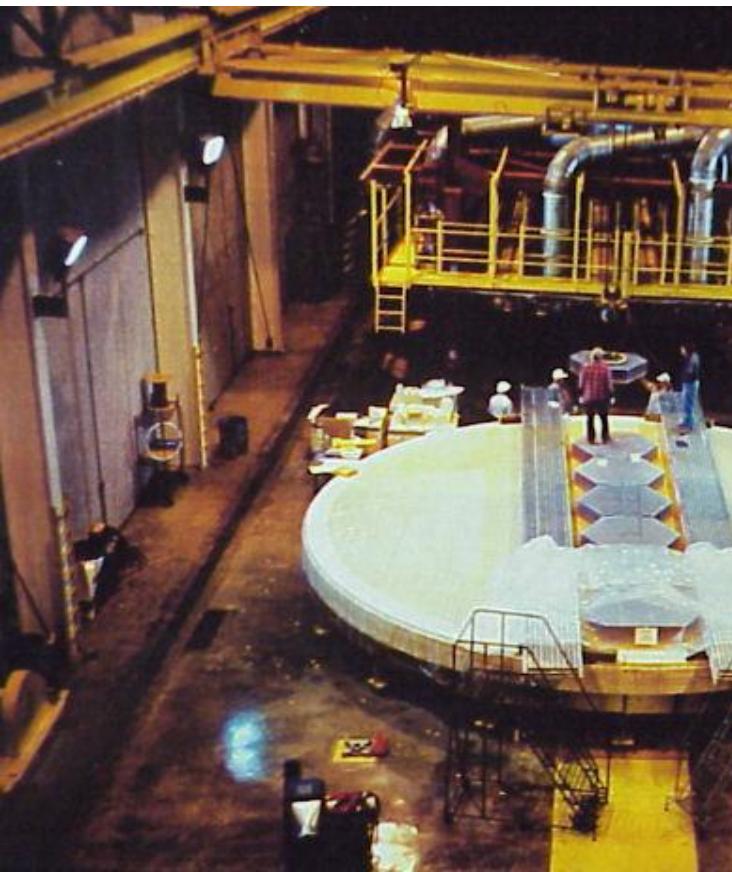
1989年10月14日

# Active mirror control experiment on a 62cm mirror supported by 12 actuators

```
le = coeffD.088
90T23CT+12U05 62cmB22=3
ane 88 - 15
it in nano-meter
n Z= -653 Max Z= 659
0= 0 A40= -9 A60= -2
1= 0 A31= -14 A51= 2
1= 0 B31= 14 B51= -7
2= 27 A42= -8 A62= -9
2= -779 B42= 49 B62= -11
3= 9 A53= -8
3= 6 B53= 3
4= -8 A64= 8
4= 18 B64= -8
5=
5=
5= 14
5= -13
5= -2 EL= 0
= -117 02= -107 03= -129
= 116 05= 262 06= 229
= 45 08= -377 09= -152
= 124 11= -160 12= 278
ntour pitch = 66
ntour lines = 20
```

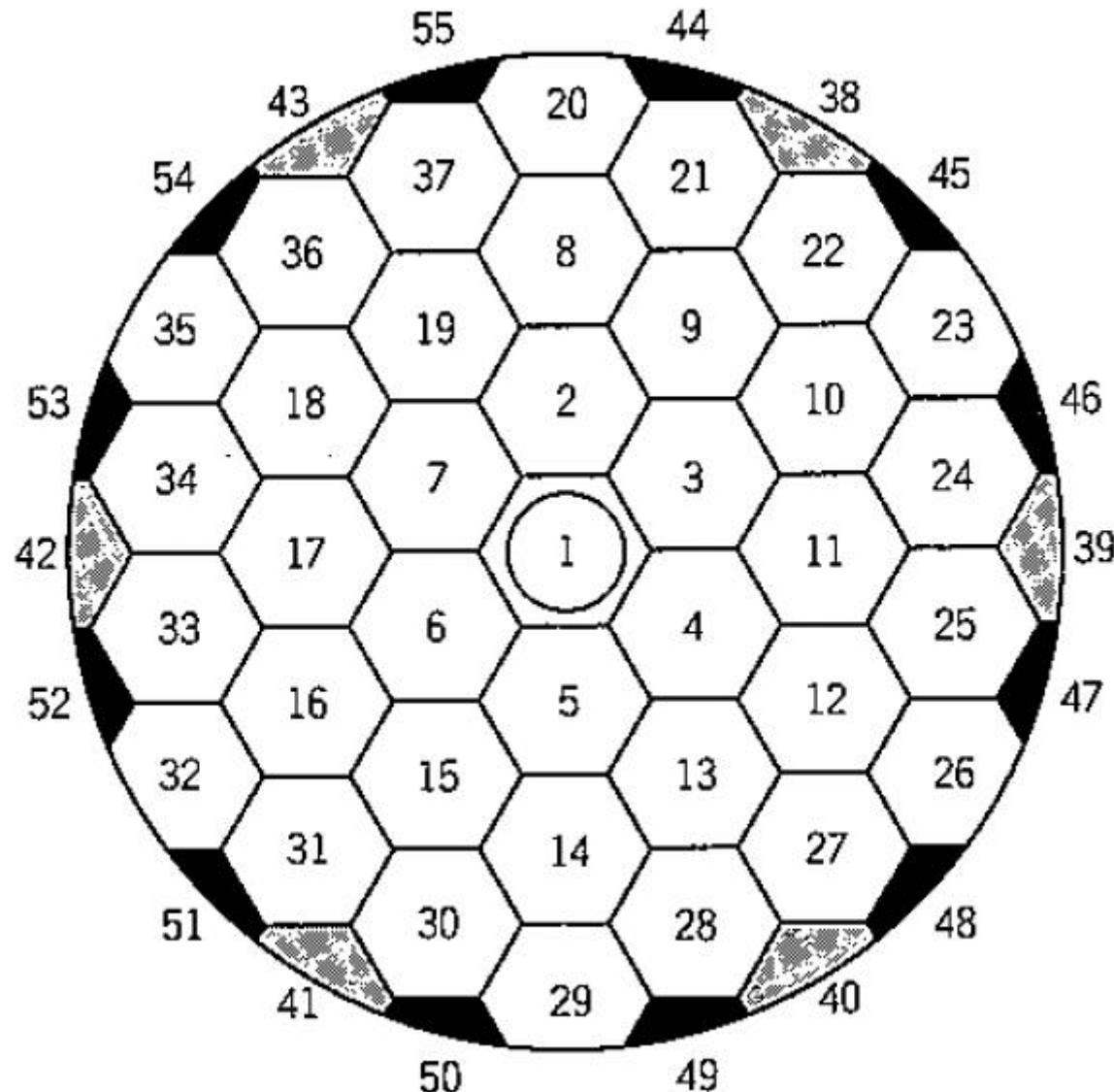


# Corning ULE



**First mathematical problem :**  
**Find the best configuration to distribute 55 segments with different CTE to form a monolithic mirror with minimum thermal deformation**

**~ 55! combinations to examine.**



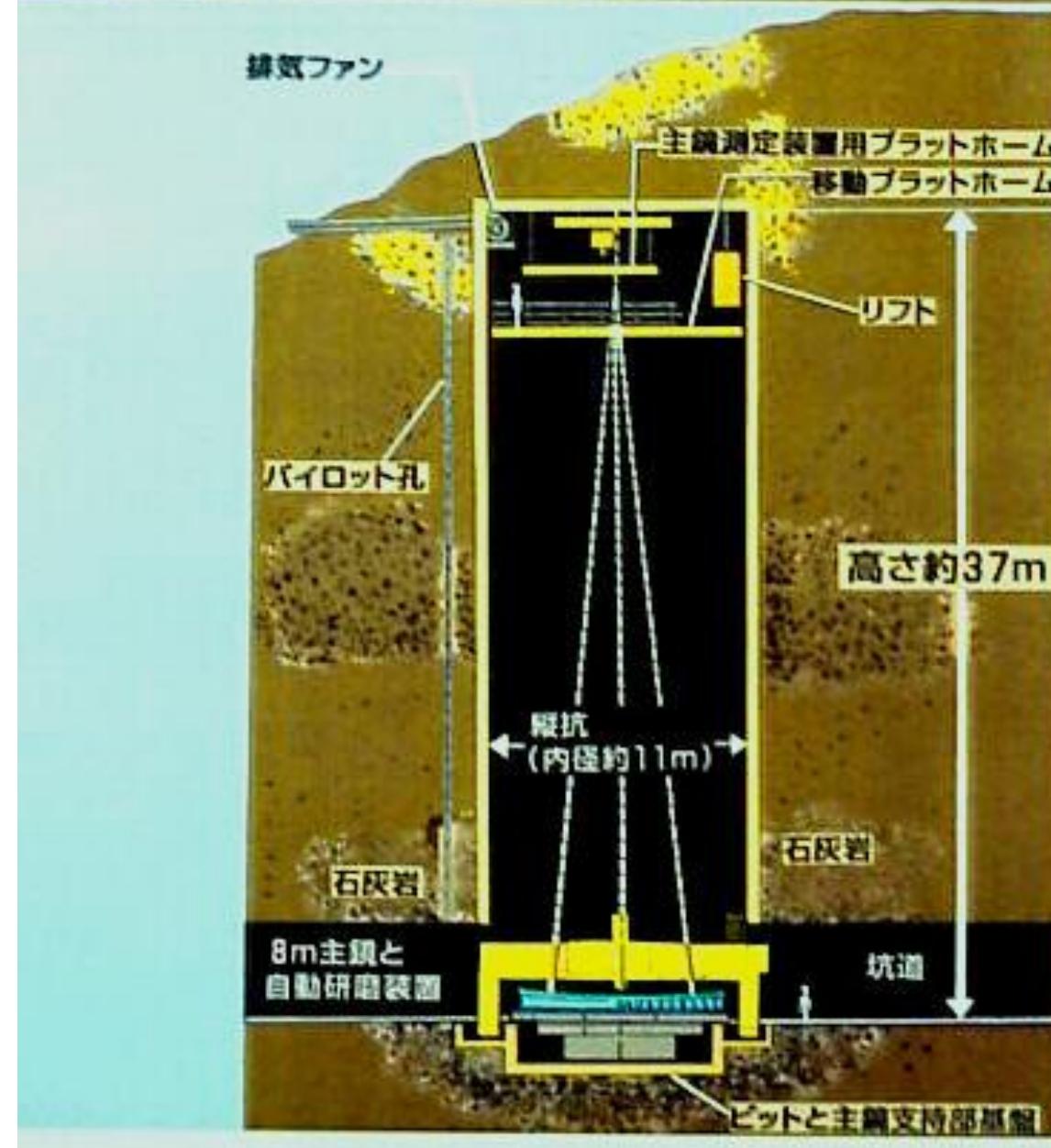
# Glass transportation to Pittsburgh



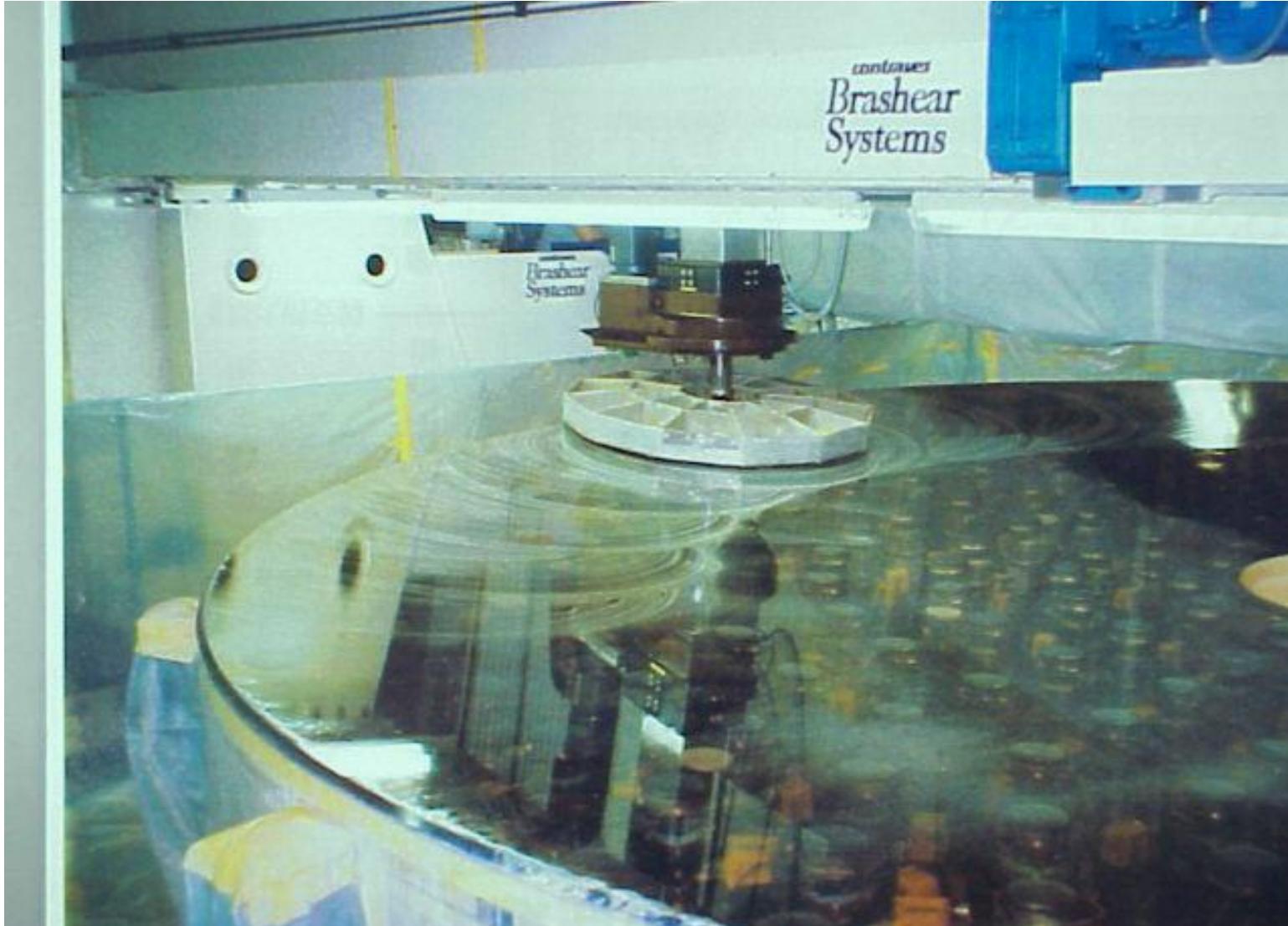
# すばる望遠鏡主鏡研磨設備

Contraves  
underground  
optics facility  
needed 30m  
span

no wind shake  
no diurnal tilt  
no traffic  
vibration



# Four years of patient polishing

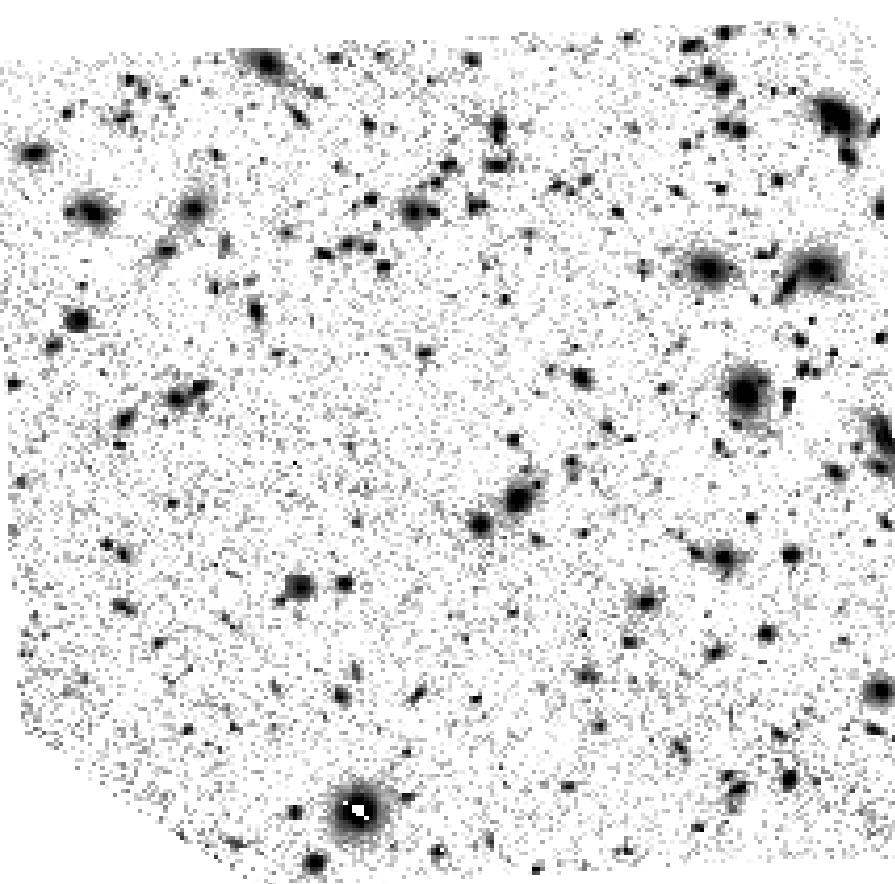




Aug. 1998  
13nm rms figure error achieved

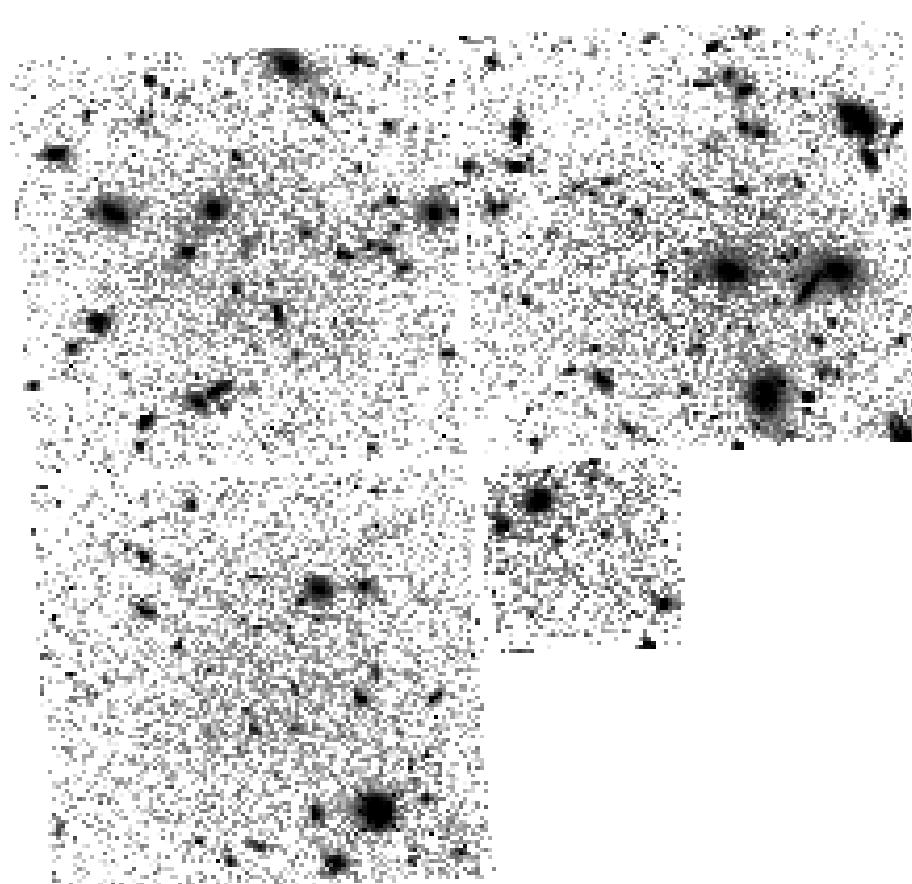
# First light sensitivity comparison

**Subaru R**  
**3600 s**



すばる 60分, 28等級

**HST F702W**  
**4200 s**

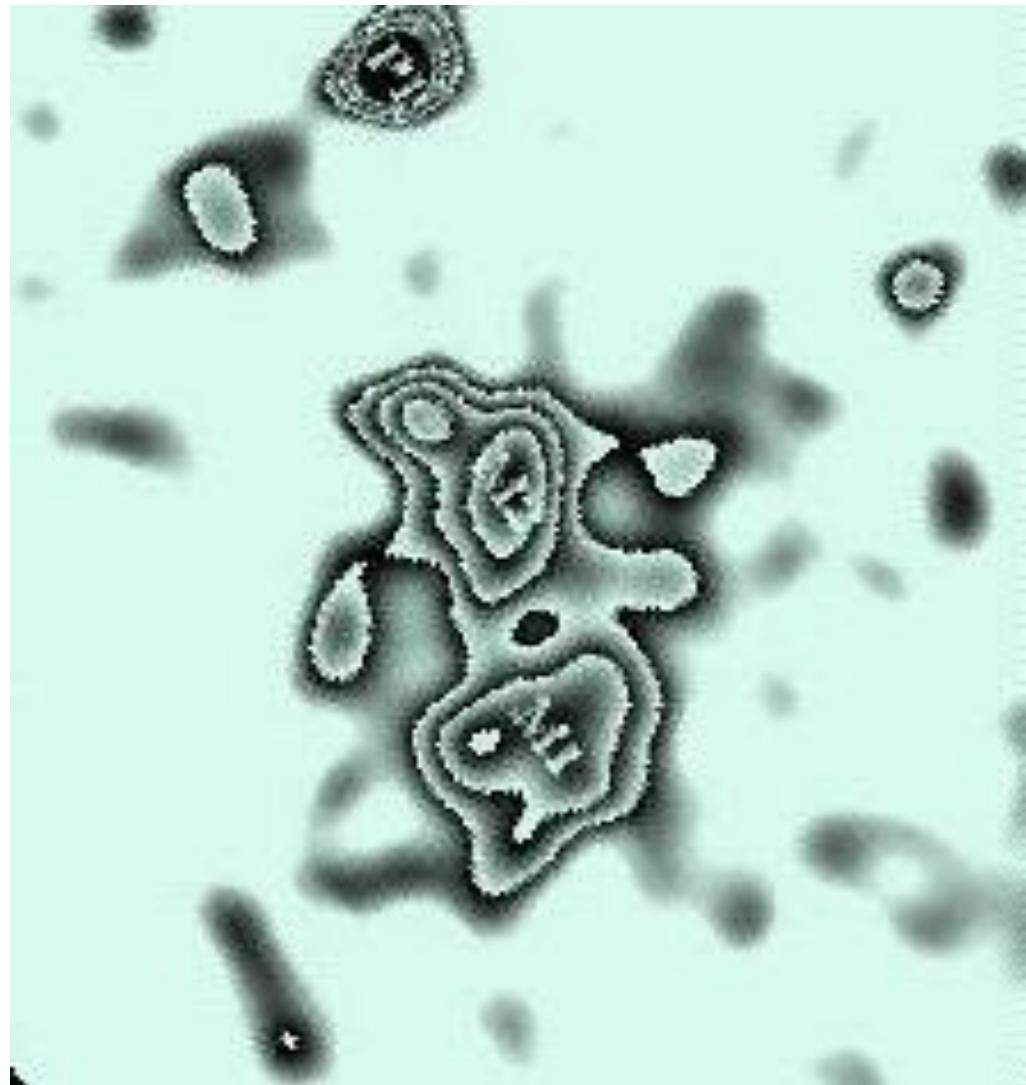
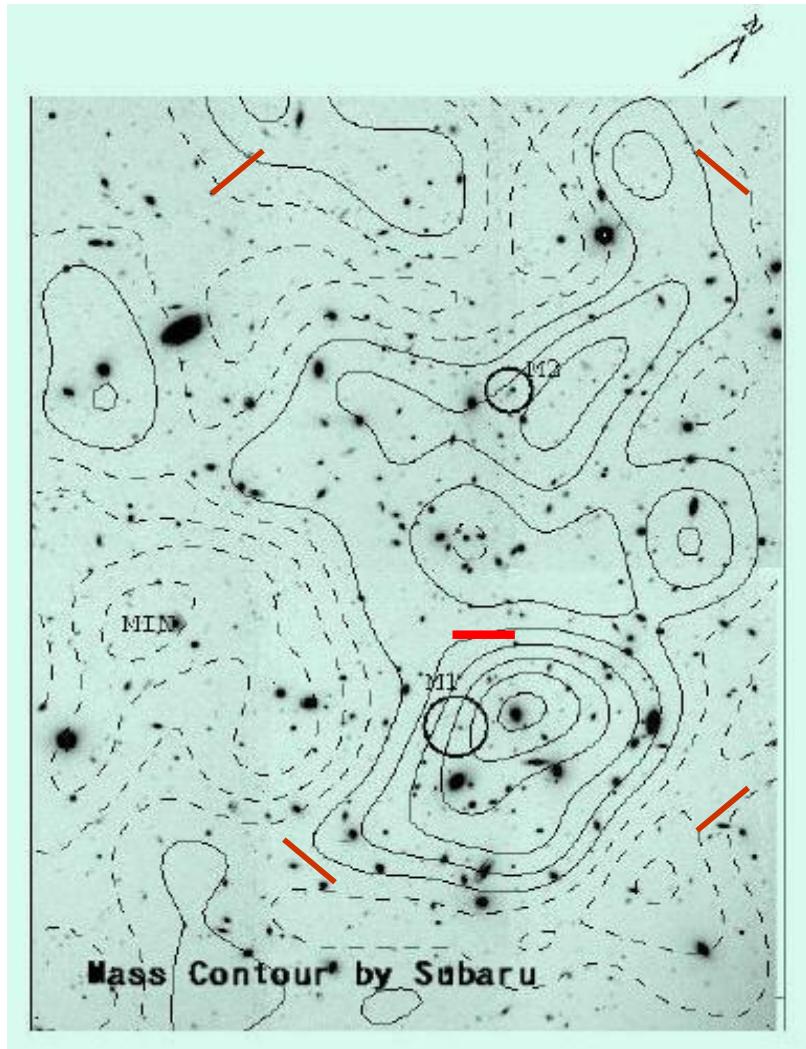


勝負は引き分け

ハッブル 70分 28等級

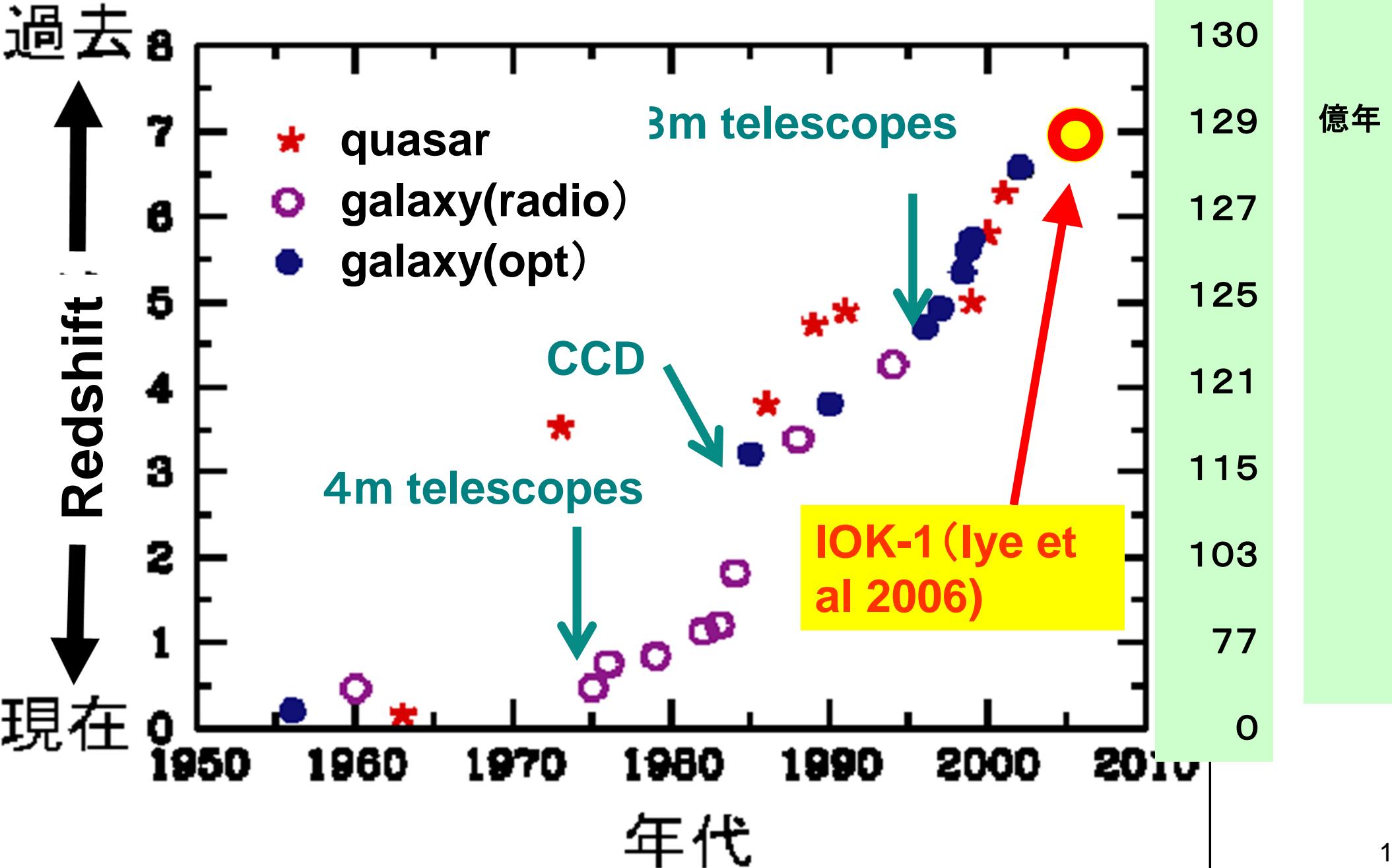
# Weak lensing by a cluster of galaxies

Miyazaki in Iye et al.(2000)



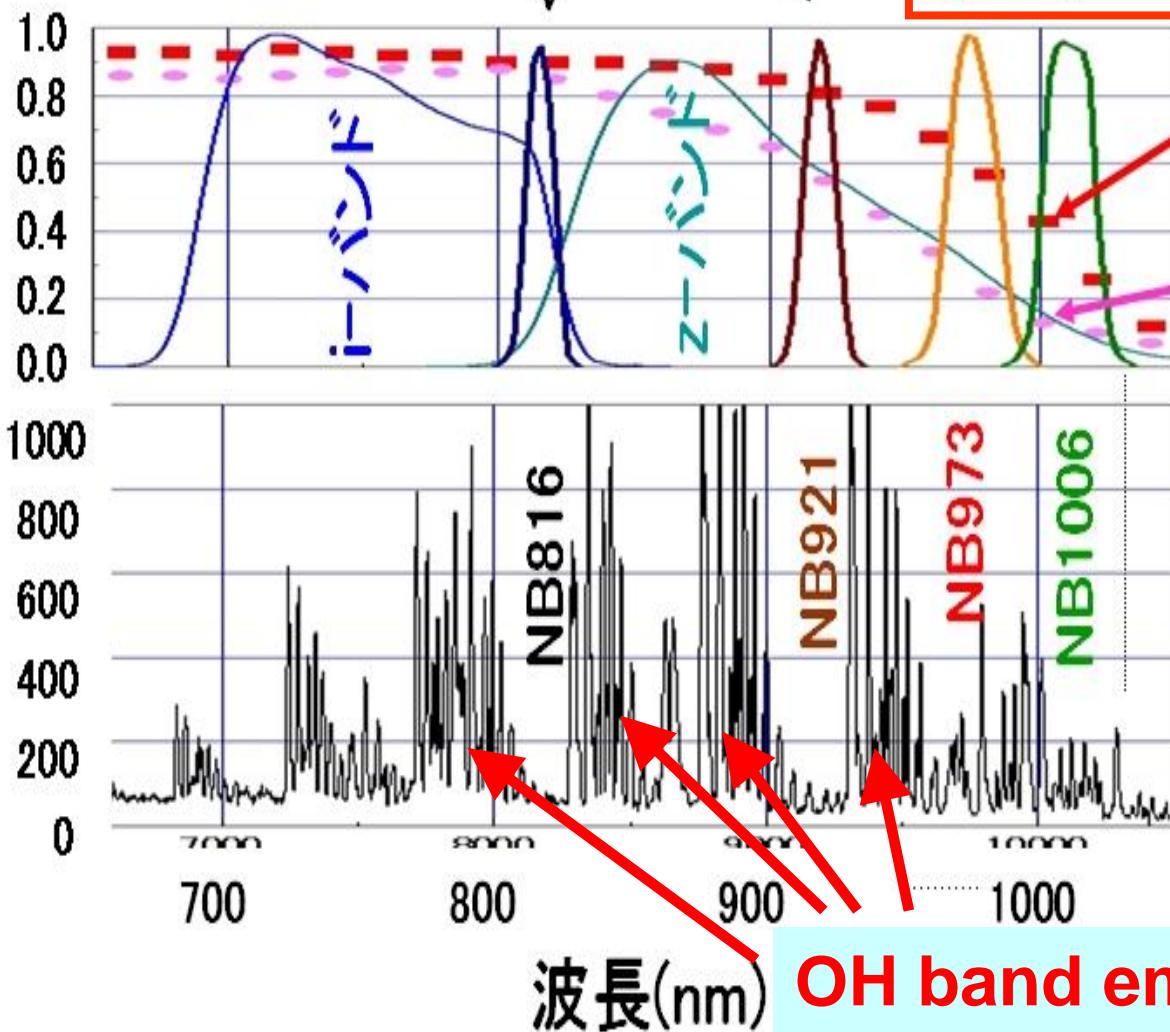
- Looking far is looking back in time

# Most distant galaxies in history

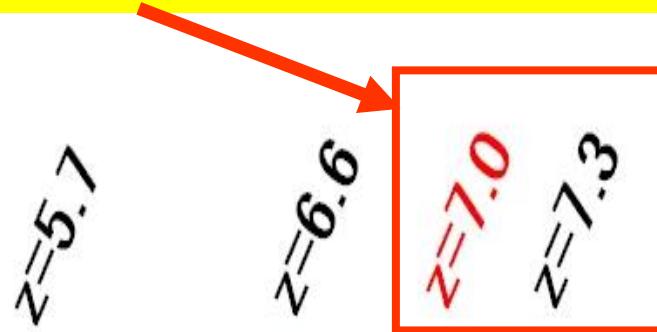


# NB filters for SuprimeCam

夜光強度



夜空の暗い窓



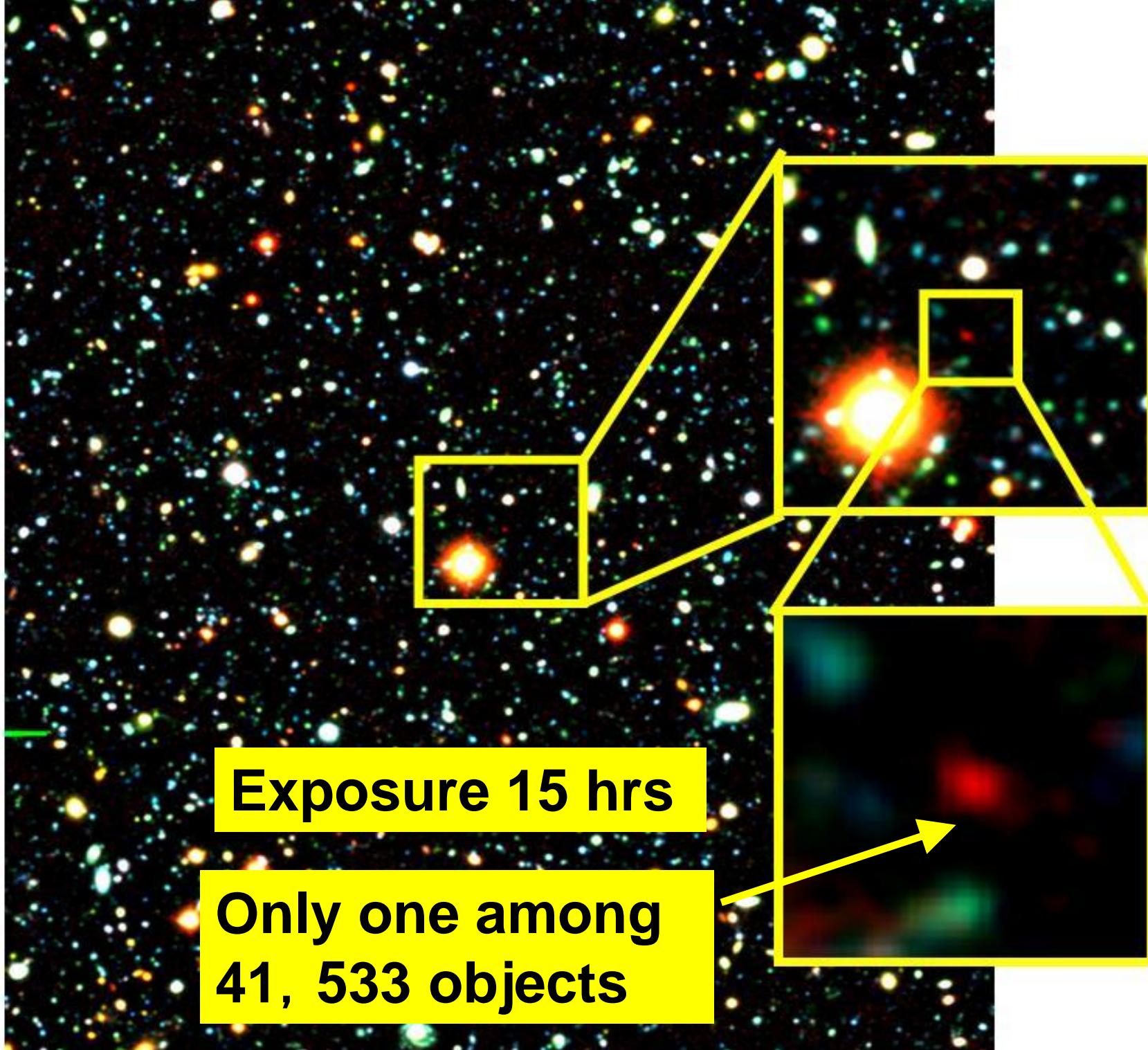
New  
CCD

効率4倍

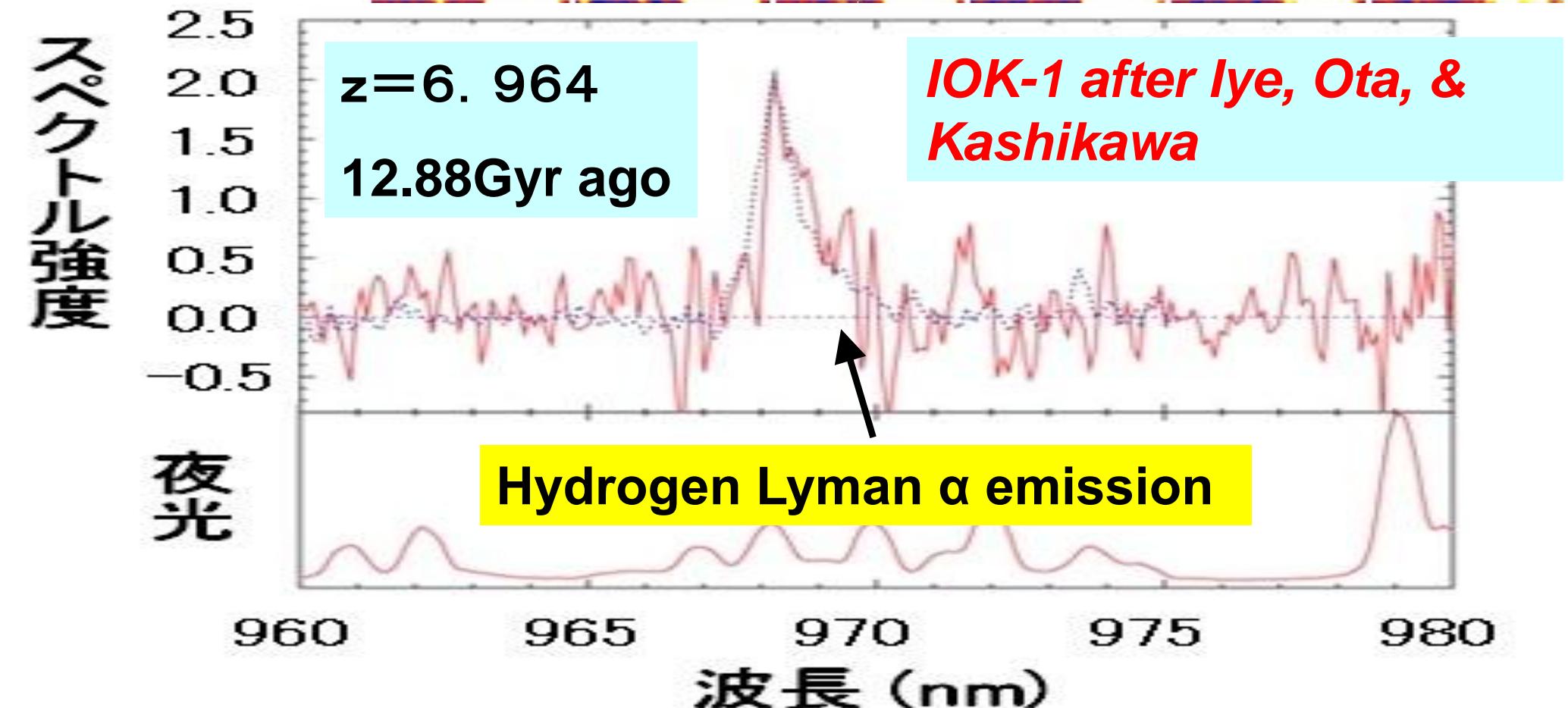
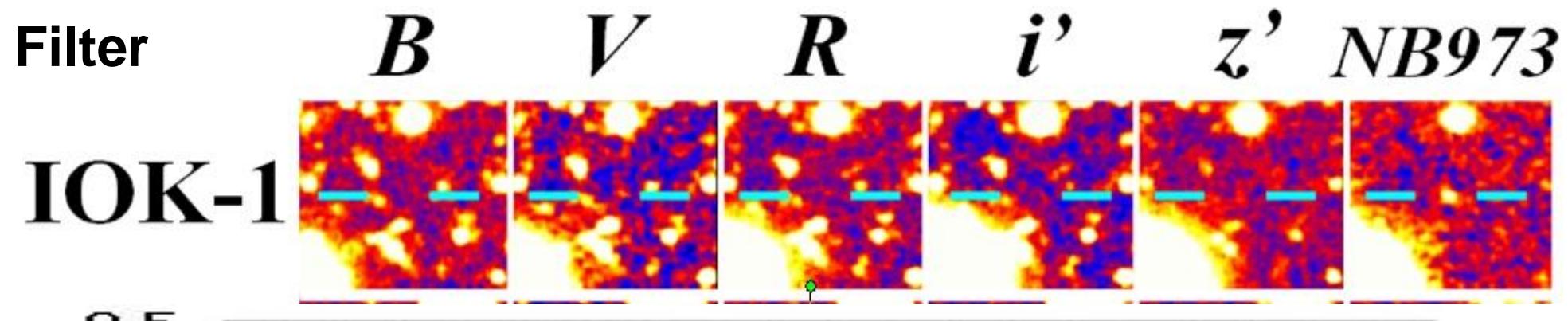
Old CCD

NB973





# The most distant galaxy at 12.88Gyr IOK-1



# High-z galaxy list

10/21で1498日目

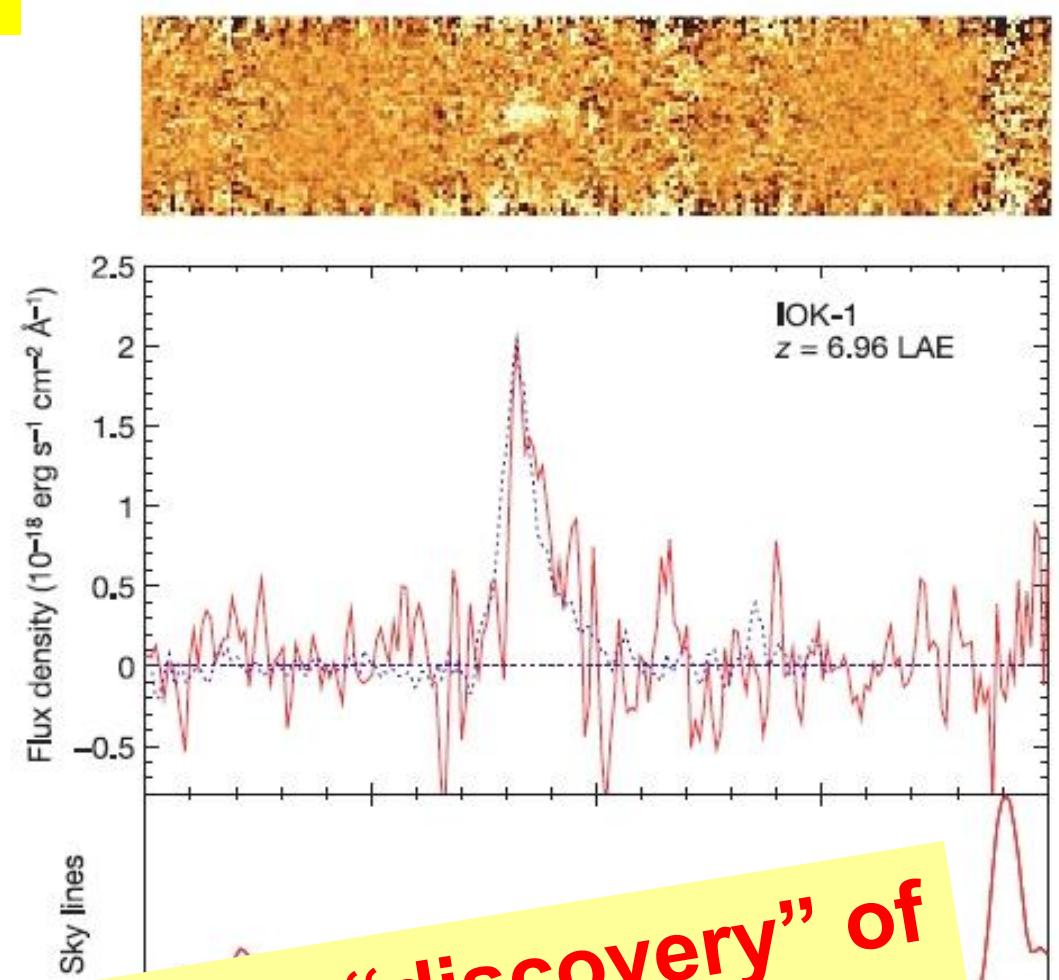
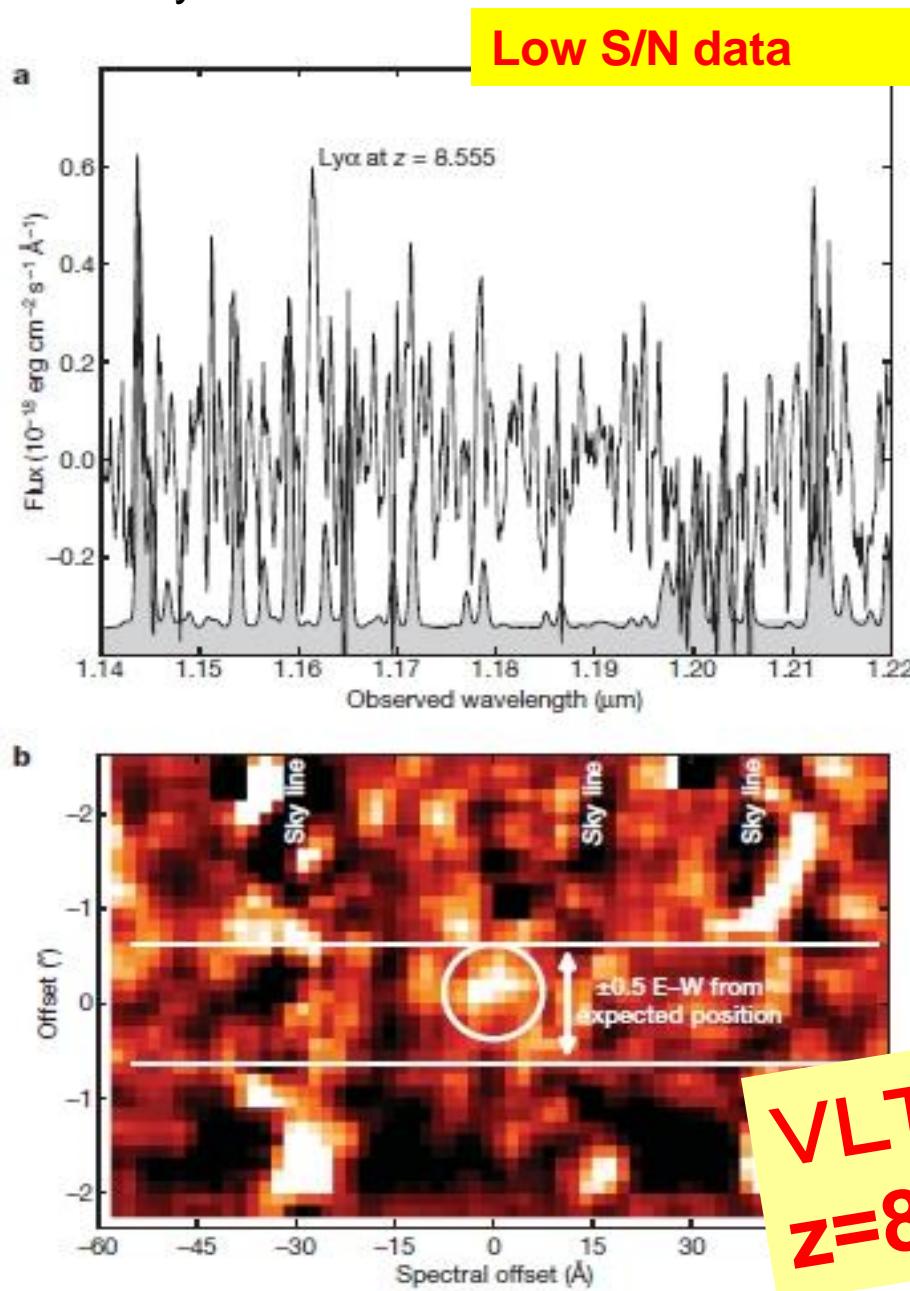
Table 1: 赤方偏移が確定した最遠銀河ベスト 12(2009 年 10 月 1 日時点).

順位	銀河名	座標	赤方偏移	億年 #	論文	出版日
1	IOK-1	J132359.8+272456	6.964	128.8	家ほか	2006 年 9 月 14 日
2	SDF ID1004	J132522.3+273520	6.597	128.2	谷口ほか	2005 年 2 月 25 日
3	SDF ID1018	J132520.4+273459	6.596	128.2	柏川ほか	2006 年 4 月 25 日
4	SXDF Himiko		6.595	128.2	大内ほか	2008 年 7 月 25 日
5	SDF ID1030	J132357.1+272448	6.589	128.2	柏川ほか	2006 年 4 月 25 日
6	SDF ID91163	J132343.4+272954.5	6.587	128.2	柏川ほか	2009 年 2 月
6	SDF ID91988	J132342.2+272644.5	6.587	128.2	柏川ほか	2009 年 2 月
6	SDF ID71101	J132450.7+272159.7	6.587	128.2	柏川ほか	2009 年 2 月
9	SDF ID1	<b>Tens of LAE at z=6.6 but only one LAE at z= 7.0.</b>				
10	SDF ID1					年 2 月 25 日
11	SDF ID1					年 2 月 25 日
12	SDF ID157057	J132419.3+274124.8	6.568	128.2	柏川ほか	2009 年 2 月
13	HCM-6A	J023954.7-013332	6.560	128.2	Hu ほか	2002 年 4 月 1 日

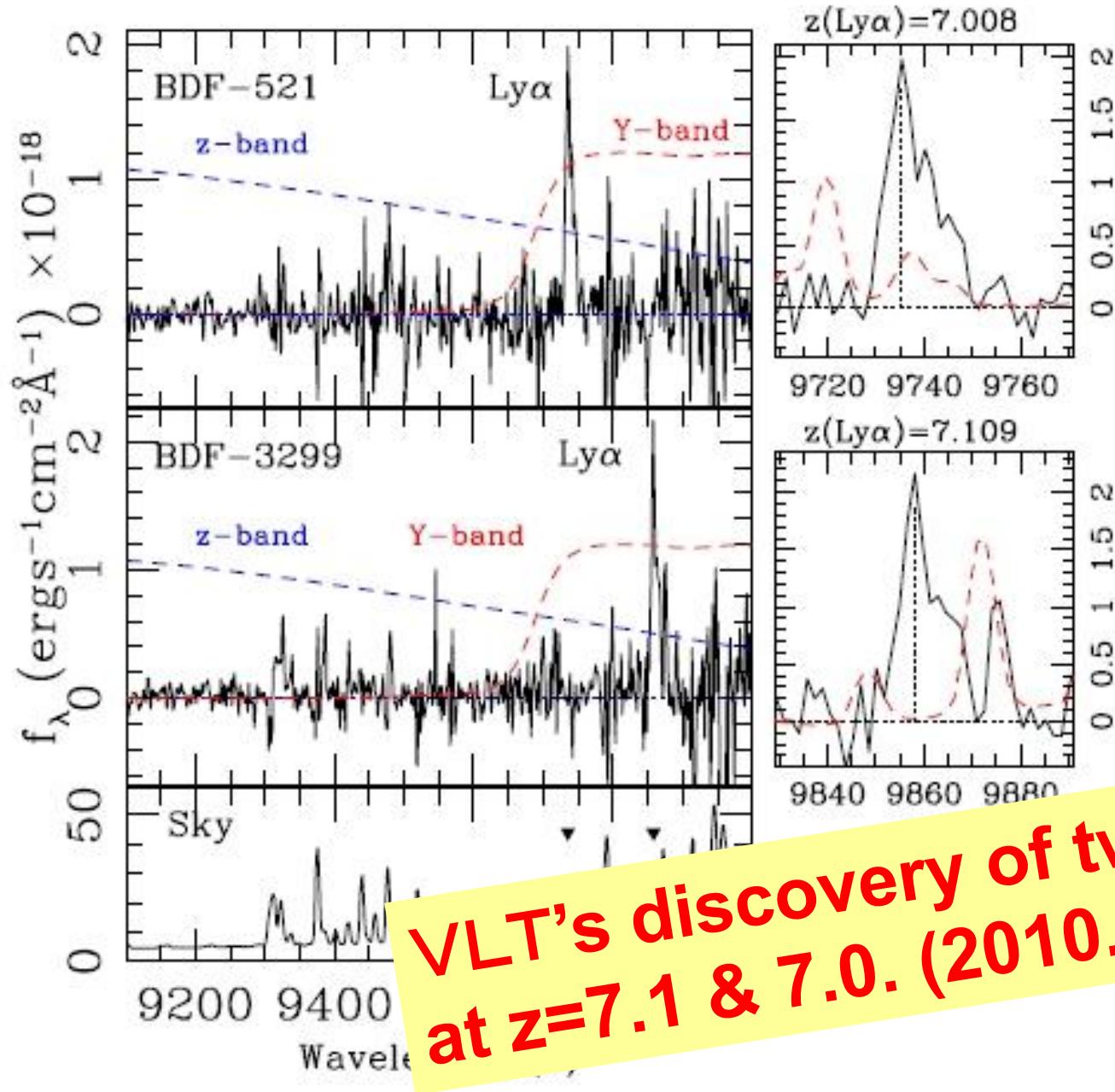
# 年齢は宇宙年齢が 136.6 億年となるモデルに基づいて算出.

Lehnert et al.(2010)  $z=8.55?$   
UDFy-38135539

Iye et al.(2006)  $z=6.96$   
IOK-1



VLT reports “discovery” of  
 $z=8.6$  galaxy? (2010.10.21)



VLT's discovery of two LAEs  
at  $z=7.1$  &  $7.0$ . (2010.11.30)

# High redshift galaxies

Table 1: Top 10 most distant galaxies with measured redshift(Dec 9 2010)

Rank	ID	Coordinates	Redshift	0.1Gyr	Paper	Date
?	UDFy-38135539	J033237.79-274600.0	8.555?	128.29	Lehnert et al.	Oct.21, 2010
1	BDF-3299	J222846.0-350959	7.109	128.29	Vanzella et al.	Nov.24 2010
2	BDF-521	J222746.6-350707	7.008	128.29	Vanzella et al.	Nov.24 2010
3	IOK-1	J132359.8+272456	6.964	128.85	Iye et al.	Sep. 14, 2006
4	SXDF N-71598	J021844.7-043637	6.621	128.29	Ouchi et al.	in preparation
5	SXDF W-30717	J021654.6-045557	6.617	128.28	Ouchi et al.	in preparation
6	SDF ID1004	J132522.3+273520	6.597	128.25	Taniguchi et al.	Feb. 25, 2005
7	SDF ID1018	J132520.4+273459	6.596	128.25	Kashikawa et al.	Apr. 25, 2006
8	SXDF Himiko	J021757.6-050844	6.595	128.25	Ouchi et al.	Jul. 25, 2008
9	SXDF w-31790	J021653.9-050601	6.590	128.24	Ouchi et al.	in preparation
10	SDF ID1030	J132357.1+272448	6.589	128.24	Kashikawa et al.	Apr. 25, 2006
11	SXDF C-106098	J021703.5-045619	6.589	128.24	Ouchi et al.	in preparation
12	SDF ID91163	J132343.4+272954.5	6.587	128.23	Kashikawa et al.	Feb. 2009
13	SDF ID91988	J132342.2+272644.5	6.587	128.23	Kashikawa et al.	Feb. 2009
14	SDF ID71101	J132450.7+272159.7	6.587	128.23	Kashikawa et al.	Feb. 2009
15	SXDF W-36217	J021650.4-050545	6.586	128.23	Ouchi et al.	in preparation
16	SDF ID1007	J132432.5+271647	6.580	128.22	Taniguchi et al.	Feb. 25, 2005
17	SDF ID1008	J132518.8+273043	6.578	128.22	Taniguchi et al.	Feb. 25, 2005
18	SDF ID1001	J132418.3+271455	6.578	128.22	Kodaira et al.	Apr. 25, 2003
19	SXDF C-22057	J021820.7-051110	6.575	128.21	Ouchi et al.	in preparation
20	SXDF W-25755	J021658.3-045557	6.573	128.21	Ouchi et al.	in preparation
21	SDF ID157057	J132419.3+274124.8	6.568	128.20	Kashikawa et al.	Feb. 2009
22	SXDF C-34609	J021819.4-050901	6.563	128.19	Ouchi et al.	in preparation
23*	HCM-6A	J023954.7-013332	6.560	128.19	Hu et al.	Apr. 1, 2002

# Many brilliant Japanese Supernovae burst in Early Universe Studies

- N.Kashikawa
- M.Ouchi
- K.Shimasaku
- K.Ota
- N.Yoshida
- K.Omukai
- A.Inoue
- M.Kobayashi

**Japan Strategy**

**Subaru SupremeCam**

V

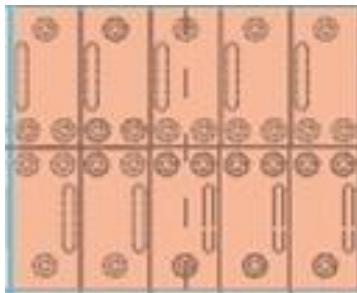
**HyperSuprimeCam**

V

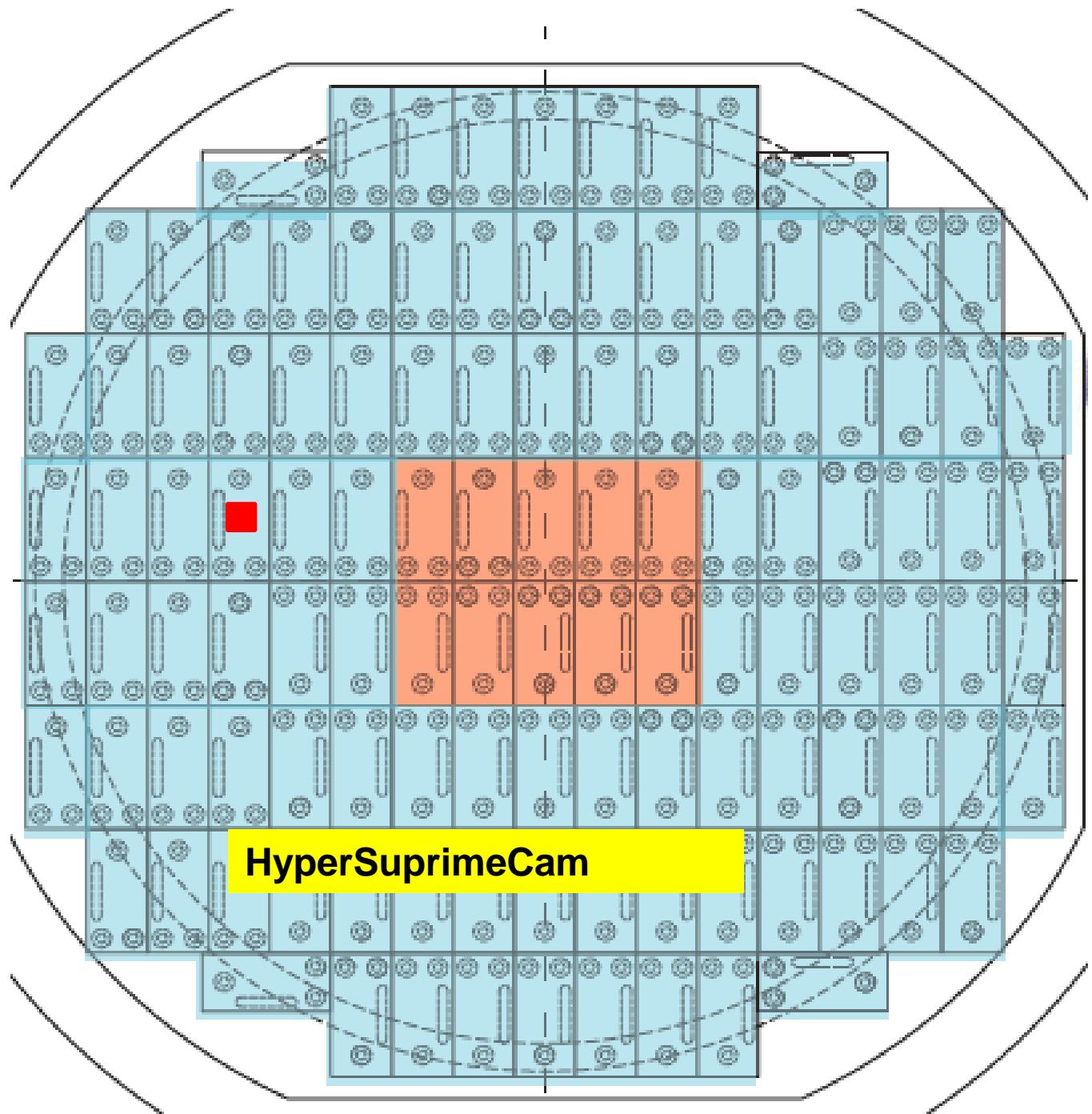
**TMT**

# 視野比べ

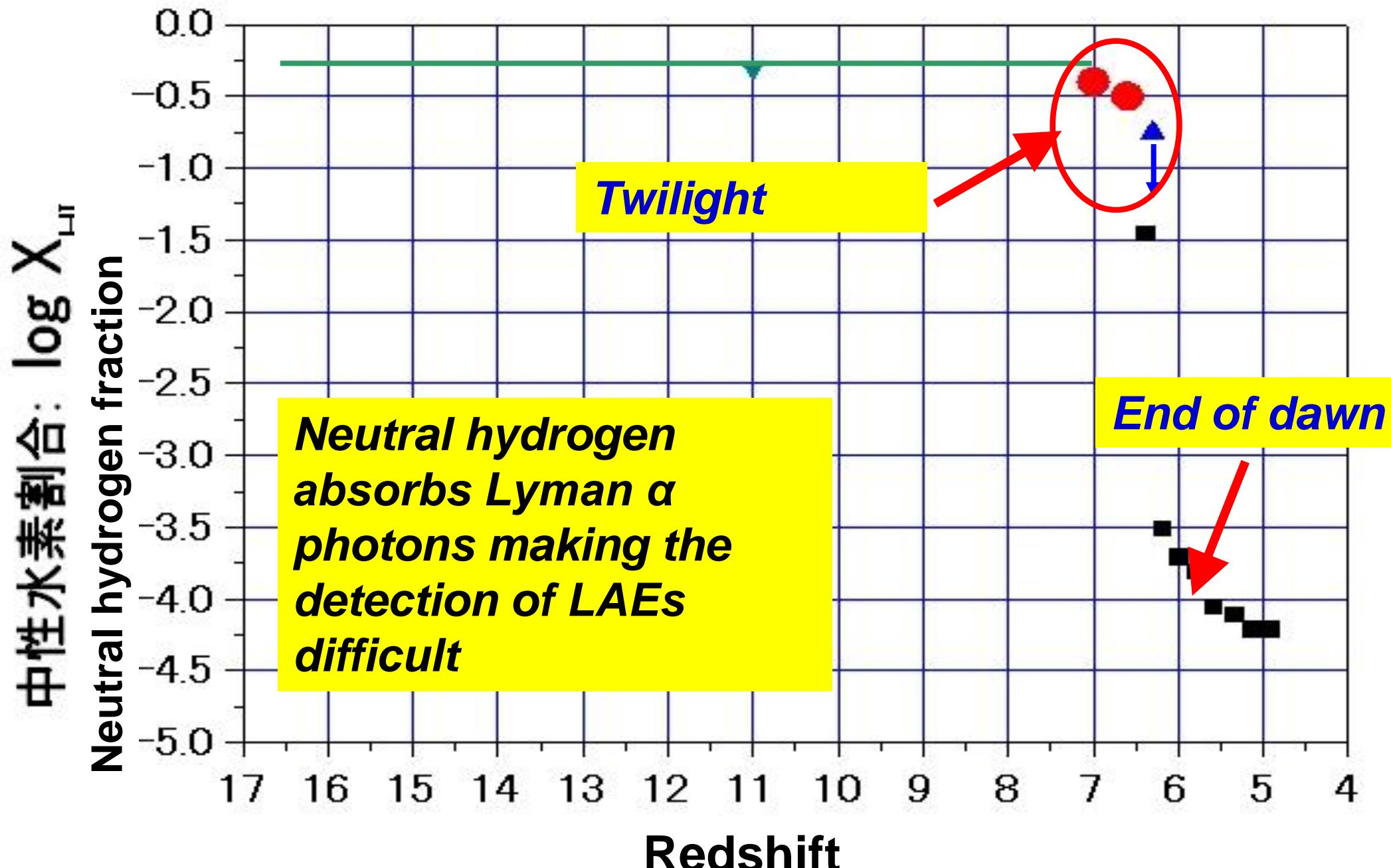
■ HST WFC3



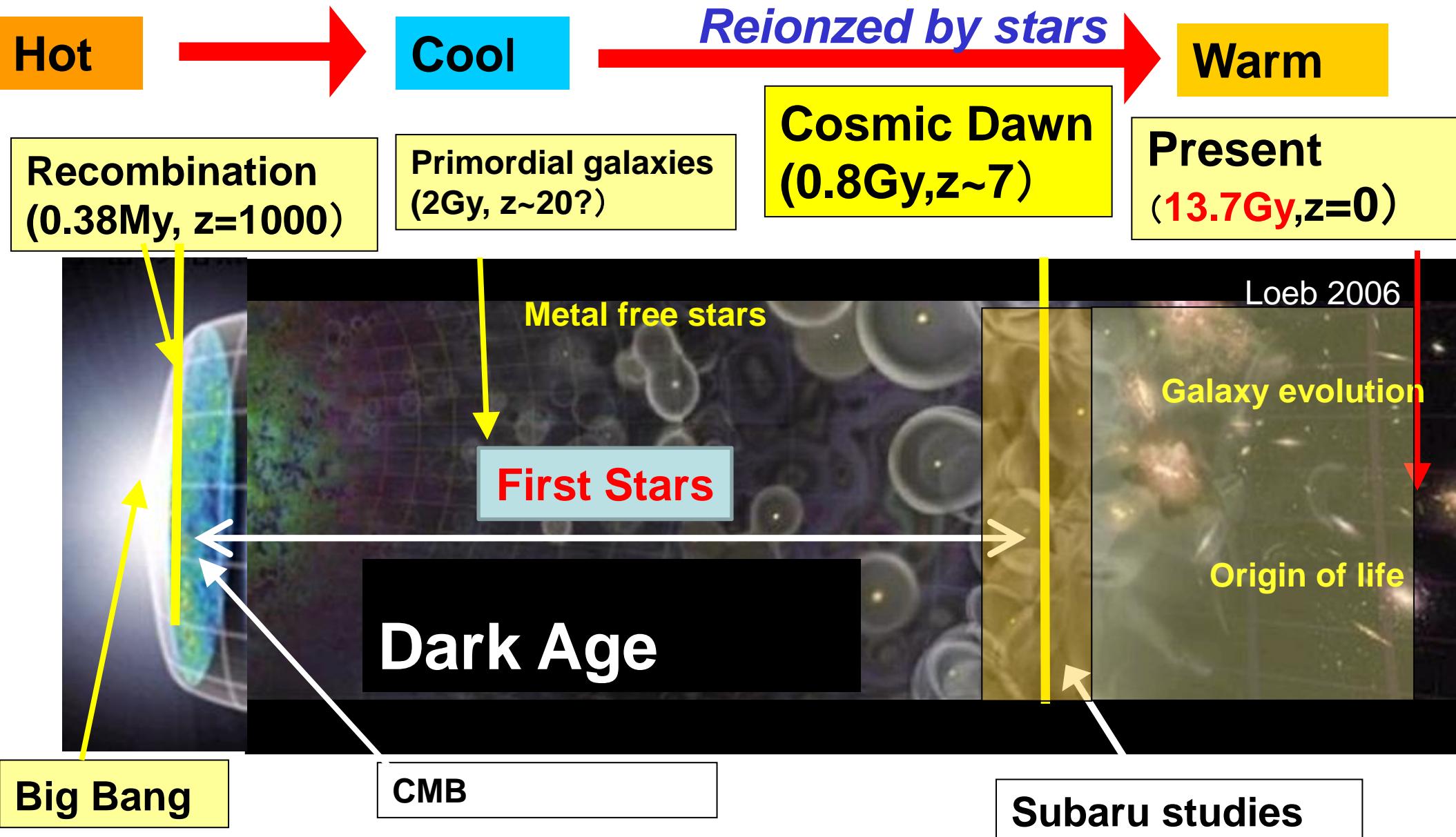
SuprimeCam



# Cosmic Reionization History



# Looking at Cosmic Dawn



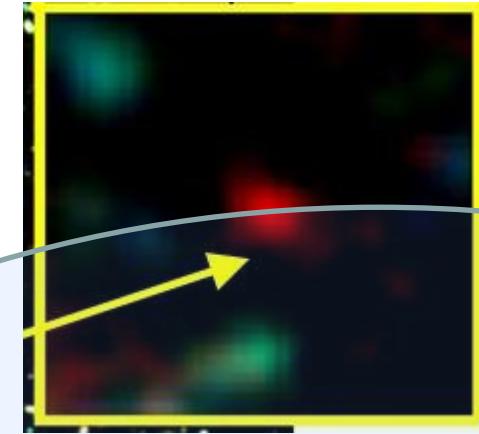
# Laser guide star

# Adaptive Optics



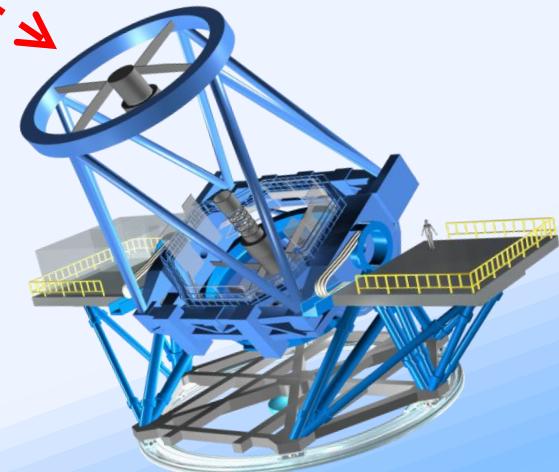
光子  
(photon)

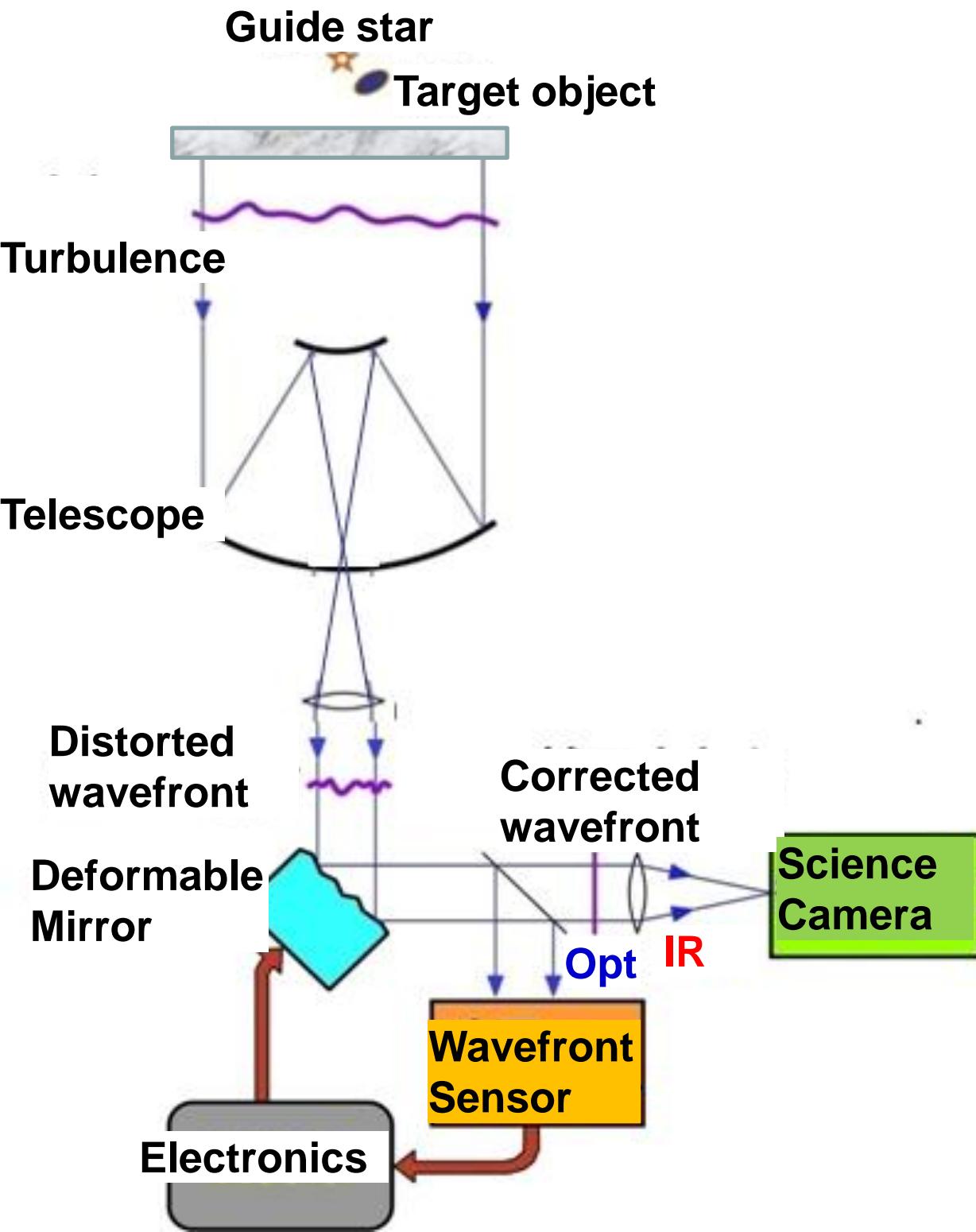
12.9 Gyr  
=40, 0000, 0000, 0000, 0000sec



Last 0. 001sec

The last milli second of photon's long journey  
blurs its beauty



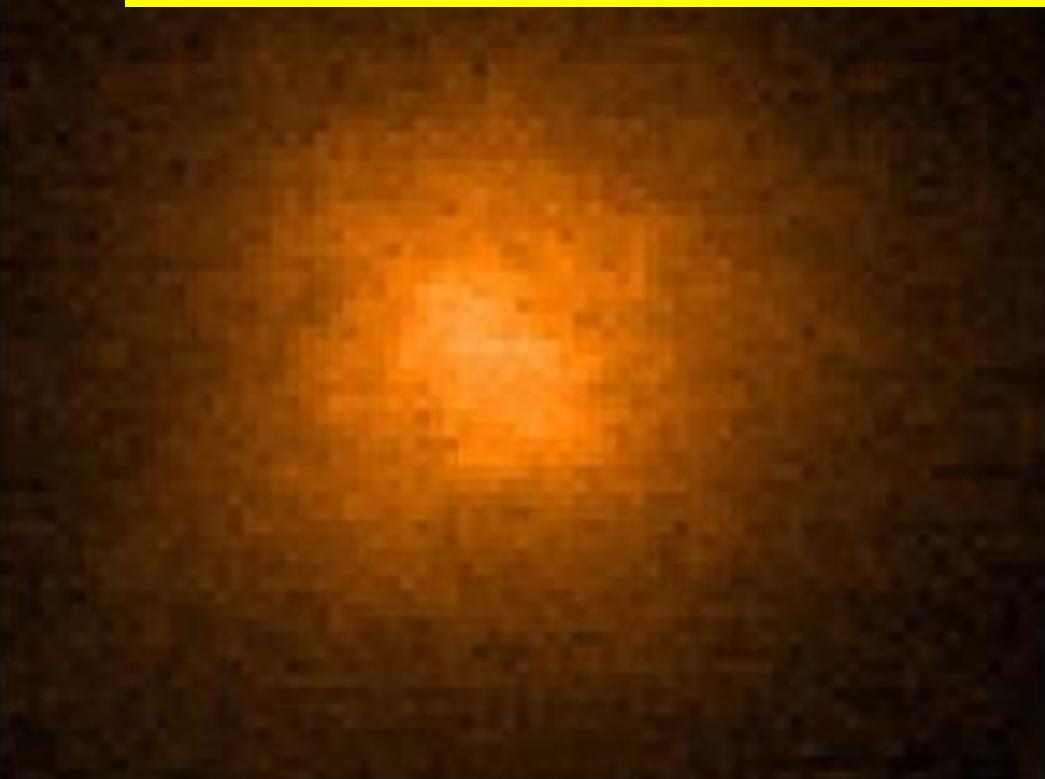


## Adaptive Optics

Wavefront sensor  
Deformable mirror  
Real time control

# First light of AO188(2006/10/9)

**AO188 improved the PSF of 400M\$ Subaru Telescope at an additional 7M\$ investment.**



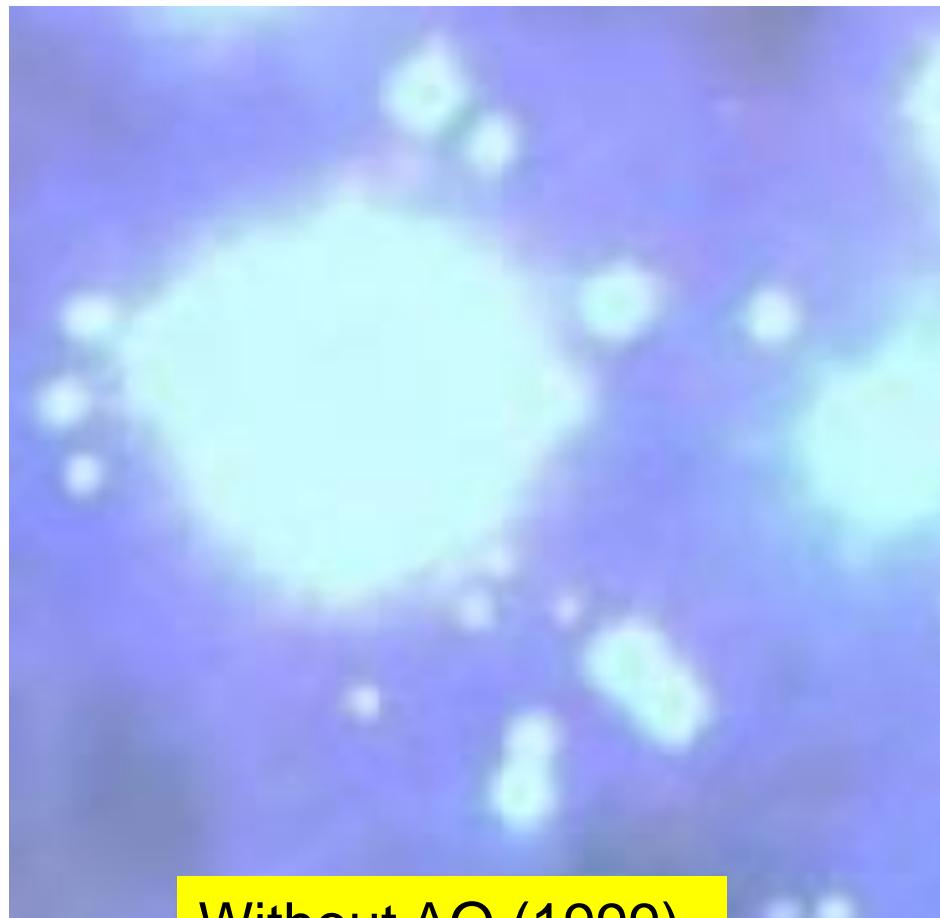
**AO off (0.6arcsec)**

**2.2 $\mu$ m**

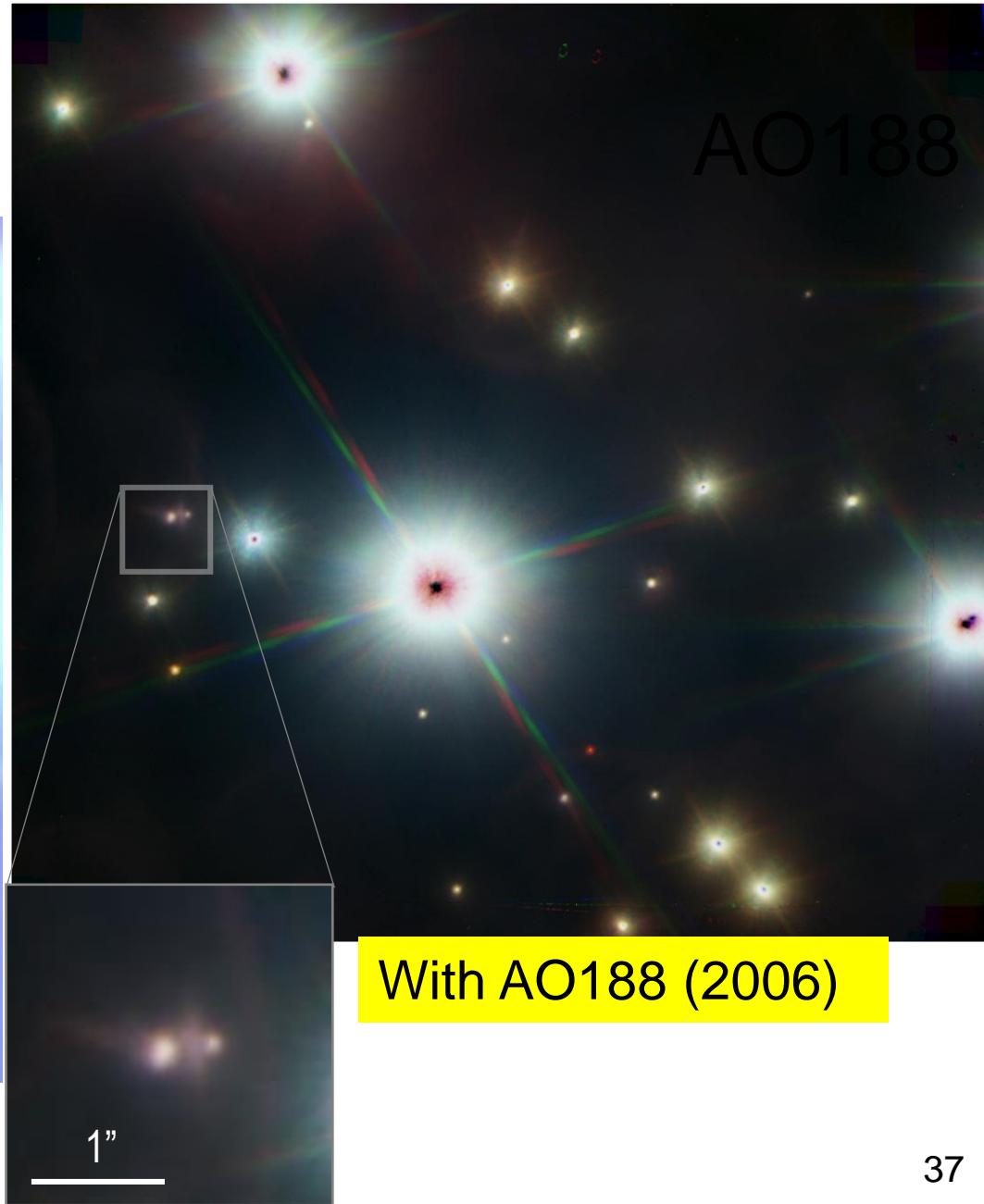


**AO on (0.063arcsec)**

# Orion Trapezium



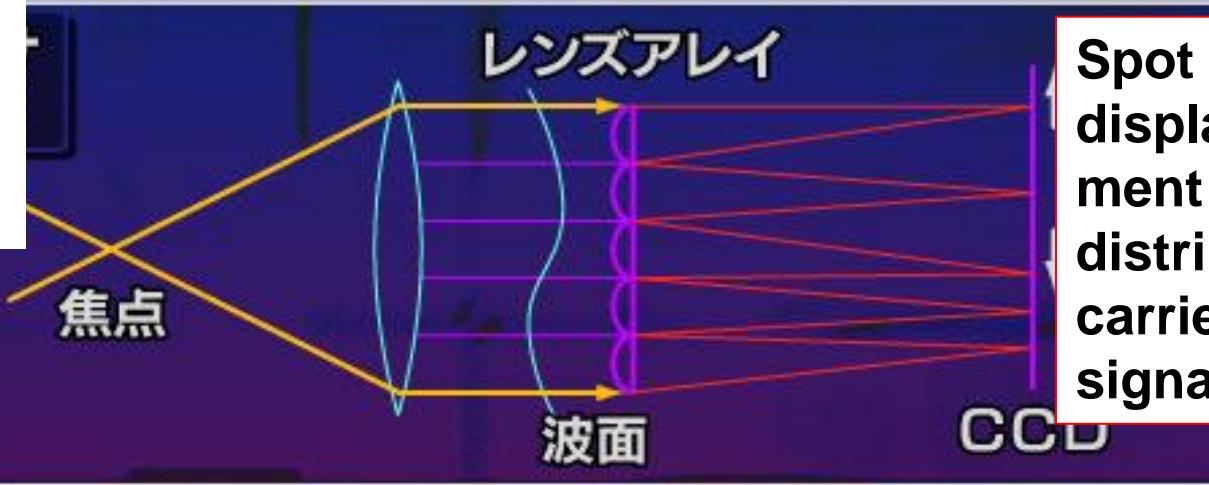
Without AO (1999)





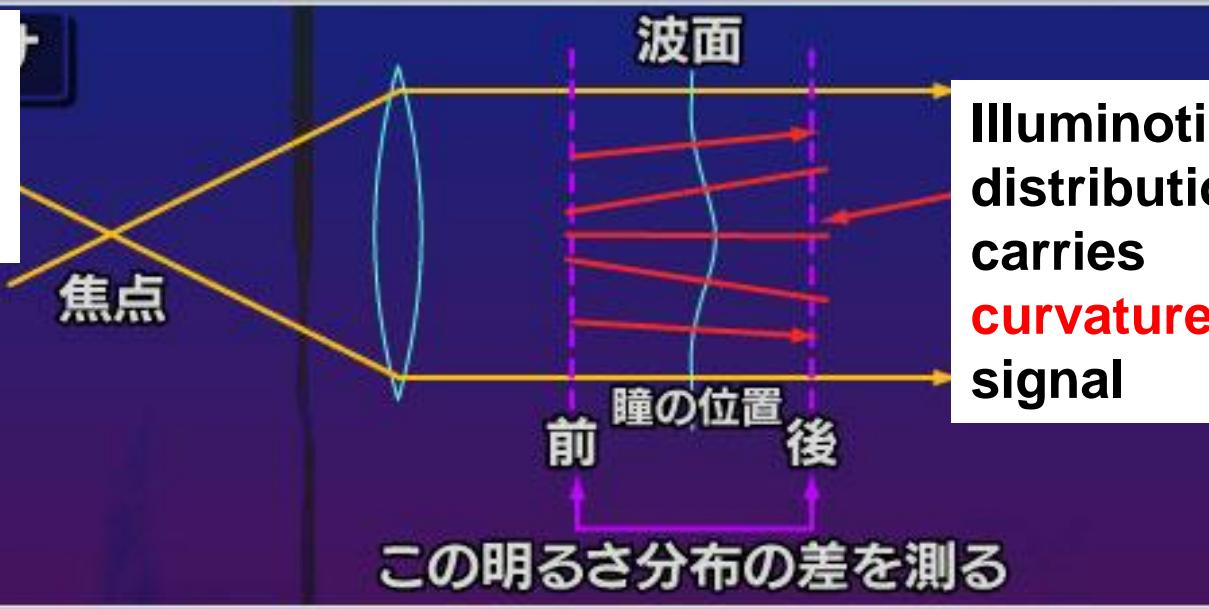
# Wavefront sensors

Shack-Hartmann  
wavefront sensor:  
1<sup>st</sup> order derivative



Spot  
displace-  
ment vector  
distribution  
carries **tilt**  
signal

Curvature  
wavefront sensor:  
2<sup>nd</sup> order derivative

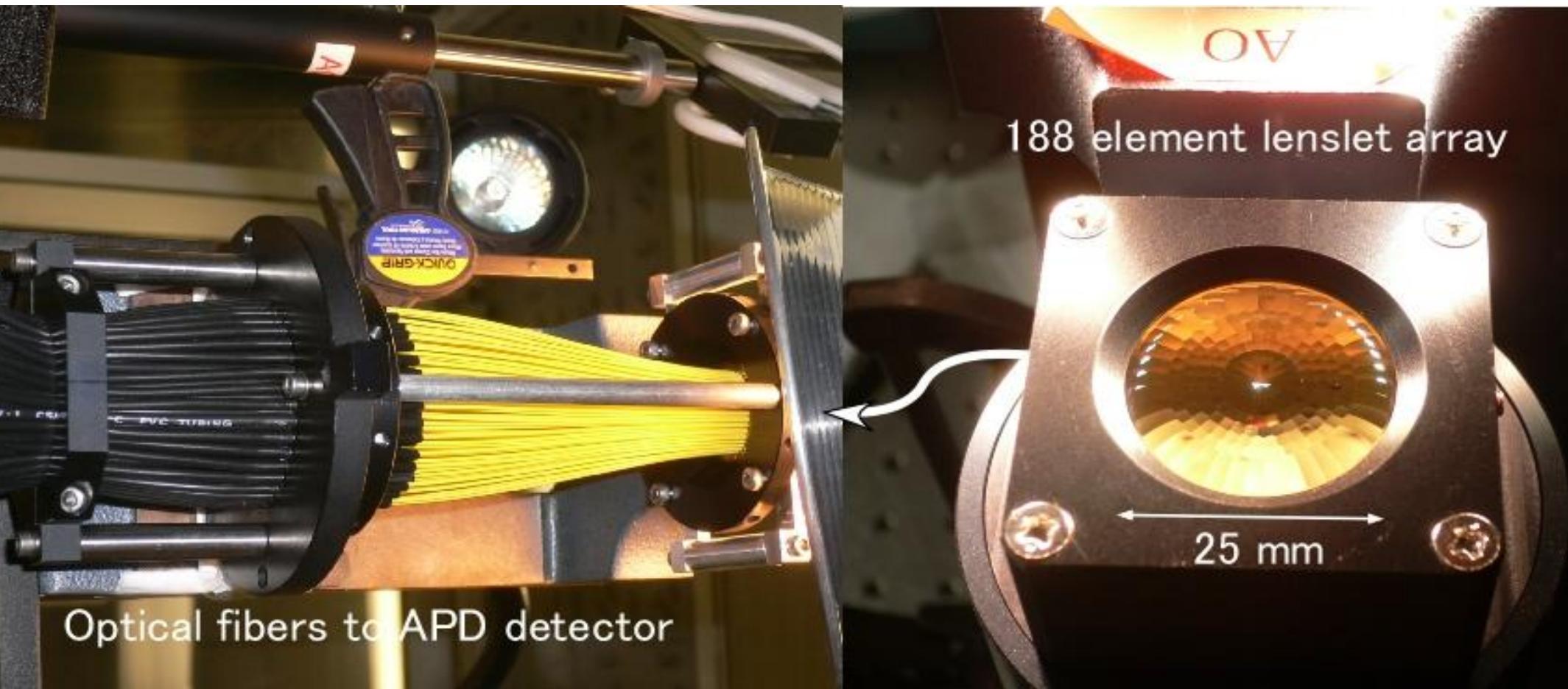


Illumination  
distribution  
carries  
**curvature**  
signal

この明るさ分布の差を測る

# Wavefront sensor

wavefront sensing at 188 subapertures



# Functioning Principles of Deformable mirrors

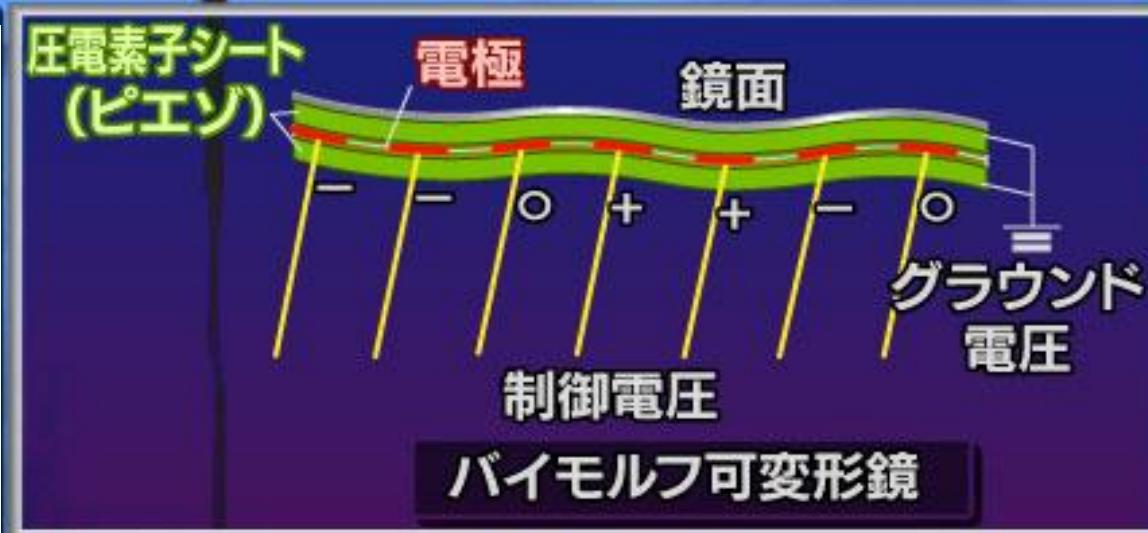
Stacked piezo rods:

Used with Shack-Hartmann sensor for high order correction



Bi-morph piezo sheet :

Used with curvature sensor for medium/low order correction

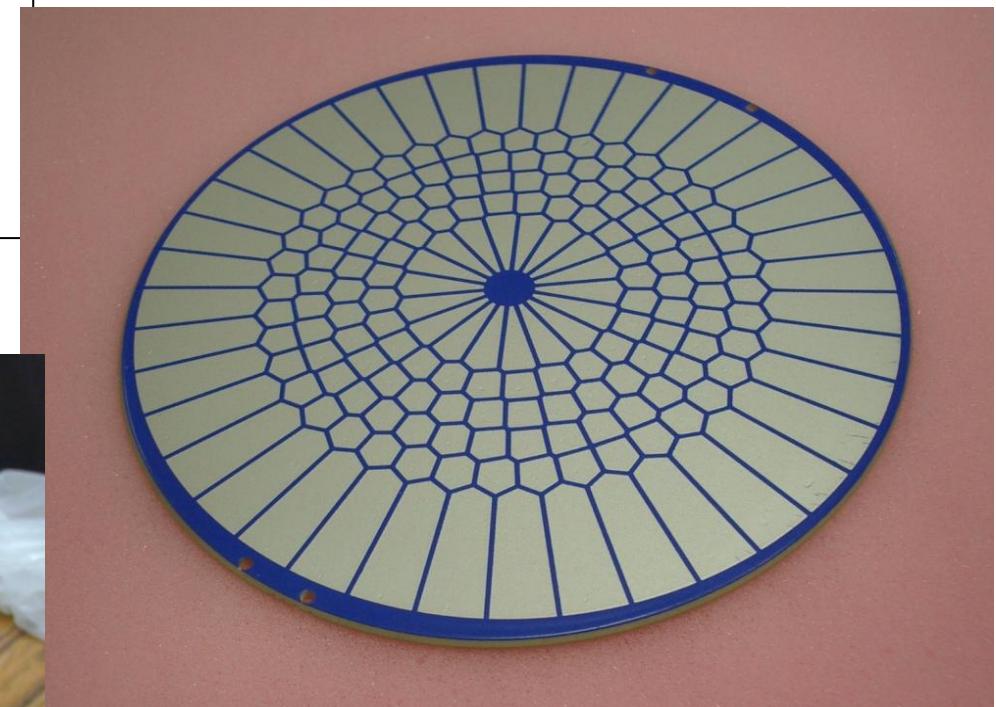


# Deformable mirror with 188 electrodes.

Diameter : 90 mm

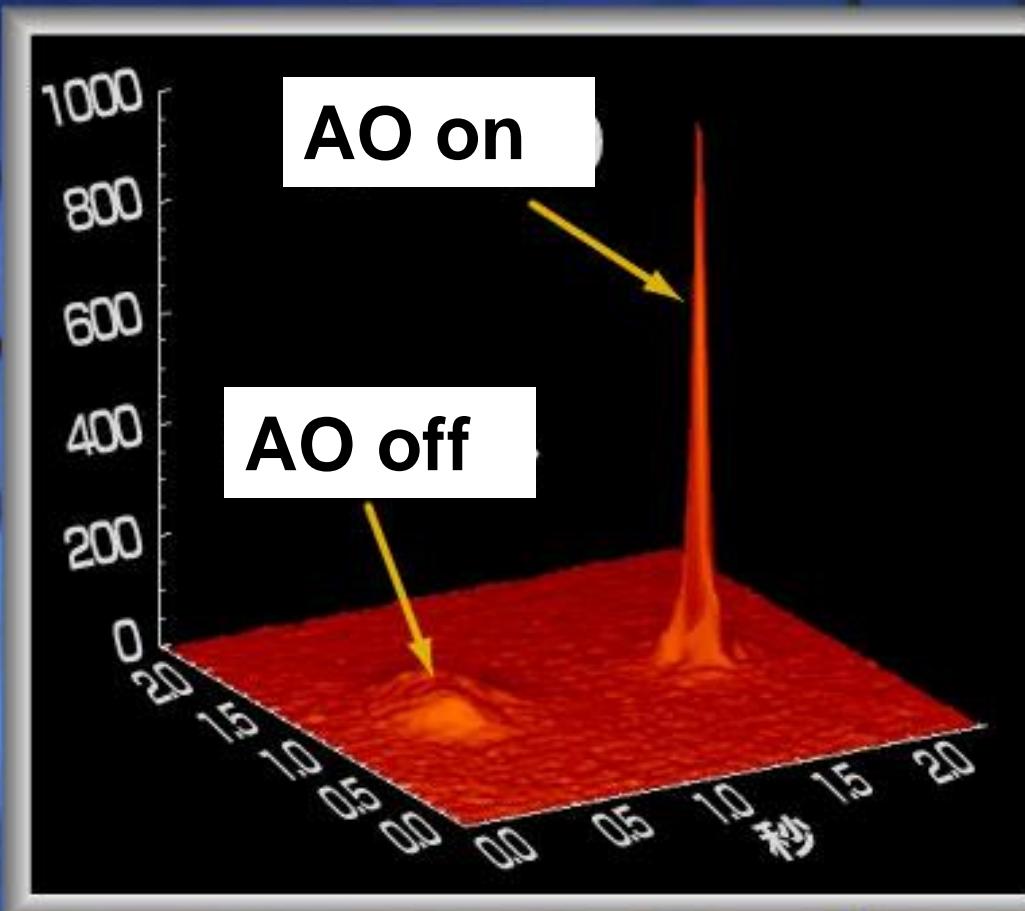
Speed : 200Hz

Price : 200k\$ (CILAS)



**188 Electrode distribution  
Designed by NAOJ**

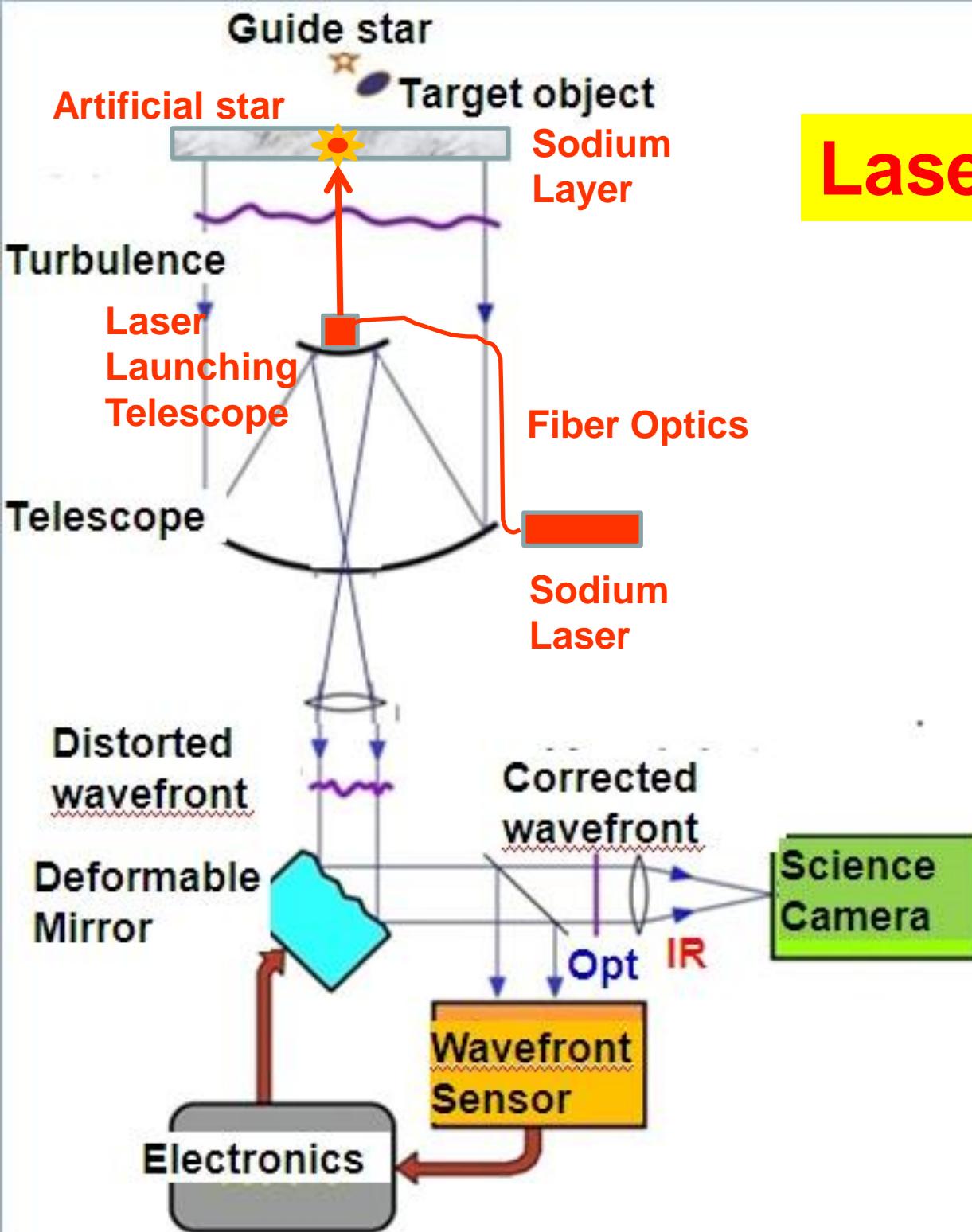
# Image improvement in the AO188 first light (2006)



Band : 2.2 micron

Seeing : 0.4 arcsec

Corrected Image :  
0.07 arcsec



## Laser guide AO system

Generate an artificial guide star by illuminating the sodium layer at 90km height using a special 6W laser shining at 589nm.

# Sum frequency sodium laser



概念図

Nd:YAGレーザー

1064 nm

Nd:YAGレーザー

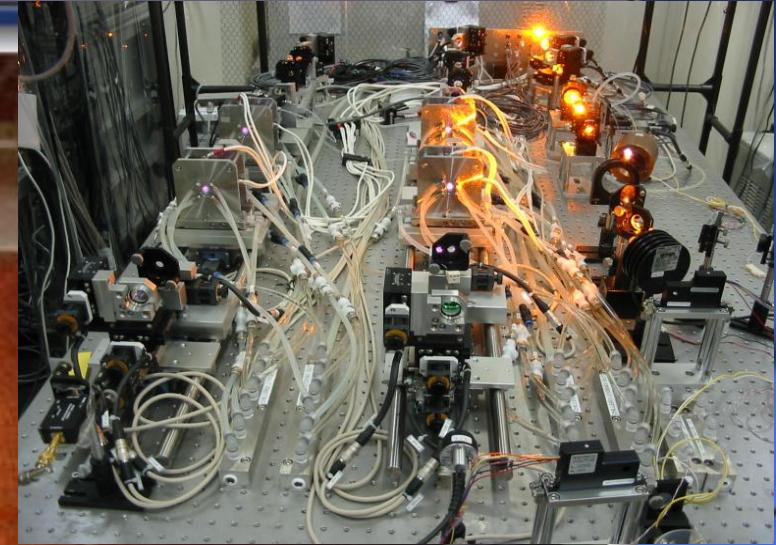
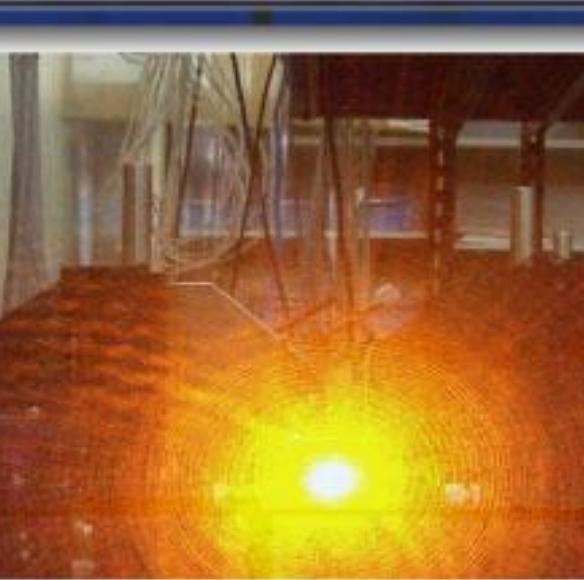
1319 nm

Non linear optics crystal

$$\frac{1}{1064 \text{ nm}} + \frac{1}{1319 \text{ nm}} = \frac{1}{589 \text{ nm}}$$

589 nm

ナトリウム  $D_2$  線共鳴波長

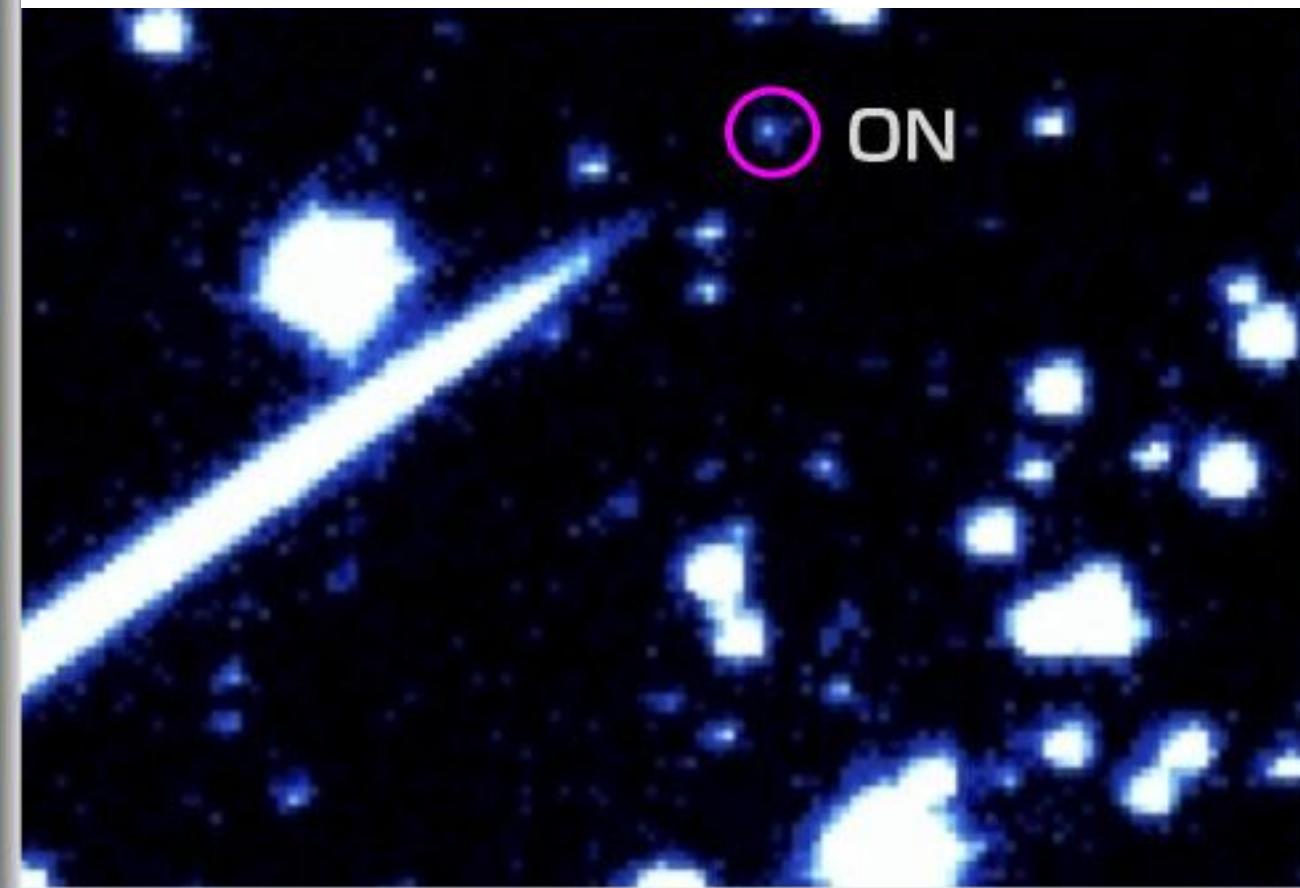


RIKEN/NAOJ

## GS Generation Verification Experiment(Oct 2005)

tuning the frequency to sodium D2 line at 0.1ppm level

Laser guide star shining at 90km height



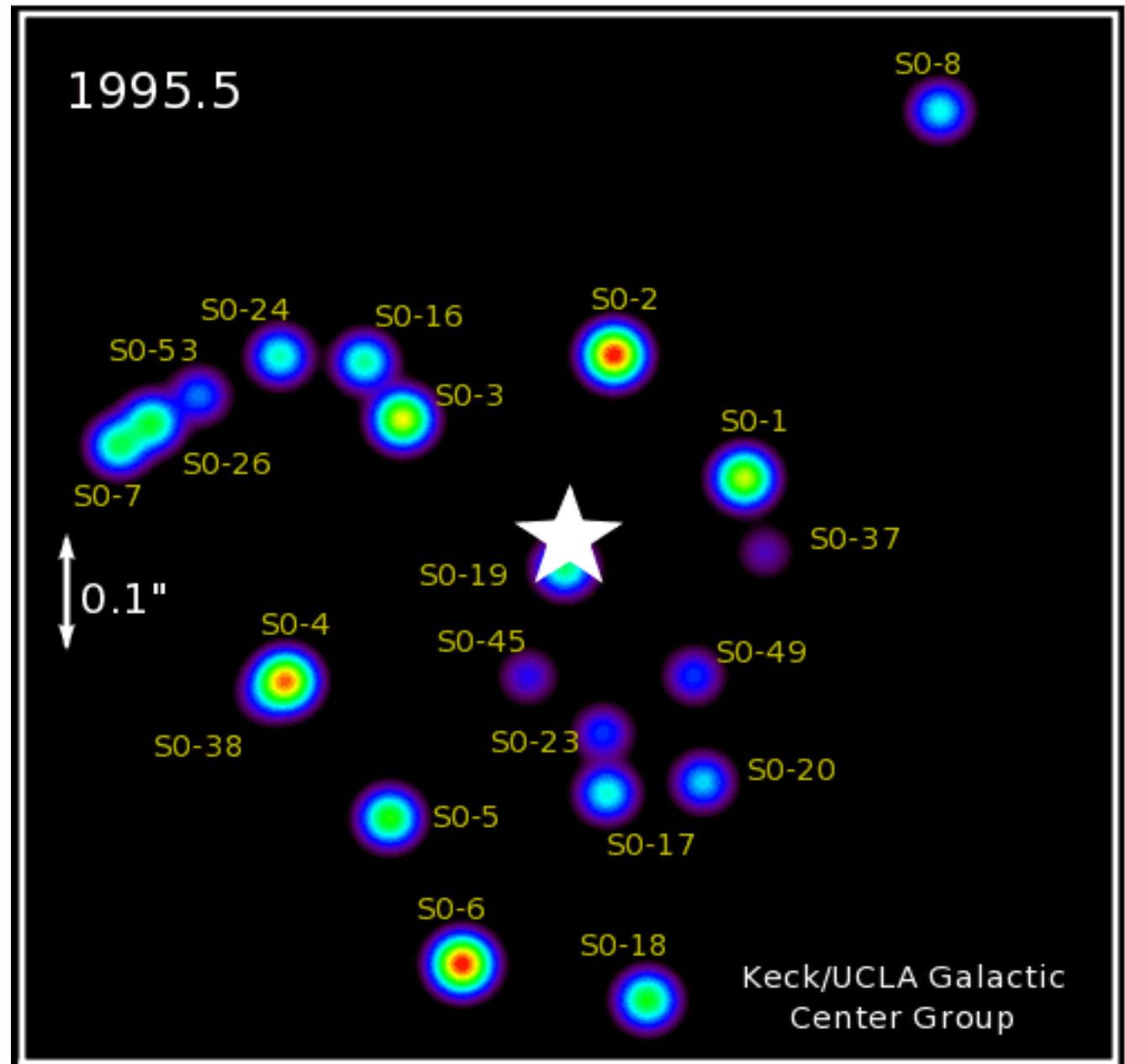
No star wars.



T.Fuse 2009

# Orbits of stars around SMBH in GC

**Orbital motion of stars around Sgr A\* indicates a BH of about million times the mass of the Sun**

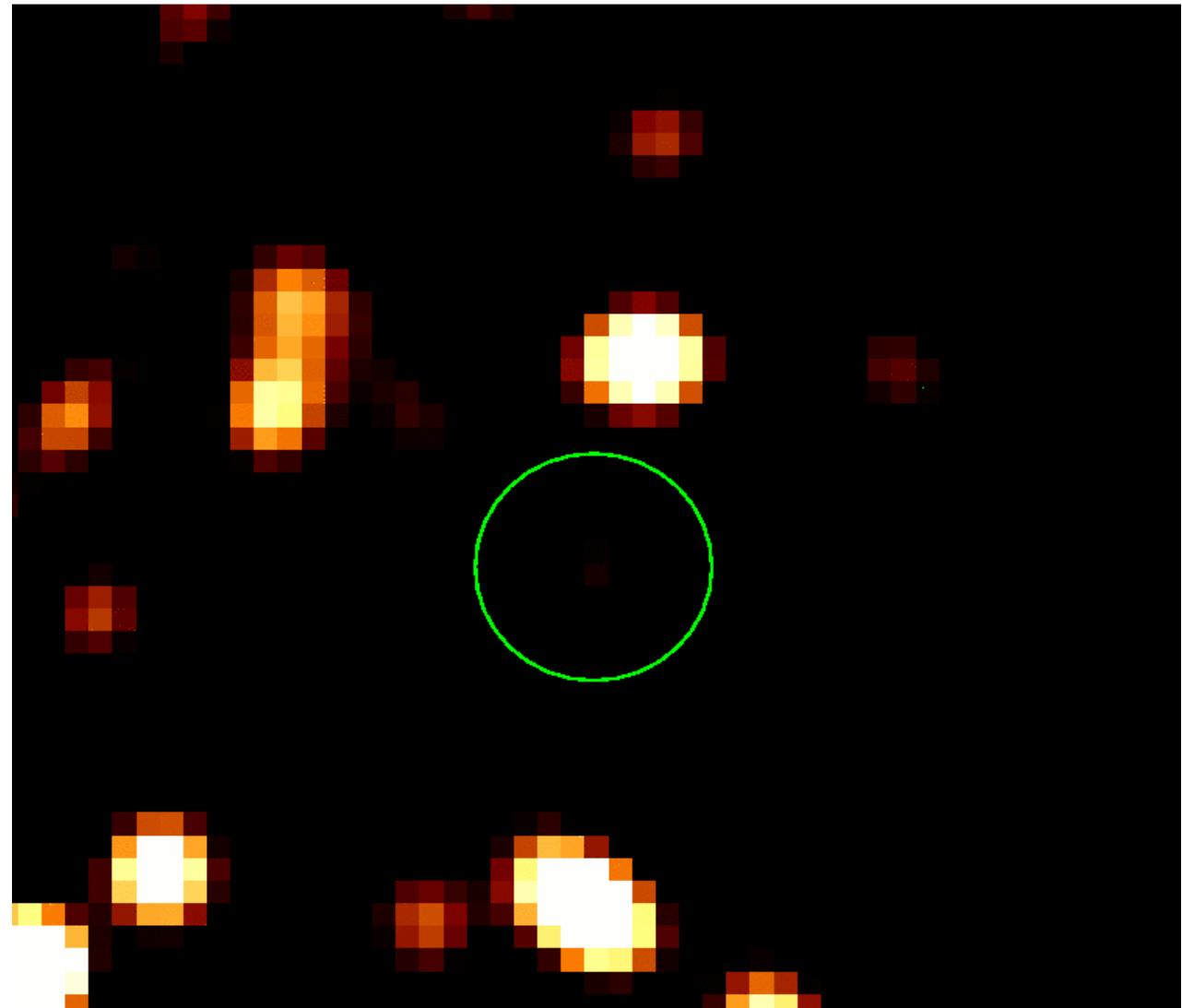


# Flickering of infalling gas onto BH in the GC

**Subaru AO36**

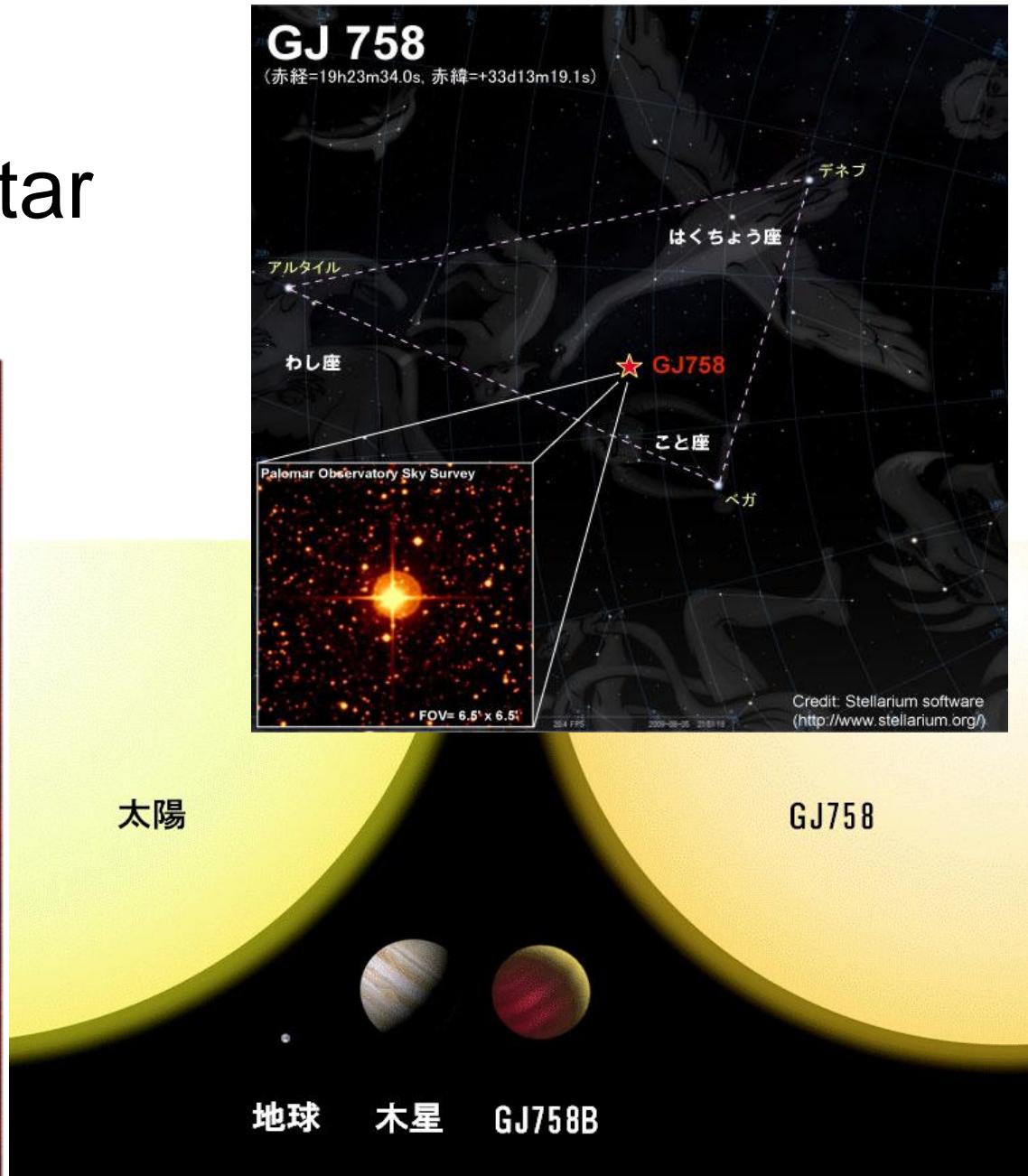
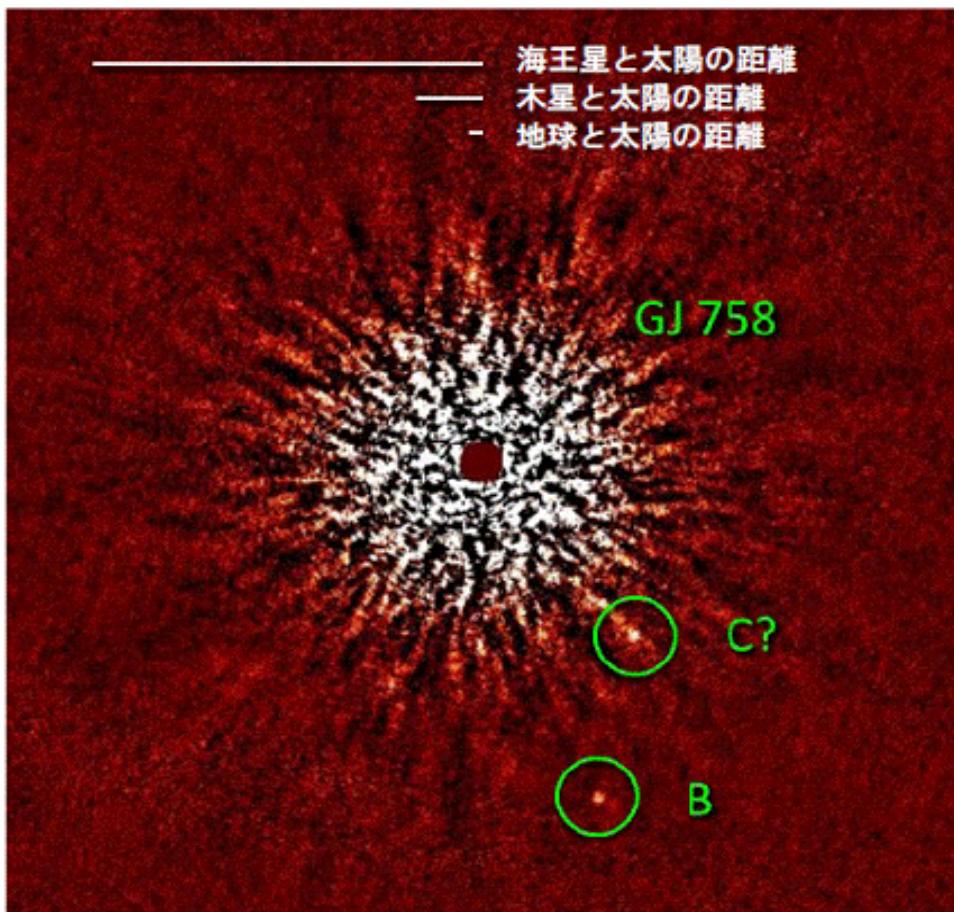
**2008 May 28**

**Nishiyama et al.  
(2009)**



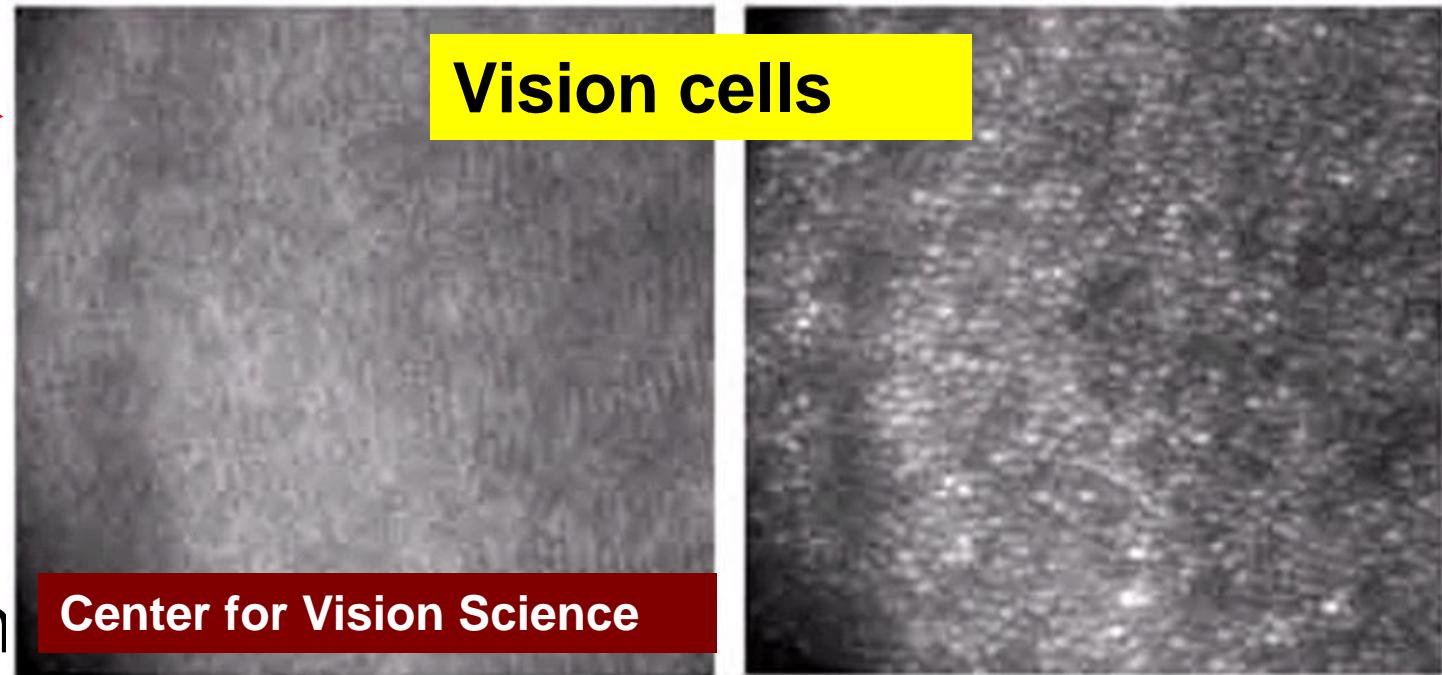
# Exoplanet candidates GJ758B,C

- Jupiter like planet around a sun like star



# Other applications of AO

- Remote sensing
- Retinal imaging
- Laser surgery
- Laser Communication
- Laser cutter
- Laser nuclear fusion
- Guiding lightening
- Purification of Uranium



# *Thirty Meter Telescope*

# Thirty Meter Telescope

1.5m segments x 492

First stars and cosmic dawn

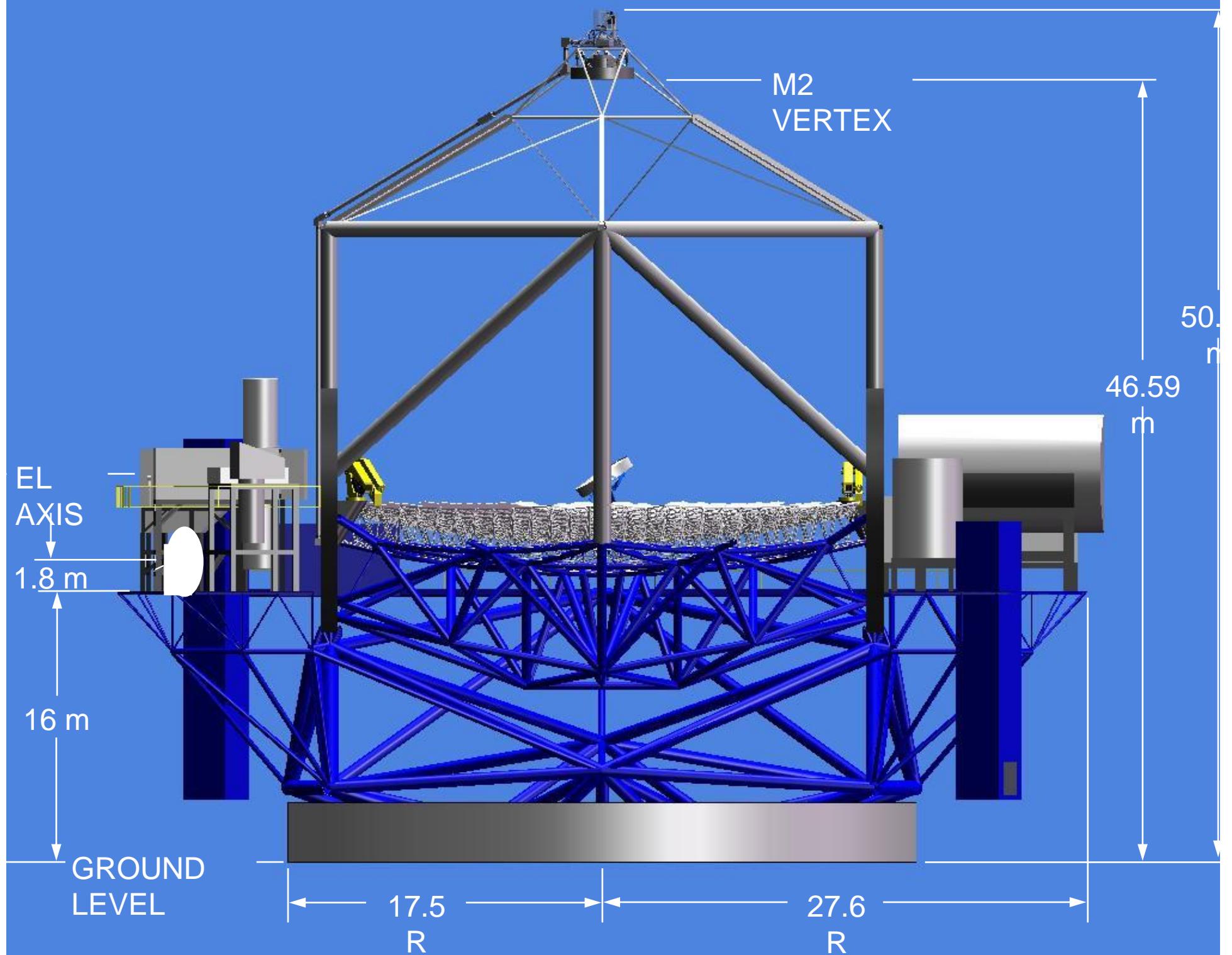
2<sup>nd</sup> Earth, Life in other planets

Dark energy, Cosmology constants

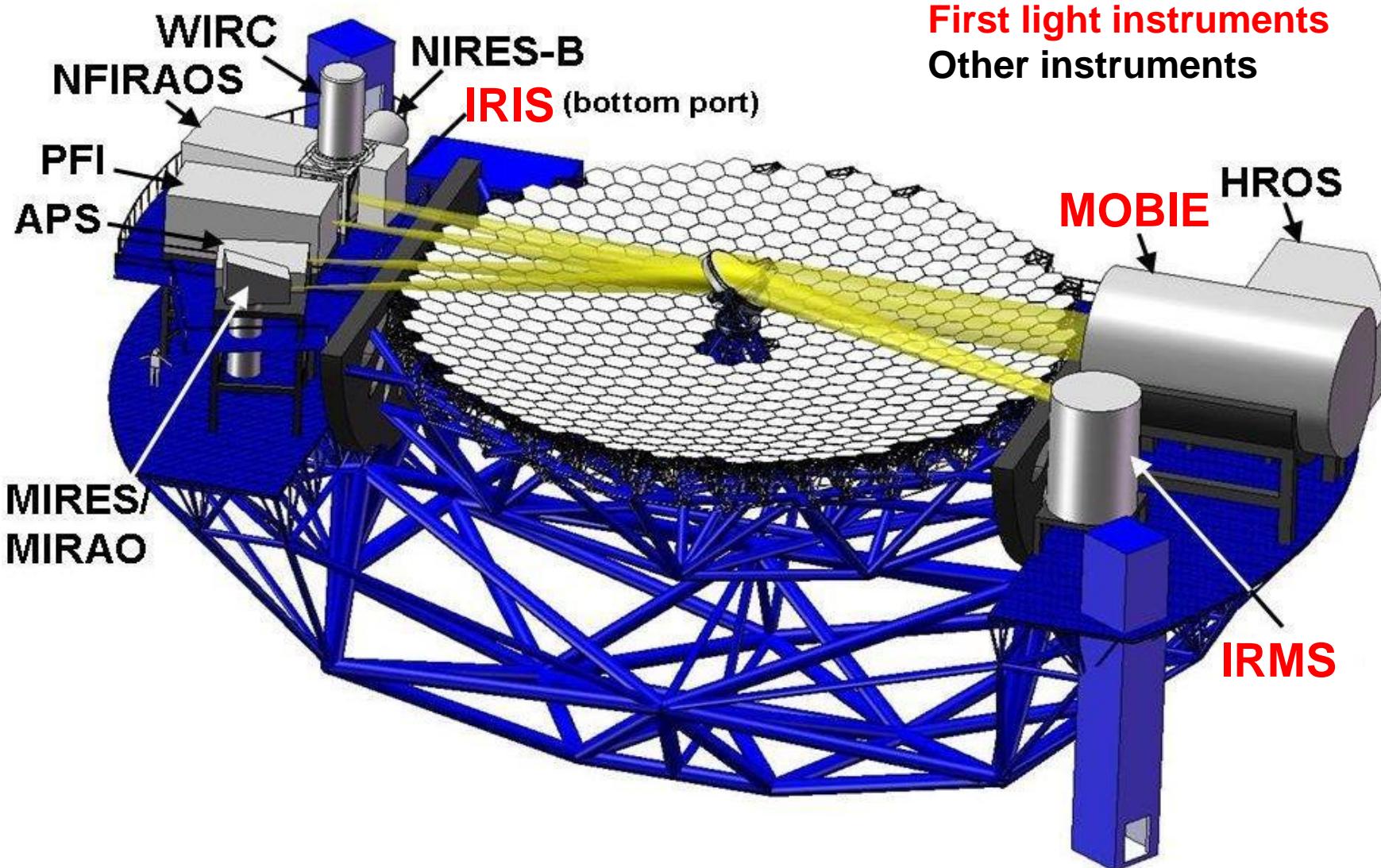
Construction 1.3B\$  
FL 2019 End?



NAOJ, Caltech, Univ.California  
ACURA, AURA, NAOC, India



# Instrument layout



# Evolution of telescopes

Masanori Iye(NAOJ)

Okayama 1.88m



8.2m Subaru



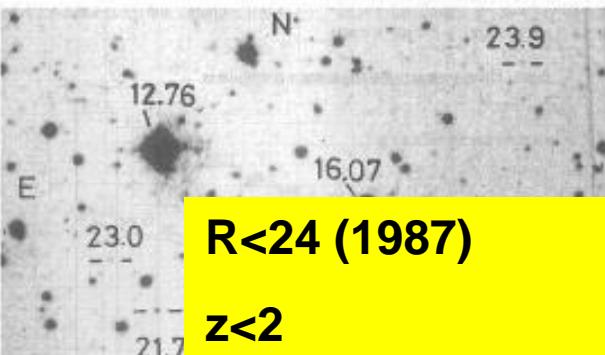
30m TMT



No. 777

学内広報

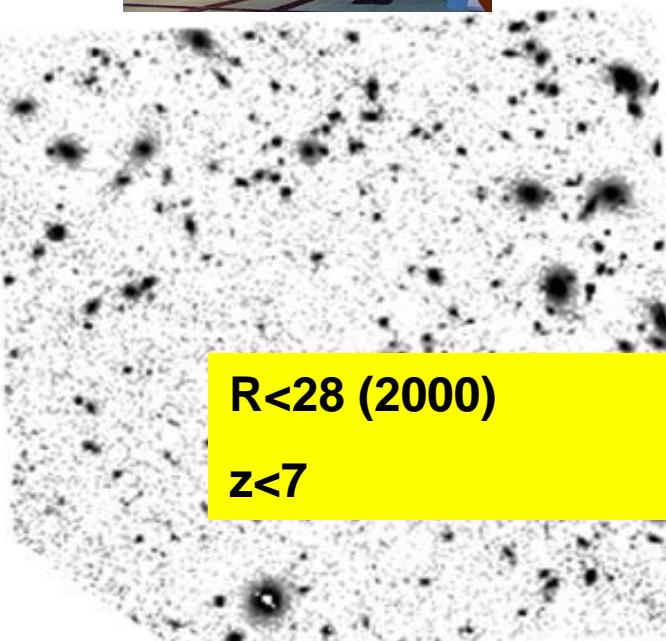
1987. 12. 21  
東京大学広報委員会



24等星を撮られたCCDカメラ

東京天文台山天体物理観測所で完成した准望遠鏡用CCDカメラ(写真右)を、夜空の暗い本音鏡測所のシット環道路に取り付けて、1987年5月に北極極域を観測し、24等級の星を撮影する事初めて達成することことができた(写真左)。CCD(固体撮像素子)カメラは高量子効率、低噪音の光検出装置であり、測光精度が良いため、従来の2P真乾板による観測に比べて2等級以上優る(平均分光率1以下の場合の)天体が撮影できるようになった。

北極極域は、天の川から最も離れた方向で星の分布密度が小さいため、深い宇宙を撮るために



Subaru/S-Cam 60min, R=28, 0''.5

*Light collecting power =x13*

*Spatial resol'n =x3.6*

*Efficiency =x180*

R<33? (2019?)

Z< 17?

# R~33 A firefly on the dark side of the Moon. (~10mW), 1 hour exposure on TMT with LGSAO



# New frontier for TMT

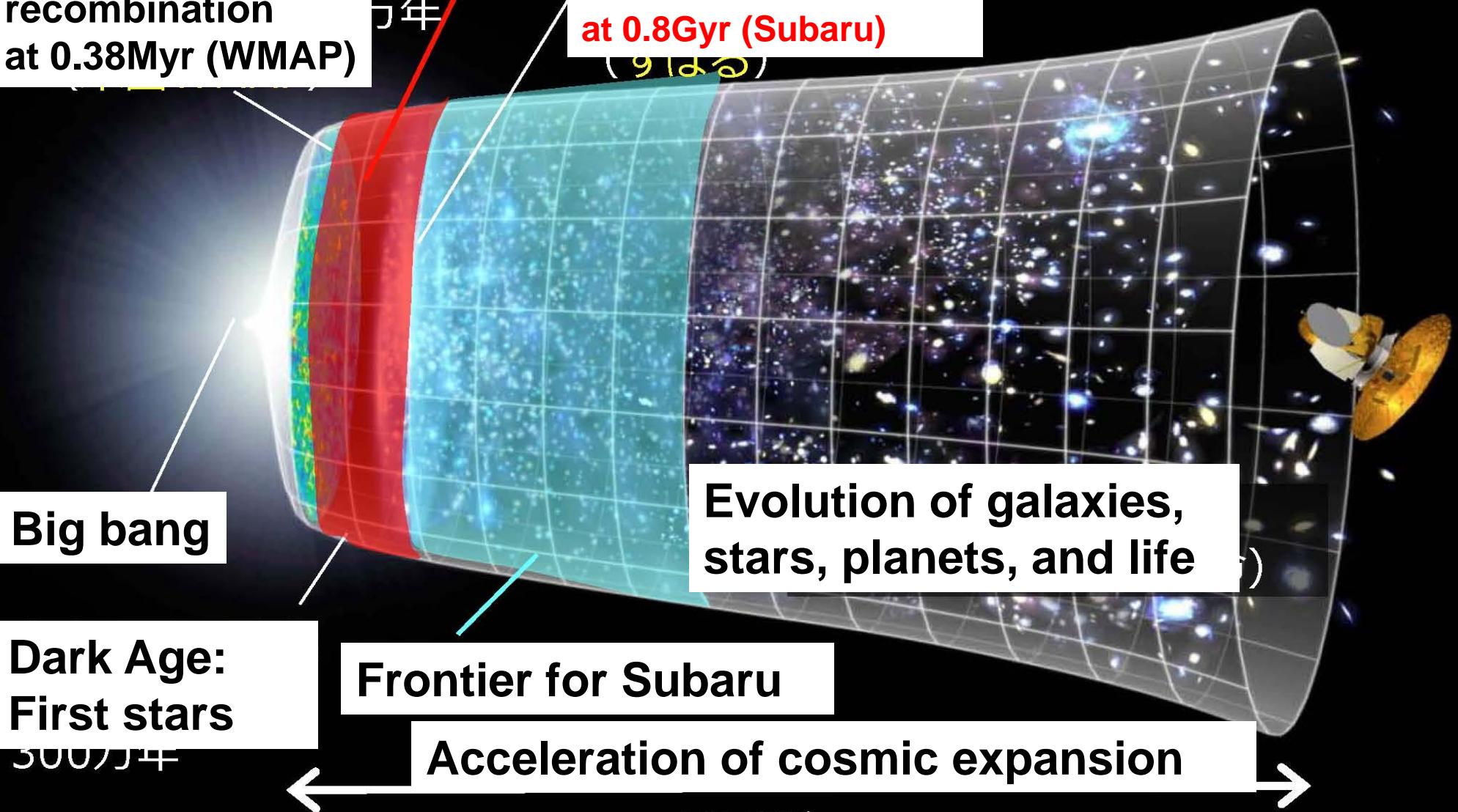
Cosmic  
recombination  
at 0.38Myr (WMAP)

行年

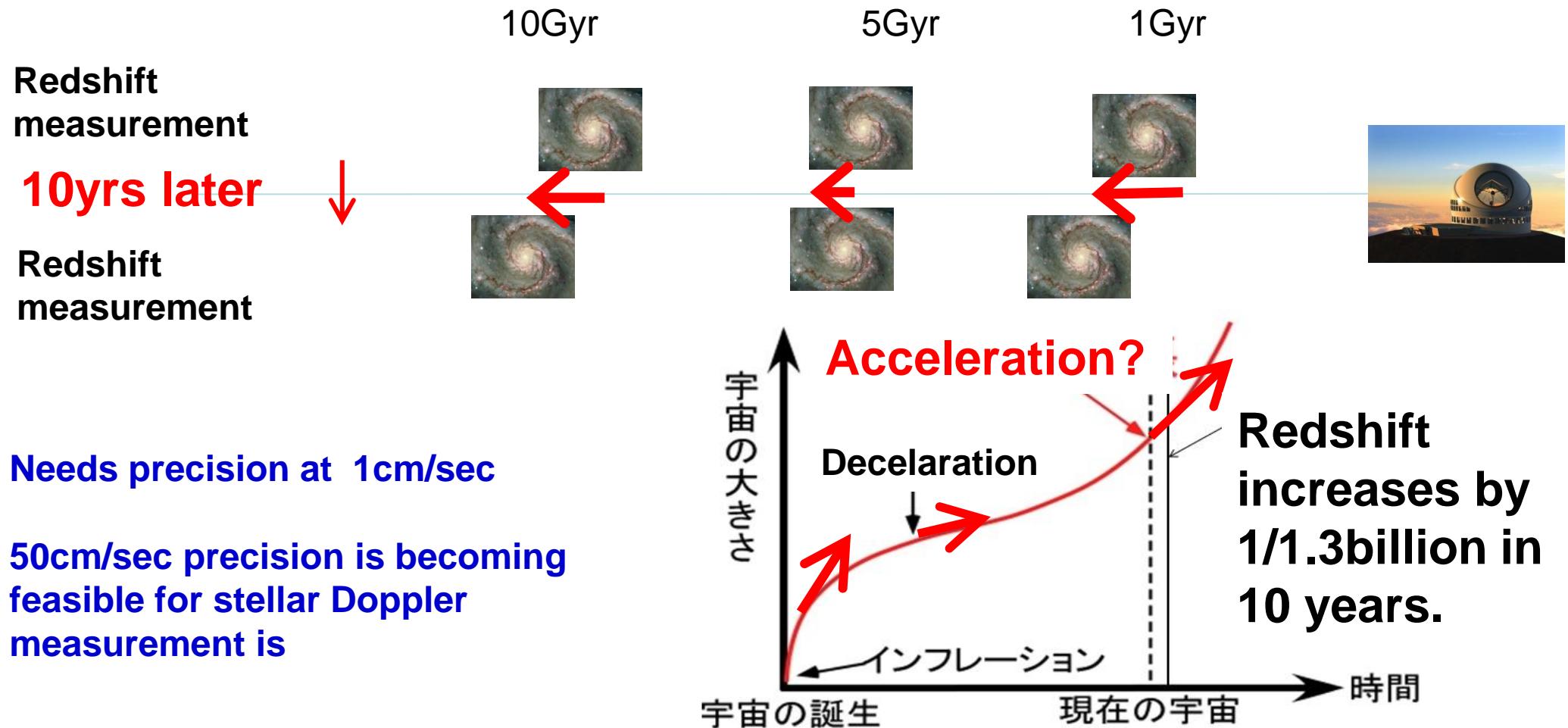
最遠方の銀河

Cosmic re-ionization  
at 0.8Gyr (Subaru)

(9はる)

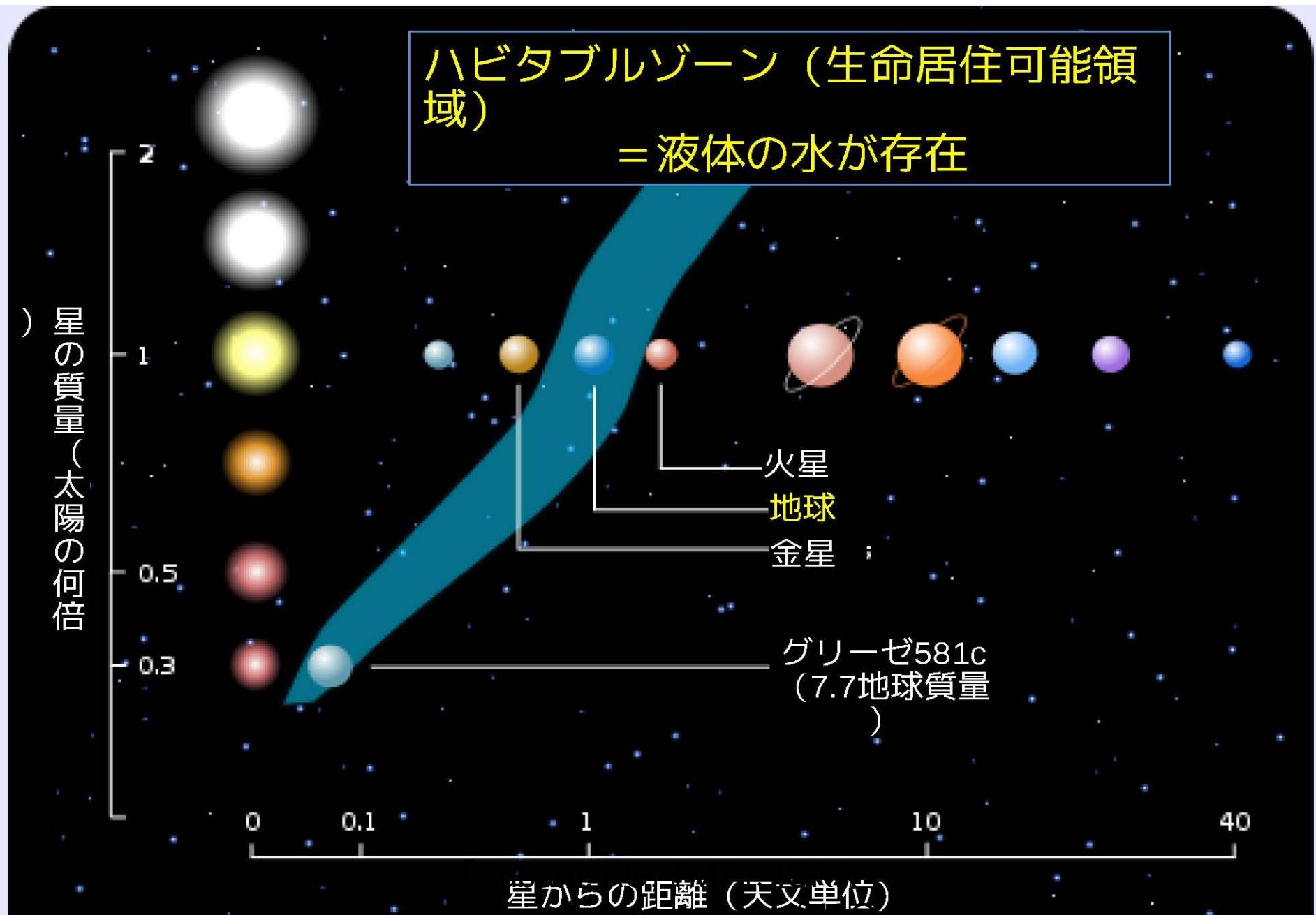


# Direct measurement of cosmic expansion history to quantify the dark energy property

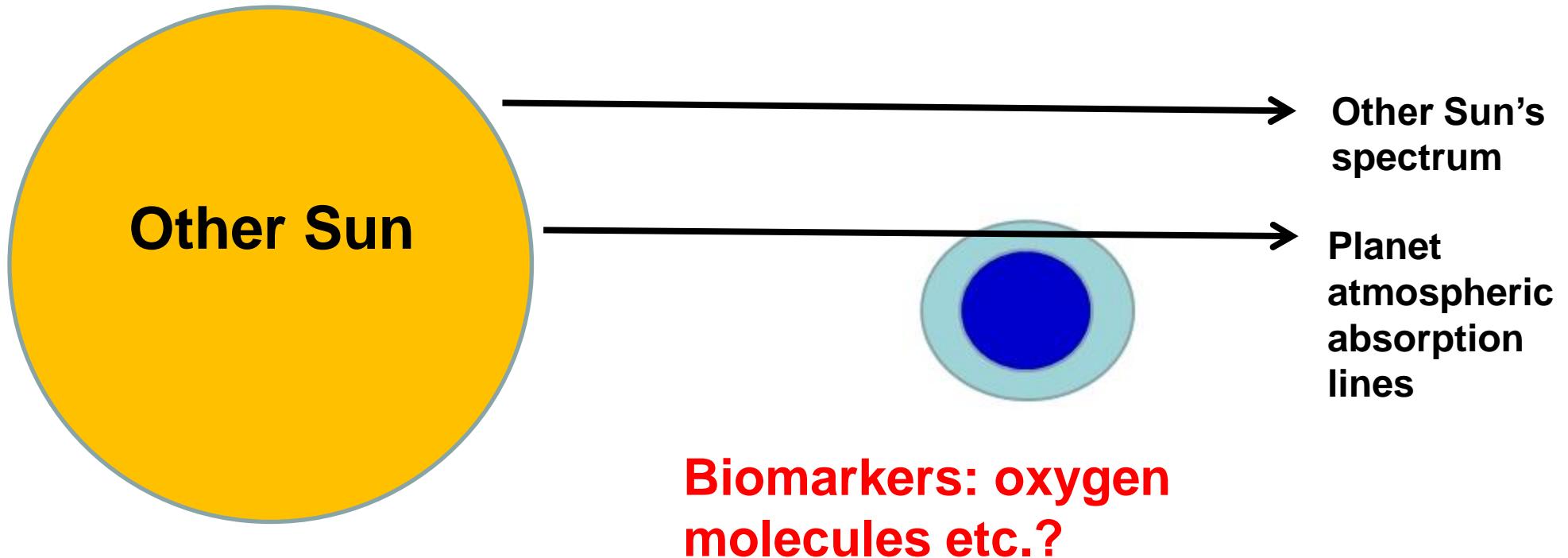


# Planets in habitable zone

## Search for biomarkers in the atmosphere



# Composition analysis of atmosphere





250M\$  
Gordon Moore (Intel)

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