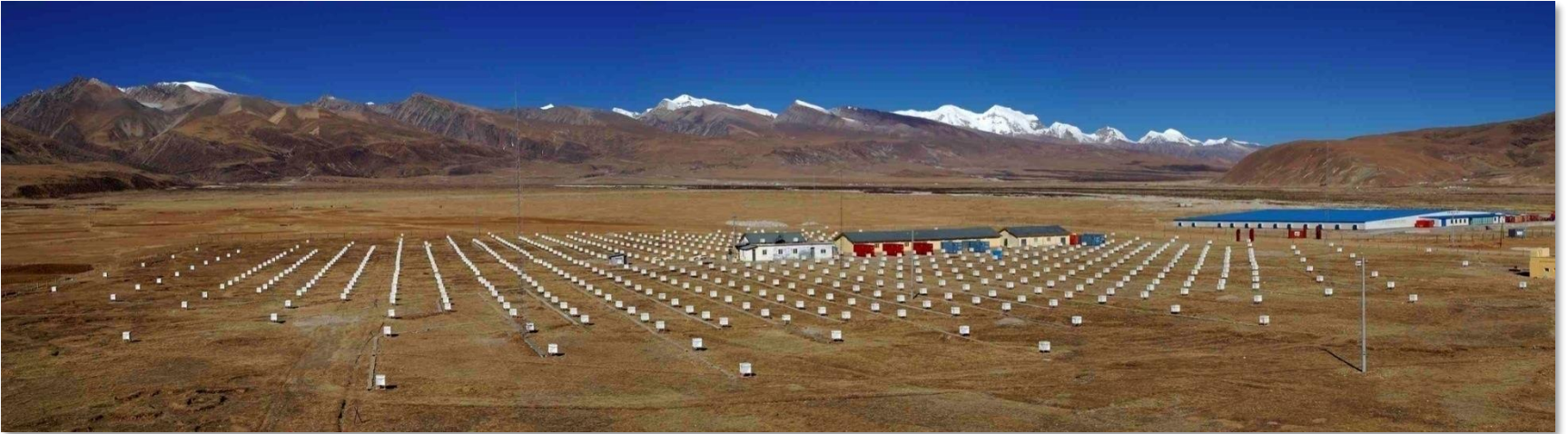




Probing the Universe from the Tibetan Highland

Masato TAKITA (ICRR, Univ. of Tokyo)
For the Tibet AS γ Collaboration

@Utokyo New York Event, February 12, 2022
Online



Outline

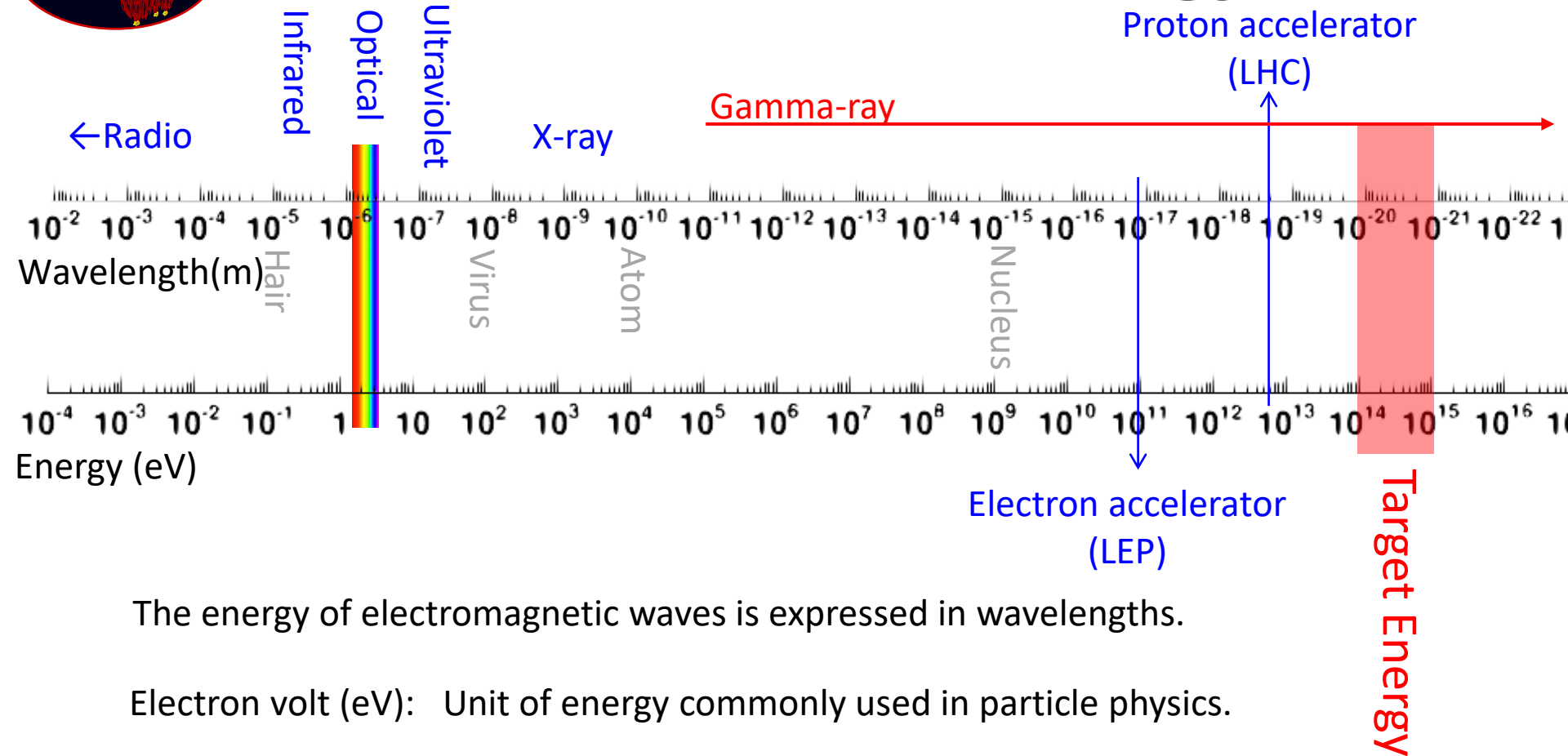
- Tibet AS γ Experiment (Multi-purpose)
- Gamma-ray Observation for cosmic-ray origin
- Sun Shadow Observation for solar magnetic field

Skipped in this talk are:

- Chemical composition in cosmic rays
- Tests of hadronic interaction models
- Cosmic-ray anisotropy measurement
- Summary



Observational Energy



The energy of electromagnetic waves is expressed in wavelengths.

Electron volt (eV): Unit of energy commonly used in particle physics.

$10^{15} \text{ eV} = 1 \text{ Peta electron volt (PeV)}$

→ 1000 trillion times more energetic than visible light!

Tibet AS_γ Experiment



The Tibet AS γ Collaboration



Roughly 100 researchers

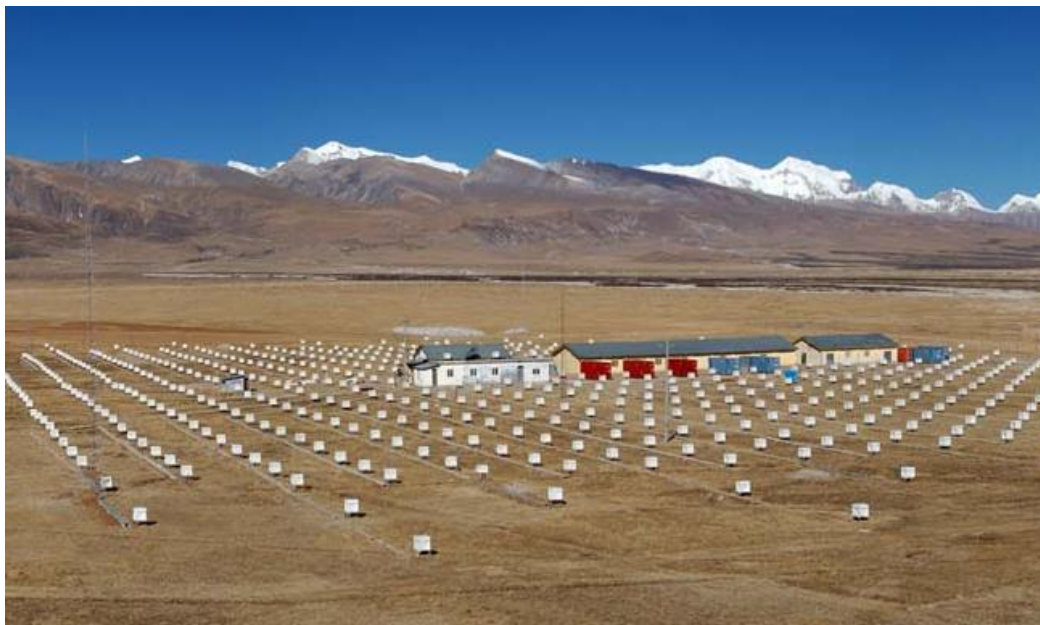
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(1) Department of Physics, Hirosaki University, Hirosaki 036-8561, Japan (2) Department of Physics, Shinshu University, Matsumoto 390-8621, Japan (3) School of Astronomy and Space Science, Nanjing University, Nanjing 210093, China (4) Key Laboratory of Particle Astrophysics, Institute of High Energy Physics, Chinese Academy of Sciences, Beijing 100049, China (5) National Astronomical Observatories, Chinese Academy of Sciences, Beijing 100012, China (6) Department of Mathematics and Physics, Tibet University, Lhasa 850000, China (7) Department of Physics, Hebei Normal University, Shijiazhuang 050016, China (8) University of Chinese Academy of Sciences, Beijing 100049, China (9) Institute of Frontier and Interdisciplinary Science and Key Laboratory of Particle Physics and Particle Irradiation (MOE), Shandong University, Qingdao 266237, China (10) Institute of Modern Physics, SouthWest Jiaotong University, Chengdu 610031, China (11) Faculty of Engineering, Yokohama National University, Yokohama 240-8501, Japan (12) Faculty of Engineering, Kanagawa University, Yokohama 221-8686, Japan (13) Faculty of Education, Utsunomiya University, Utsunomiya 321-8505, Japan (14) Faculty of Systems Engineering, Shibaura Institute of Technology, Omiya 330-8570, Japan (15) Institute for Cosmic Ray Research, University of Tokyo, Kashiwa 277-8582, Japan (16) Institute of Space and Astronautical Science, Japan Aerospace Exploration Agency (ISAS/JAXA), Sagami-hara 252-5210, Japan (17) National Center for Space Weather, China Meteorological Administration, Beijing 100081, China (18) School of Information Science and Engineering, Shandong Agriculture University, Taian 271018, China (19) Department of Astronomy, School of Physical Sciences, University of Science and Technology of China, Hefei, Anhui 230026, China (20) Department of Physics and Tsinghua Centre for Astrophysics (THCA), Tsinghua University, Beijing 100084, China (21) Tsinghua University-National Astronomical Observatories of China (NAOC) Joint Research Center for Astrophysics, Tsinghua University, Beijing 100084, China (22) Department of Astronomy, Tsinghua University, Beijing 100084, China (23) College of Industrial Technology, Nihon University, Narashino 275-8575, Japan (24) National Institute of Informatics, Tokyo 101-8430, Japan (25) National Institute of Information and Communications Technology, Tokyo 184-8795, Japan (26) Department of Mechanical and Electrical Engineering, Shangdong Management University, Jinan 250357, China (27) College of Science, China University of Petroleum, Qingdao 266555, China (28) Tokyo Metropolitan College of Industrial Technology, Tokyo 116-8523, Japan (29) Department of Physics, Konan University, Kobe 658-8501, Japan (30) Shonan Institute of Technology, Fujisawa 251-8511, Japan (31) Research Institute for Science and Engineering, Waseda University, Tokyo 169-8555, Japan (32) Japan Atomic Energy Agency, Tokai-mura 319-1195, Japan (33) Key Laboratory of Dark Matter and Space Astronomy, Purple Mountain Observatory, Chinese Academy of Sciences, Nanjing 210034, China



Tibet AS γ Experiment

Yangbajing, Tibet, China at the altitude of 4300m



**Surface air shower array 65,700m²
scintillation detectors**

Observation regardless day and night
with wide field of view.



**Underground muon detectors
3400m² (Before filling water)**

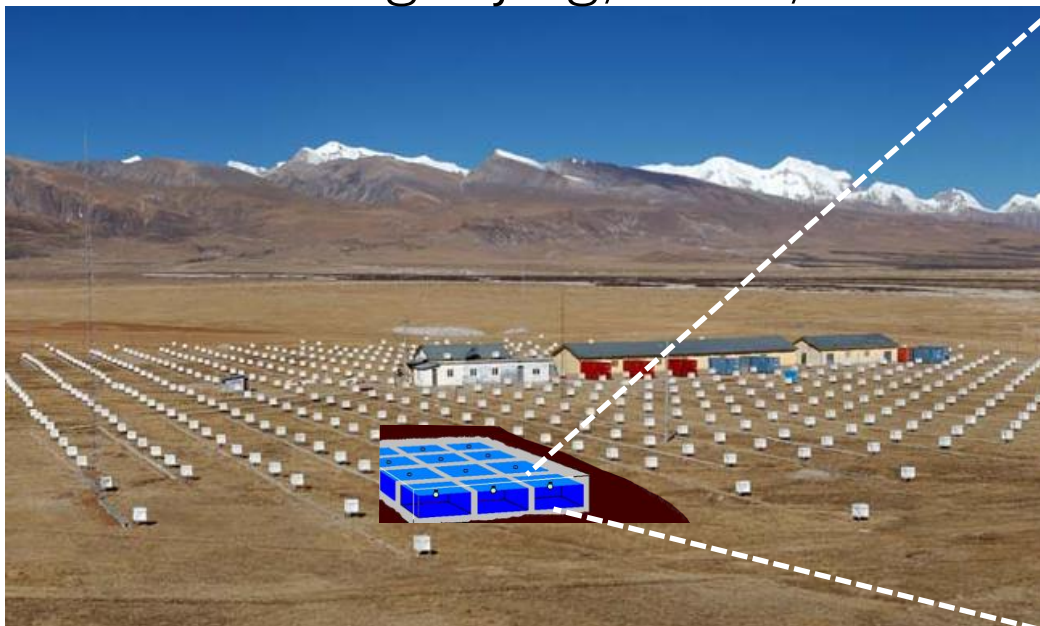
Low-cost large-area muon detector
→ utilize to discriminate gamma ray
signals from cosmic ray background.

Key technology! in this study



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Detection Principle

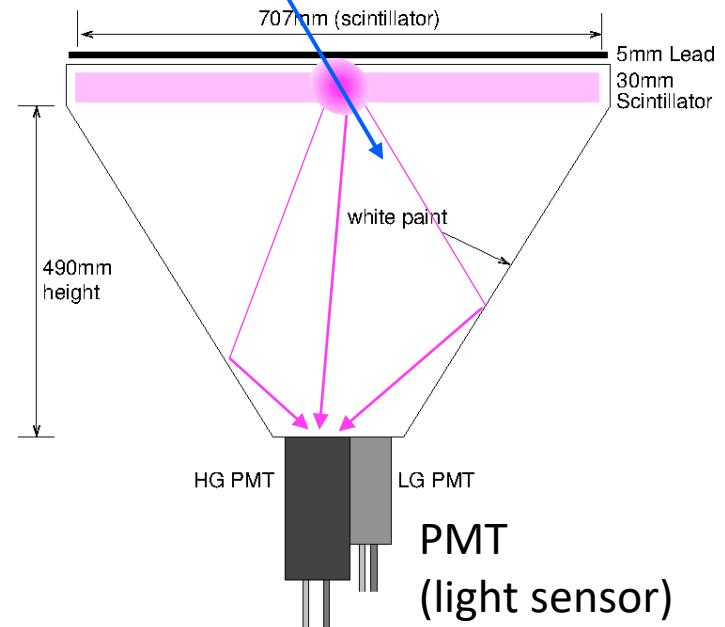
Gamma ray/Cosmic ray

Interact with Nitrogen

A lot of secondary particles = **Air shower**

Observe secondary particles by each detector

Scintillator:
Emit fluorescence lights
when particles pass it



Determine the direction & energy of the gamma ray

- Angular resolution 0.2 deg
- Energy resolution 20% @0.1PeV

Need another technique to separate γ rays from cosmic rays.

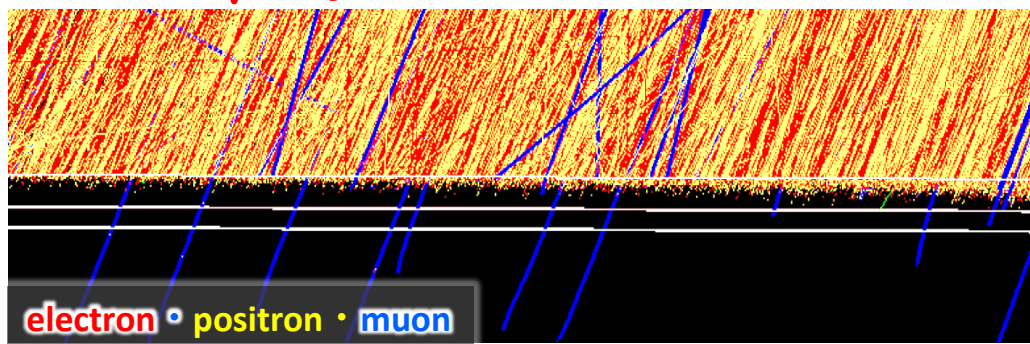


Gamma-Ray Selection

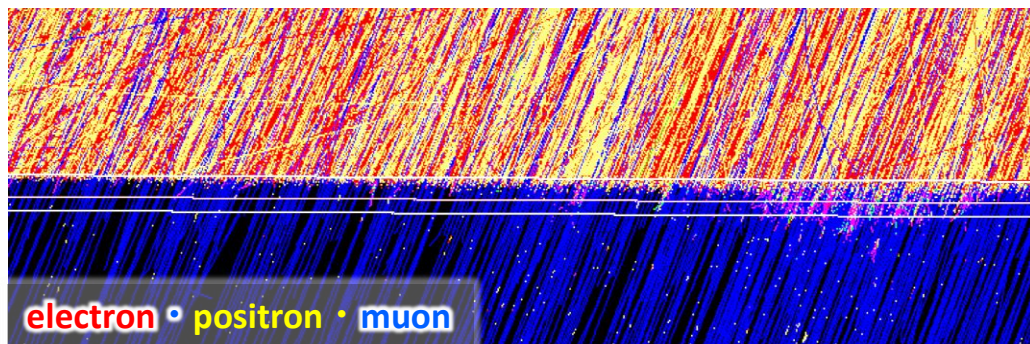
γ -ray $\rightarrow$ poor muons

Muons can penetrate underground

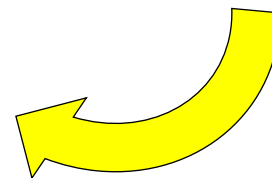
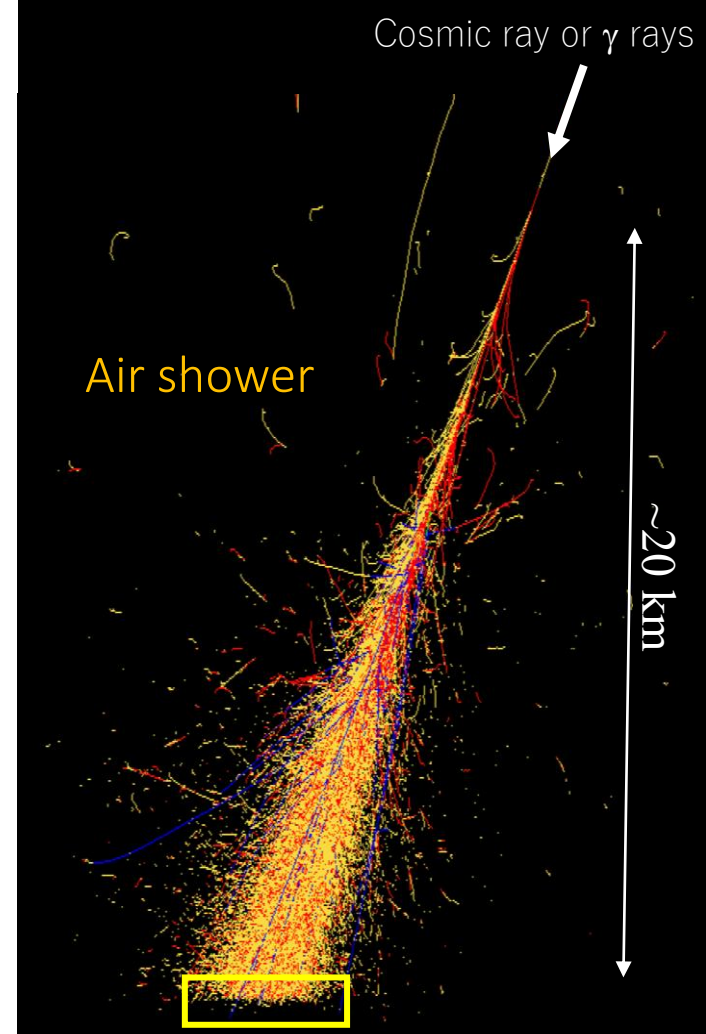
0.2PeV γ -ray



0.2PeV Cosmic ray (Noise)



$\rightarrow$ Underground muon detectors



Enlarged view

Gamma-ray Observation



Cosmic-Ray Background Rejection

Look for a needle in a haystack

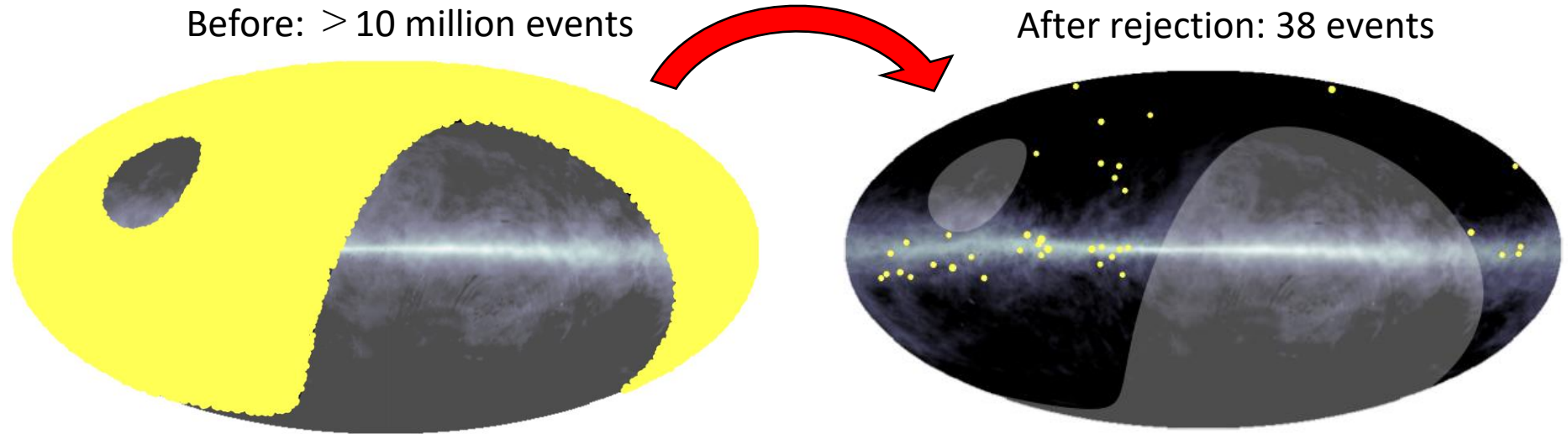
Expected cosmic-ray noise
= γ -ray signals x 100,000

→ Selecting muon poor events
measured by the underground
detectors → Key technology!

Reducing cosmic-ray background noise to
one part per million @0.4 PeV

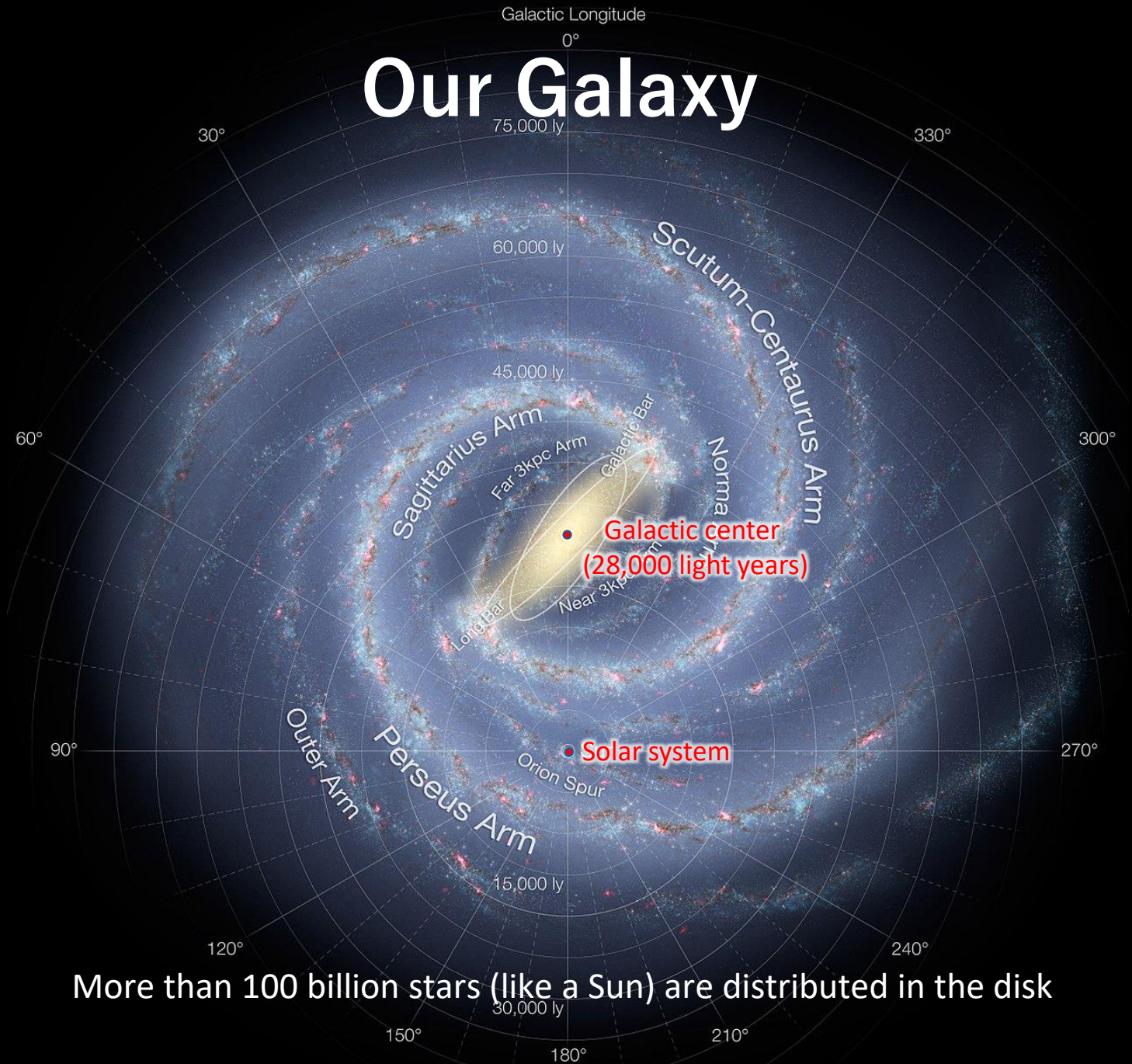
Before: > 10 million events

After rejection: 38 events



First Detection of 0.1-1 PeV gamma ray from the Milky Way

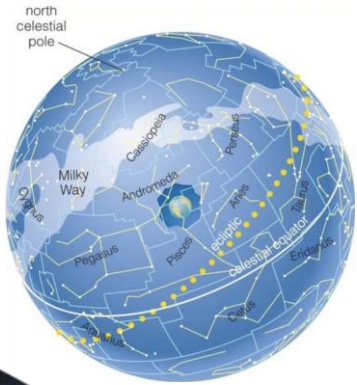
Our Galaxy



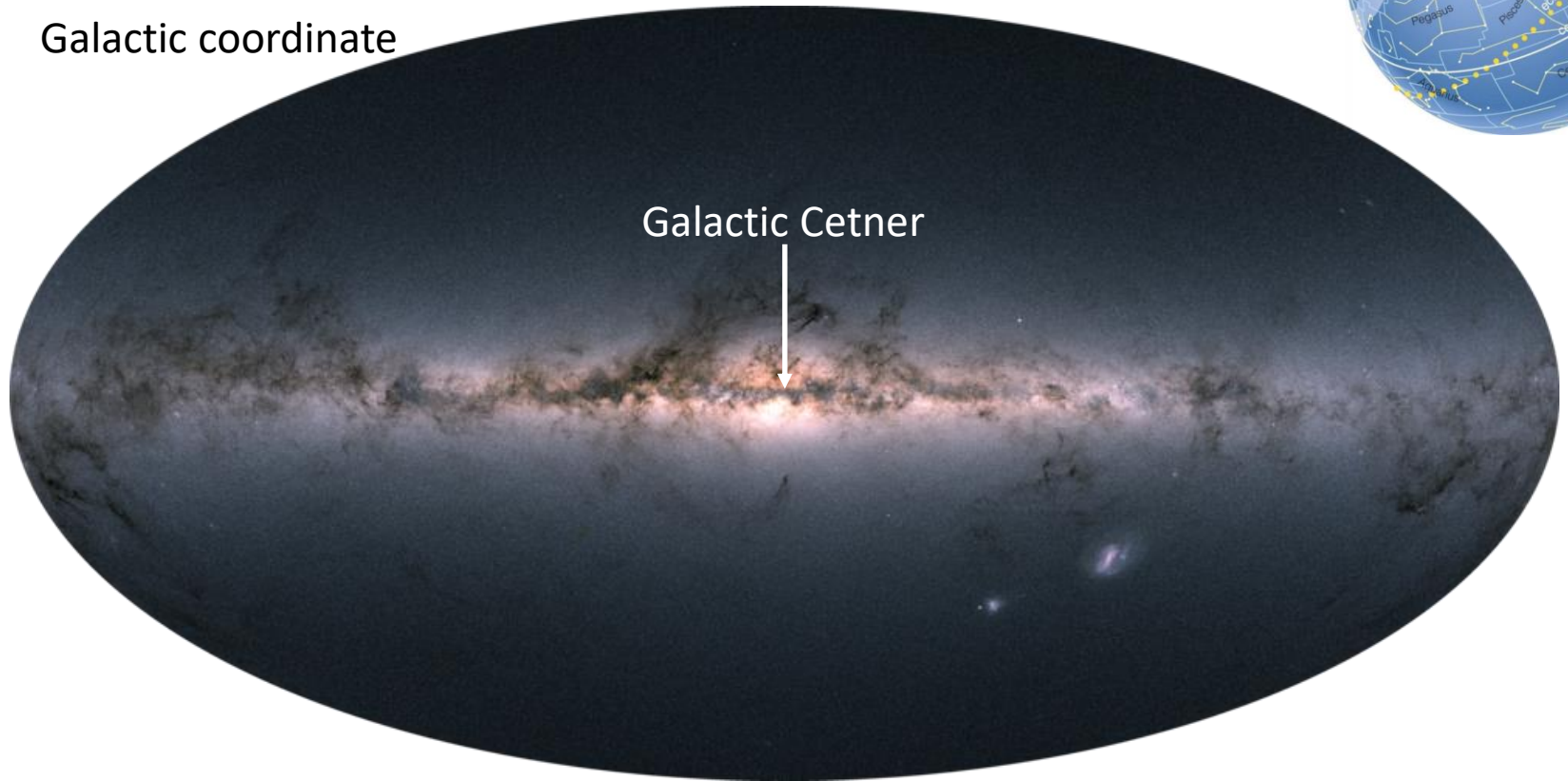
More than 100 billion stars (like a Sun) are distributed in the disk

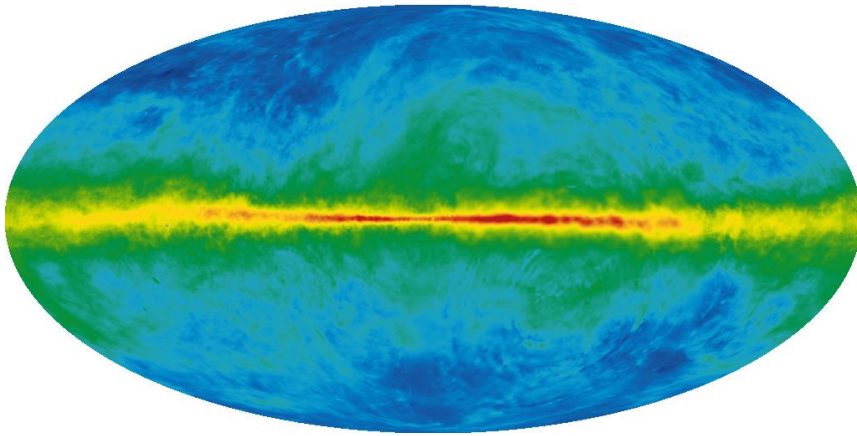


Milky Way by Optical Observation



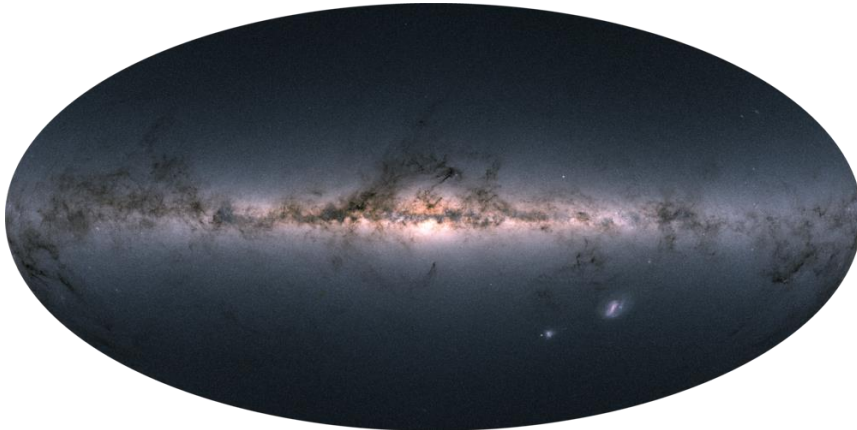
Galactic coordinate





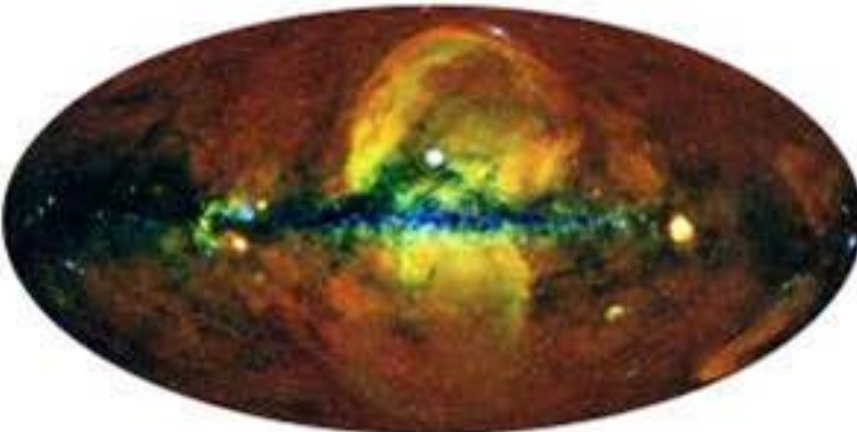
Radio: Atomic Hydrogens

Radio (21cm) HI Map
Hartmann et al. (1997)
Dickey & Lockman (1990)



Optical: Stars

Copyright: ESA/Gaia/DPAC



X-Ray : Hot plasma

Copyright: eROSITA (MPE/IKI)

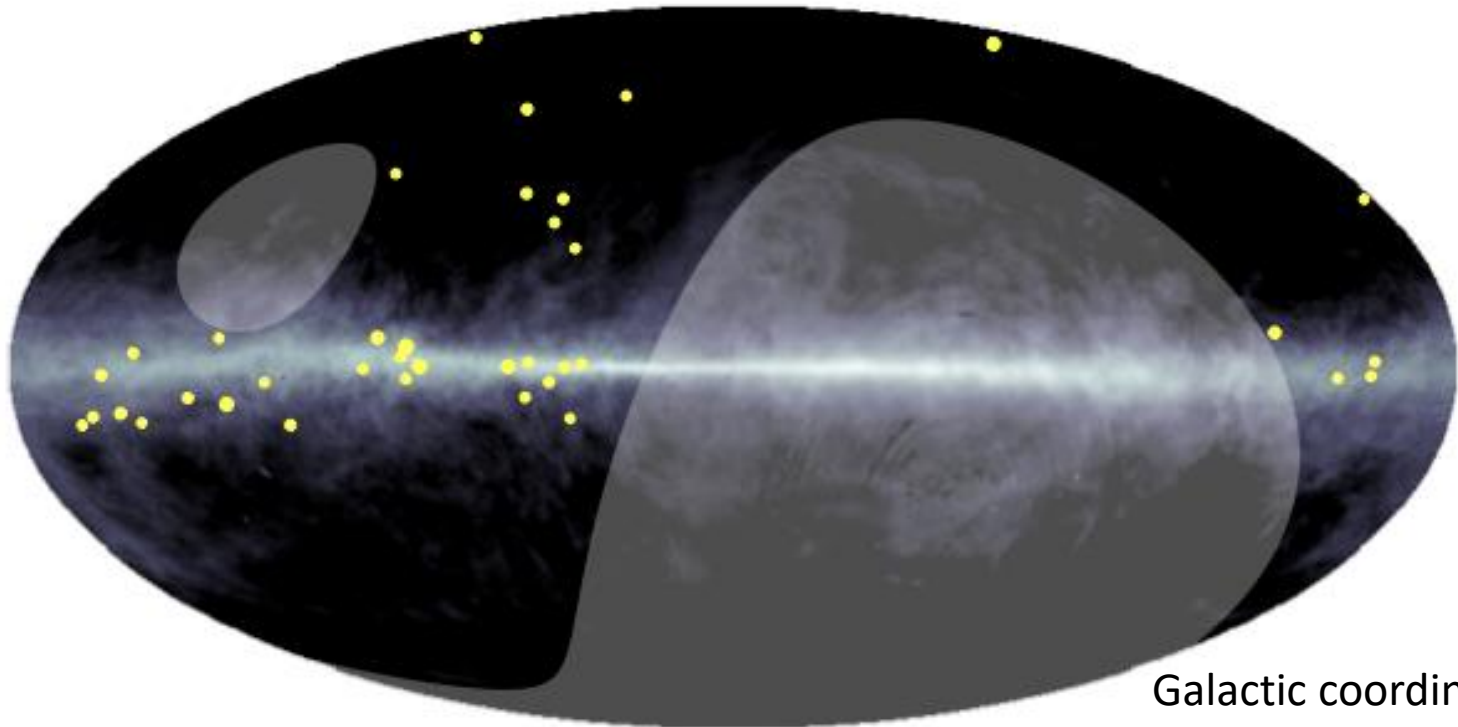


Highest Energy γ Rays from the Milky Way

23 ultra-high-energy gamma-ray events along the Milky Way!

(2 years of data during period between 2014 and 2017)

The highest energy 0.957 PeV (~ 1 PeV) gamma ray is detected



0.4 – 1 PeV (sub-PeV) gamma rays

& Atomic hydrogen (HI) distribution

HI data available at https://lambda.gsfc.nasa.gov/product/foreground/fg_combnh_map.cfm

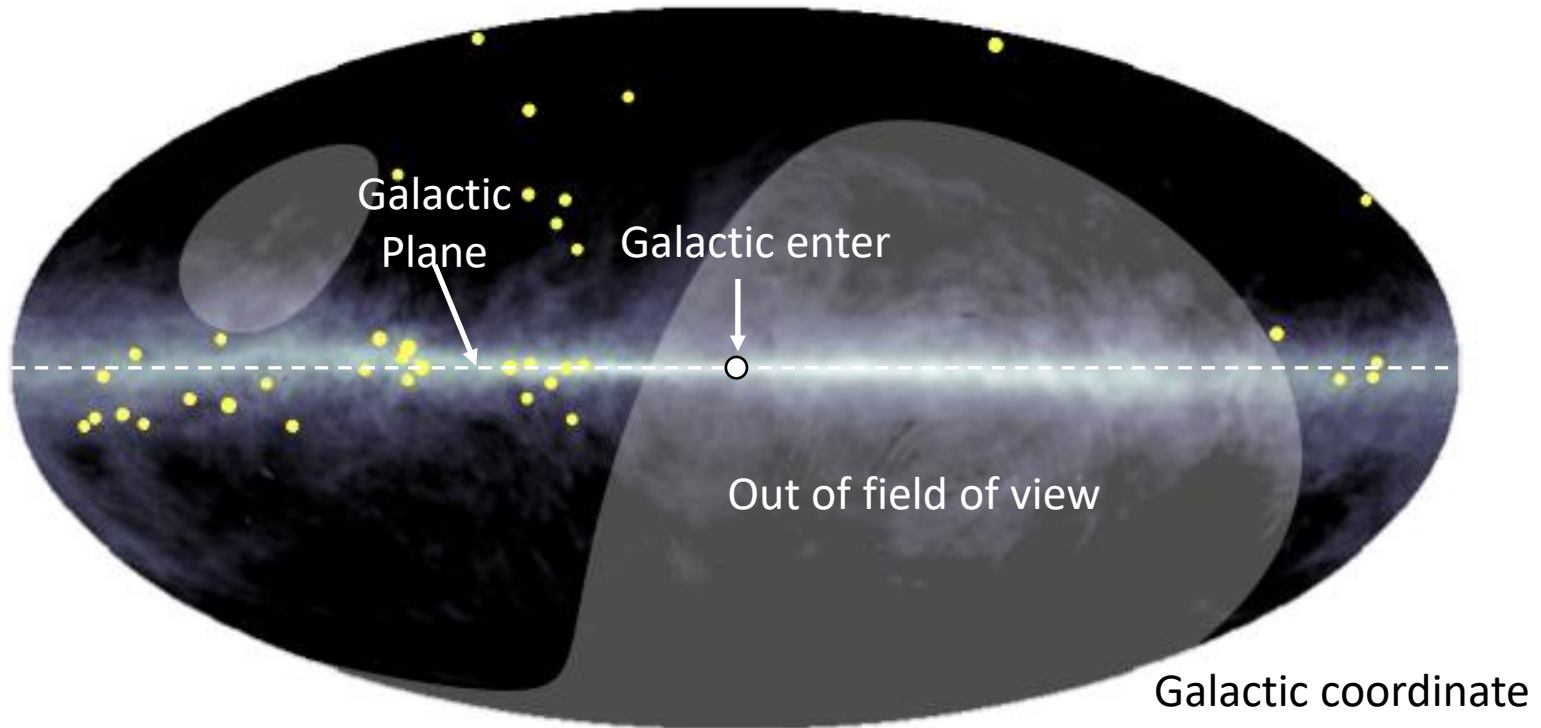


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→ Gamma rays from inside of our Galaxy

§ Scientific Interpretation

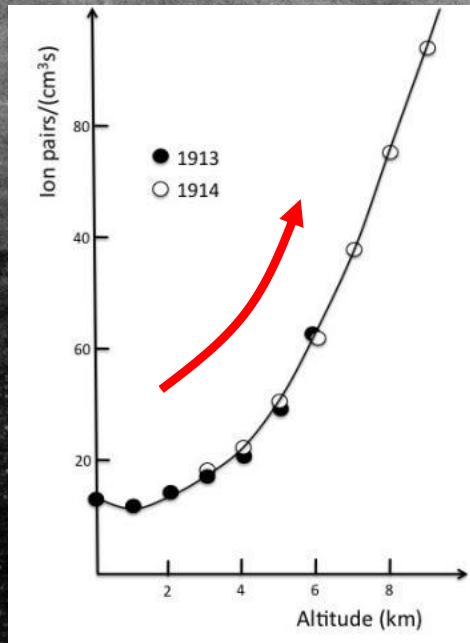
Cosmic rays were discovered by Victor Hess in 1912

(Nobel prize in 1936)



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Cosmic rays...

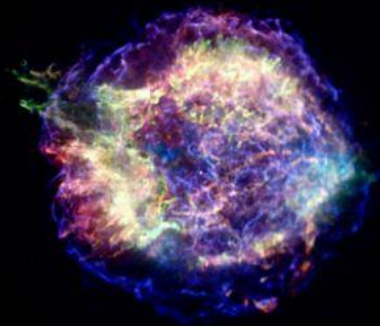
High energy particles (=protons)
Coming from space

Cosmic rays...

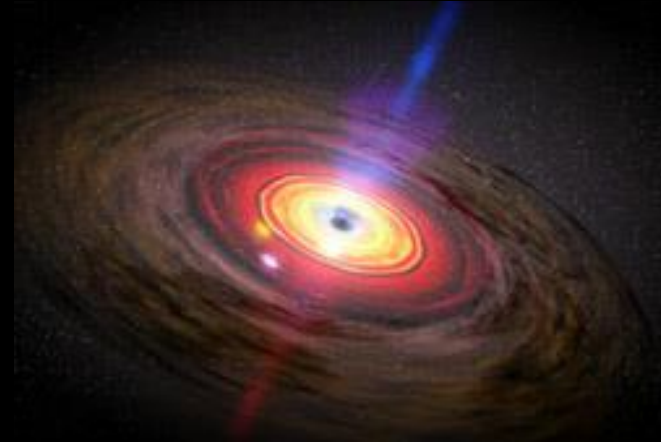
Where do they come from?



Sun



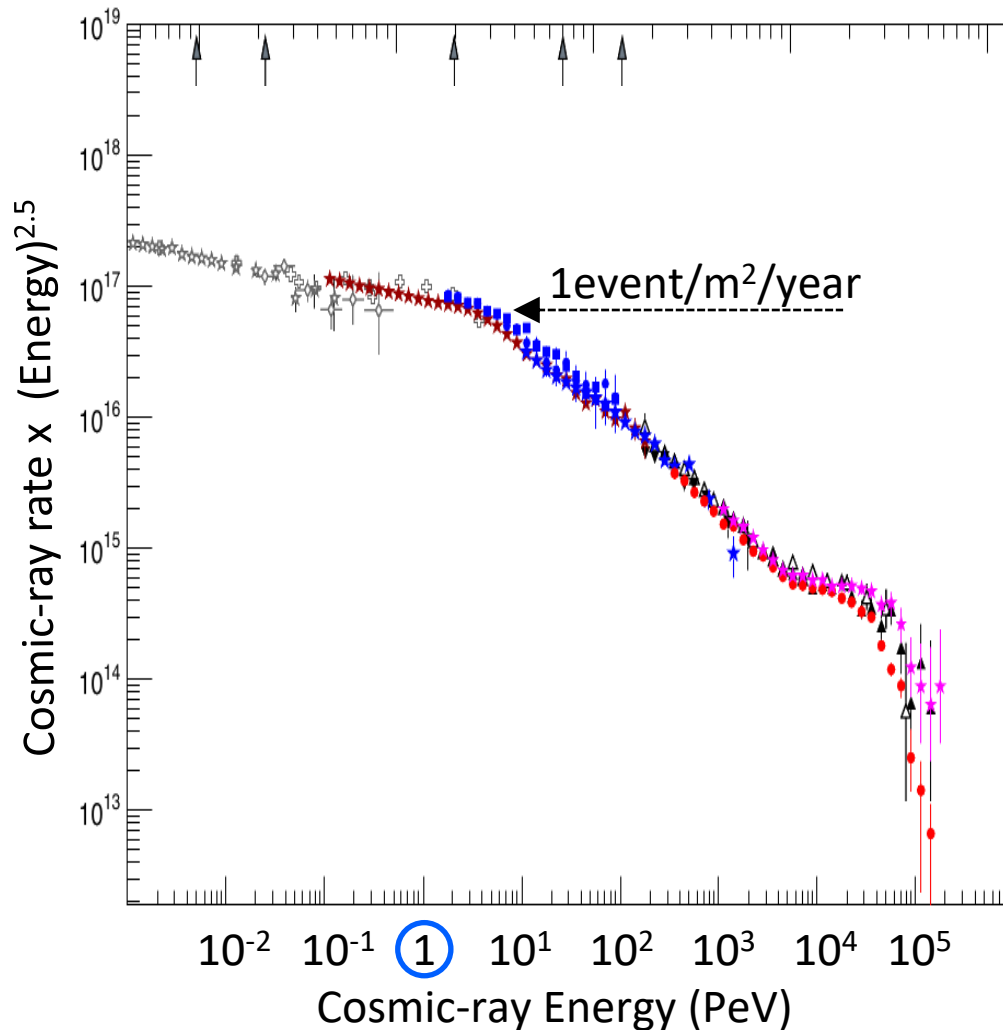
Supernova remnant?
(after star explosion)



Active galaxy??
(massive blackhole)



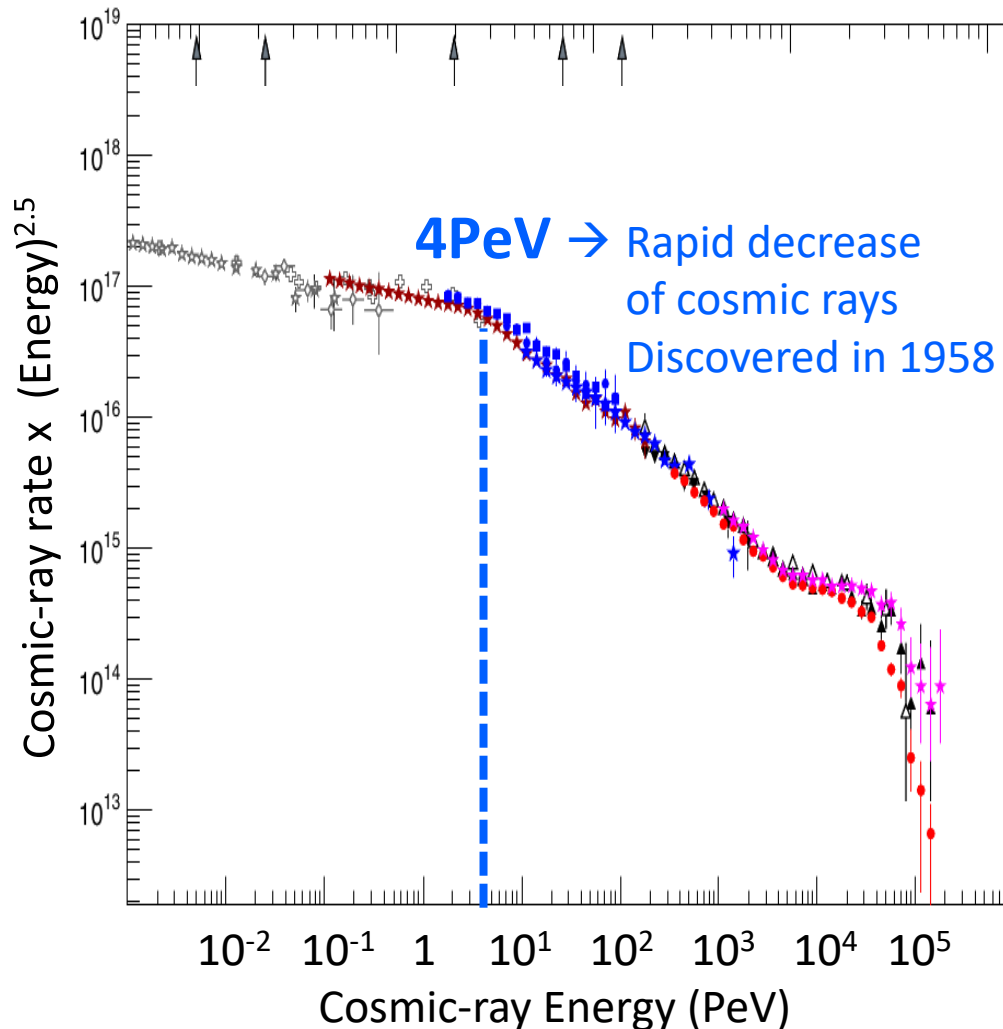
Cosmic Ray Rate & Energies



- ❖ Wide energy range
- ❖ Main component is proton
- ❖ Rate decreases to 1/100 when energy is 10 times higher



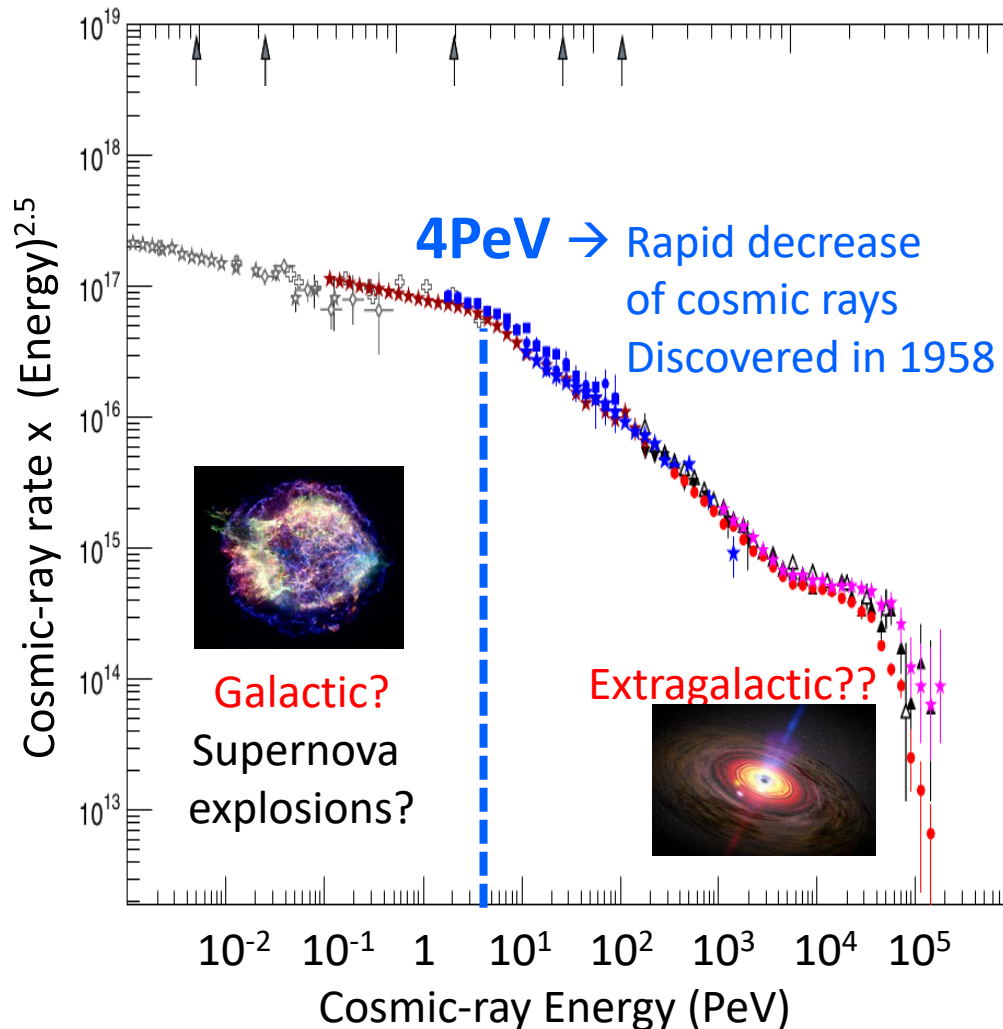
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Cosmic Ray Rate & Energies

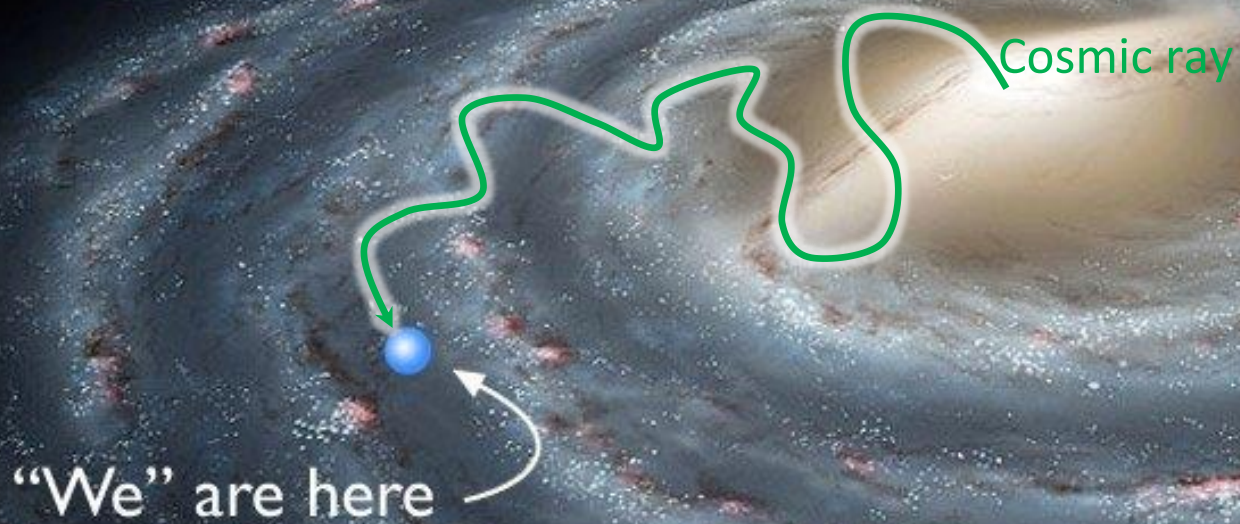


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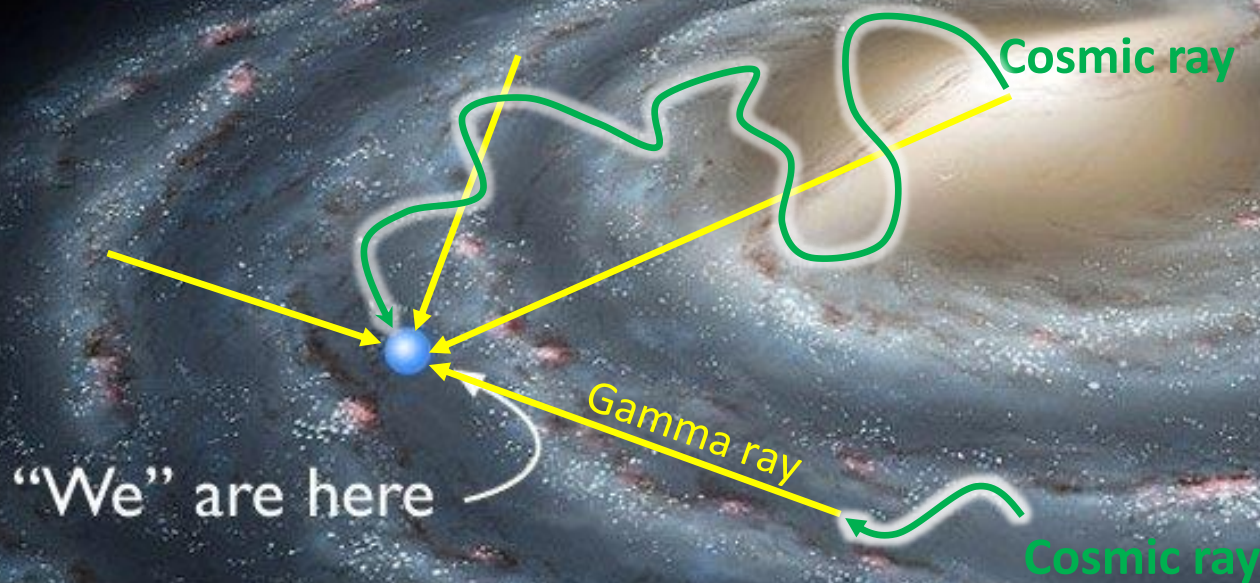
As an open question,
Did/Do “PeVatrons” really
exist in our Galaxy?

PeVatron: Cosmic superaccelerators
accelerating CR to Peta electron volt

Cosmic rays with electric charge are bent by the magnetic field in the universe, and lose their directions



Cosmic rays with electric charge are bent by the magnetic field in the universe, and lose their directions



Cosmic rays interact with interstellar gas, and produce γ rays
Gamma rays go straight unaffected by magnetic field,
pointing back to the sources.

History of Cosmic-Ray Physics

1912 Discovery of cosmic rays (Hess)

1934 Supernova hypothesis as cosmic-ray origin (Baade & Zwicky)

1949 Acceleration theory of cosmic rays (E. Fermi)

1958 Discovery of rapid decrease in cosmic-ray flux (Kulikov & Khristiansen)

**→ Cosmic rays below a few PeV are/were produced in our Galaxy? =
PeVatrons**

1970s-1990s Detection of γ rays from Milky Way (OSO-3, SAS-2, COS-B, EGRET)

→ 1/10000th PeV γ rays induced by cosmic rays in our Galaxy

2013 Detection of 1/1000th PeV γ rays induced by cosmic rays in SNR
(Fermi-LAT satellite + Ground γ -ray telescope)

Despite exhaustive searches for **PeVatrons** over last 20 years,
researchers have no conclusive evidence yet.

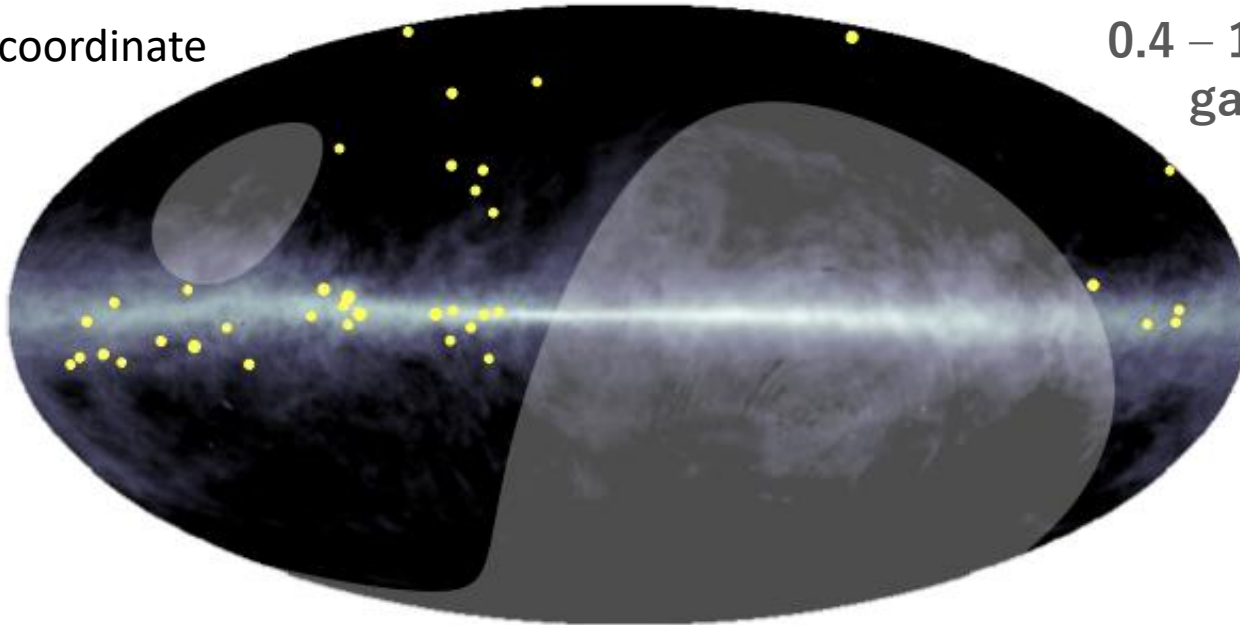
→ A mystery for 60 years in cosmic-ray physics



Scientific Interpretation

Galactic coordinate

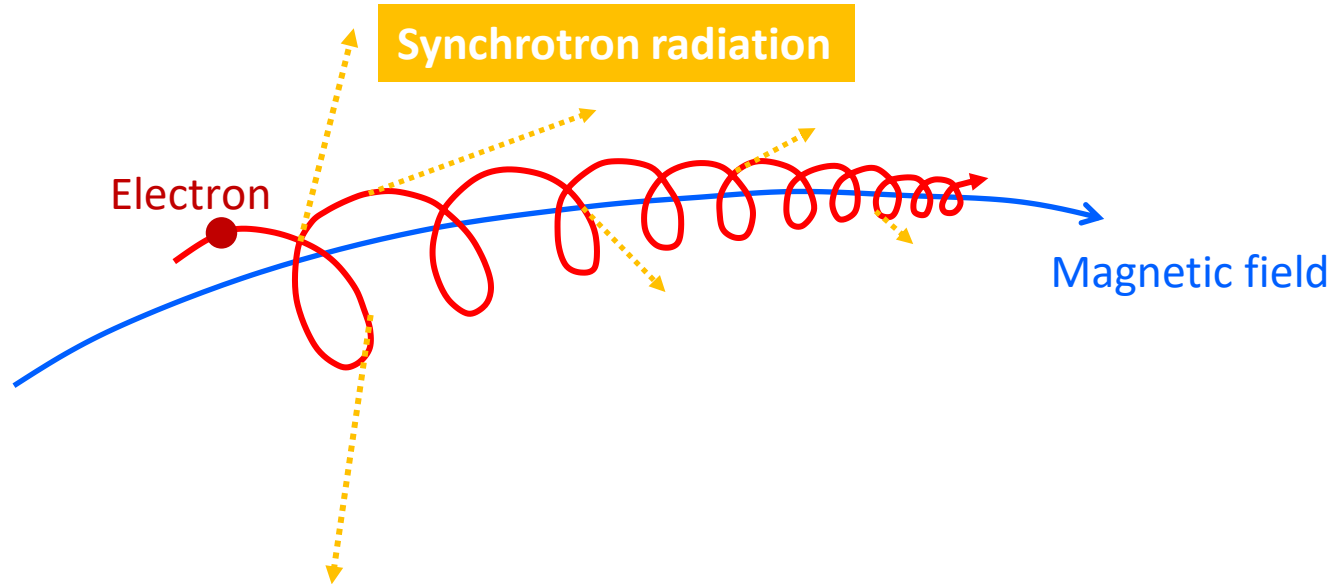
0.4 – 1 PeV
gamma rays



- ✓ Highest-energy gamma rays, including **one Peta-electron-volt**, are discovered to be spread out along the Milky Way, **rather from known gamma-ray objects**.
- Gamma rays coming directly from the objects are difficult to identify cosmic-ray proton origin, since they could be gamma rays produced by collisions of “electrons” with photons.



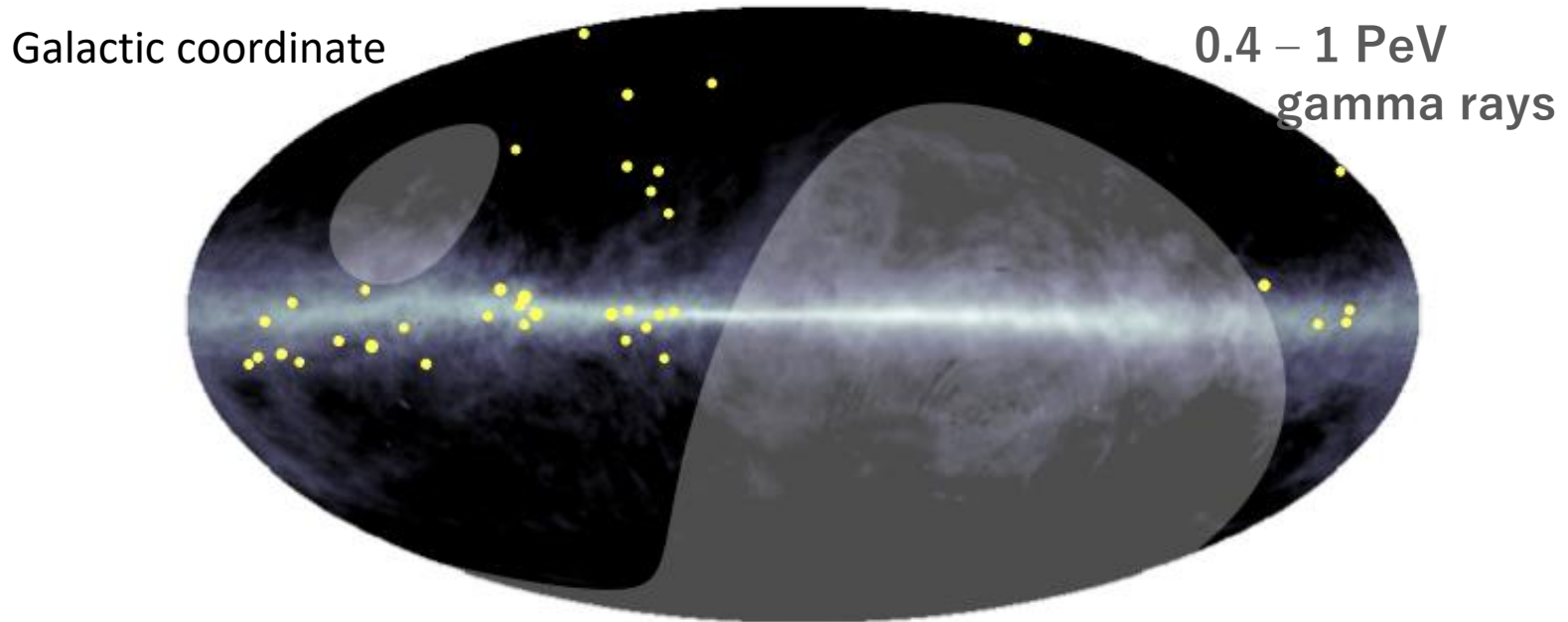
Electron origin? vs Proton origin?



- ✓ Gamma rays are coming isolated from known gamma-ray sources.
 - **Electrons** lose their energy quickly, so they **should stay near the object**.
 - **Protons** don't lose energy and **can escape farther from the object**.



Electron origin? vs Proton origin?



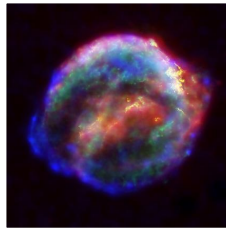
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 - **Electrons** lose their energy quickly, so they **should stay near the object**.
 - **Protons** don't lose energy and **can escape farther from the object**.

Strong evidence for sub-PeV γ rays induced by cosmic rays



PeVatrons & Cosmic-Ray Pool

Supernova (SN) explosion occurs every 30 years in our Galaxy



In SN remnants, cosmic rays are accelerated within 10,000 years after explosion → **“PeVatron”**



Cosmic rays have been trapped by galactic magnetic field more than a few million years, forming **cosmic-ray pool**
= Accumulated for 100,000 SN explosions

Cosmic rays in galactic magnetic field
3 – 30 PeV

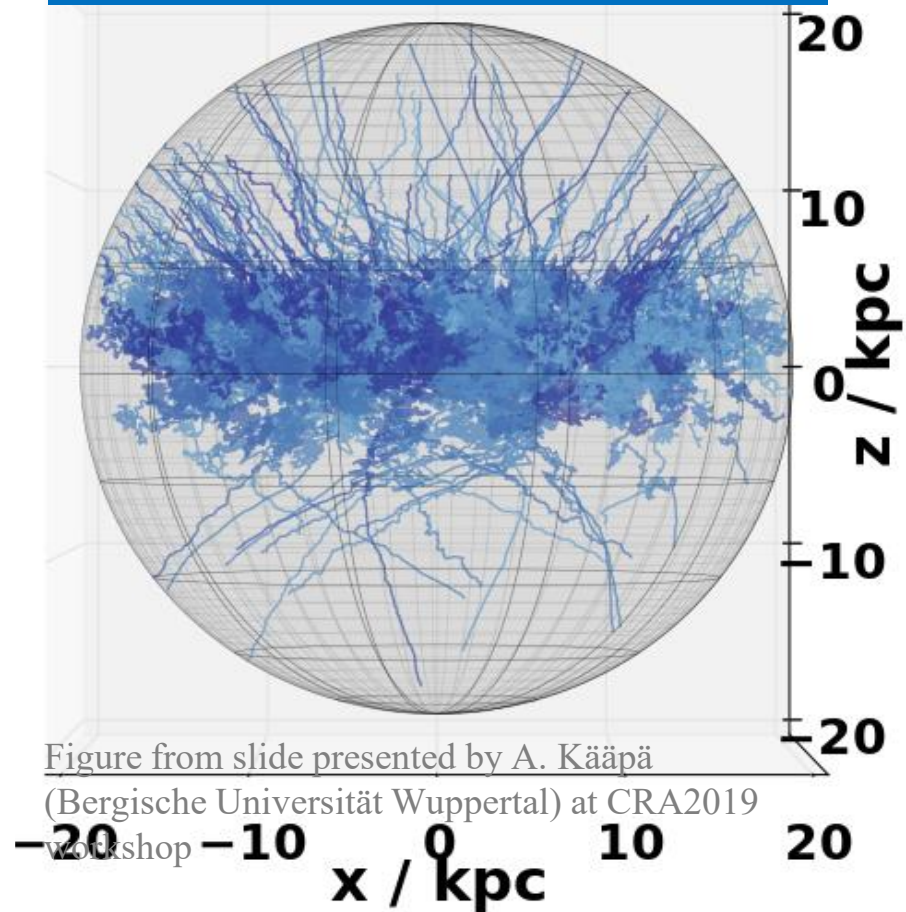
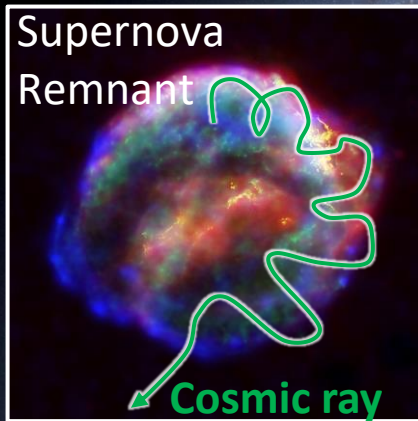
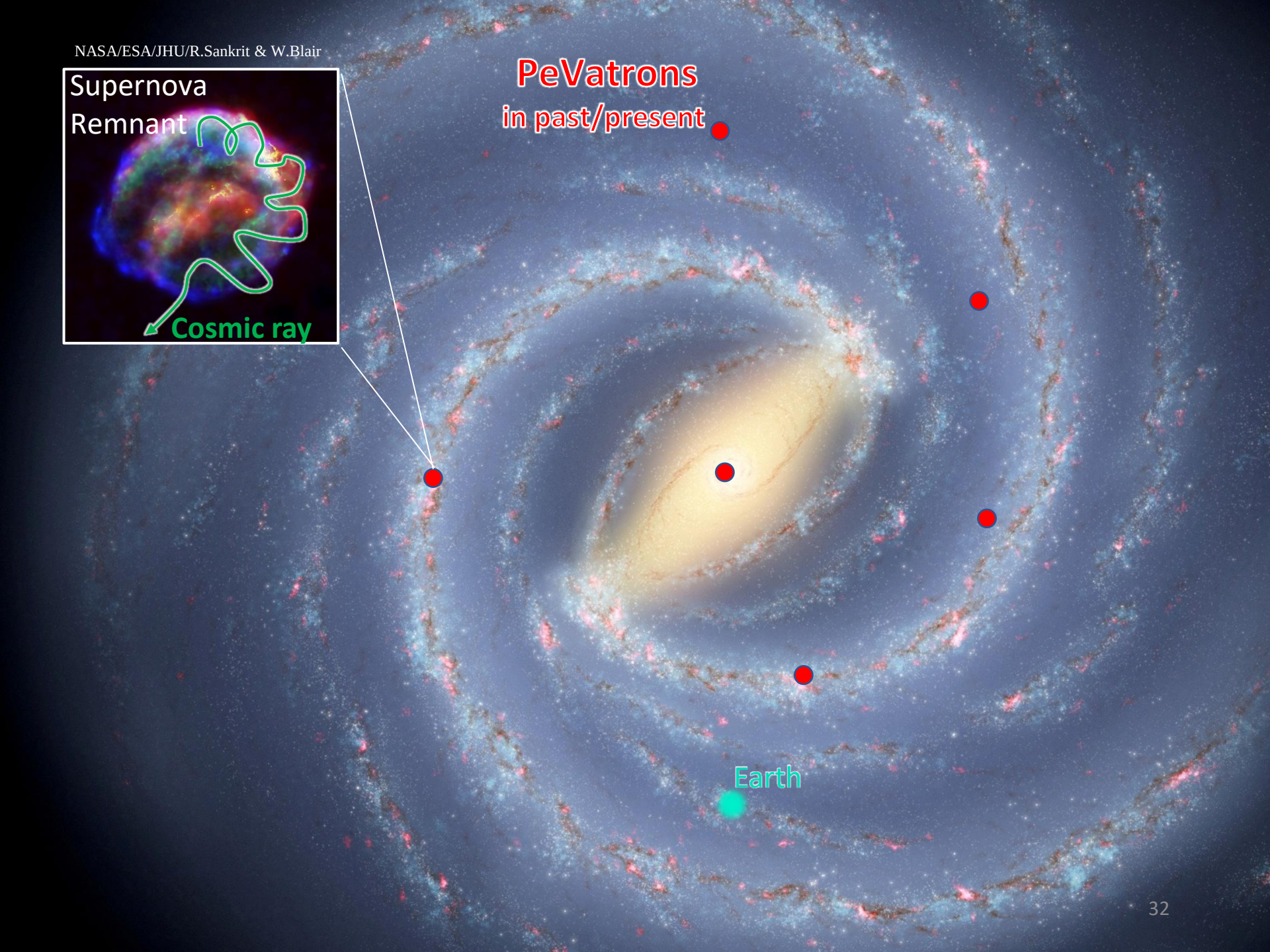


Figure from slide presented by A. Kääpä
(Bergische Universität Wuppertal) at CRA2019

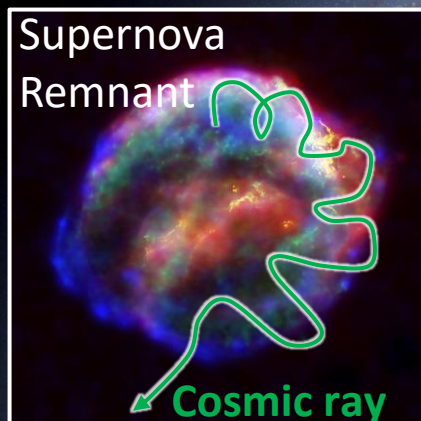


PeVatrons
in past/present



The main image is a detailed view of a spiral galaxy, likely the Milky Way, showing its central bulge and spiral arms. Several red dots are placed along the spiral arms, representing the locations of PeVatrons. A cyan dot is placed near the bottom center, representing Earth. The text 'PeVatrons in past/present' is written in red above the top right red dot. Two white lines connect the top left red dot to the inset image.

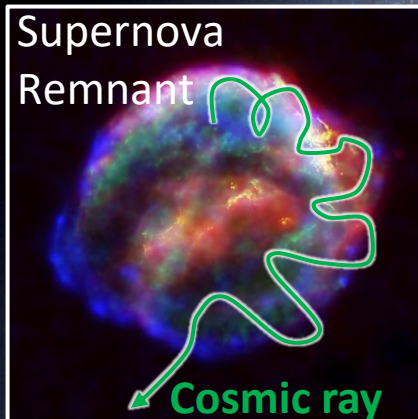
Earth



PeVatrons
in past/present

PeV cosmic ray

Earth



PeVatrons
in past/present

PeV cosmic ray

0.1PeV gamma ray

Earth

Cosmic rays interact with
interstellar gas, and produce γ rays

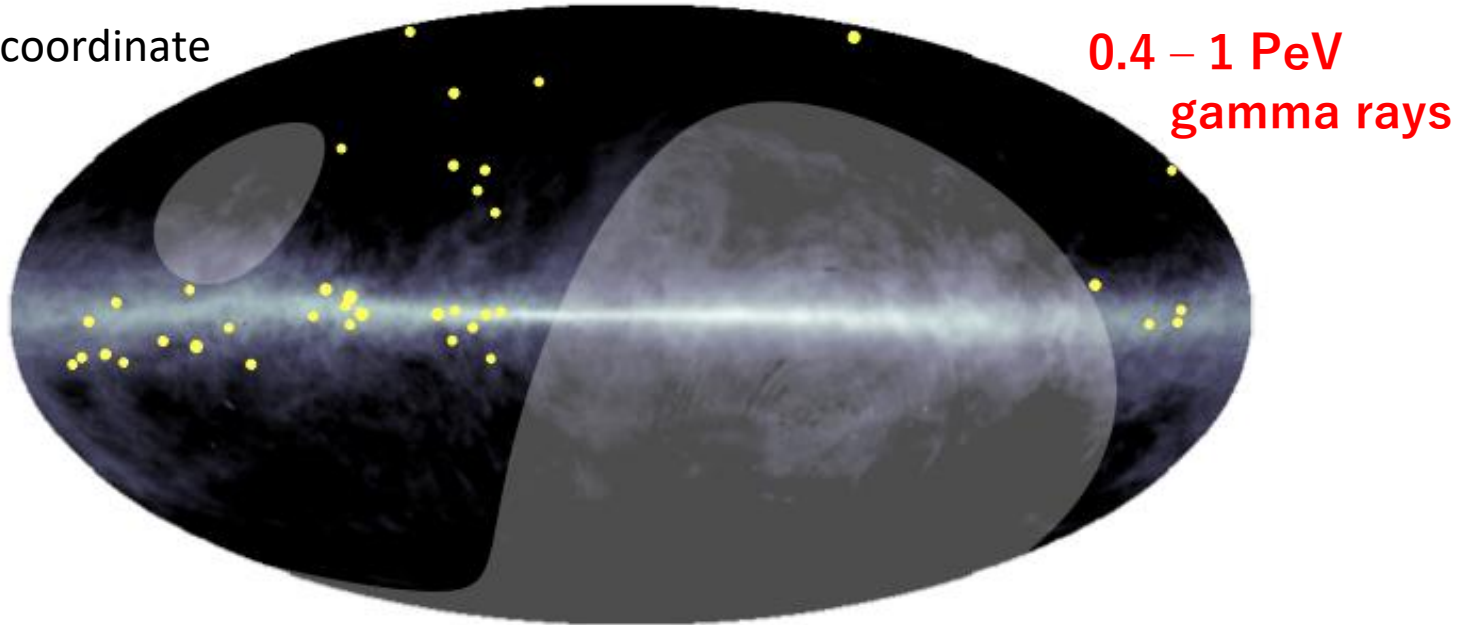
$$p + p \rightarrow X's + \pi^{\pm} + \pi^0 \rightarrow 2\gamma$$

(γ -ray energy has 10% of cosmic rays)



Scientific Interpretation

Galactic coordinate



- ✓ This is **the first compelling evidence for existence of PeVatrons**, in the past and/or present Galaxy, which accelerate protons up to the Peta electron volt (PeV) region.



Scientific Interpretation

High-energy
cosmic rays

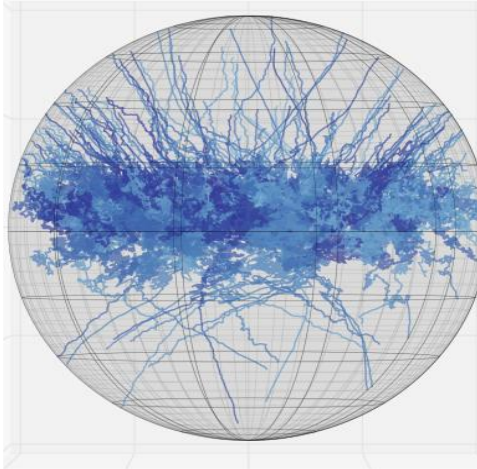
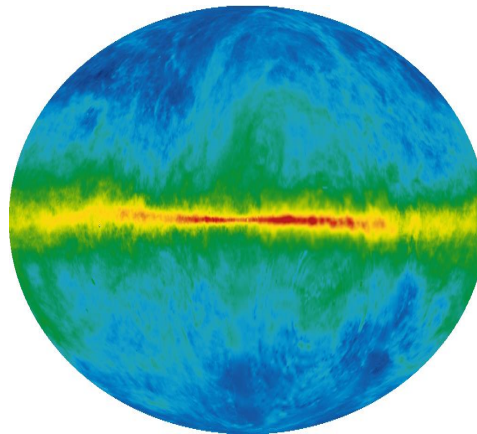


Figure from slide presented by A. Kääpä (Bergische Universität Wuppertal) at CRA2019 workshop

Interstellar
matter

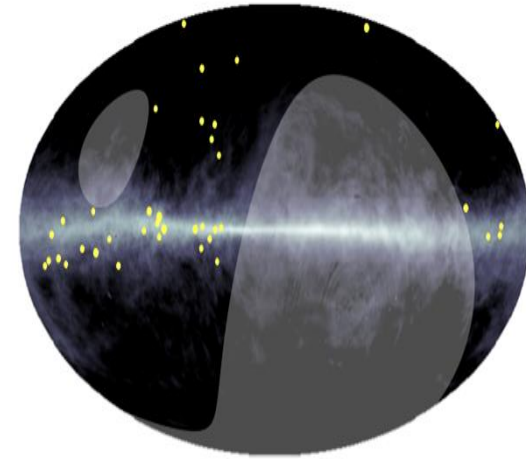
×



Radio (21cm) HI Map
Hartmann et al. (1997)
Dickey & Lockman (1990)

=

High-energy
gamma rays



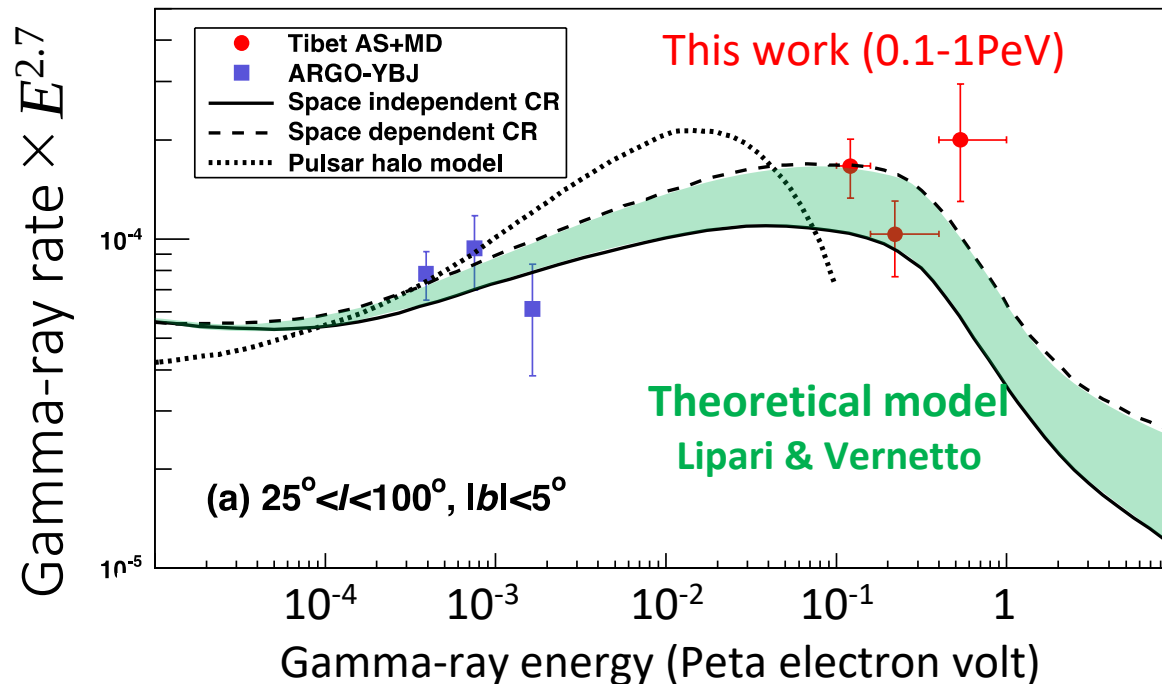
This Work

- ✓ This work proves a theoretical model that cosmic rays produced by PeVatrons are trapped in the Galactic magnetic field for millions of years, **forming a pool of cosmic rays.**



Scientific Interpretation

The measured γ -ray rates are consistent with the expected one from cosmic-ray pool scenario assuming the cosmic-ray rate observed on Earth.



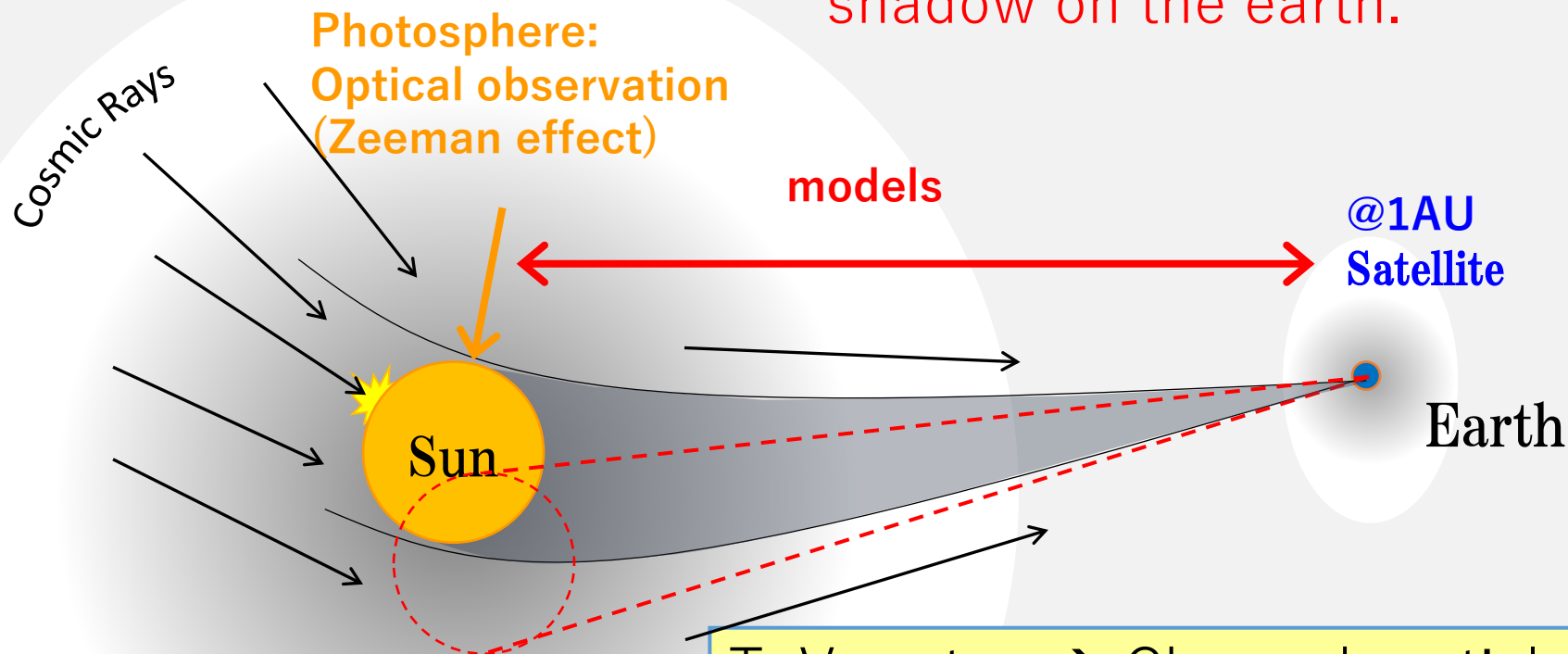
- ✓ It is verified that all the high-energy cosmic rays propagated to Earth can be explained by the **cosmic-ray pool produced by Pevatrons in the past/present Galaxy.**

Sun Shadow Observation



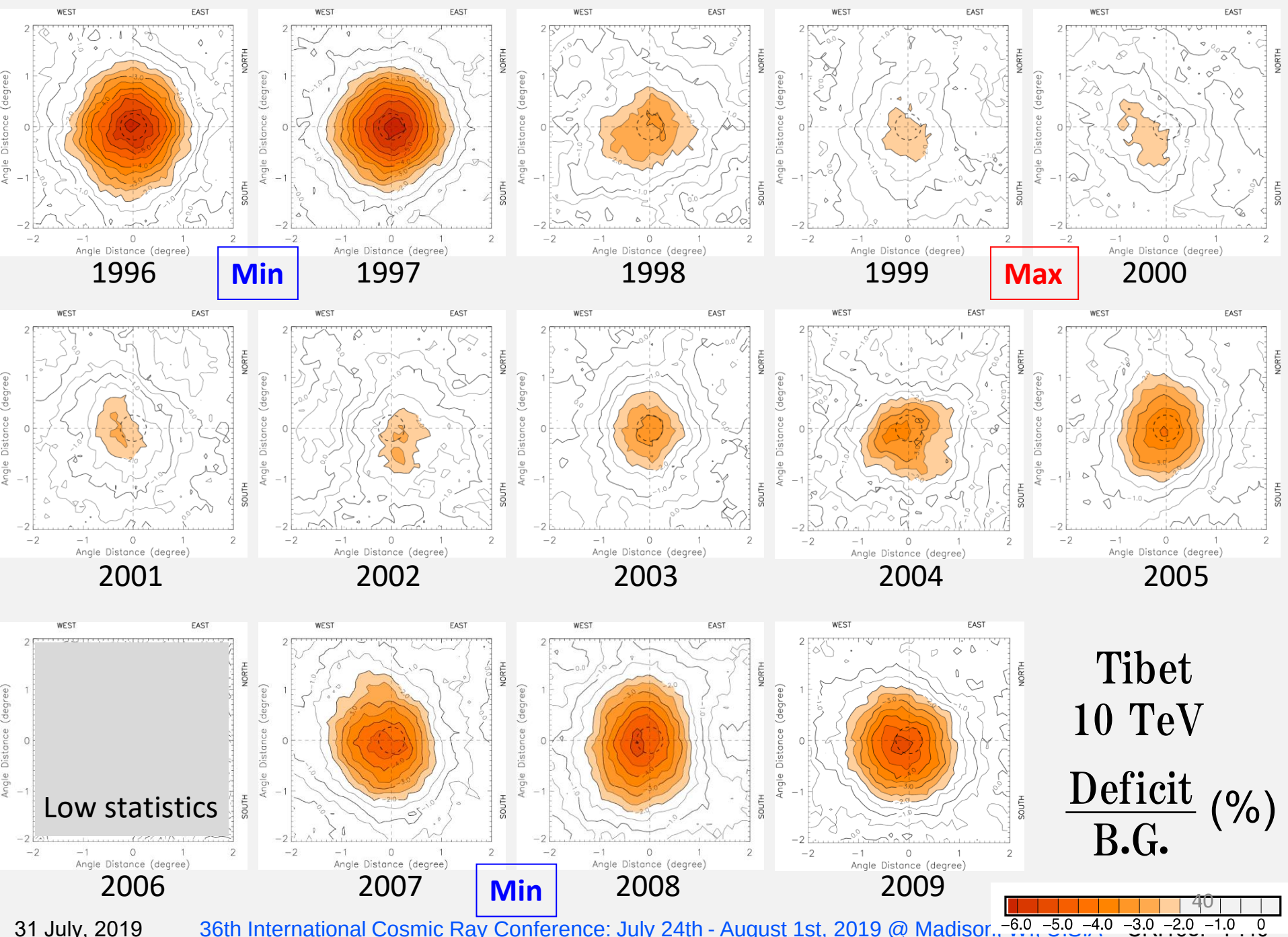
Sun Shadow

Sun blocks VHE cosmic rays, and cast the cosmic-ray shadow on the earth.



TeV proton $\rightarrow$ Charged particle
Gyroradius
 $\sim 7.4\text{AU}$ ($B=30\mu\text{G}$ near the earth)
 $\sim 0.16R_{\odot}$ ($B=300\text{mG}$ near the sun)

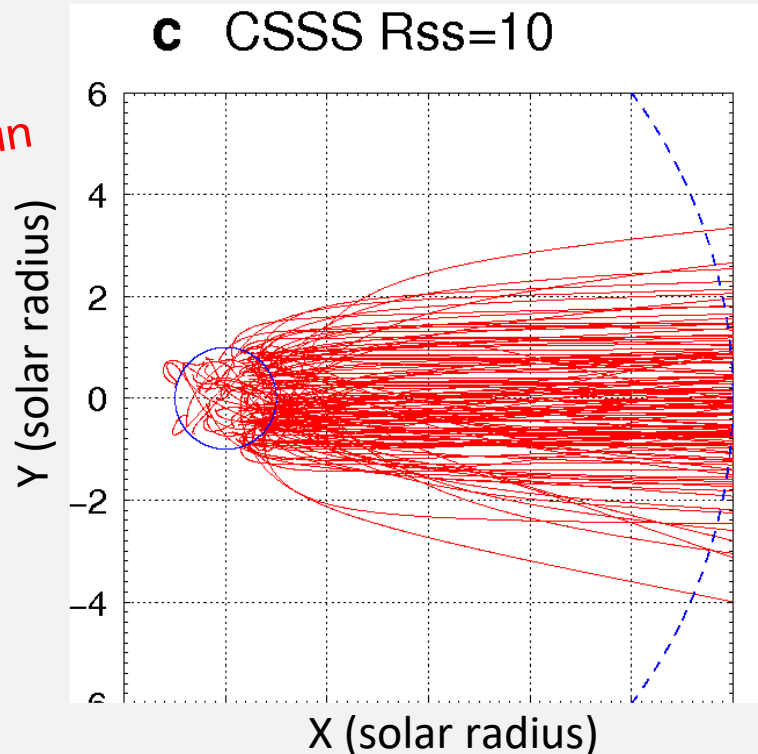
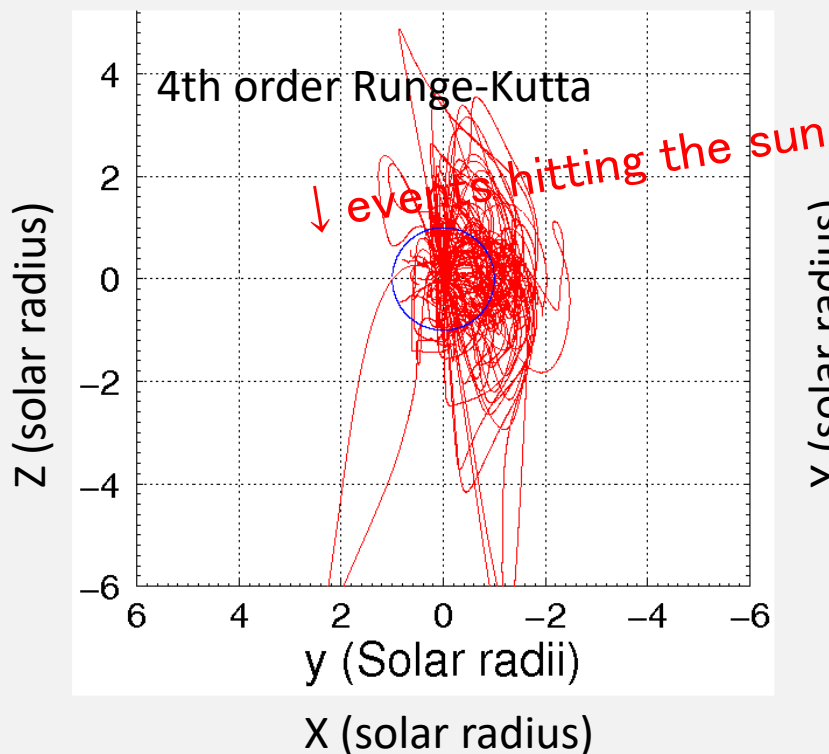
$\rightarrow$ Sun shadow provides unique method to probe solar MFs experimentally.





MC Simulation of Sun Shadow

- Anti-particles are shot back to the Sun from the Earth assuming CR spectra, compositions, detector responses
- Particle trajectories are traced assuming the solar MFs between the Sun and the Earth
- > Events hitting the sun reconstruct the Sun shadow





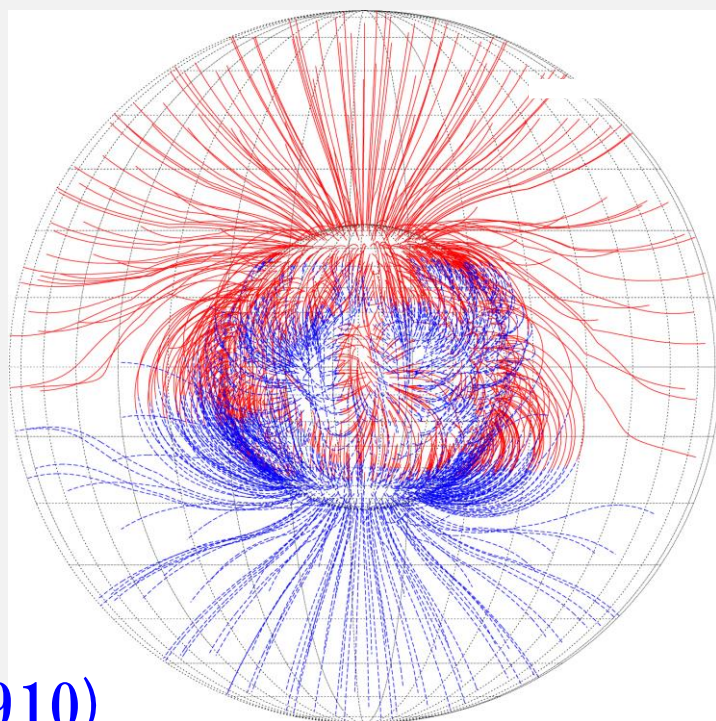
Time Dependent Coronal MF

Corona -> potential field models (*PSSS and CSSS*)

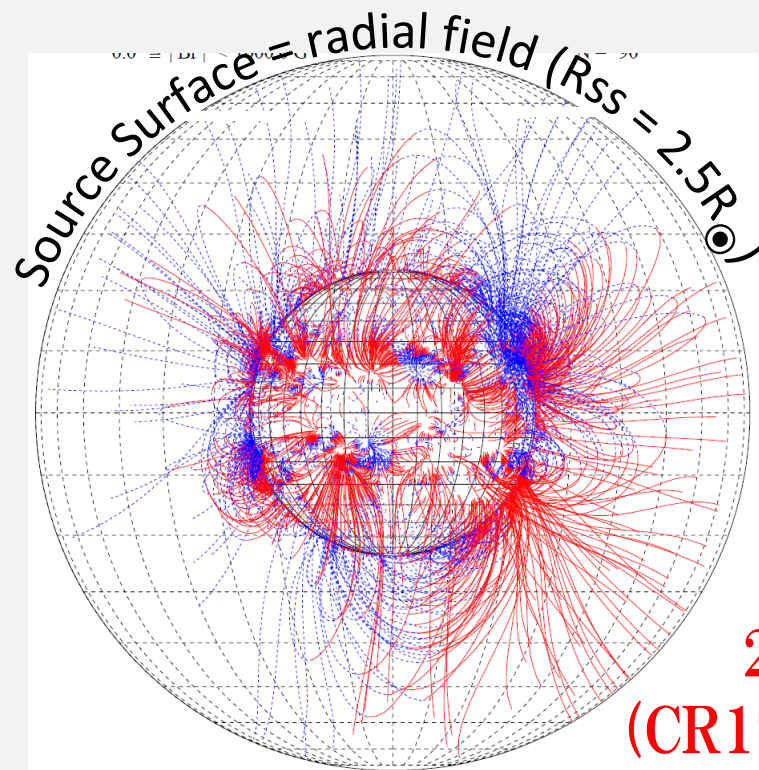
Extrapolated from the photospheric MFs measured by Kitt Peak (KPVT/SOLIS) in each Carrington rotation (~27 days)

IMF -> Parker spiral model with latitude dependence of the solar wind velocity taken into account.

Geomag. -> Dipole model



1996
(CR1910)

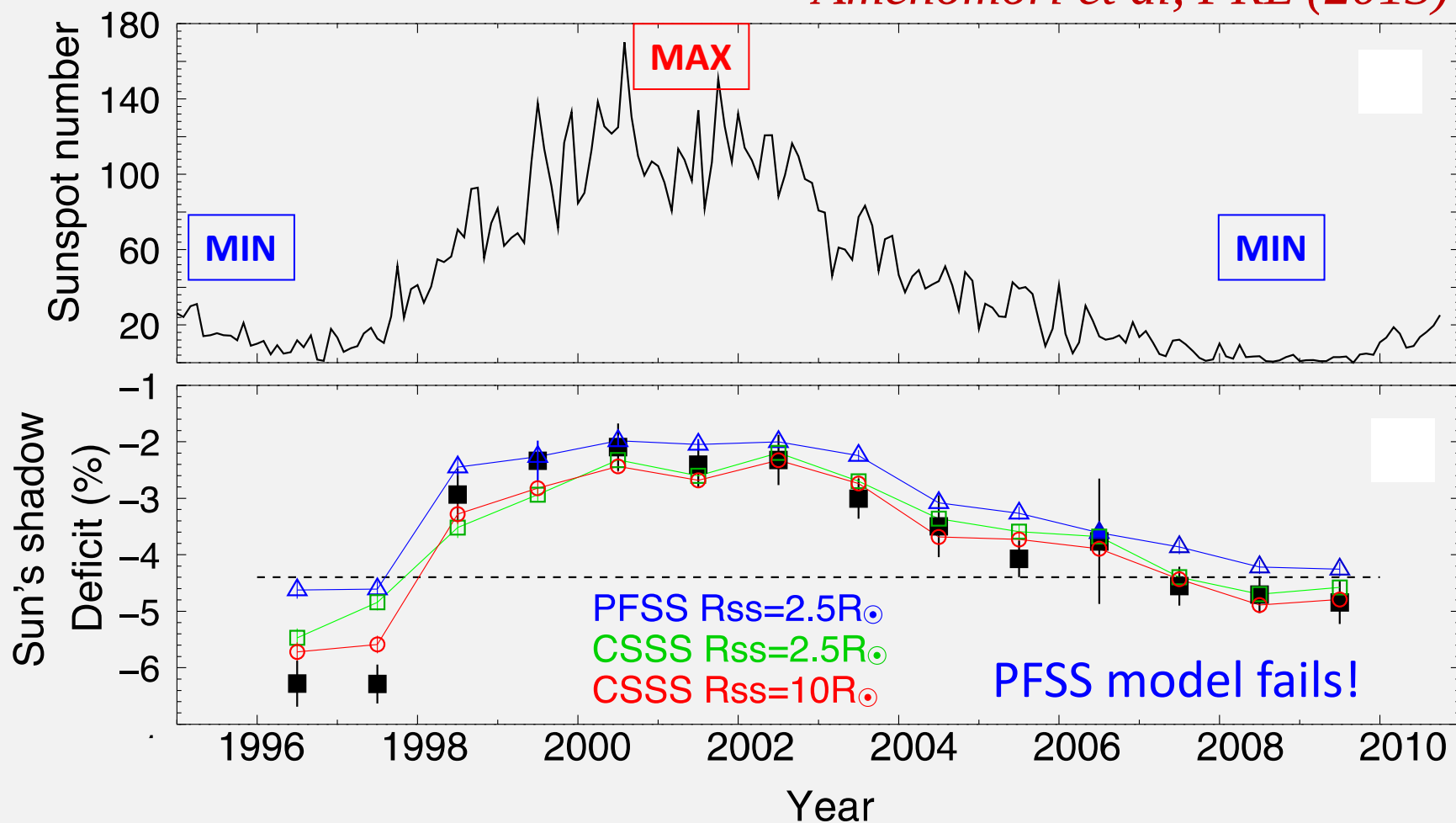


2001
(CR1978)



Evaluation Solar MF models

Amenomori et al, PRL (2013)



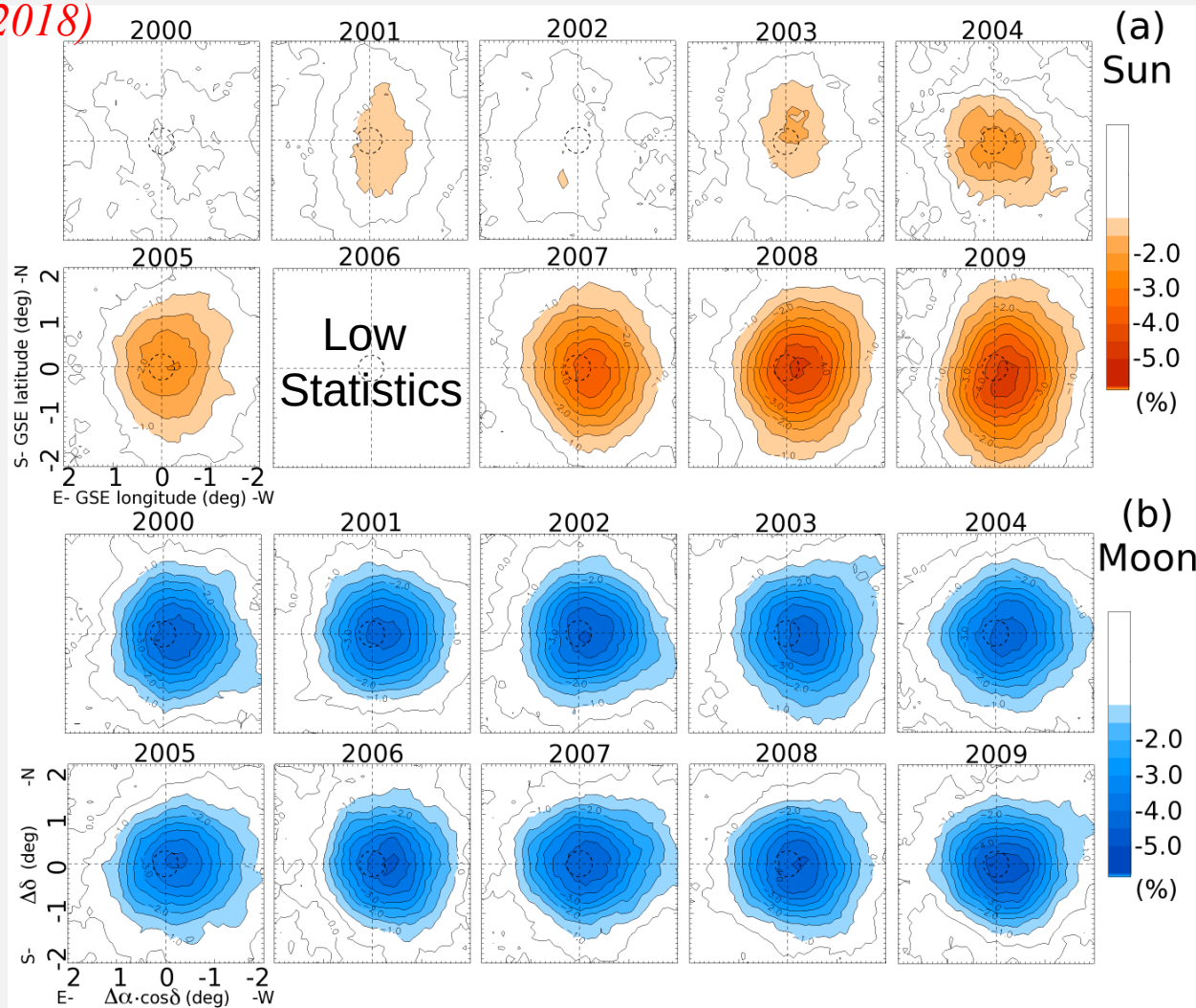
- ✓ Discovery of a clear anti-correlation of the deficits
- ✓ First demonstration to evaluate the coronal MF models



→ A clear solar-cycle variation of the deficits
CRs are scattered by solar magnetic field.

Amenomori et al., ApJ, 860,13 (2018)

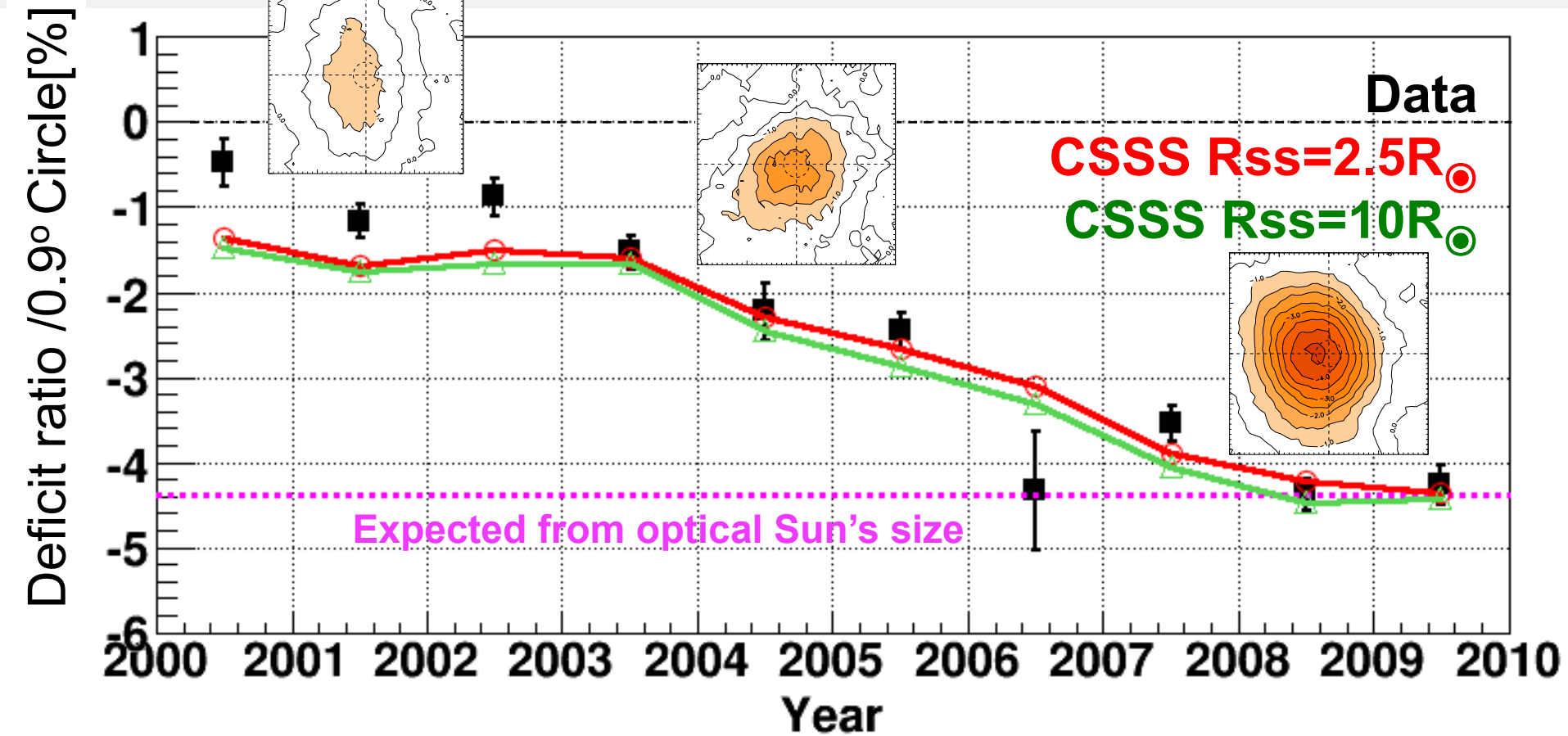
Sun
2000-
2009
Tibet-III
($>3\text{TeV}$)
Moon



→ Shift westward by geomagnetic field
Detector stability calibration



Deficit – Obs/MC All Data - 3 TeV



χ^2 test :

$$\chi^2 / \text{dof} = 32.1 / 10 (3.4\sigma)$$

$$\chi^2 / \text{dof} = 46.9 / 10 (4.8\sigma)$$

*only stat. error

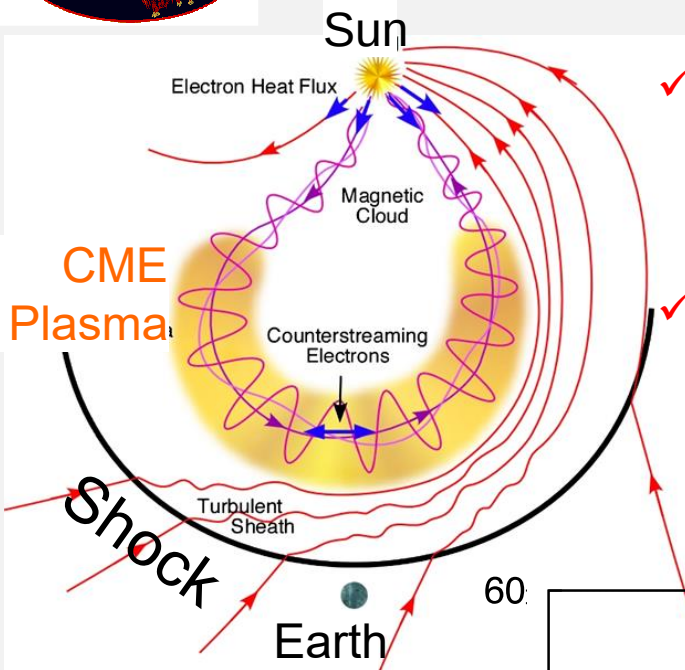
CSSS does not reproduce well at the solar maximum

Influence of CMEs?



CME Catalog

Richardson & Cane, Solar Phys (2010)



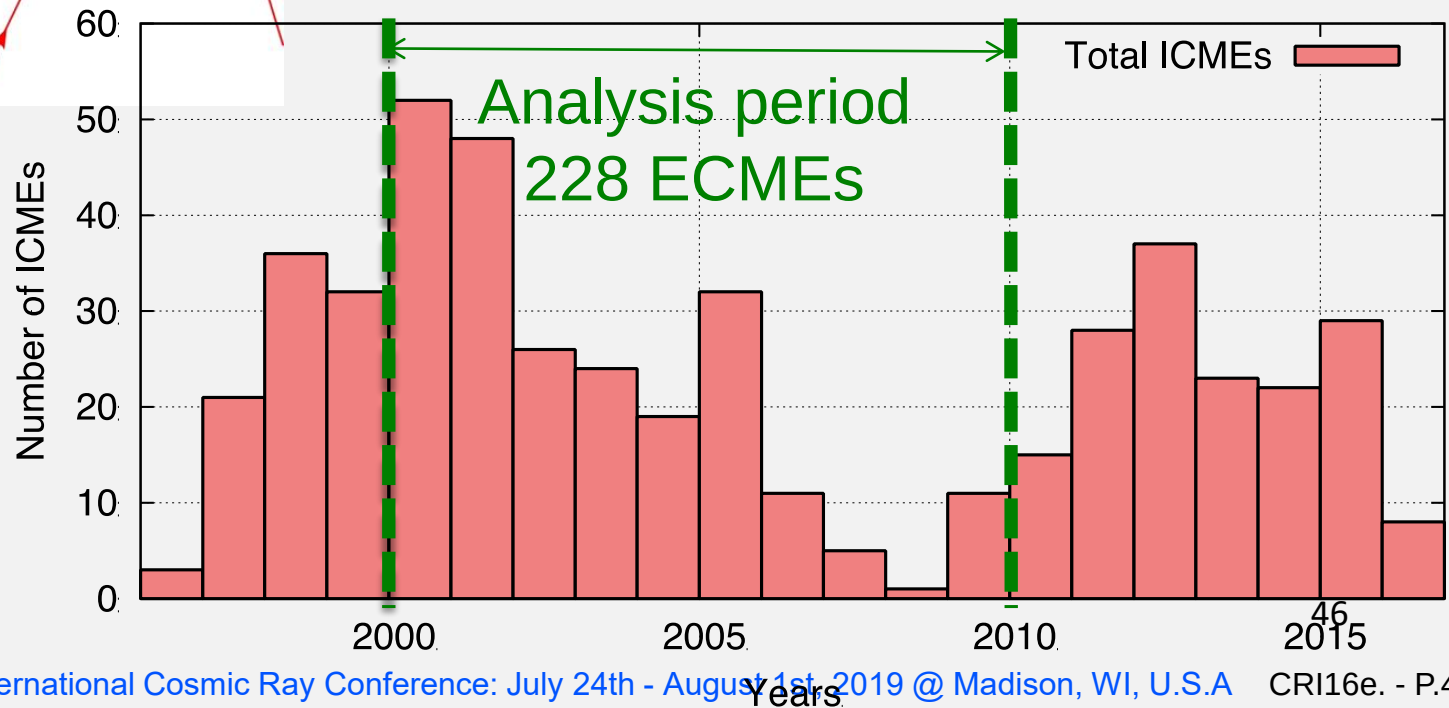
✓ Richardson & Cane Catalog

- CME plasma information observed at the earth
- Only Earth-directed CMEs (ECMEs) are listed

✓ Exclude transit periods of ECMEs from the analysis

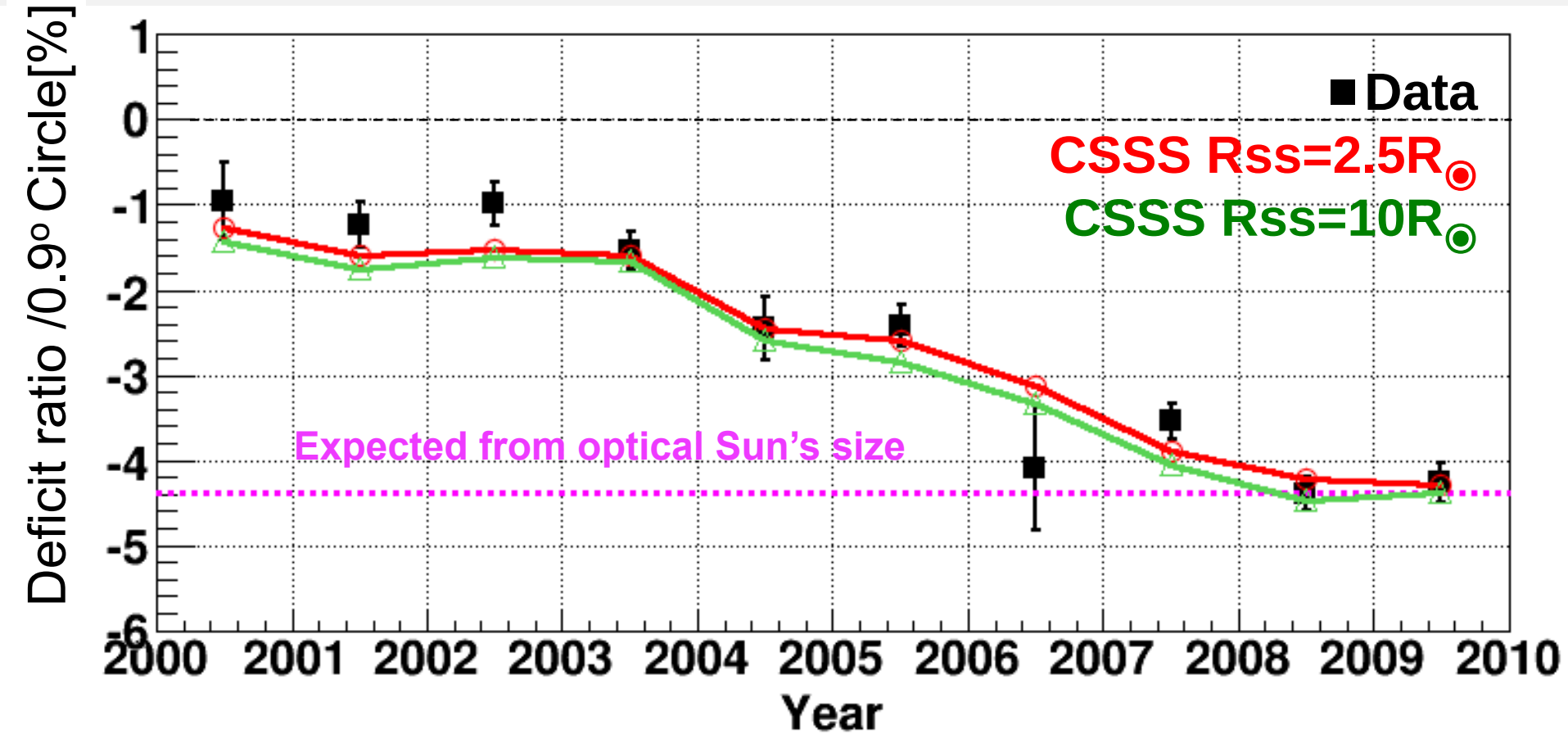
- ECME start is the eruption time at the Sun
- ECME end is plasma end at the earth.
- ECME transit period is ± 4 days

**Coronal
Mass
Ejection**





Deficit – Obs/MC Excluding ECMEs - 3 TeV



χ^2 test :

$\chi^2 / \text{dof} = 12.2 / 10$ (0.6σ)

$\chi^2 / \text{dof} = 21.0 / 10$ (2.0σ)

*only stat. error

Exclude ECMEs → CSSS works

First evidence for influence of
ECMEs on the Sun shadow at 3
TeV

47

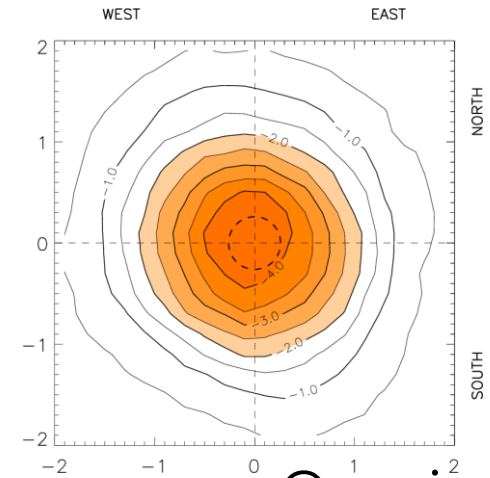
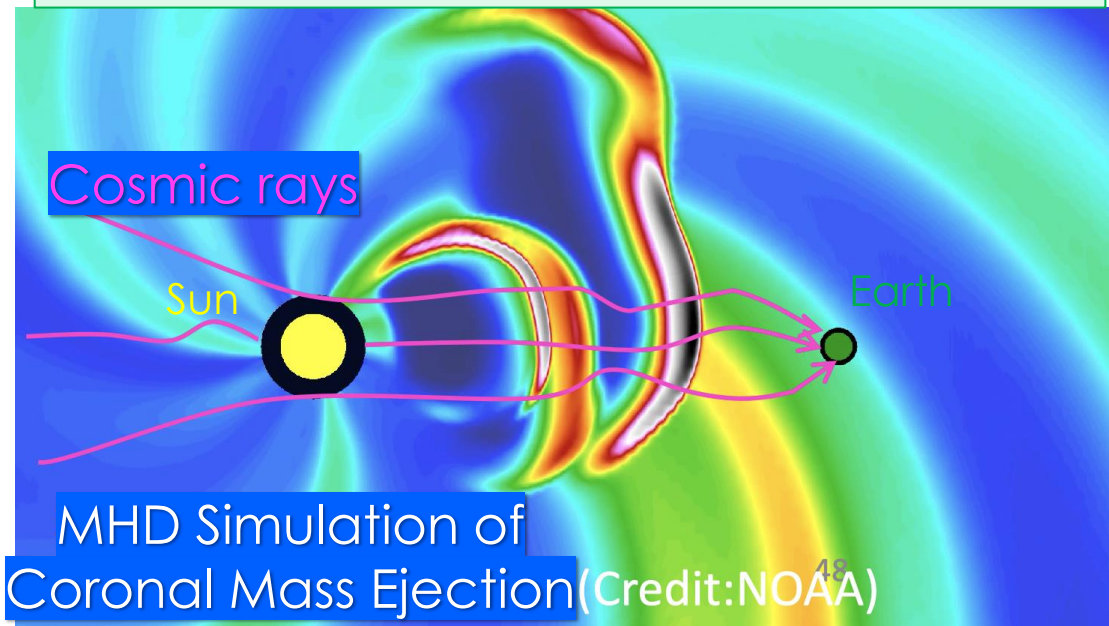
Forecast of CME-> Space Weather

Magnetic Field between Sun and Earth is indirectly observed by cosmic rays!

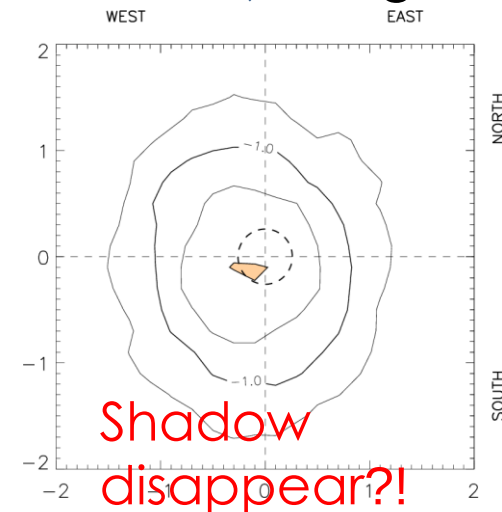
CME takes ~4days from Sun to Earth.

→vs. Cosmic Rays takes ~ 8min.

The Sun Shadow could forecast space weather in principle!



CME occurs ! Cosmic rays scattered by magnetic field





Summary

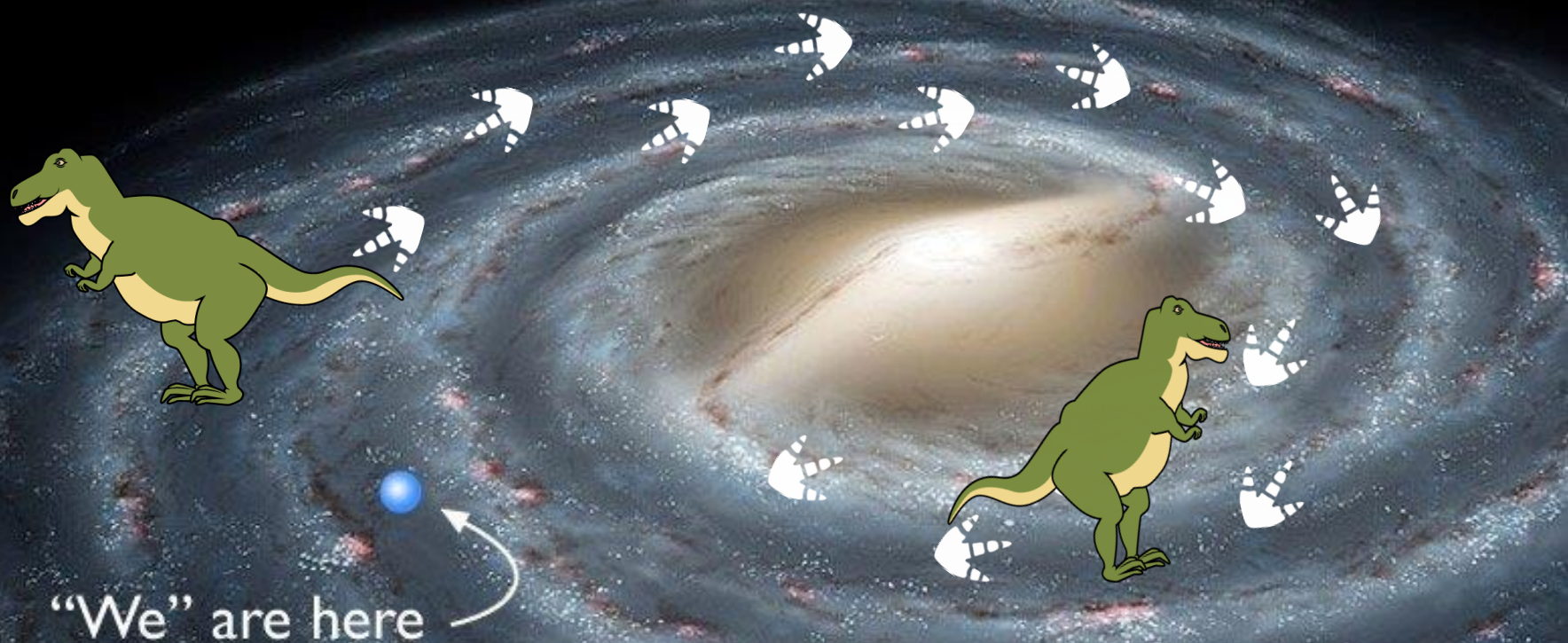
We are Unraveling 60-Year-Old Mystery of Cosmic Ray Origin,

- ✓ Highest energy sub-PeV (0.1-1 PeV) gamma rays
from the Milky Way galaxy
- ✓ Conclusive evidence for existence of PeVatrons
in past/present Milky Way galaxy
- ✓ Experimental evidence for the theoretical model of
high-energy “cosmic-ray pool” in the Milky Way galaxy.

The Sun shadow in Cosmic Rays could forecast Space Weather.



Thank you for your attention!



Dinosaur = PeVatron
Footprints = sub-PeV γ rays