

Recent results from Telescope Array

CosPA 2013 Cosmic Rays 11/14

Takeshi Okuda (Ritsumeikan University)
for Telescope Array Collaboration

Outline

- Recent Physics Result
 - Energy Spectrum
 - Mass Composition
 - Photon Limit
 - Anisotropy
- Experiment for Extended Energy Region
 - TA Low Energy Extension
 - TA High Energy Extension
- TA Affiliate Experiment



Telescope Array Collaboration

T. Abu-Zayyad^a, M. Allen^a, R. Anderson^a, R. Azuma^b, E. Barcikowski^a, J. W. Belz^a, D. R. Bergman^a, S. A. Blake^a, R. Cady^a, M. J. Chae^c, B. G. Cheon^d, J. Chiba^e, M. Chikawa^f, W. R. Cho^g, T. Fujii^h, M. Fukushima^{h,i}, K. Goto^j, W. Hanlon^a, Y. Hayashi^j, N. Hayashida^k, K. Hibino^k, K. Honda^l, D. Ikeda^h, N. Inoue^m, T. Ishii^l, R. Ishimori^b, H. Itoⁿ, D. Ivanov^{a,o}, C. C. H. Jui^a, K. Kadota^p, F. Kakimoto^b, O. Kalashev^q, K. Kasahara^r, H. Kawai^s, S. Kawakami^j, S. Kawana^m, K. Kawata^h, E. Kido^h, H. B. Kim^d, J. H. Kim^a, J. H. Kim^d, S. Kitamura^b, Y. Kitamura^b, V. Kuzmin^q, Y. J. Kwon^g, J. Lan^a, J.P. Lundquist^a, K. Machida^l, K. Martensⁱ, T. Matsuda^t, T. Matsuyama^j, J. N. Matthews^a, M. Minamino^j, K. Mukai^l, I. Myers^a, K. Nagasawa^m, S. Nagatakiⁿ, T. Nakamura^u, H. Nanpei^j, T. Nonaka^h, A. Nozato^f, S. Ogio^j, S. Oh^c, M. Ohnishi^h, H. Ohoka^h, K. Oki^h, T. Okuda^v, M. Onoⁿ, A. Oshima^j, S. Ozawa^r, I. H. Park^w, M. S. Pshirkov^x, D. C. Rodriguez^a, G. Rubtsov^q, D. Ryu^y, H. Sagawa^h, N. Sakurai^j, A. L. Sampson^a, L. M. Scott^o, P. D. Shah^a, F. Shibata^l, T. Shibata^h, H. Shimodaira^h, B. K. Shin^d, T. Shirahama^m, J. D. Smith^a, P. Sokolsky^a, R. W. Springer^a, B. T. Stokes^a, S. R. Stratton^{a,o}, T. A. Stroman^a, M. Takamura^e, A. Taketa^z, M. Takita^h, Y. Tameda^k, H. Tanaka^j, K. Tanaka^{aa}, M. Tanaka^t, S. B. Thomas^a, G. B. Thomson^a, P. Tinyakov^{q,x}, I. Tkachev^q, H. Tokuno^b, T. Tomida^{ab}, S. Troitsky^q, Y. Tsunesada^b, K. Tsutsumi^b, Y. Uchihori^{ac}, F. Urban^x, G. Vasiloff^a, Y. Wada^m, T. Wong^a, H. Yamaoka^t, K. Yamazaki^j, J. Yang^c, K. Yashiro^e, Y. Yoneda^j, S. Yoshida^s, H. Yoshii^{ad}, R. Zollinger^a, Z. Zundel^a

^aUniversity of Utah, ^bTokyo Institute of Technology, ^cEwha Womans University, ^dHanyang University, ^eTokyo University of Science,

^fKinki University, ^gYonsei University, ^hInstitute for Cosmic Ray Research, Univ. of Tokyo,

ⁱKavli Institute for the Physics and Mathematics of the Universe (WPI), Todai Institutes for Advanced Study, the University of Tokyo,

^jOsaka City University, ^kKanagawa University, ^lUniv. of Yamanashi, ^mSaitama University, ⁿAstrophysical Big Bang Laboratory, RIKEN,

^oRutgers University, ^pTokyo City University, ^qInstitute for Nuclear Research of the Russian Academy of Sciences, ^rWaseda University,

^sChiba University, ^tInstitute of Particle and Nuclear Studies, KEK, ^uKochi University, ^vRitsumeikan University, ^wSungkyunkwan University,

^xUniversite Libre de Bruxelles, ^yChungnam National University, ^zEarthquake Research Institute, University of Tokyo,

^{aa}Hiroshima City University, ^{ab}Advanced Science Institute, RIKEN, ^{ac}National Institute of Radiological Science, ^{ad}Ehime University



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~120 collaborators in 5 countries
Japan, USA, Korea, Russia, Belgium







^aUniversity of Utah, ^bTokyo Institute of Technology, ^cEwha Womans University, ^dHanyang University, ^eTokyo University of Science, ^fKinki University, ^gYonsei University, ^hInstitute for Cosmic Ray Research, Univ. of Tokyo, ⁱKavli Institute for the Physics and Mathematics of the Universe (WPI), Todai Institutes for Advanced Study, the University of Tokyo, ^jOsaka City University, ^kKanagawa University, ^lUniv. of Yamanashi, ^mSaitama University, ⁿAstrophysical Big Bang Laboratory, RIKEN, ^oRutgers University, ^pTokyo City University, ^qInstitute for Nuclear Research of the Russian Academy of Sciences, ^rWaseda University, ^sChiba University, ^tInstitute of Particle and Nuclear Studies, KEK, ^uKochi University, ^vRitsumeikan University, ^wSungkyunkwan University, ^xUniversite Libre de Bruxelles, ^yChungnam National University, ^zEarthquake Research Institute, University of Tokyo, ^{aa}Hiroshima City University, ^{ab}Advanced Science Institute, RIKEN, ^{ac}National Institute of Radiological Science, ^{ad}Ehime University

Telescope Array Site



Utah State in USA
39.3N, 112.9W, Alt 1400[m]

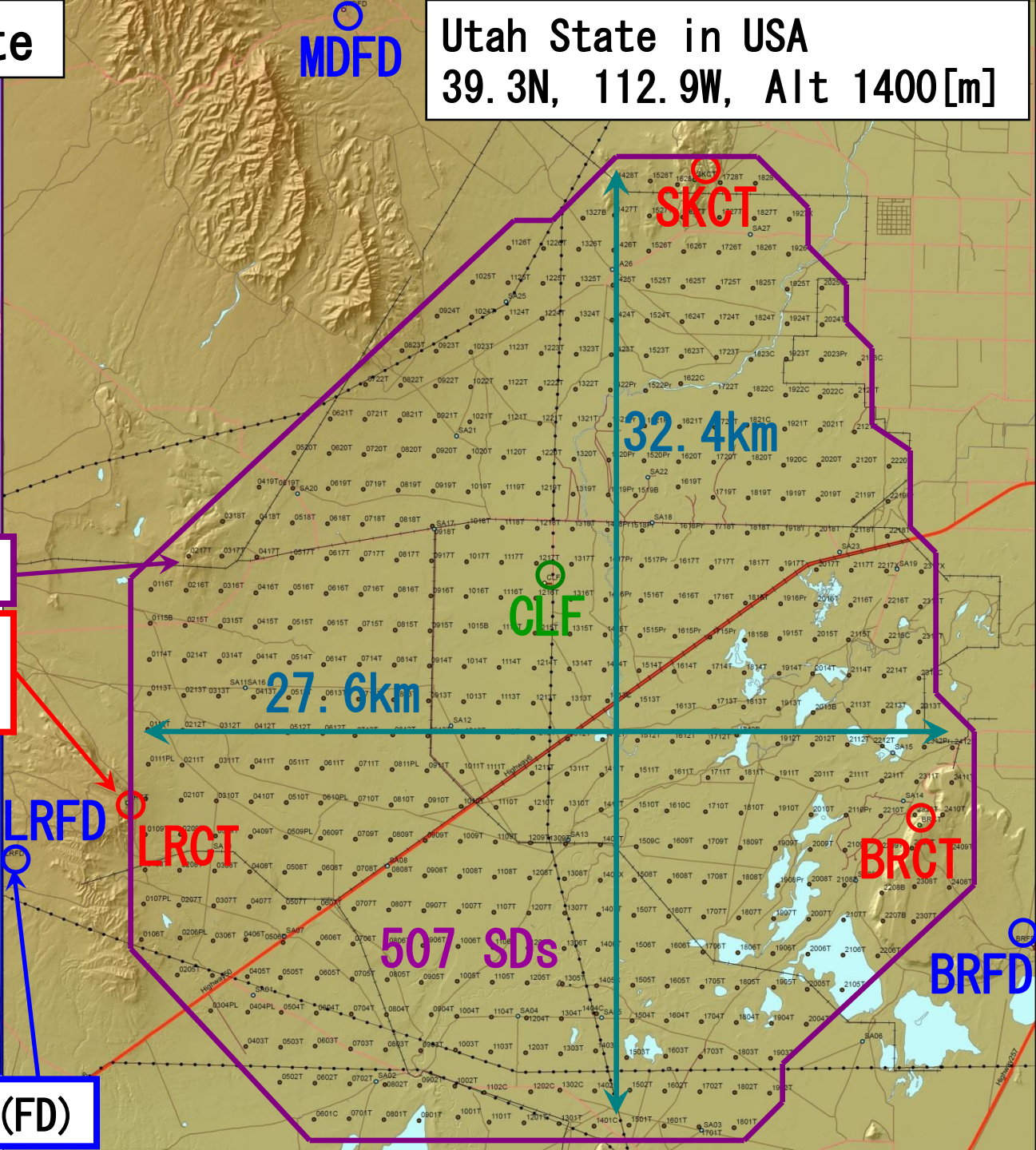


Surface Detector (SD)

Telecommunication Tower

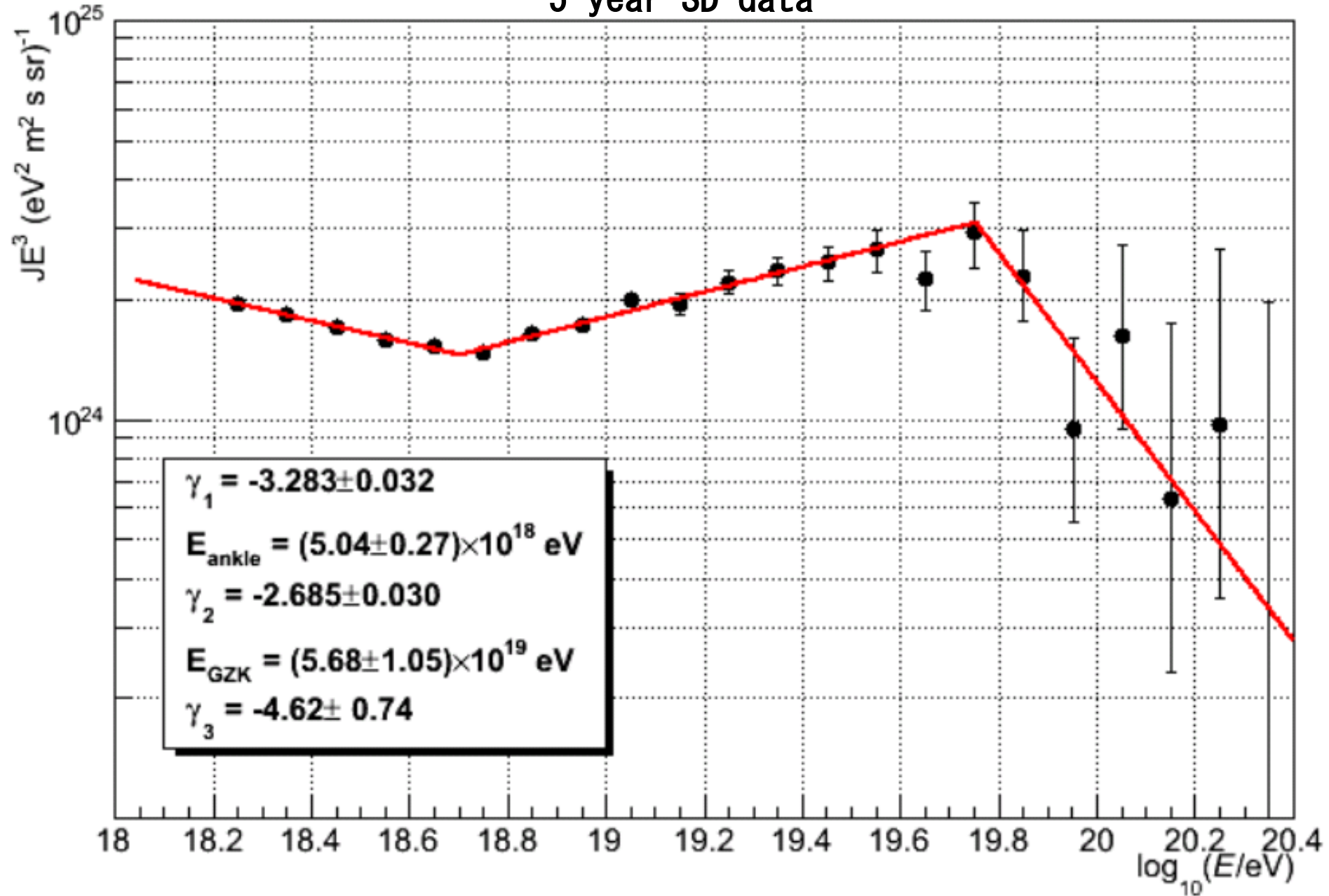


Fluorescence Detector (FD)

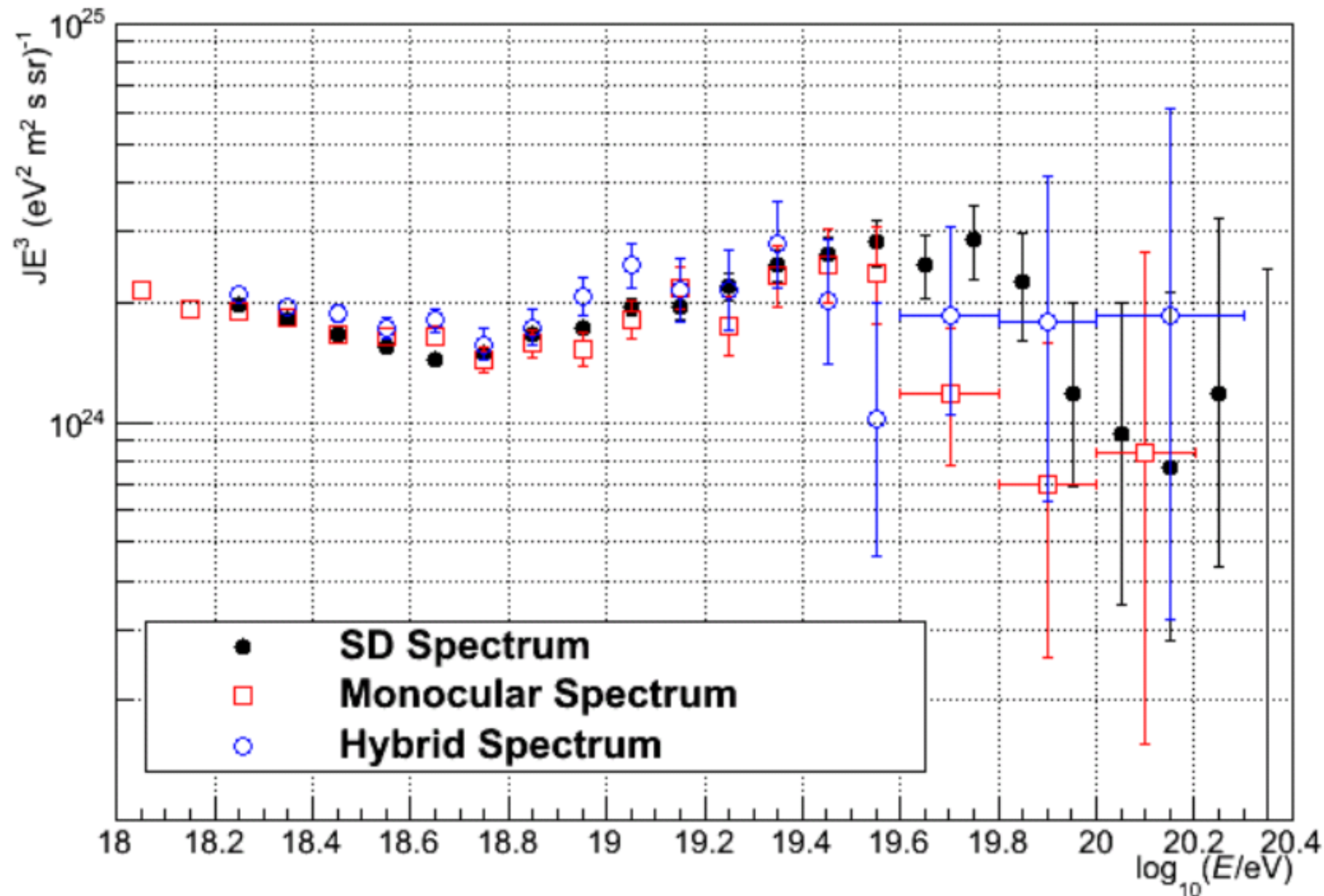


Five Year SD Spectrum

5 year SD data

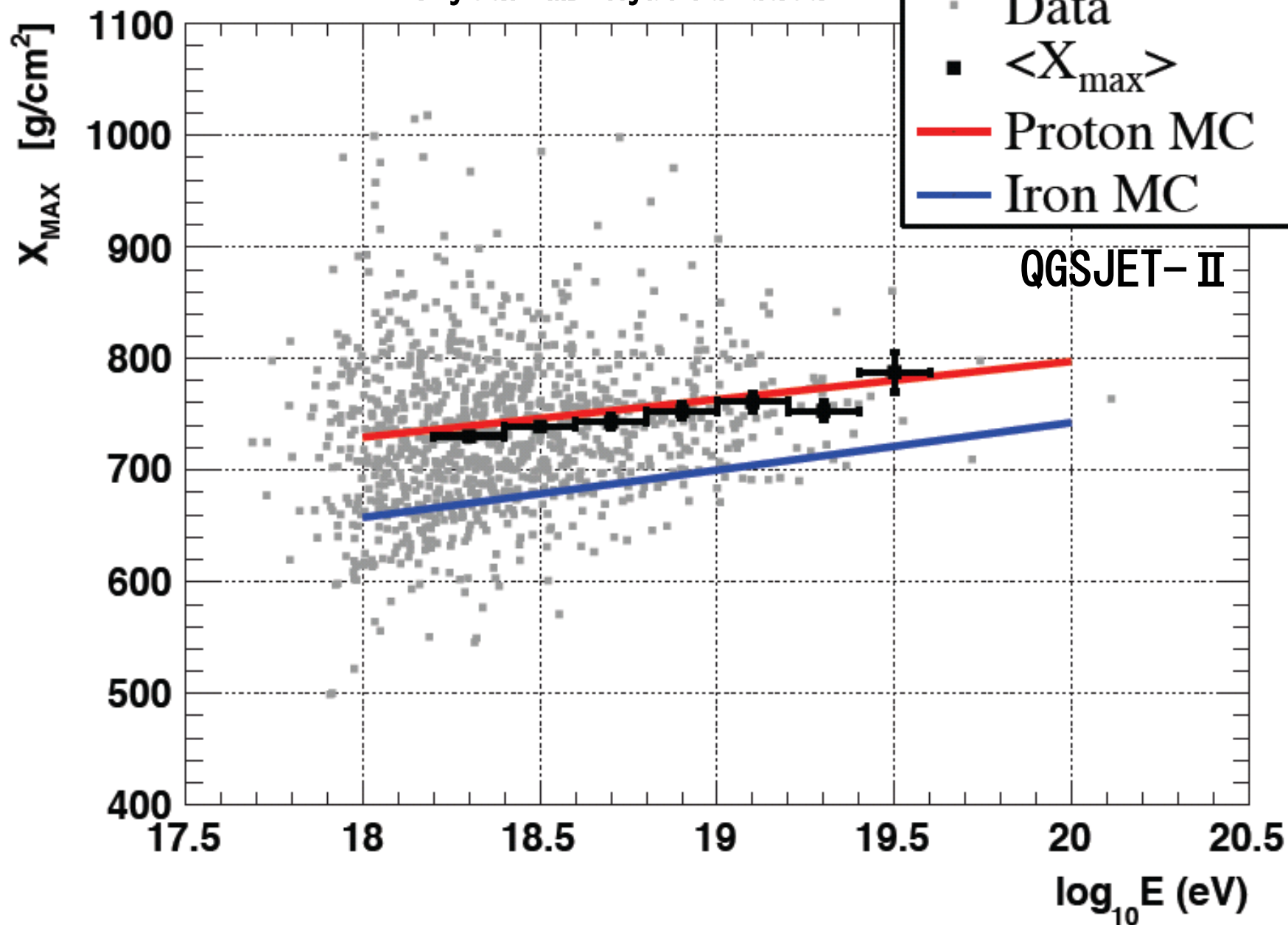


SD, Monocular and Hybrid Spectra



Mass Composition

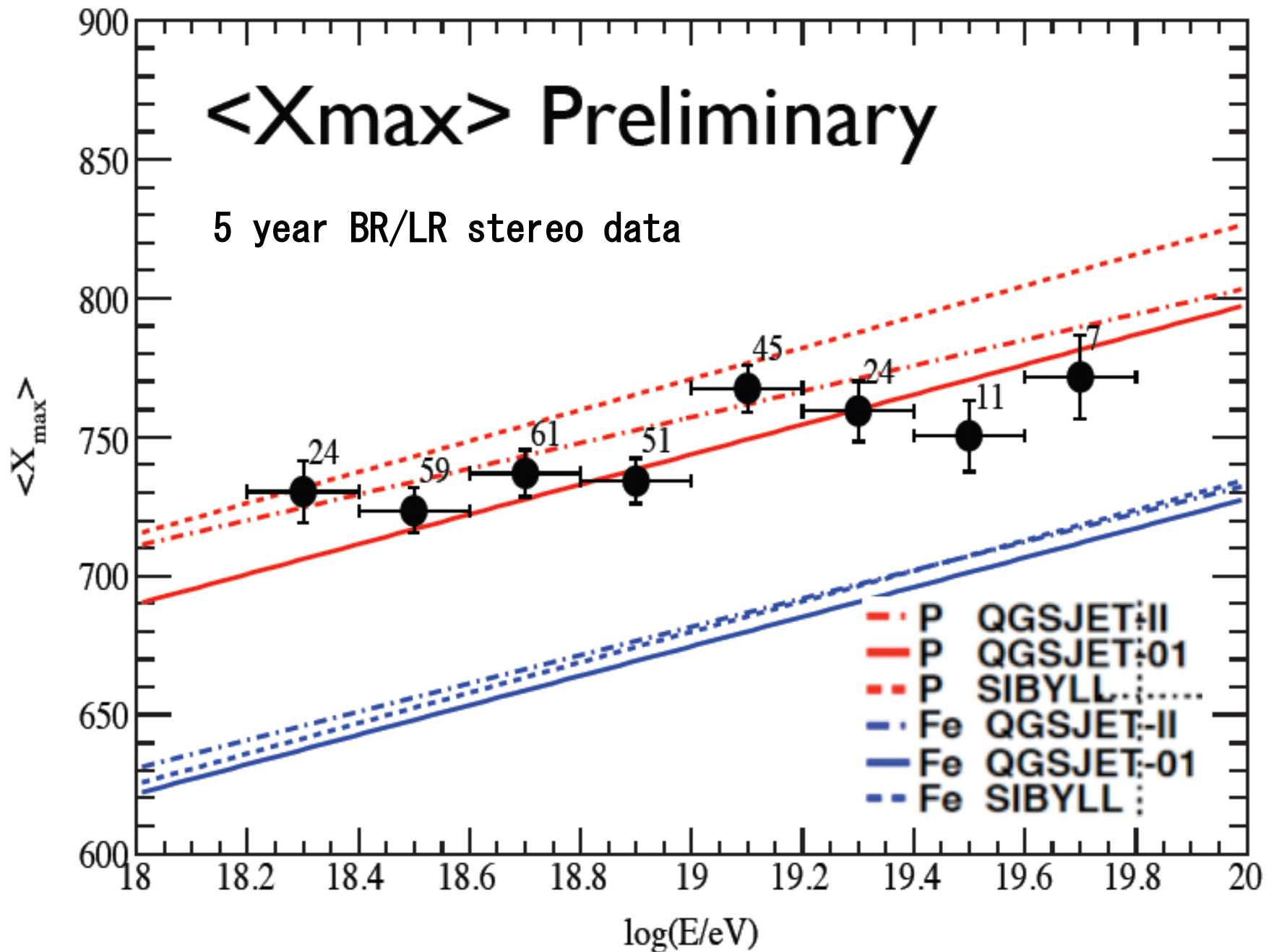
4 year MD hybrid data



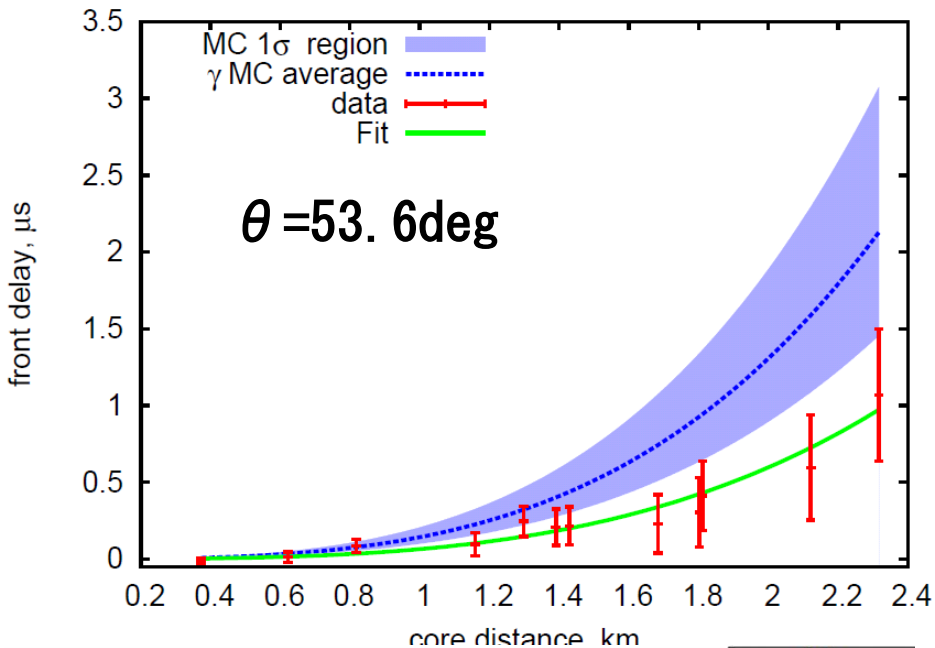
Mass Composition

$\langle X_{\max} \rangle$ Preliminary

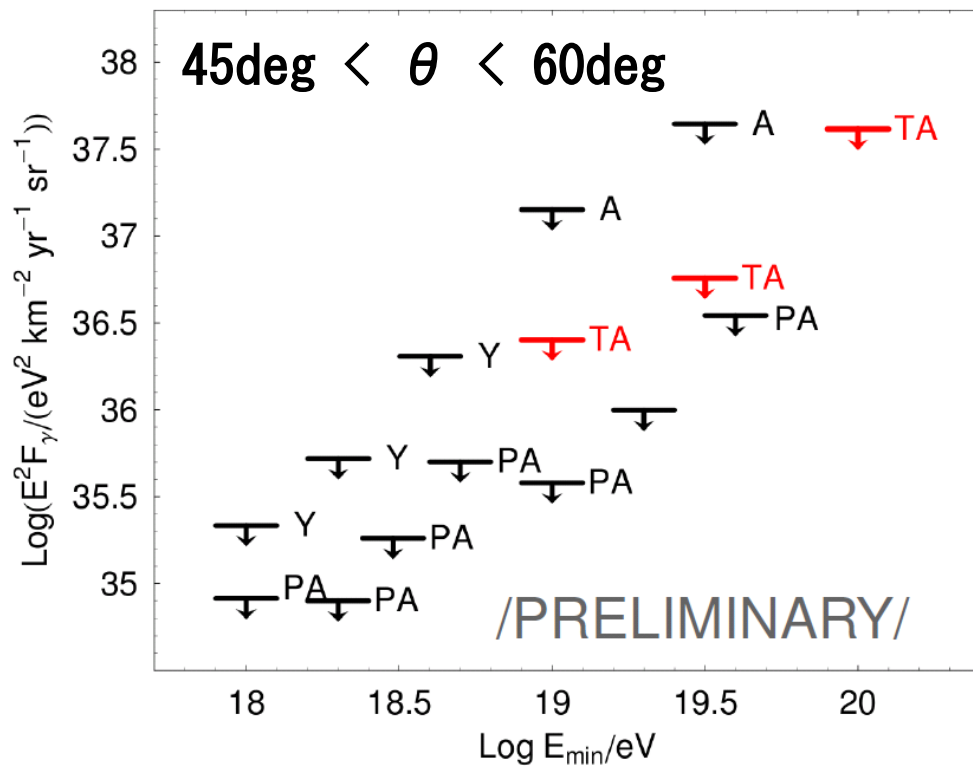
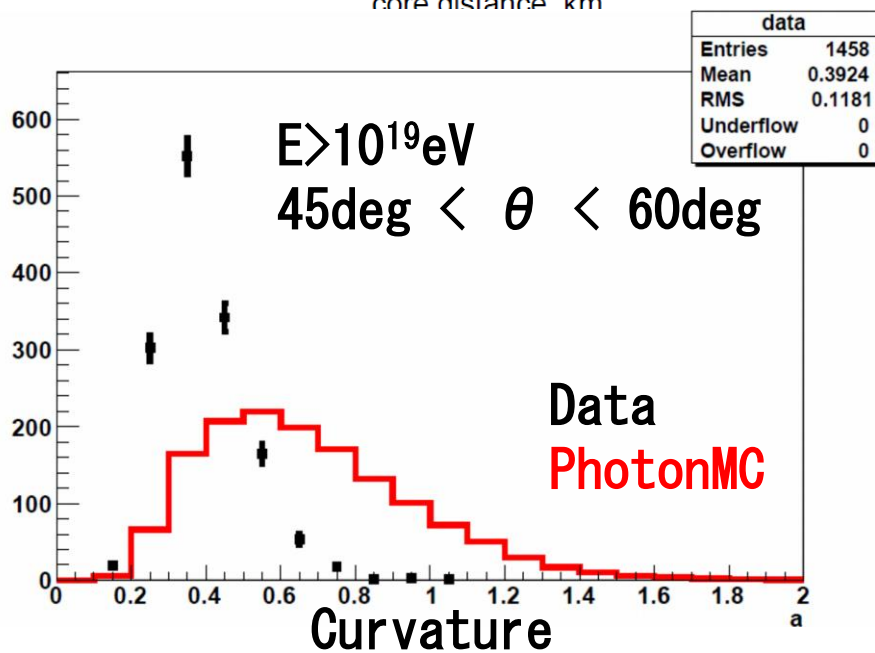
5 year BR/LR stereo data



Photon Upper Limit



The photon likelihood of the observed shower is evaluated by shower front curvature by SD. For 5 years SD data, we got the photon upper limit for which estimated energy is more than 10^{19}eV .

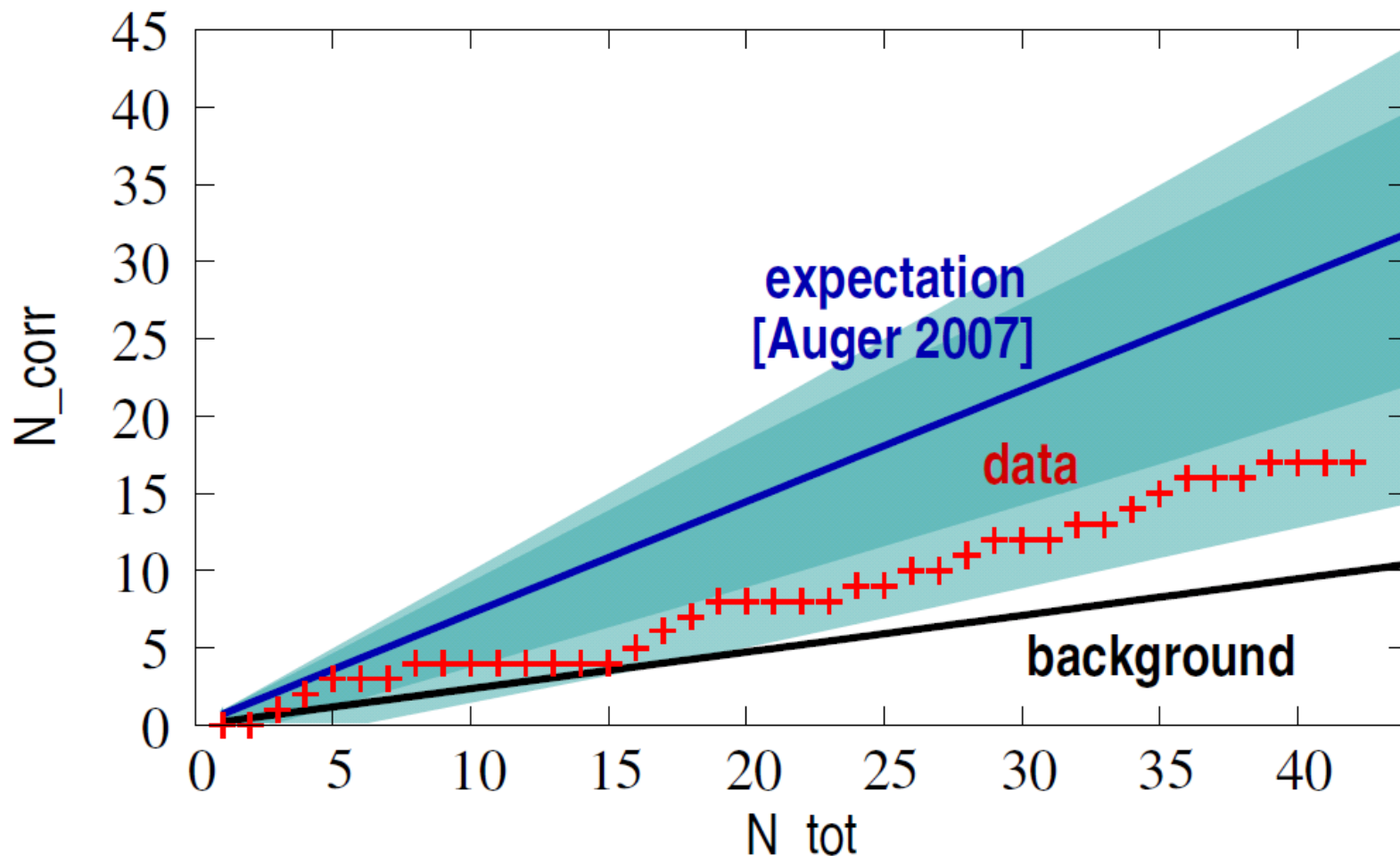


Anisotropy (Nearby AGN)

AGNs are from VCV 2006 catalog, cut by z ($0 < z \leq 0.18$).

There are 465 AGN in our field of view.

Primary Energy $> 57\text{EeV}$, Angular Separation $< 3.1^\circ$.



17 events correlate out of 42 events.

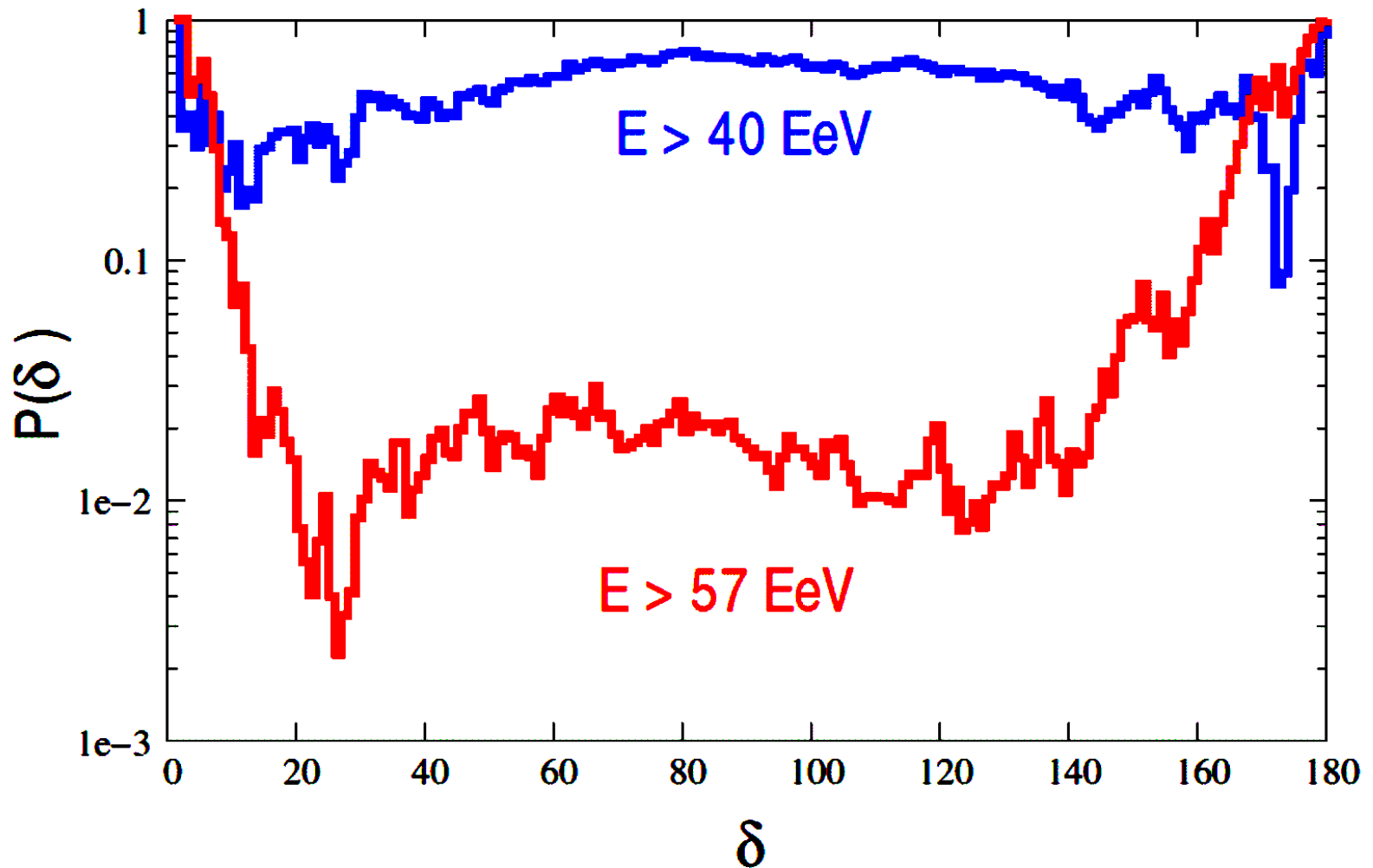
This probability is 0.014 from chance correlation.

Anisotropy (Auto-correlation)

AGASA reported clustering within 2.5° for energy $> 40\text{EeV}$.

TA found no clustering within 2.5° .

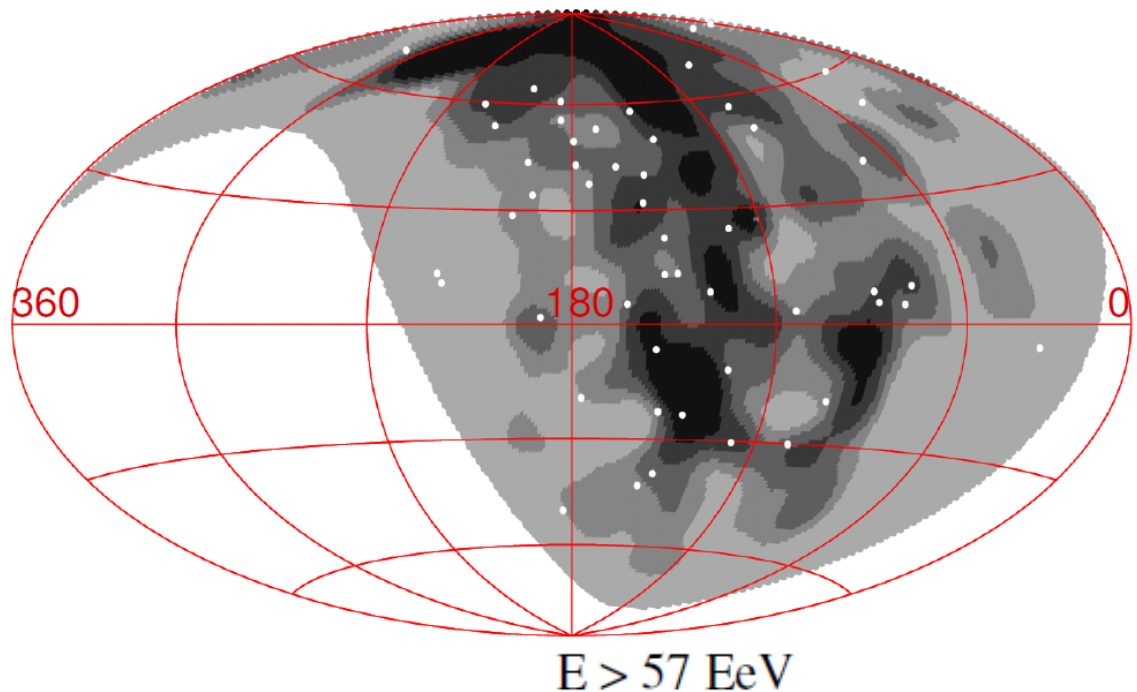
Extended search for larger angles δ and higher energy.



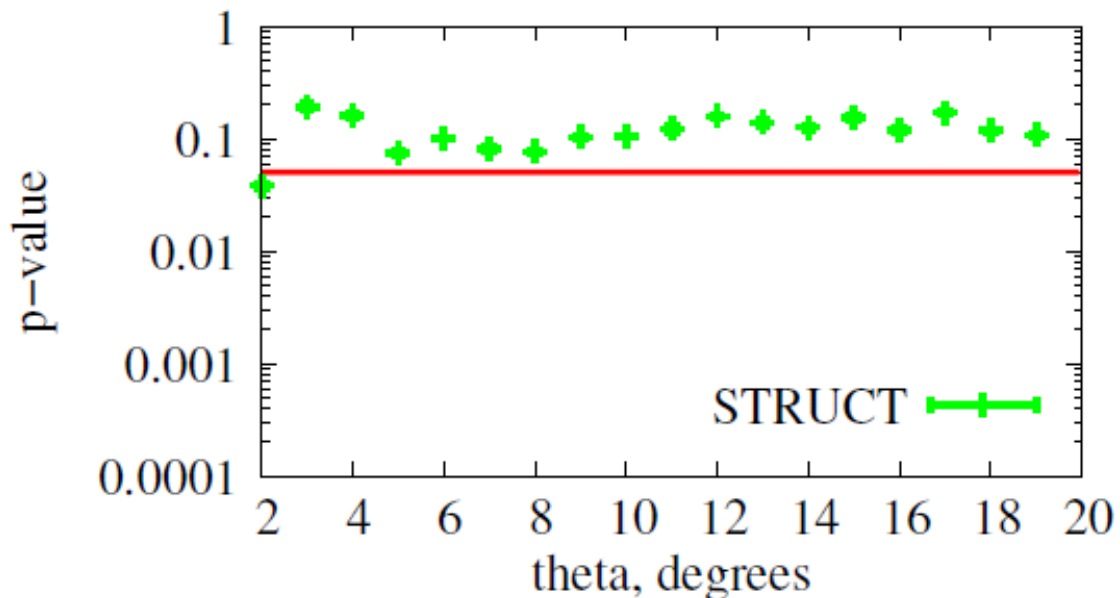
Anisotropy (LSS)

Composition measured by TA
is consistent with proton.

Data (white point)
and Model distribution
(shaded region) from LSS
within 100Mpc.



When deflection angle θ
set to 6° , p-value of
isotropy is 0.001.



Physics Summary

Significance of the five-year
SD spectrum suppression is 5.7σ .

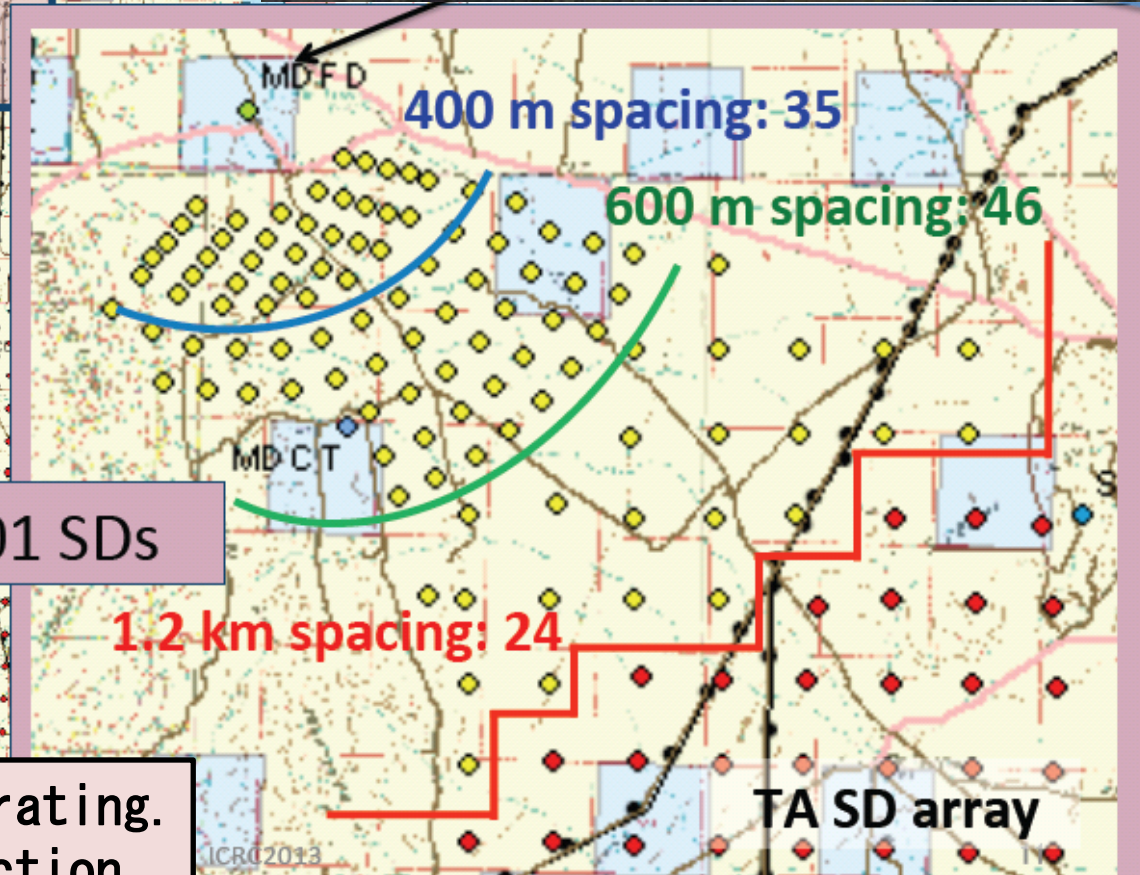
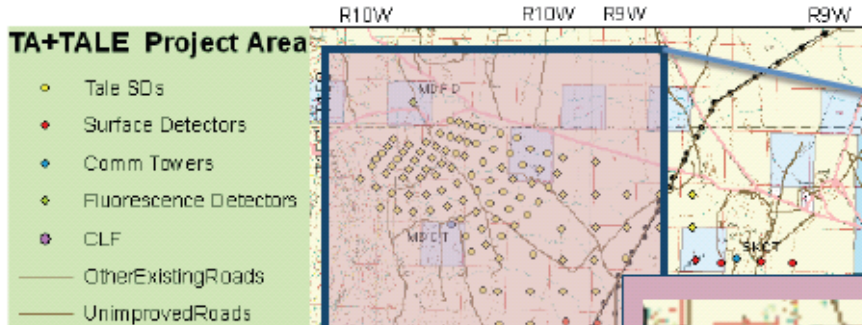
Mass composition is consistent with proton.

Arrival direction at high energy seems
inconsistent with isotropic distribution.

To Extended and Affiliate Experiment

TA Low Energy Extension (TALE)

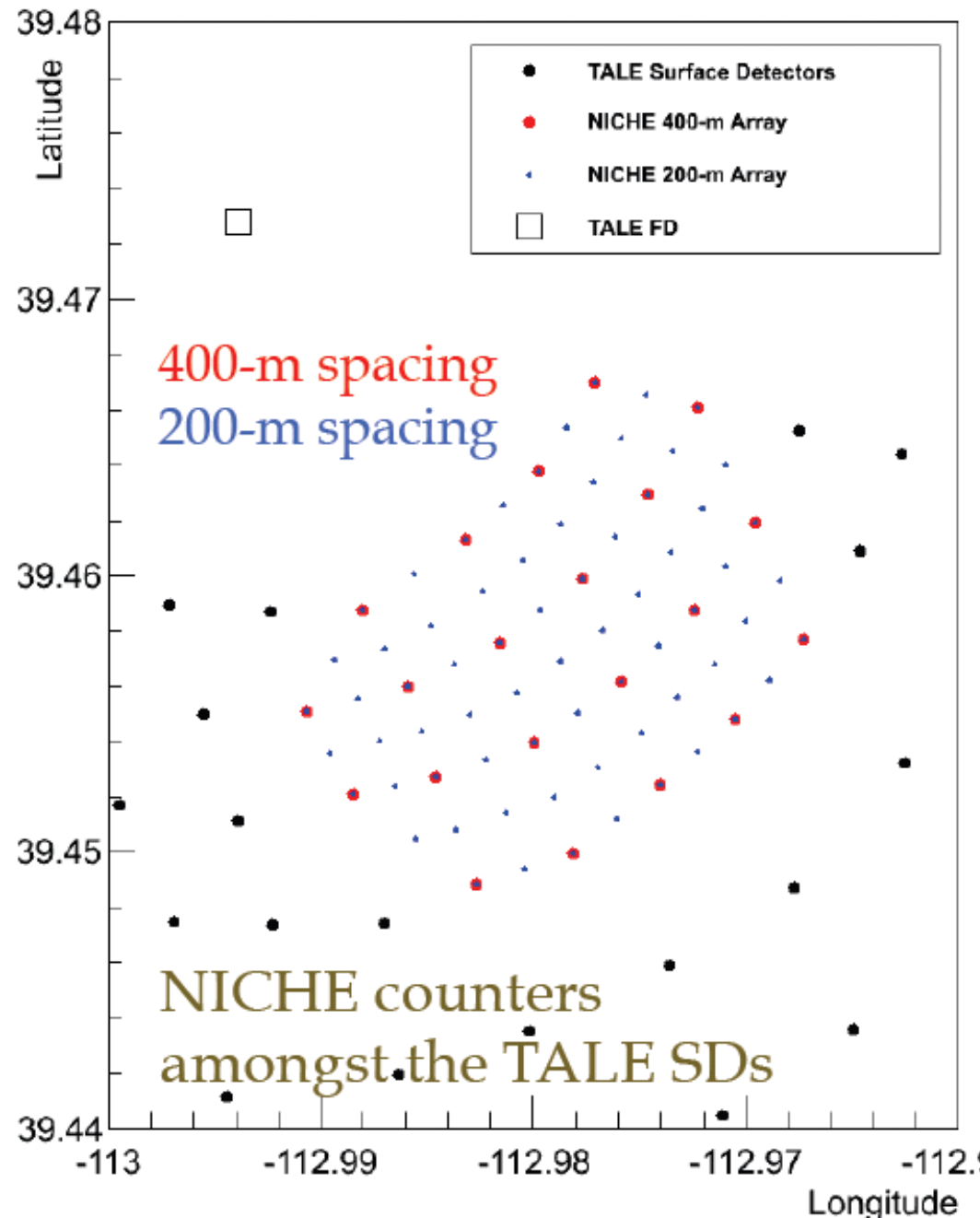
Basically it uses same method as TA, relatively specific for hybrid observation. The observing energy range is down to 10^{17} eV (LHC).



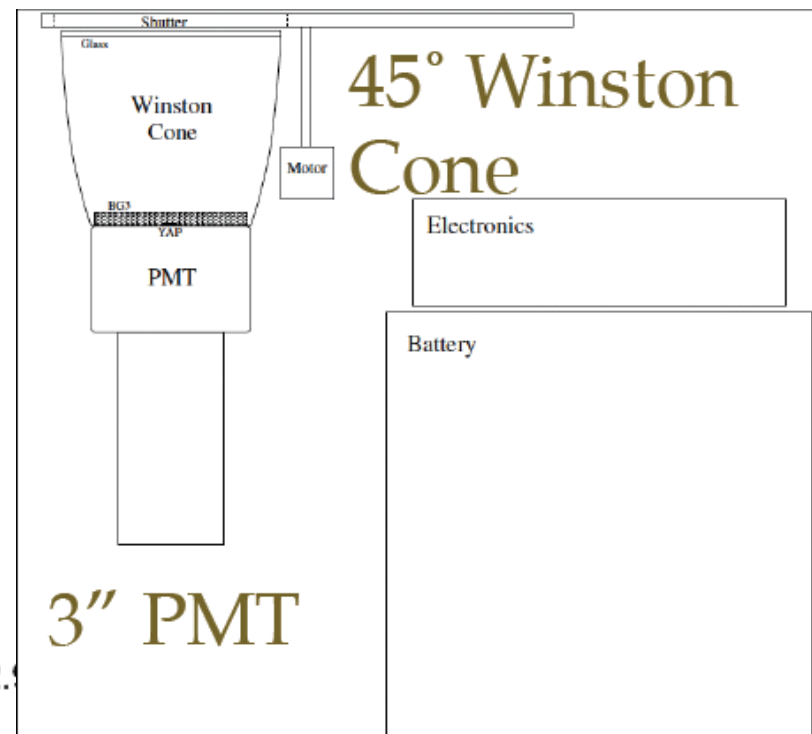
Additionally install 101 SDs

Partially installed and operating.
Now, proposing full construction.

Non-Imaging ChErenkov array (NICHE)



This is the plan of low energy extension for TALE, whose energy range down to 10^{16} eV for 200m spacing. The main purpose is composition study by Cherenkov pulse-width, working with TALE.

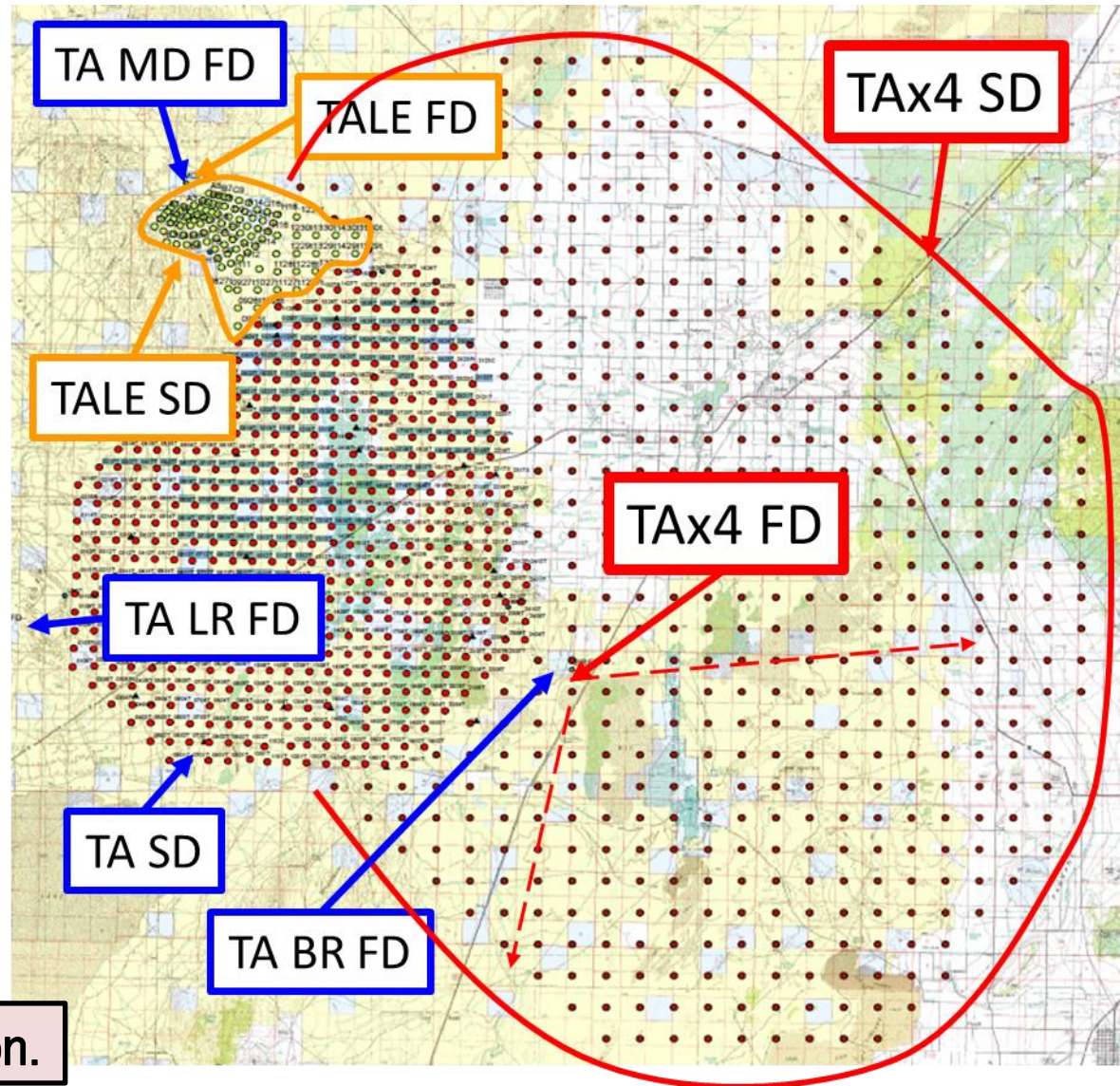


TA High Energy Extension (TAx4)

Basically it uses same method as TA, with a FD and 3 times larger coverage area by SDs. SD full area can not be covered by FD field of view.

This experiment is mainly for anisotropy study,

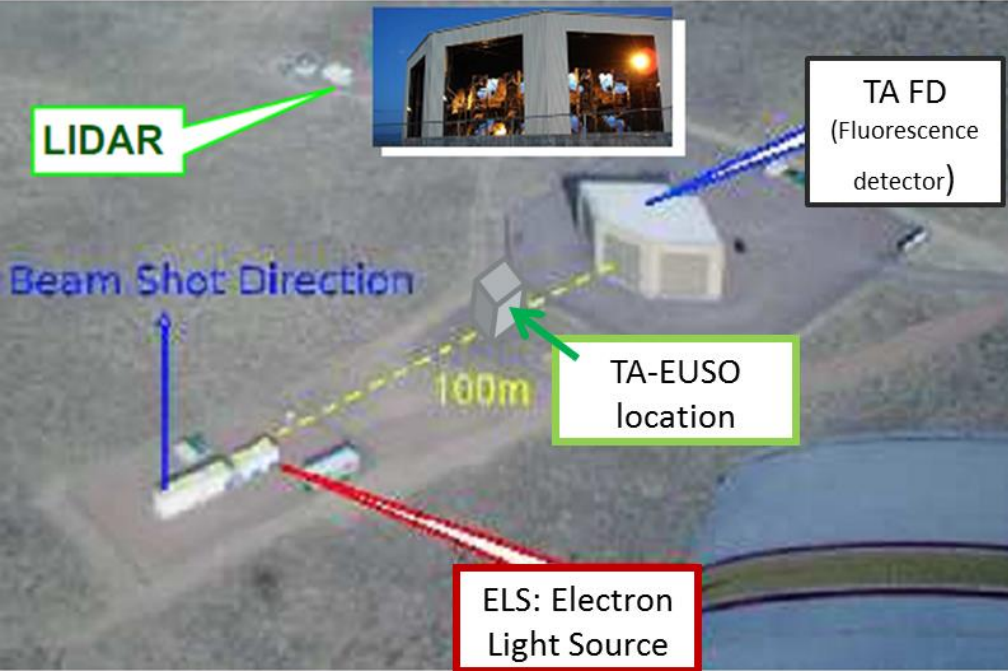
This is for detecting more cosmic rays which have energy over GZK. If they are proton, their source is within 50Mpc. From our current result of the inconsistency with isotropy at high energy region, we are halfway to catching a clue?



Now, proposing construction.

TA Affiliate Experiment

Prototype EUSO at TA

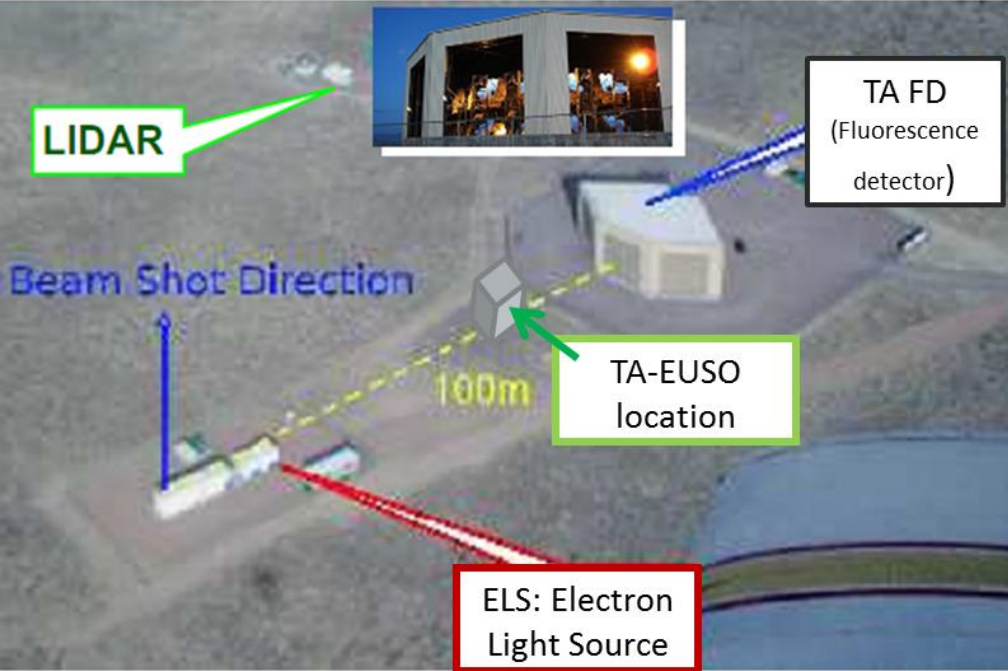


TARA



TA Affiliate Experiment

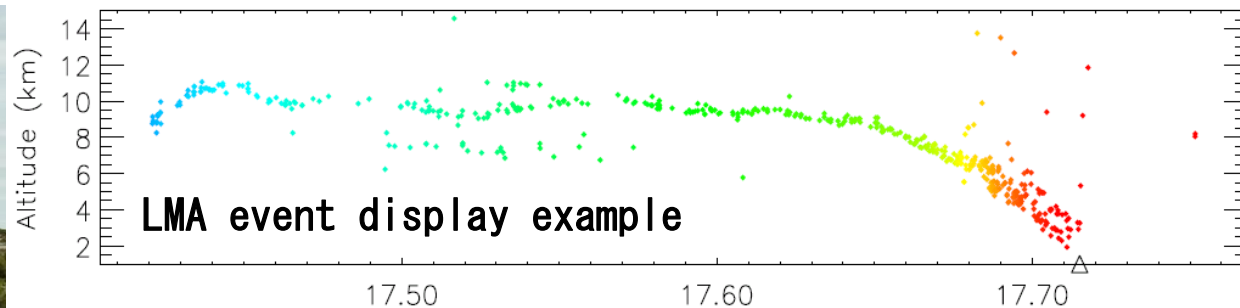
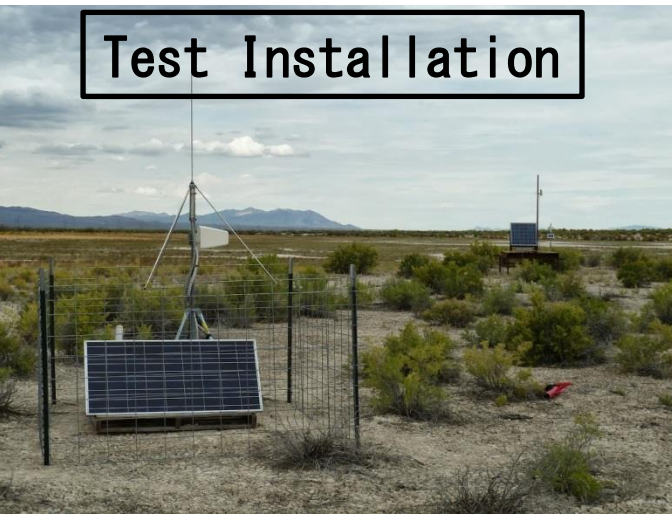
Prototype EUSO at TA



TARA



TA Lightning Mapping Array (TALMA)



LMA can measure the origin of sferics in the lightning process with high resolution of time and position.

Why TA try to install LMA?

Because T ASD observed the bursts of shower like events,
which are correlated with lightning.

From five years T ASD data,
there are 10 bursts of
shower trigger. The burst
criteria is 3 shower trigger
in 1 millisecond. The chance
expectation for five years
is less than 10^{-4} .

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Some events in 5 bursts pass
the reconstruction.

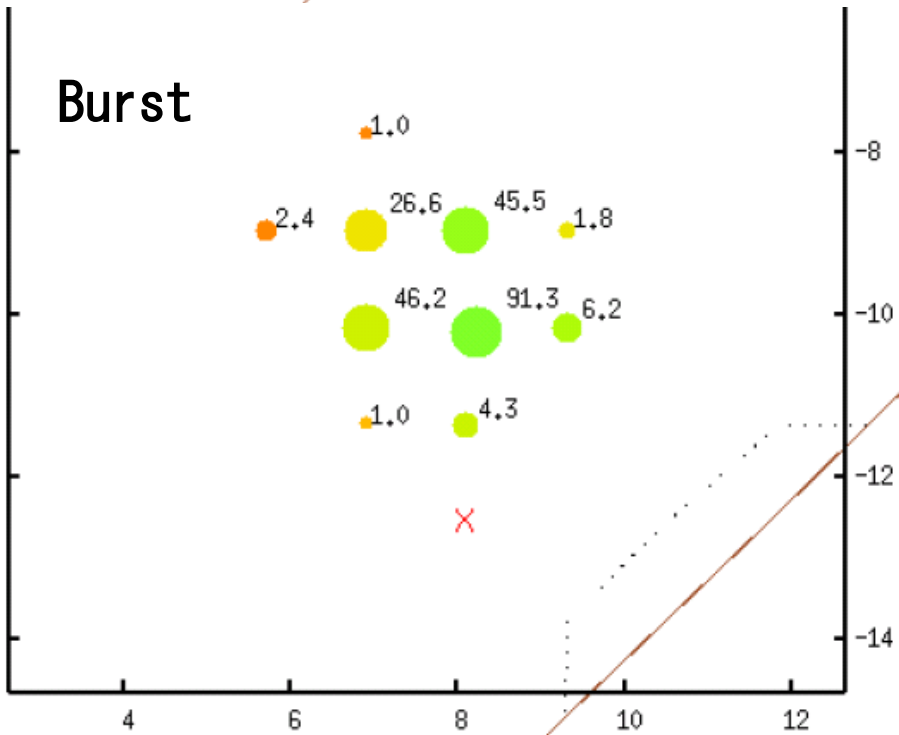
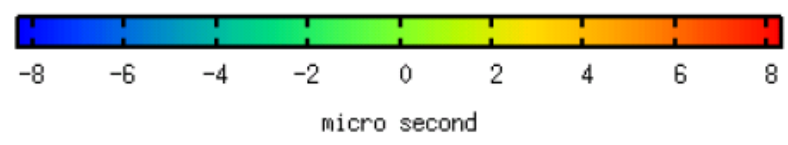
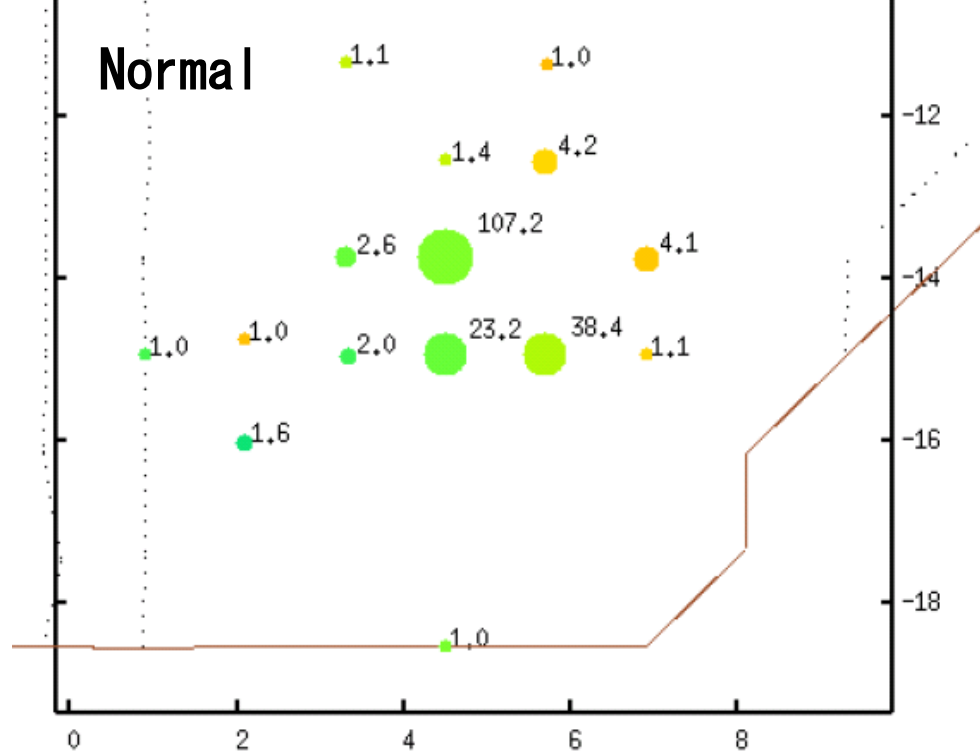
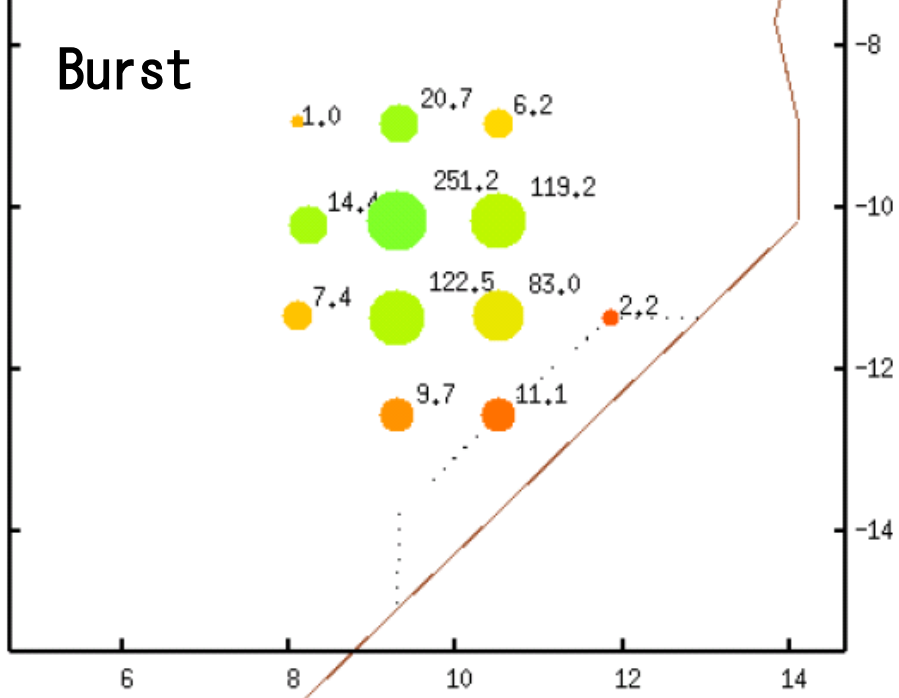
	date	time	usec	X[m]	Y[m]
AS	101004	165842	930565	11356	-7425
AS	101004	165842	930612	10478	-7368
AS	101004	165842	930835	11142	-8159
AS	110727	080615	124319	3447	1952
AS	110727	080615	124543	2897	2232
AS	110916	194056	567481	-3210	-9285
AS	110916	194056	567566	-3524	-9413
AS	120706	014911	184219	9847	-10702
AS	120706	014911	184307	7635	-9674
AS	120907	015545	380684	-8636	1254
AS	120907	015545	380755	-9857	-337
AS	120907	015545	380881	-9450	-961

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From five years T ASD data, there are 10 bursts of shower trigger. The burst criteria is 3 shower trigger in 1 millisecond. The chance expectation for five years is less than 10^{-4} .

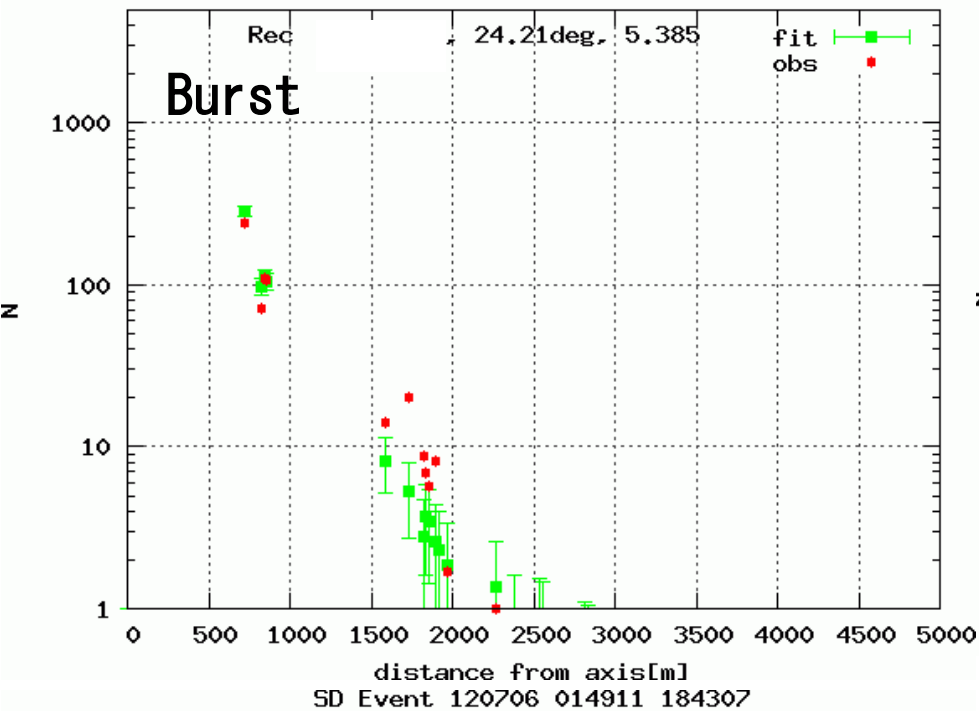
Some events in 5 bursts pass the reconstruction. These bursts are selected only by timing, but their core position is very close. In addition, these reconstructed shower front is very curved compared with normal shower. They seems start shower development at low sky.

	date	time	usec	X[m]	Y[m]
AS	101004	165842	930565	11356	-7425
AS	101004	165842	930612	10478	-7368
AS	101004	165842	930835	11142	-8159
AS	110727	080615	124319	3447	1952
AS	110727	080615	124543	2897	2232
AS	110916	194056	567481	-3210	-9285
AS	110916	194056	567566	-3524	-9413
AS	120706	014911	184219	9847	-10702
AS	120706	014911	184307	7635	-9674
AS	120907	015545	380684	-8636	1254
AS	120907	015545	380755	-9857	-337
AS	120907	015545	380881	-9450	-961

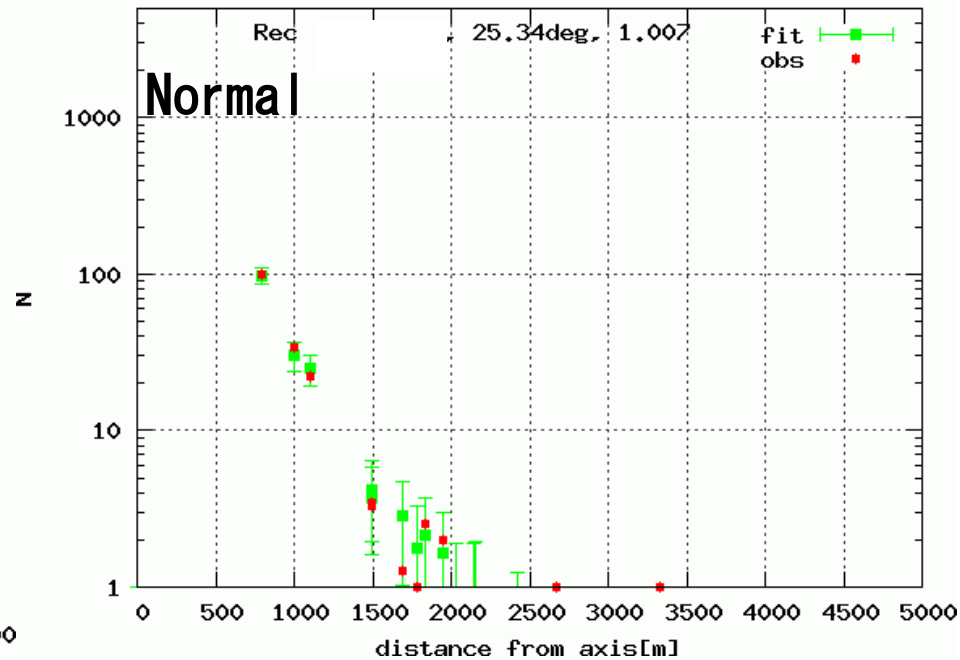
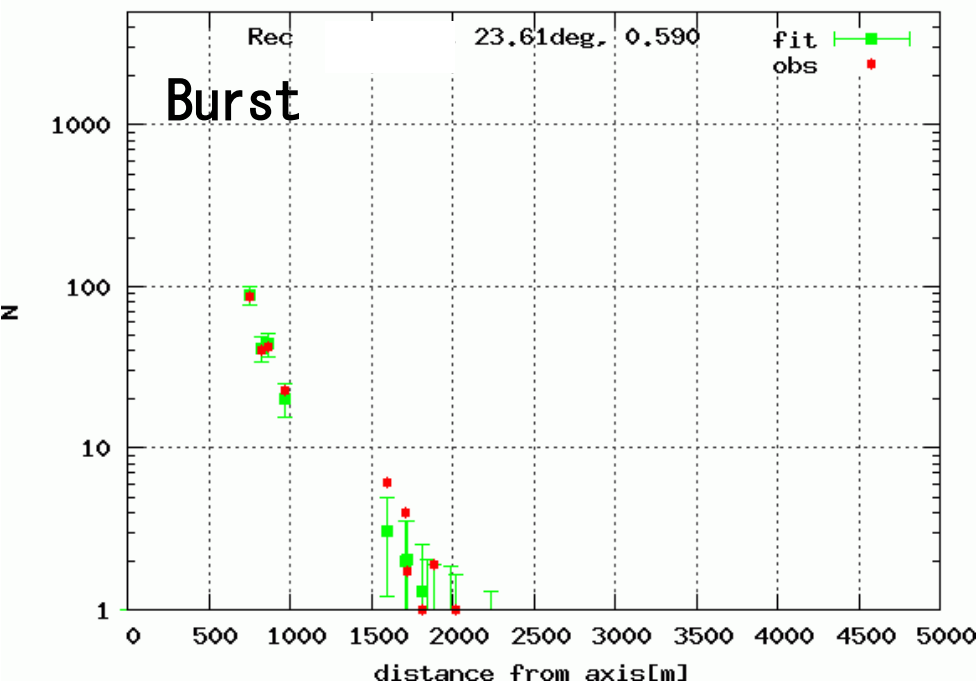


Particle Hit Map

Left plots are burst events. Right plot is a typical (normal) event whose zenith angle is almost same as left two events and its shower size is middle of left two events.

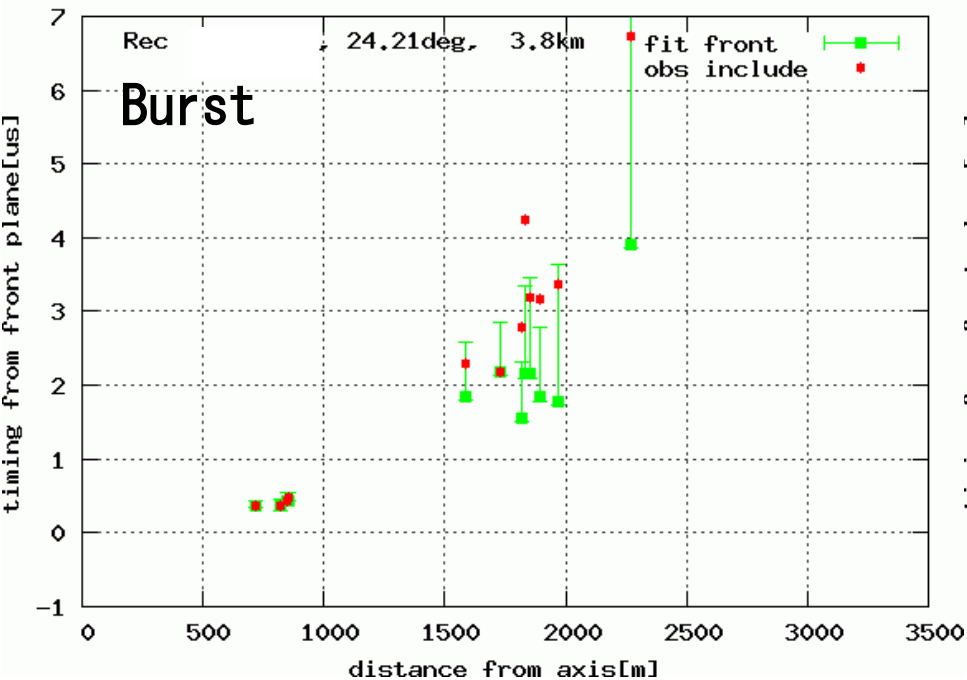


SD Event 120706 014911 184307

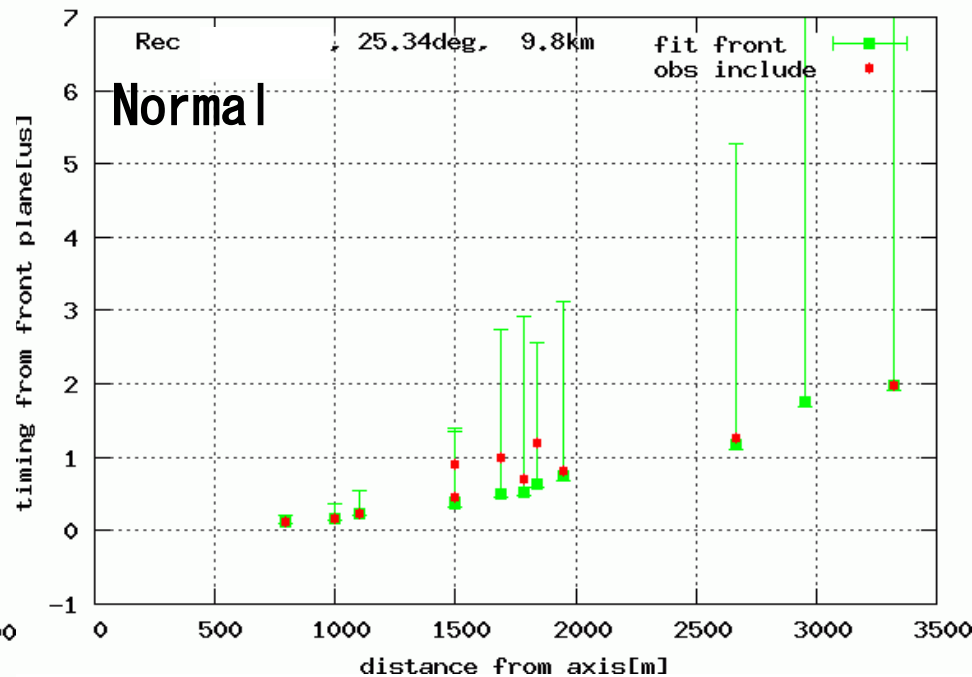
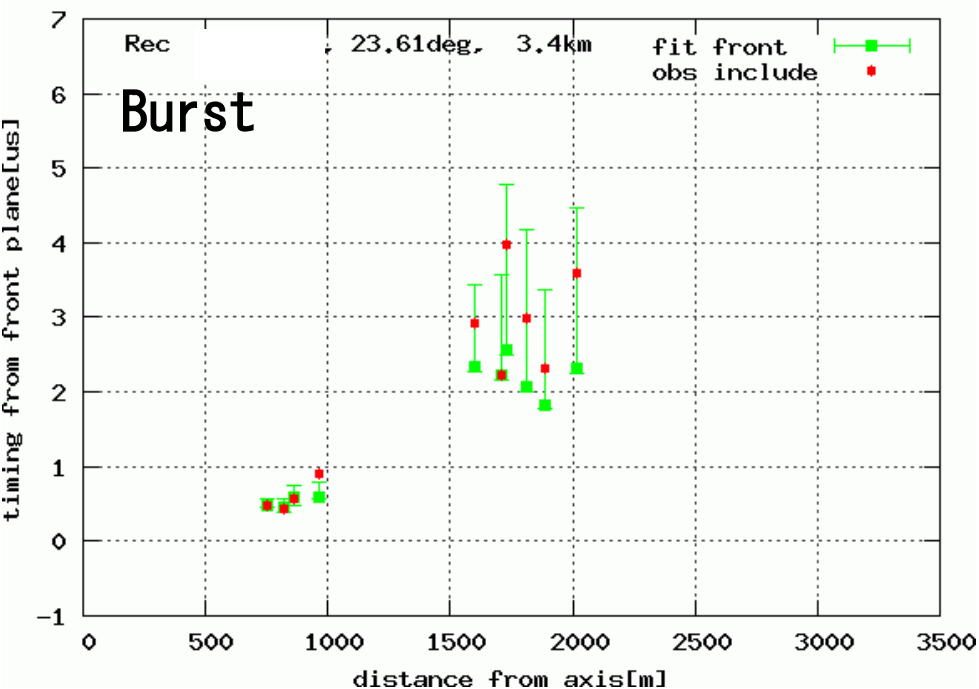


Lateral Number of Particle

Left plots are burst events.
 Right plot is a typical (normal)
 event whose zenith angle is
 almost same as left two events
 and its shower size is middle of
 left two events.



SD Event 120706 014911 184307

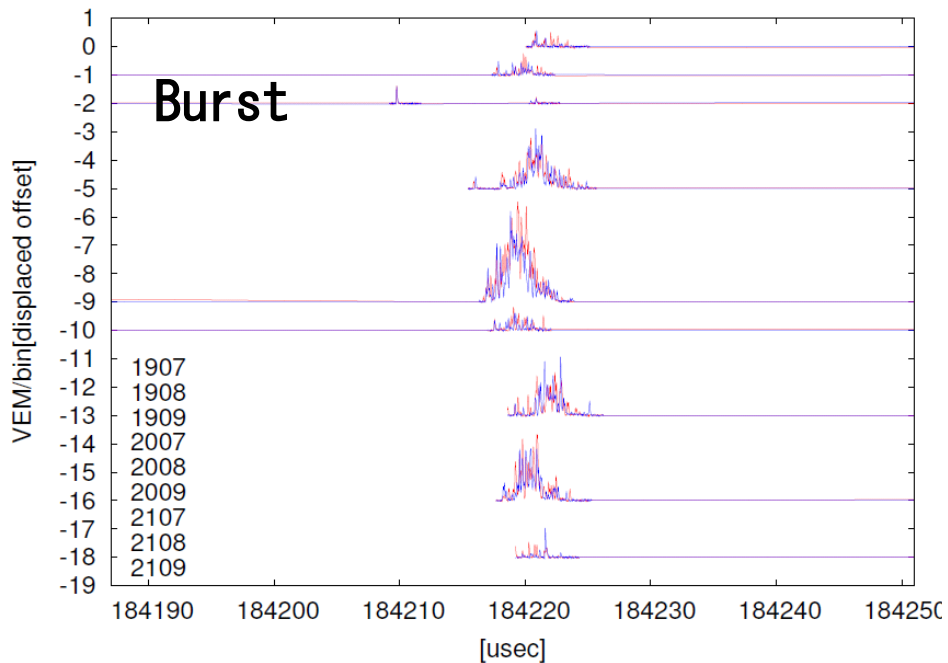


Lateral Arrival Timing

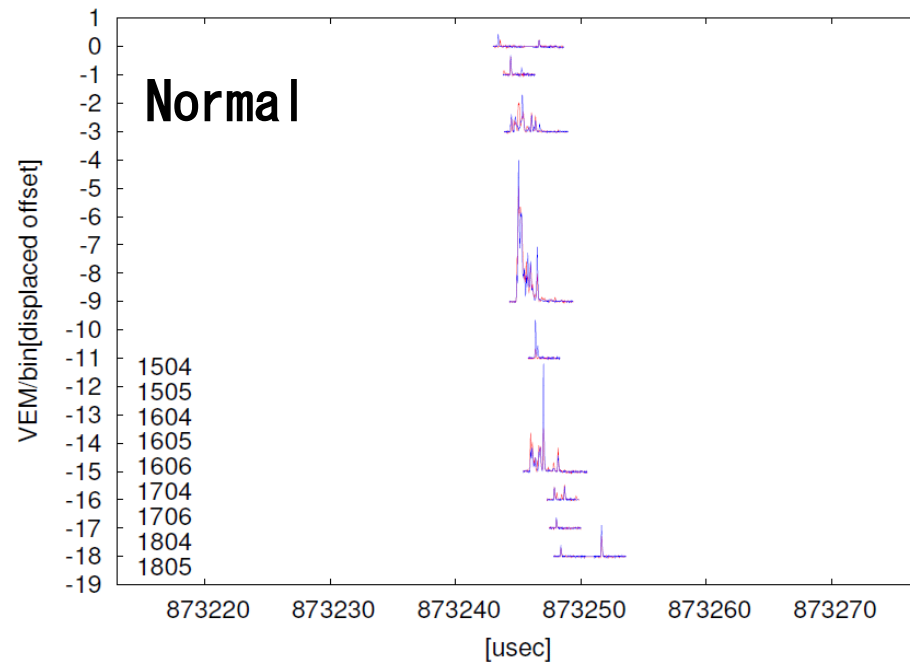
Left plots are burst events. Right plot is a typical (normal) event whose zenith angle is almost same as left two events and its shower size is middle of left two events.

Left plots' shower front is much more curved than Right plot.

wf184219



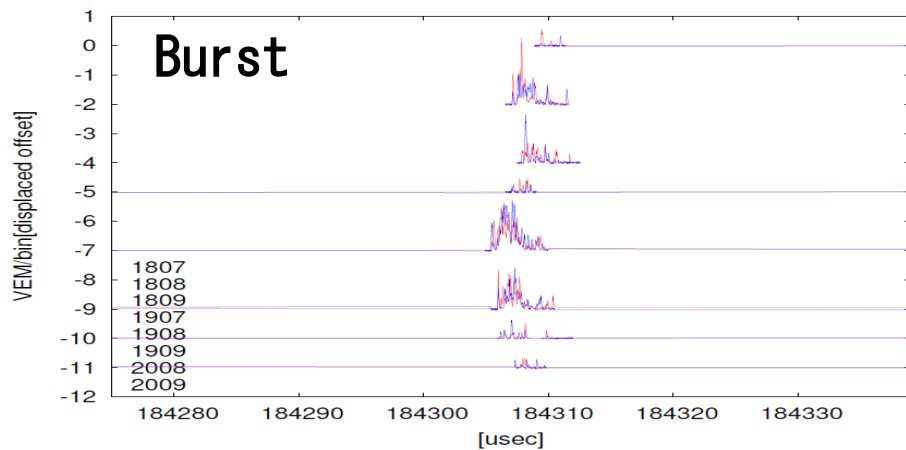
wf873245



Waveforms

Left plots are burst events.
Right plot is a typical (normal) event whose zenith angle is almost same as left two events and its shower size is middle of left two events.

wf184307



Left plots' waveforms have less sharp rising edge than Right plot.

Correlation with Lightning Data from NLDN.

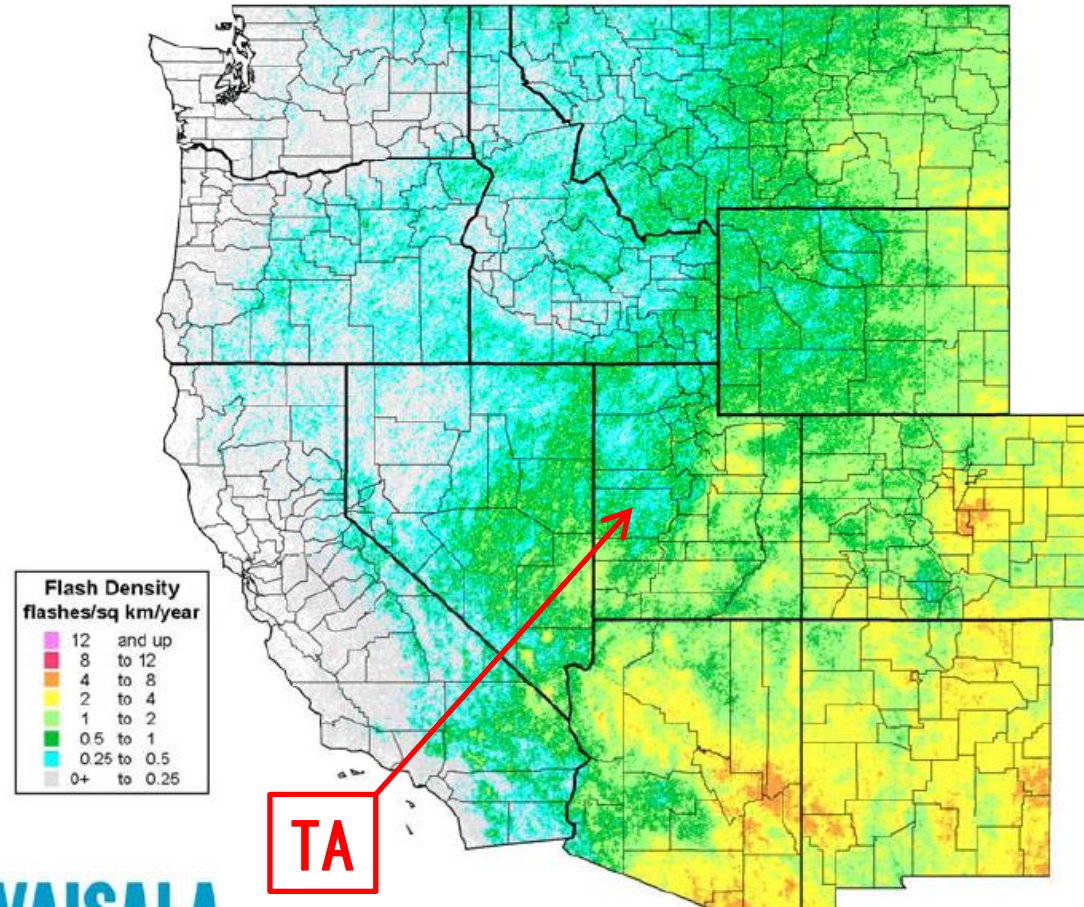
National Lightning Detection Network detects lightning by VLF. The lightning list contains lightning time, 2D coordinates, peak current and flag which indicates whether intracloud lightning or cloud-ground lightning.

NLDN is somewhat inefficient for intracloud lightning.

I defined “synchronized” lightning is the time difference within ± 1 [ms] from burst.

I defined “related” lightning is the time difference within ± 200 [ms] from burst except for synchronized lightning.

National Lightning Detection Network
2005 - 2012



AS	date	time	usec	X[m]	Y[m]	H[m]
LG	date	time	usec	X[m]	Y[m]	Flag
AS	101004	165842	930565	11356	-7425	3963
AS	101004	165842	930612	10478	-7368	4400
AS	101004	165842	930835	11142	-8159	3270
LG	101004	165842	930608	12480	-5068	C
LG	101004	165842	934058	10619	-8069	G
AS	110727	080615	124319	3447	1952	4070
AS	110727	080615	124543	2897	2232	3070
LG	110727	080615	124303	3653	2285	C
LG	110727	080615	130887	3084	1996	G
AS	110916	194056	567481	-3210	-9285	3253
AS	110916	194056	567566	-3524	-9413	3134
AS	120706	014911	184219	9847	-10702	3770
AS	120706	014911	184307	7635	-9674	3361
LG	120706	014911	184122	8997	-9670	C
AS	120907	015545	380684	-8636	1254	4446
AS	120907	015545	380755	-9857	-337	4805
AS	120907	015545	380881	-9450	-961	3361
LG	120907	015545	380675	-8942	668	C
LG	120907	015545	390411	-9635	-1952	G
LG	120907	015545	409370	-8608	-1653	G

Correlation with Lightning Data from NLDN.

There are 4 correlations in 5 reconstructed bursts.

This is good correlation with considering NLDN detection efficiency.

AS line shows T ASD event.

LG line shows NLDN event.

AS	date	time	usec	X[m]	Y[m]	H[m]
LG	date	time	usec	X[m]	Y[m]	Flag
AS	101004	165842	930565	11356	-7425	3963
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LG	120706	014911	184122	8997	-9670	C
AS	120907	015545	380684	-8636	1254	4446
AS	120907	015545	380755	-9857	-337	4805
AS	120907	015545	380881	-9450	-961	3361
LG	120907	015545	380675	-8942	668	C
LG	120907	015545	390411	-9635	-1952	G
LG	120907	015545	409370	-8608	-1653	G

Correlation with Lightning Data from NLDN.

There are 4 correlations in 5 reconstructed bursts.

This is good correlation with considering NLDN detection efficiency.

AS line shows T ASD event.

LG line shows NLDN event.

These correlations are taken only by timing, but the locations of bursts and lightning are also correlated very well.

AS	date	time	usec	X[m]	Y[m]	H[m]
LG	date	time	usec	X[m]	Y[m]	Flag
AS	101004	165842	930565	11356	-7425	3963
AS	101004	165842	930612	10478	-7368	4400
AS	101004	165842	930835	11142	-8159	3270
LG	101004	165842	930608	12480	-5068	C
LG	101004	165842	934058	10619	-8069	G
AS	110727	080615	124319	3447	1952	4070
AS	110727	080615	124543	2897	2232	3070
LG	110727	080615	124303	3653	2285	C
LG	110727	080615	130887	3084	1996	G
AS	110916	194056	567481	-3210	-9285	3253
AS	110916	194056	567566	-3524	-9413	3134
AS	120706	014911	184219	9847	-10702	3770
AS	120706	014911	184307	7635	-9674	3361
LG	120706	014911	184122	8997	-9670	C
AS	120907	015545	380684	-8636	1254	4446
AS	120907	015545	380755	-9857	-337	4805
AS	120907	015545	380881	-9450	-961	3361
LG	120907	015545	380675	-8942	668	C
LG	120907	015545	390411	-9635	-1952	G
LG	120907	015545	409370	-8608	-1653	G

Correlation with Lightning Data from NLDN.

There are 4 correlations in 5 reconstructed bursts.

This is good correlation with considering NLDN detection efficiency.

AS line shows T ASD event.

LG line shows NLDN event.

79% of the NLDN lightning data have flag “G(cloud-ground)”. However, all **synchronized** 4 lightning strokes have flag “C(intracloud)”. And all **related** subsequent strokes have flag “G(cloud-ground)”.

High Energy Radiation Shower Geometry with Lightning

Color shows the height from ground up to 3000m.

Red point shows shower core hit position.

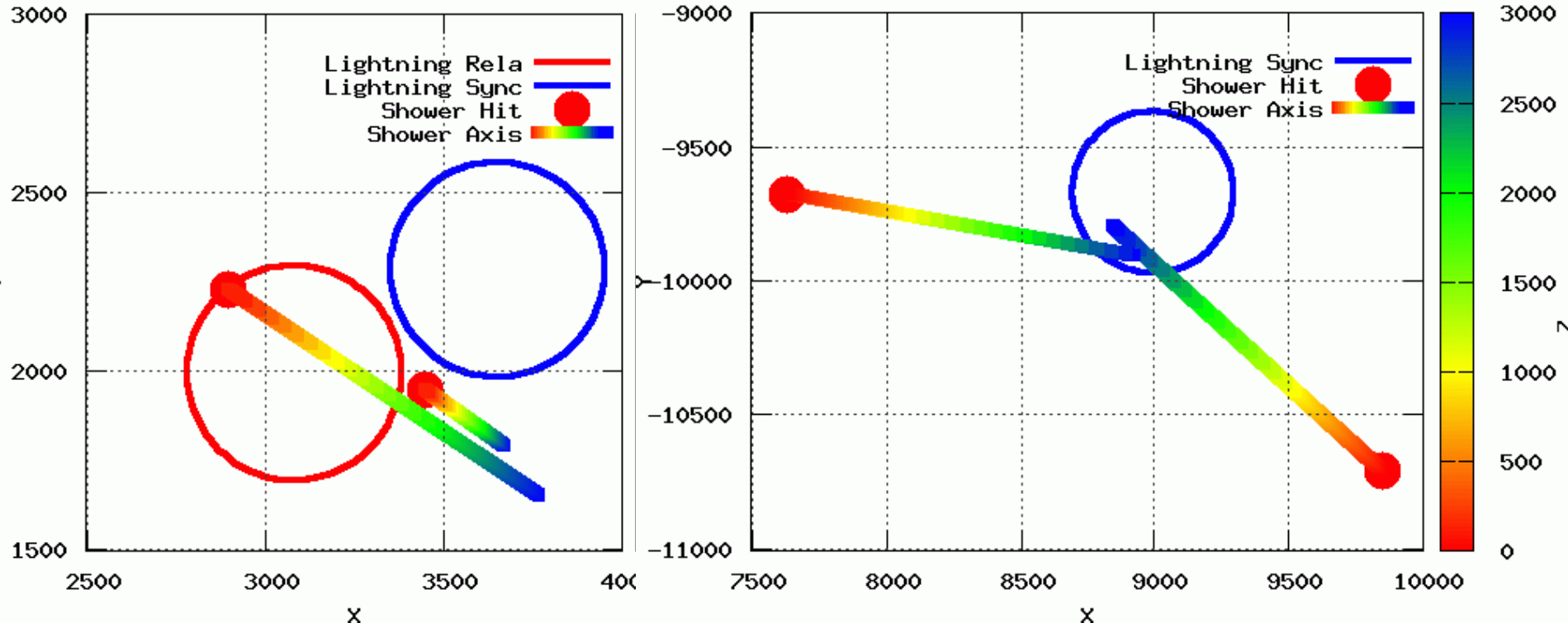
Circle shows lightning position with 300m radius.

I set “Synchronized(intracloud)” lightning at 3000m,

and “Related(cloud-ground)” lightning at 0m.

event 080615

event 014911



We are planing to install our Lightning Mapping Array (LMA) to study more detail correlation between high energy radiation and lightning.

We need fund for TALMA.

Summary

Some Extended Experiments
and
Some Affiliate Experiments
are planned and going on.

End

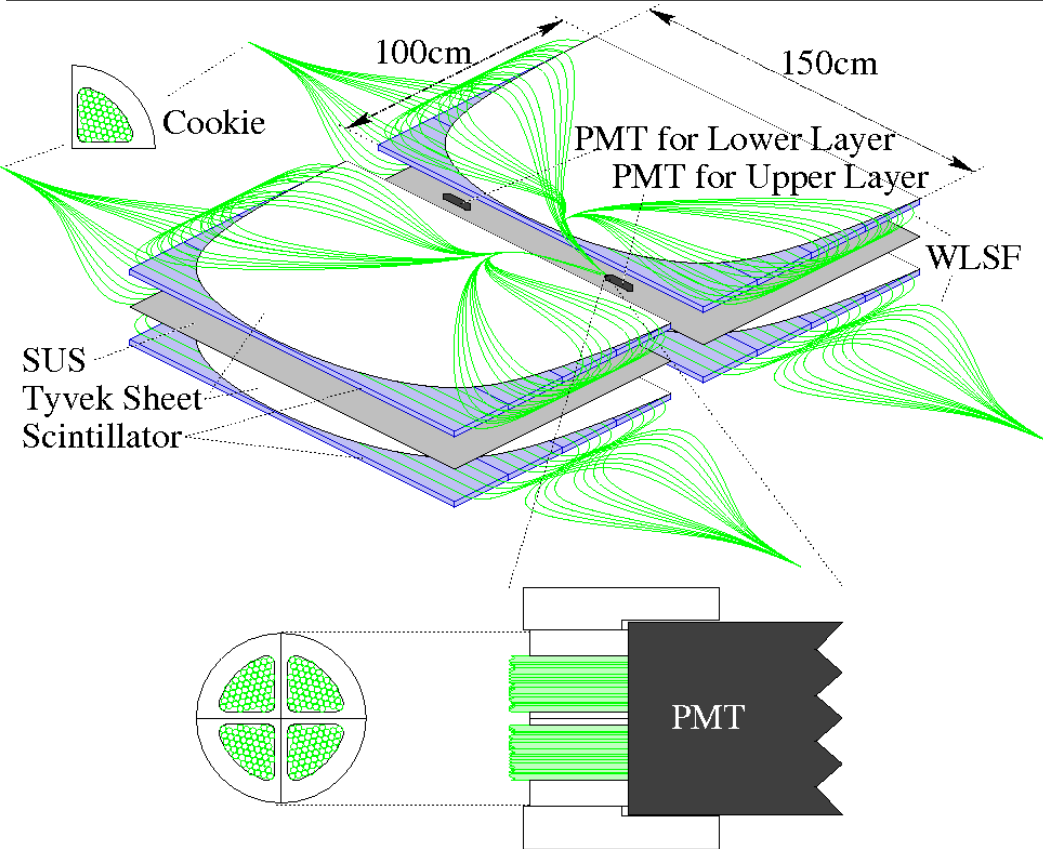
Thank You

Backup

Surface Detector

Particle Detection Part:

2 layers of $3\text{m}^2 \times 1.2\text{cm}$ plastic scintillators.
104 Wave Length Shifting Fibers for each layer.

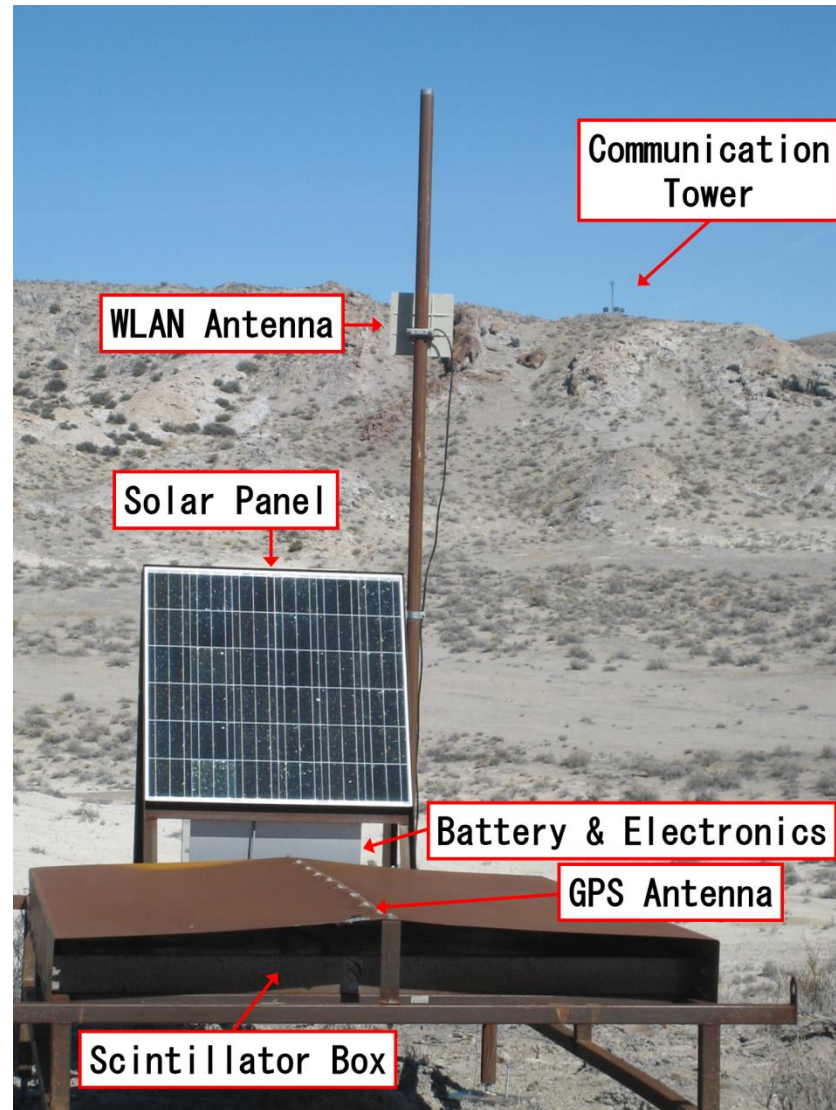


Electronics:

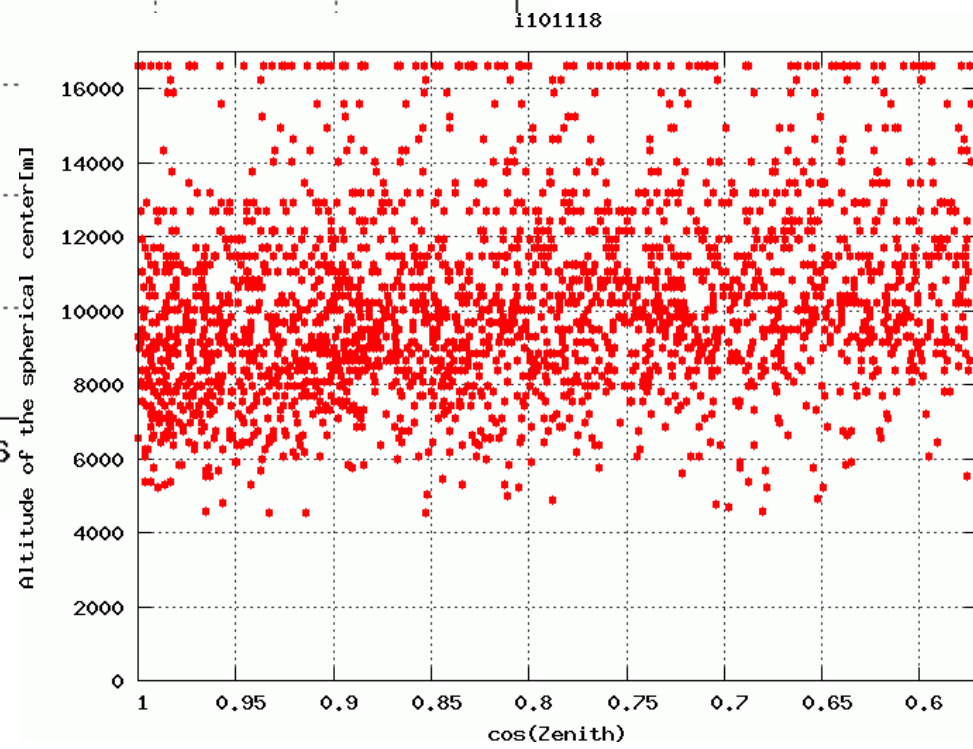
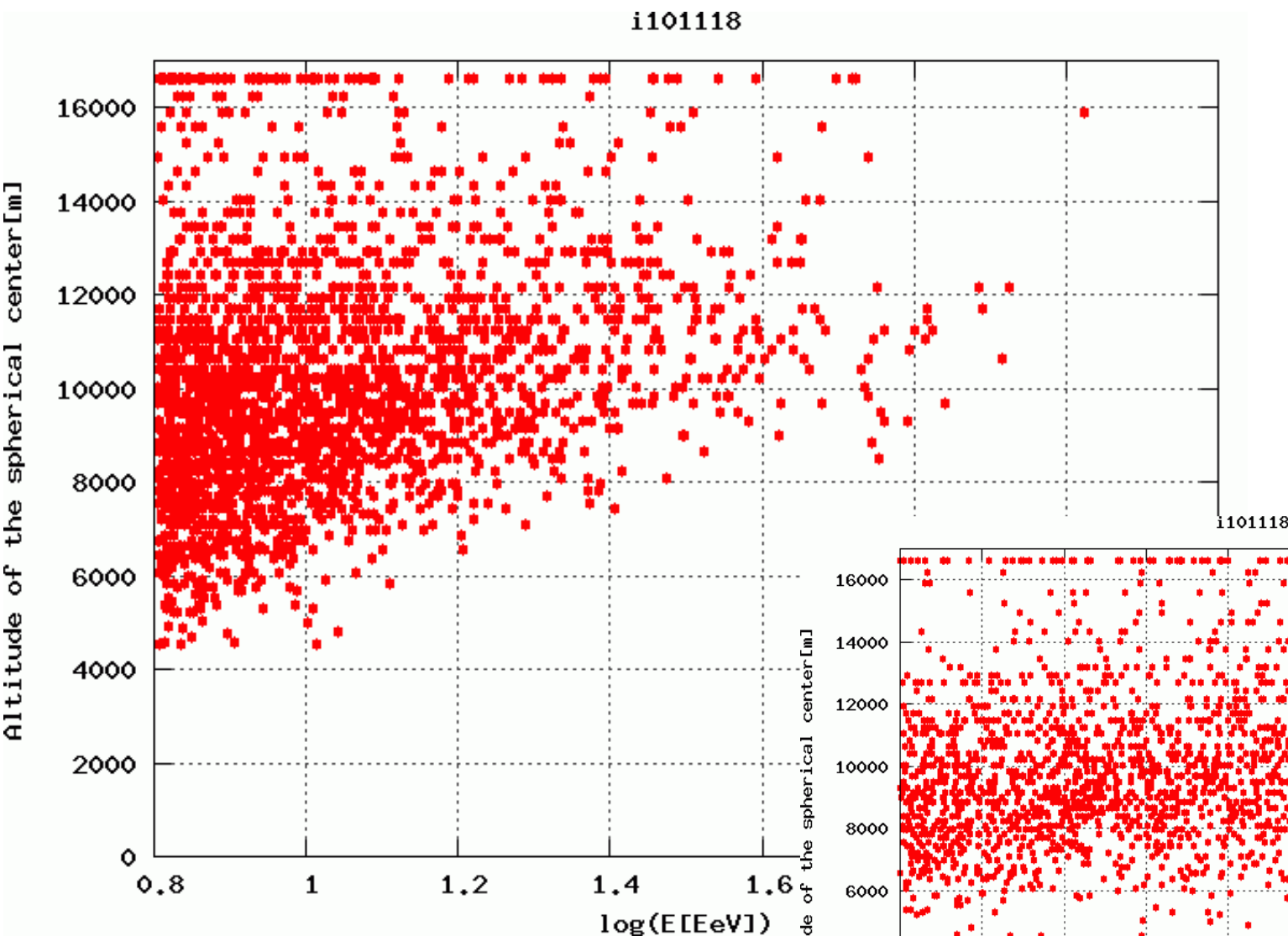
50MHz FADC
Charge Controller
GPS
Wireless LAN

Data Acquisition:

Shower Trigger is generated by adjacent 3SDs whose signal is over 3MIPs in $8\mu\text{s}$.
Waveforms for all SD within $32\mu\text{s}$ from shower trigger time.



Scatter plot of H with Energy or Zenith (080511–100907)



Events ($E > 6.36 \text{ EeV}$ & $\text{Zen} < 55 \text{ deg}$)
with H (the altitude of the
spherical center for shower front)
less than 6000 are very few.

