

High Energy Gamma ray A

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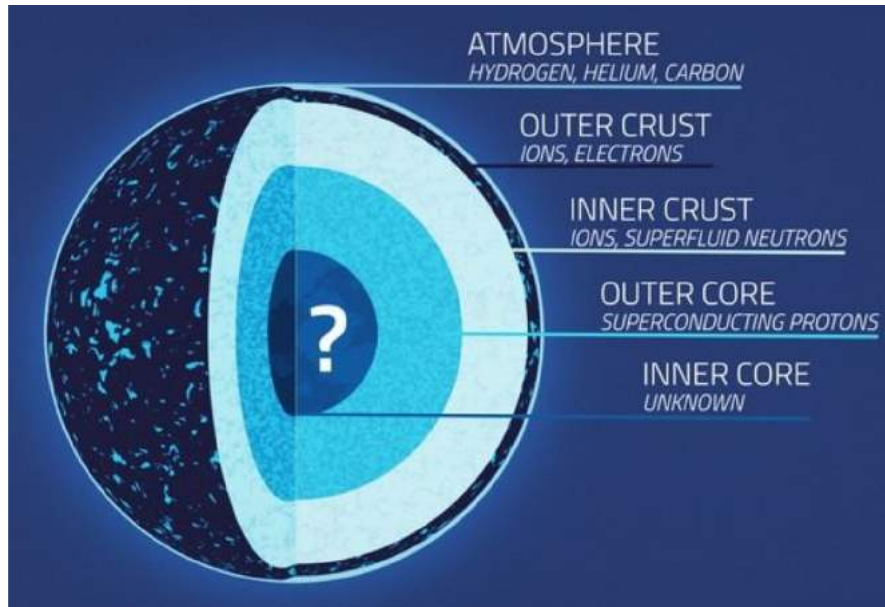
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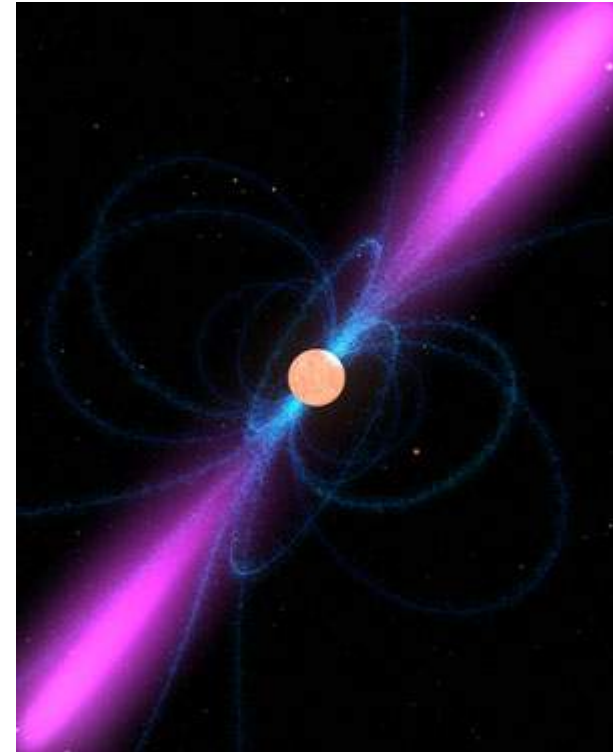
1.Introduction

Neutron stars

- Typical radius: 10 km
- Typical mass: 1 solar mass
- Density: 10^{17} kg/m^3
- Magnetic field: $10^4 - 10^{11} \text{ T}$



Credit: NASA's Goddard Space Flight Center

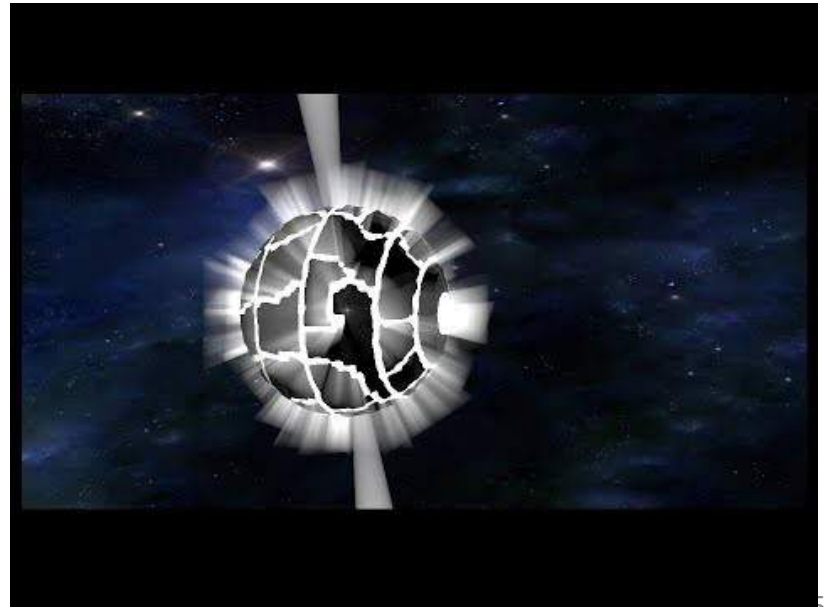


Credit: NASA

Neutron stars are really extreme objects!

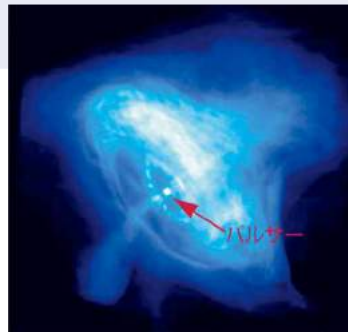
Pulsar

- Very fast rotating neutron stars
- Extraction and acceleration of particles by the powerful magnetic field
- Production of pulsed electromagnetic emission
- Loss of energy due to radiation emission: rotational spindown



We analyze

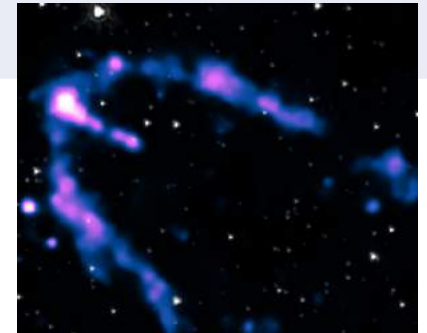
	THE CRAB	VELA	GEMINGA
Age [years]	964 (Supernova in 1054 A.D.)	~10.000	~300.000
Distance [ly]	~7180	~960	~800
Rotational period [ms]	~33	89	237



Credit: JAXA



Credit: NASA



Credit: NASA 6

The Fermi Gamma-ray Space Telescope

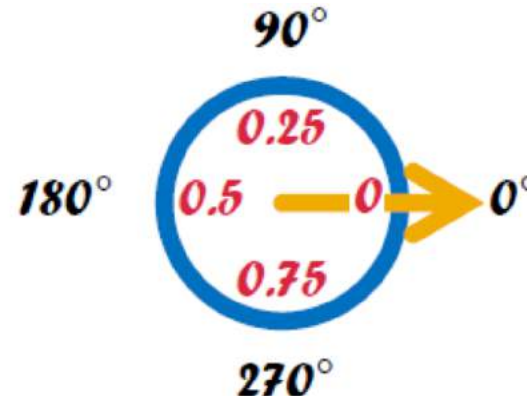
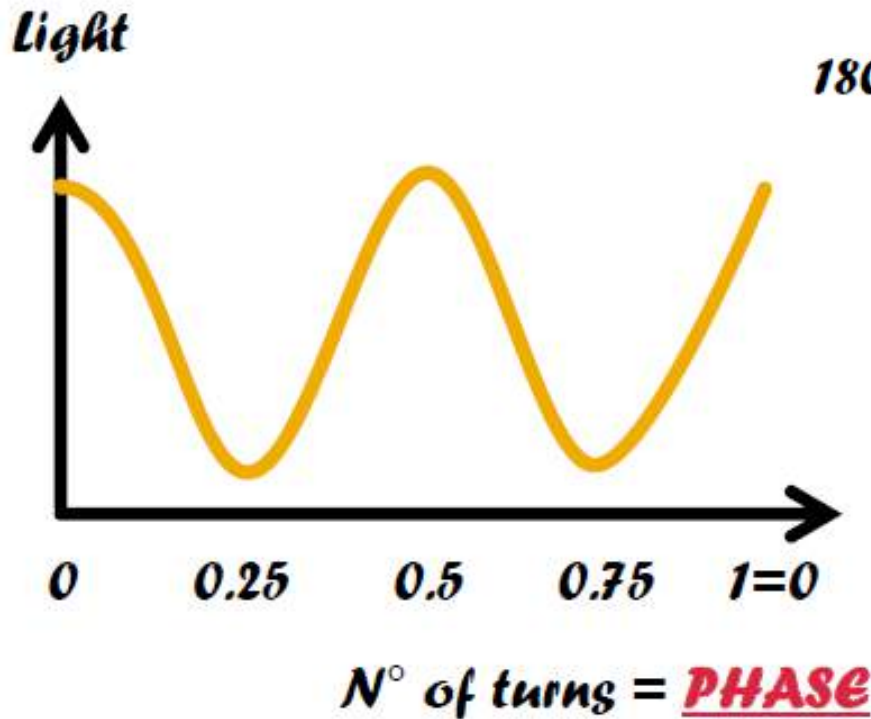


- Mission type:
Gamma-ray observatory
- Launch date:
11 June 2008, 16:05 UTC
- Atitude: 525.9 - 543.6 km
- Period: 95.33 min

It is the most sensitive
gamma-ray observatory
in orbit!

What's phaseogram?

- Phaseogram:
Sum up events that happen at the same phase
- Phase:



Physical information
on the pulsar!

Our mission

1. Learn about pulsar, Fermi, the way to analyze and etc...
2. Produce a phaseogram
3. Decide frequencies and frequencies derivatives
4. Analyze something strange



2.How to analyze

How to analyze

Download
Each Data



The
selection of
exact
photons



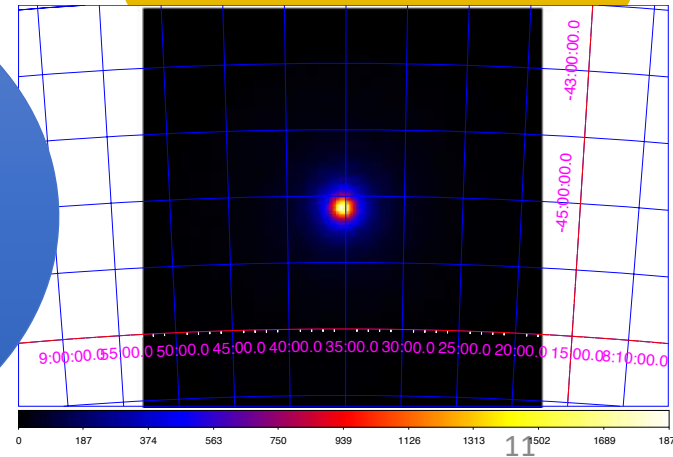
Make
Count Map
of
The Crab,
VELA and
Geminga



Energy
range

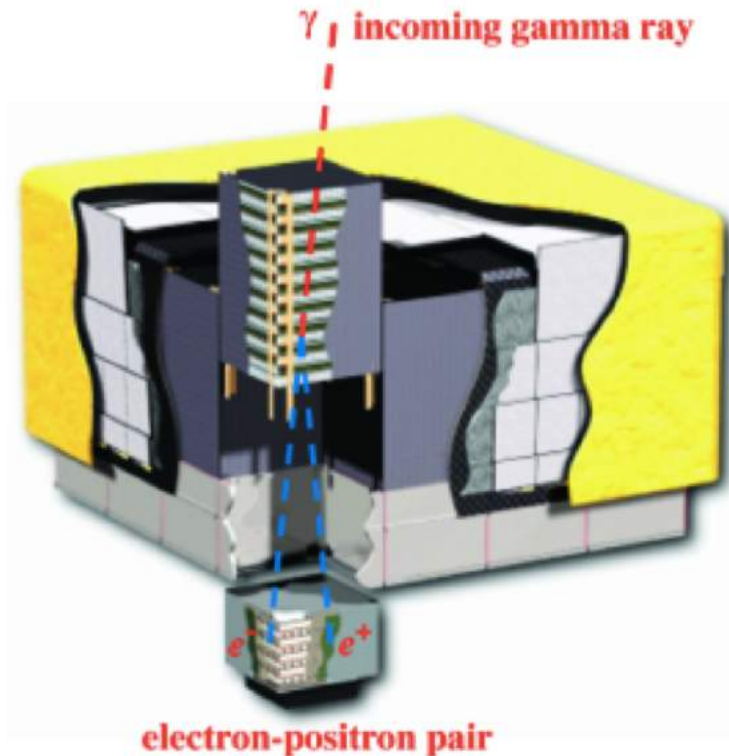
300 ~

30000 MeV

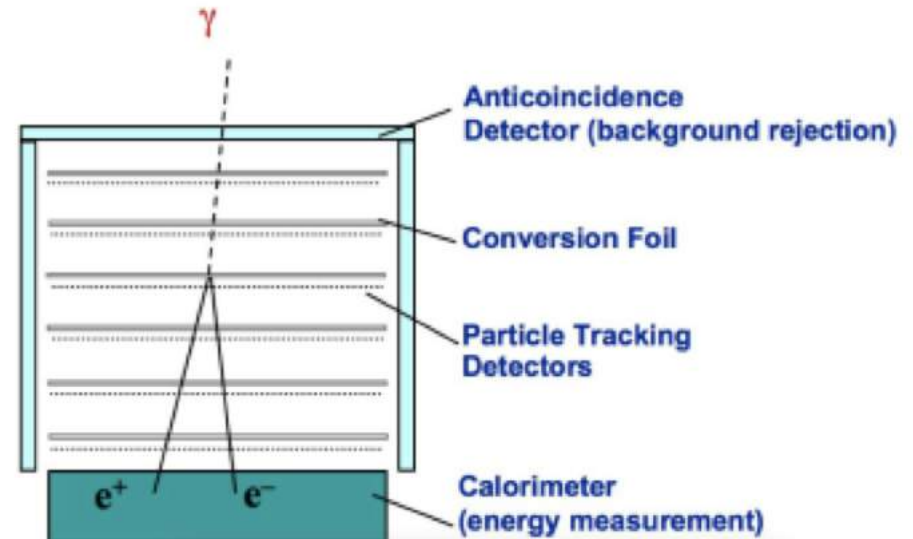


How to get data

Fermi



LAT



What can we get?

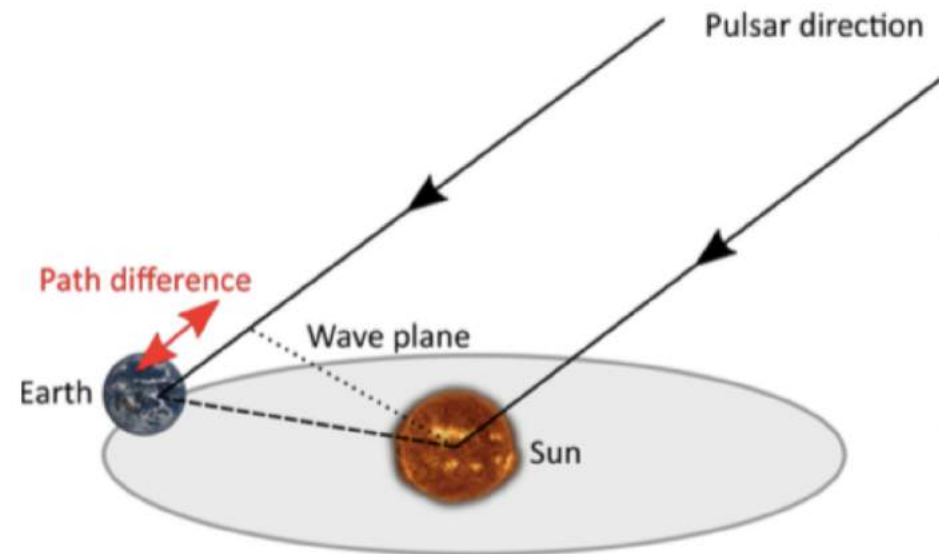
- The direction from which photons come.
- To measure amount of energy of one photon.

How to define the time

- What should we think about to define the correct time:

→ **Barycentric Collection**

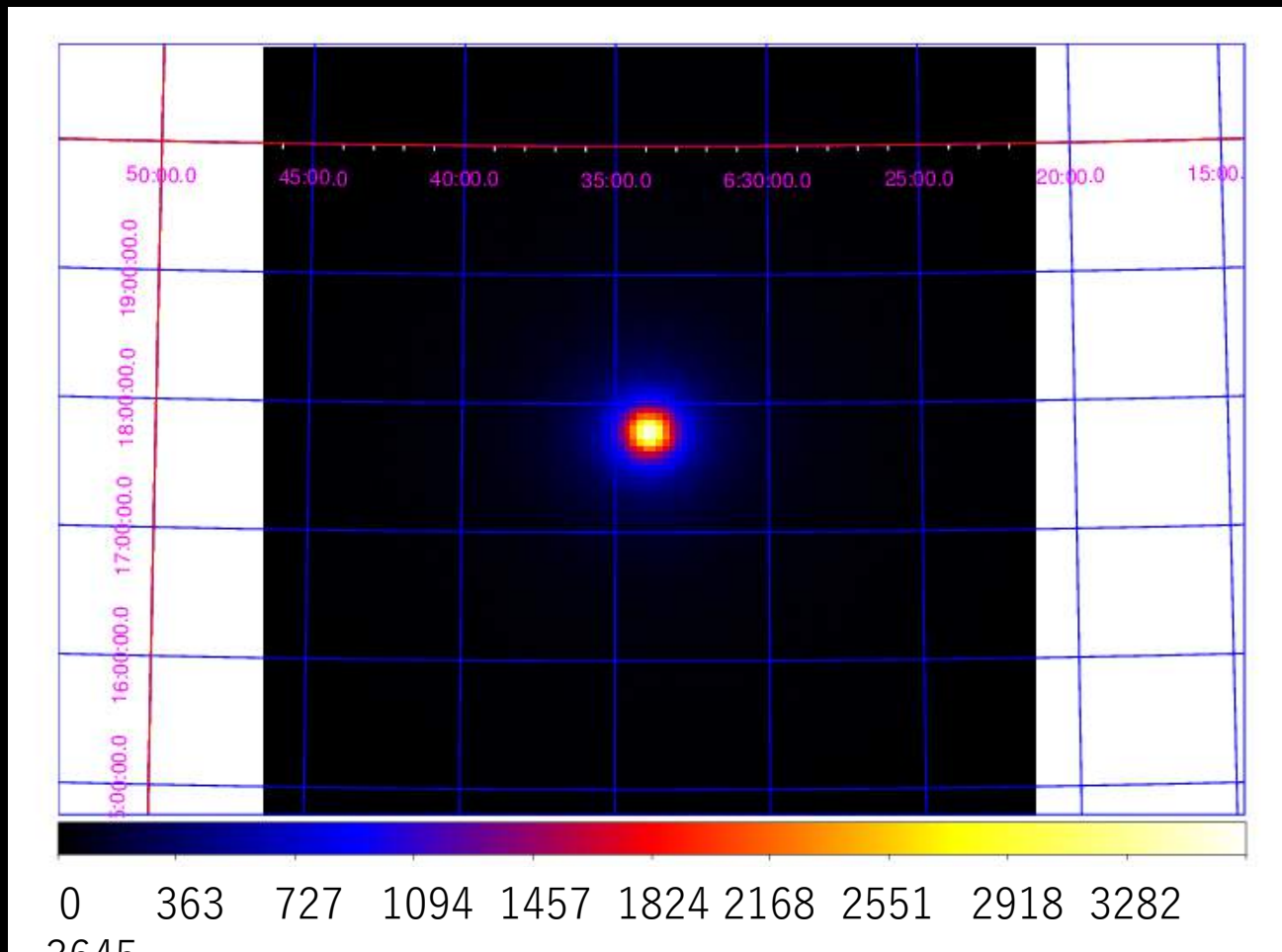
- **Geometrical** correction (Rømer delay)
- **Special relativity** correction (Einstein delay)
- **General relativity** correction (Einstein, Shapiro delay)



3.Result

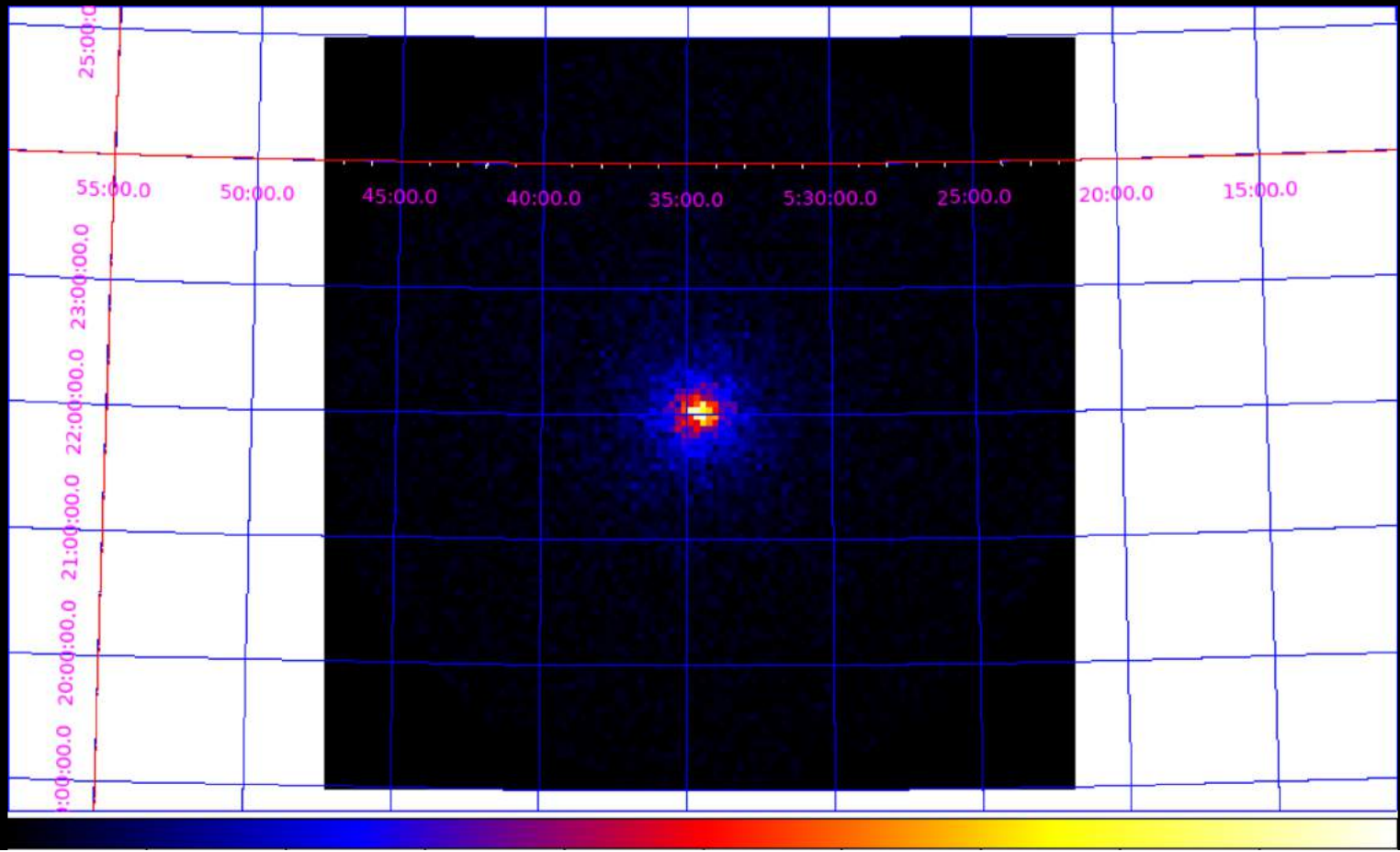
Counts Map ① Geminga

2008-08-04 ~ 2018-03-06



Counts Map ② Crab (A)

2017-01-01 ~ 2018-07-31

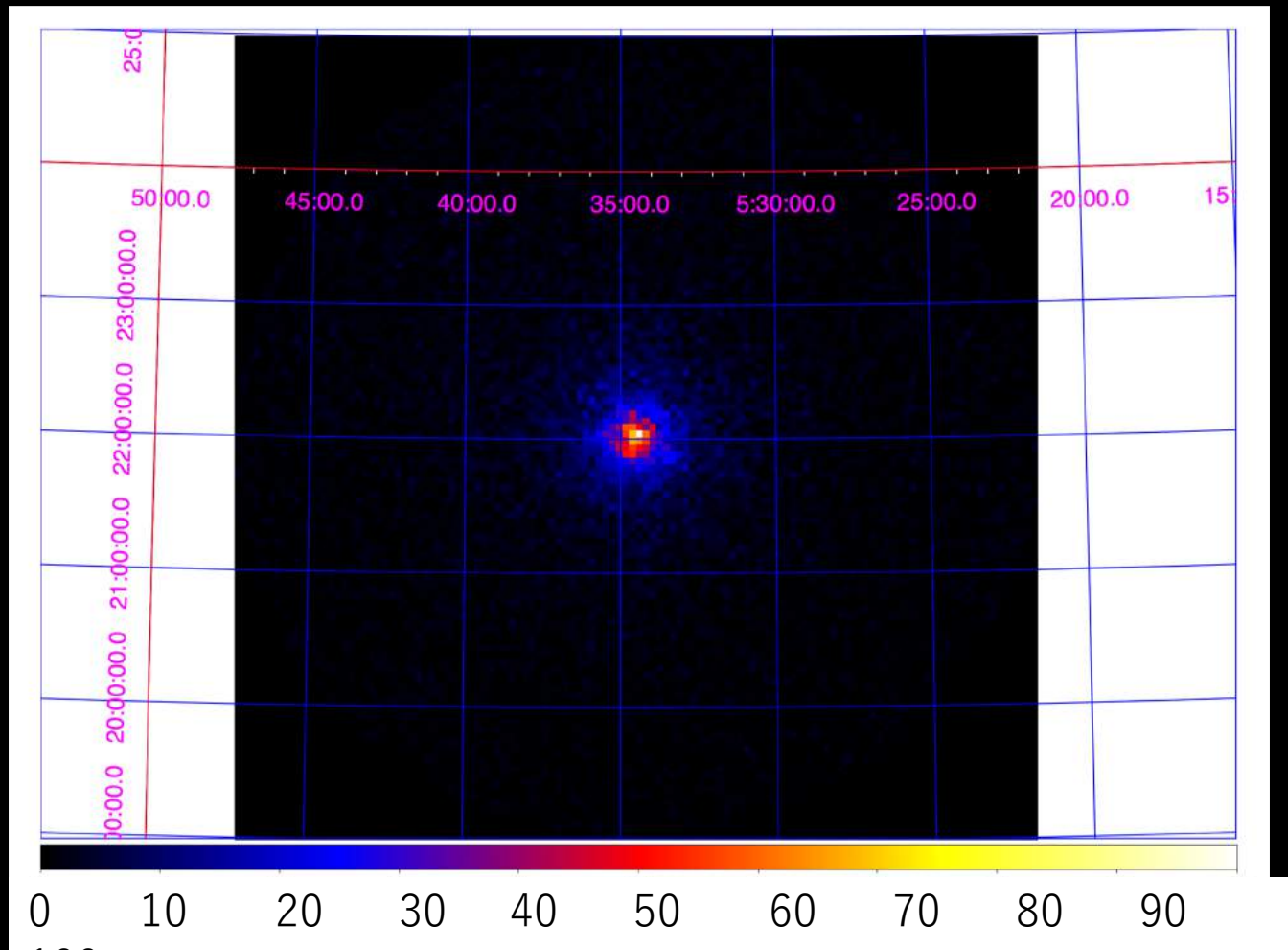


0 7.8 16 23 31 39 47 55 62 70

Counts Map

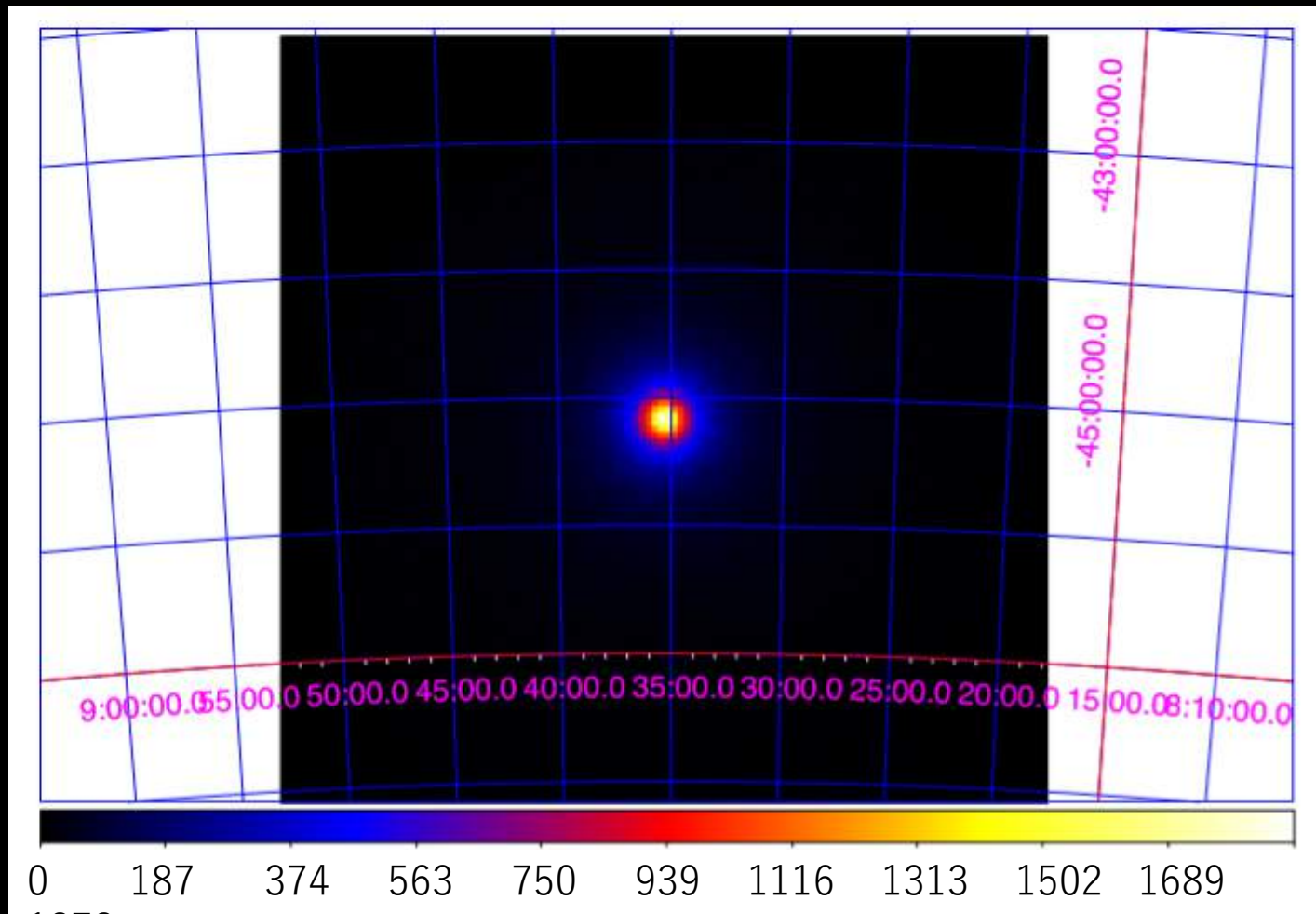
③ Crab (B)

2017-08-01 ~ 2018-03-06

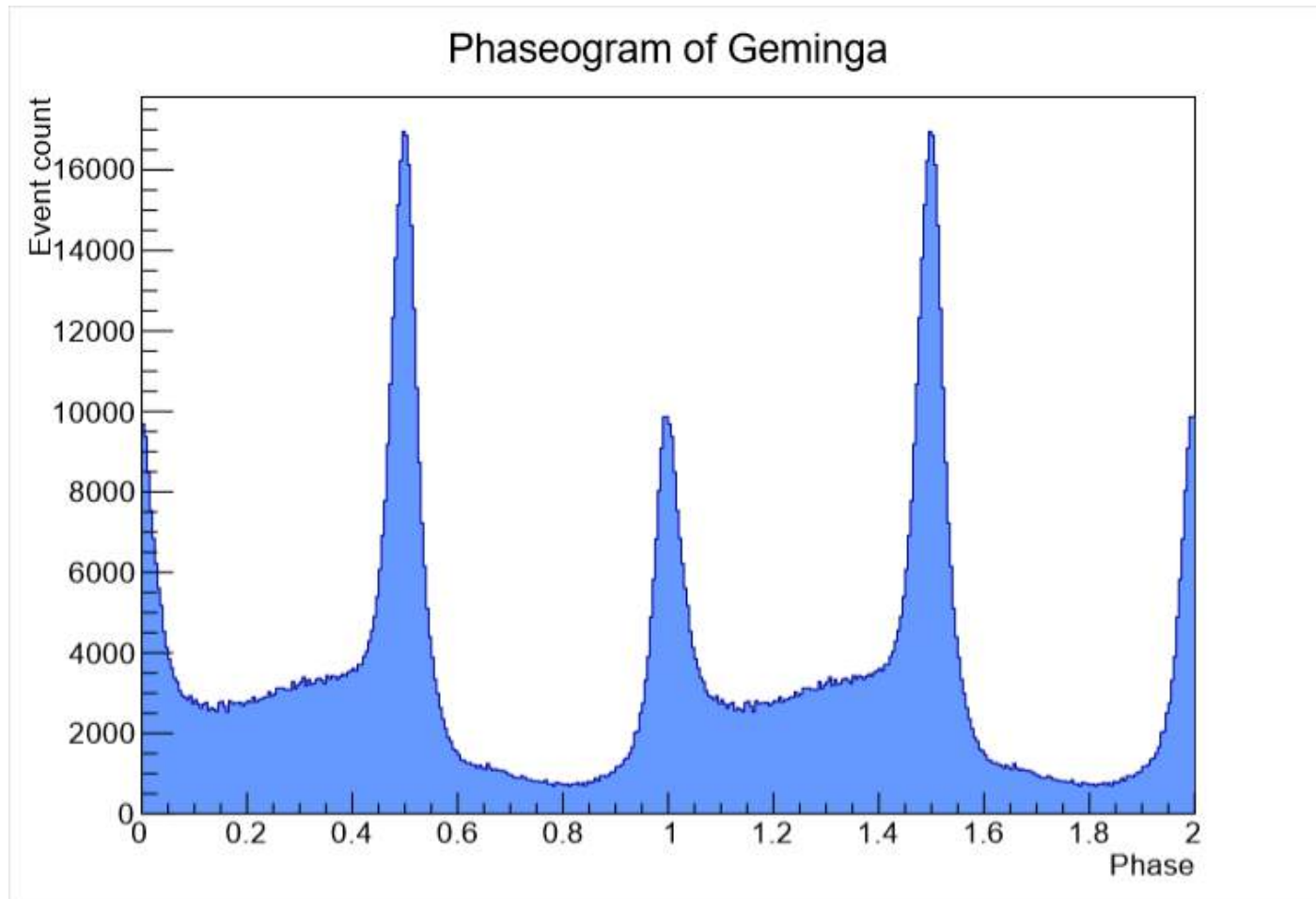


Counts Map ④ Vela

2008-11-30 ~ 2011-02-08

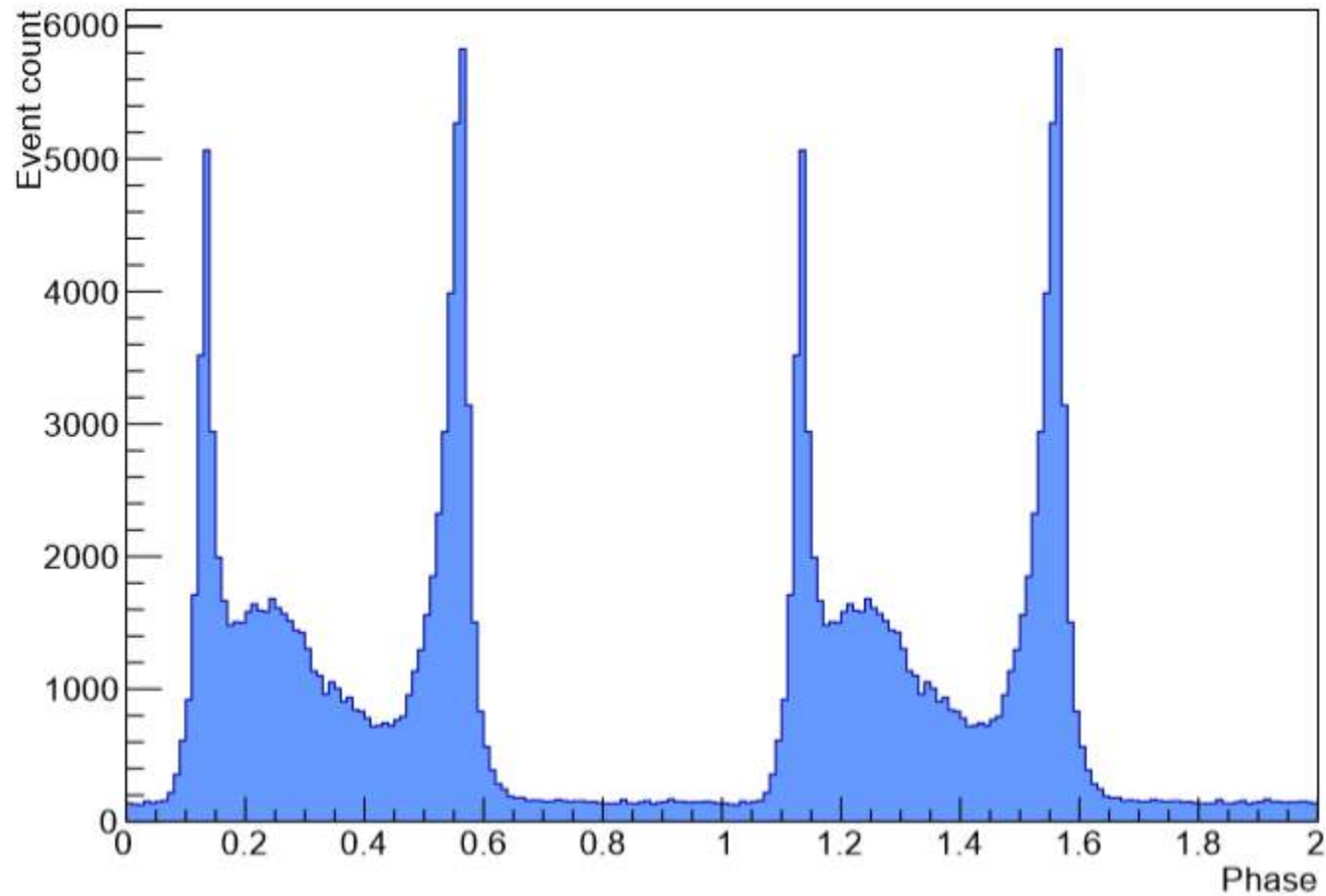


Shape of Pulse (Geminga)



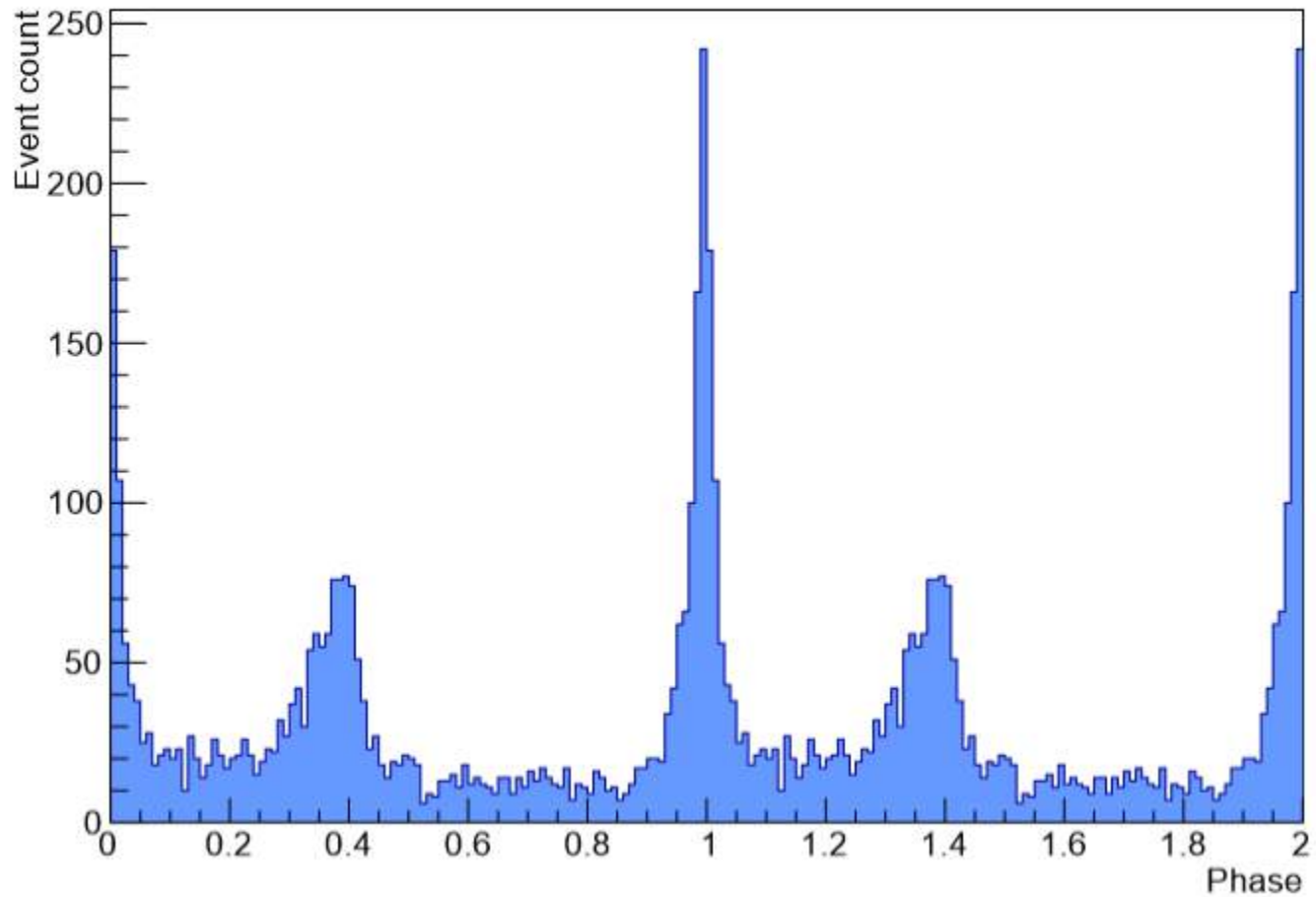
Shape of Pulse (Vela)

Phaseogram of Vela



Shape of Pulse (Crab)

Phaseogram of Crab



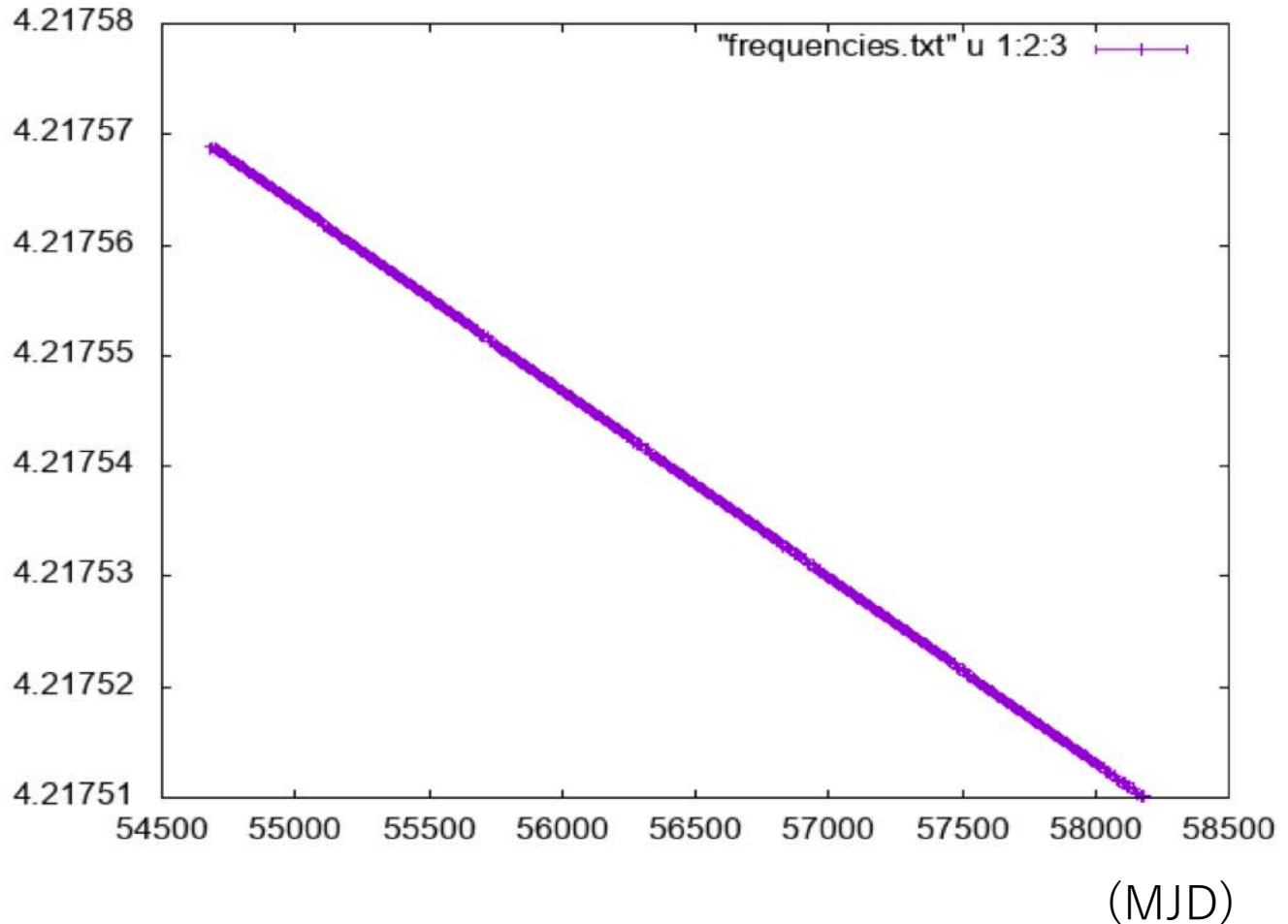
Frequency is decreasing...

Observation Time is for **about 10 Year !!**

Geminga

2008 Aug 4
– 2018 Mar 6

Frequ
ncy
[Hz]

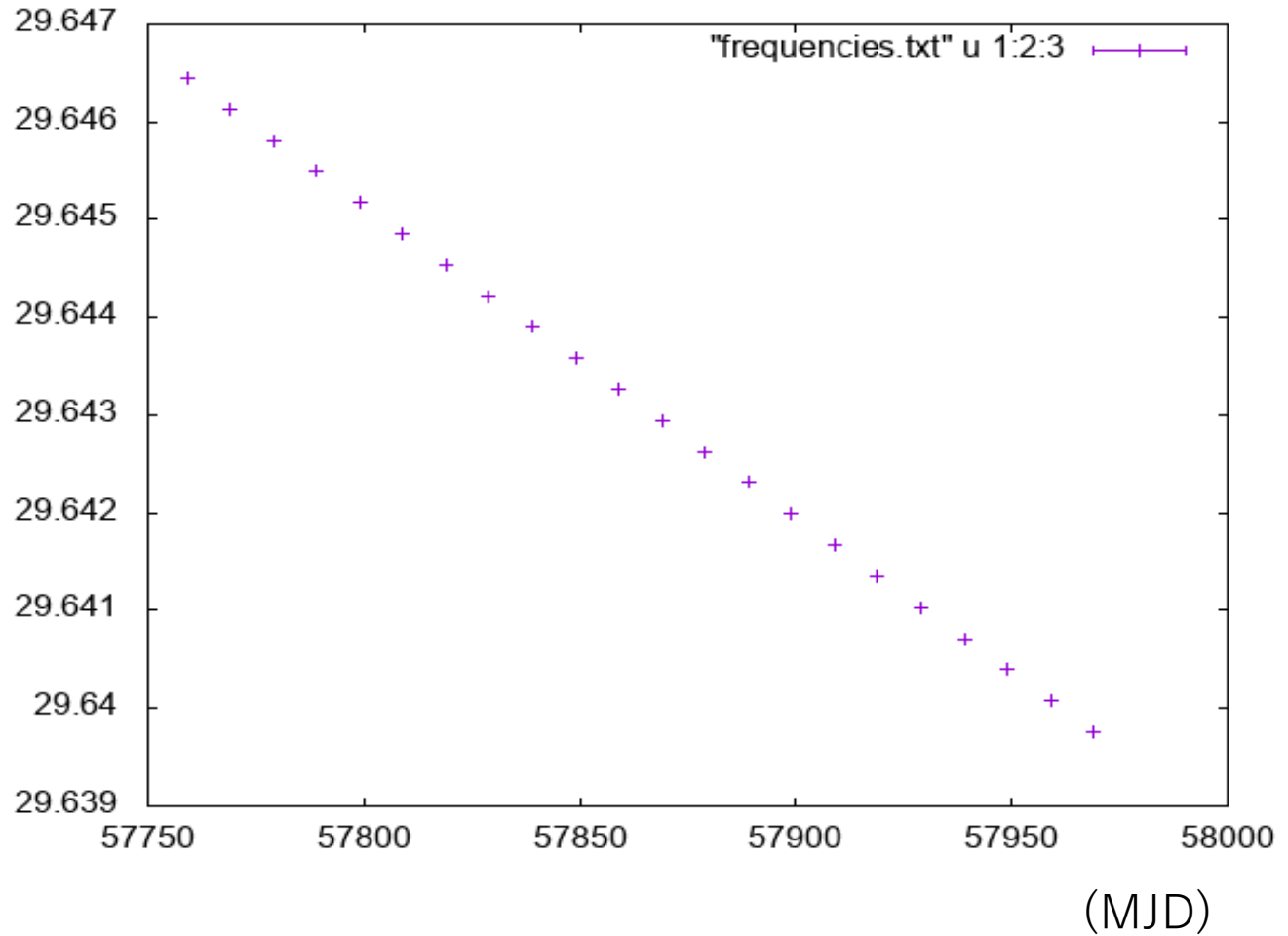


Frequency is decreasing...

Crab

2017 Jan 1
- 2017 July 31

Frequency
[Hz]

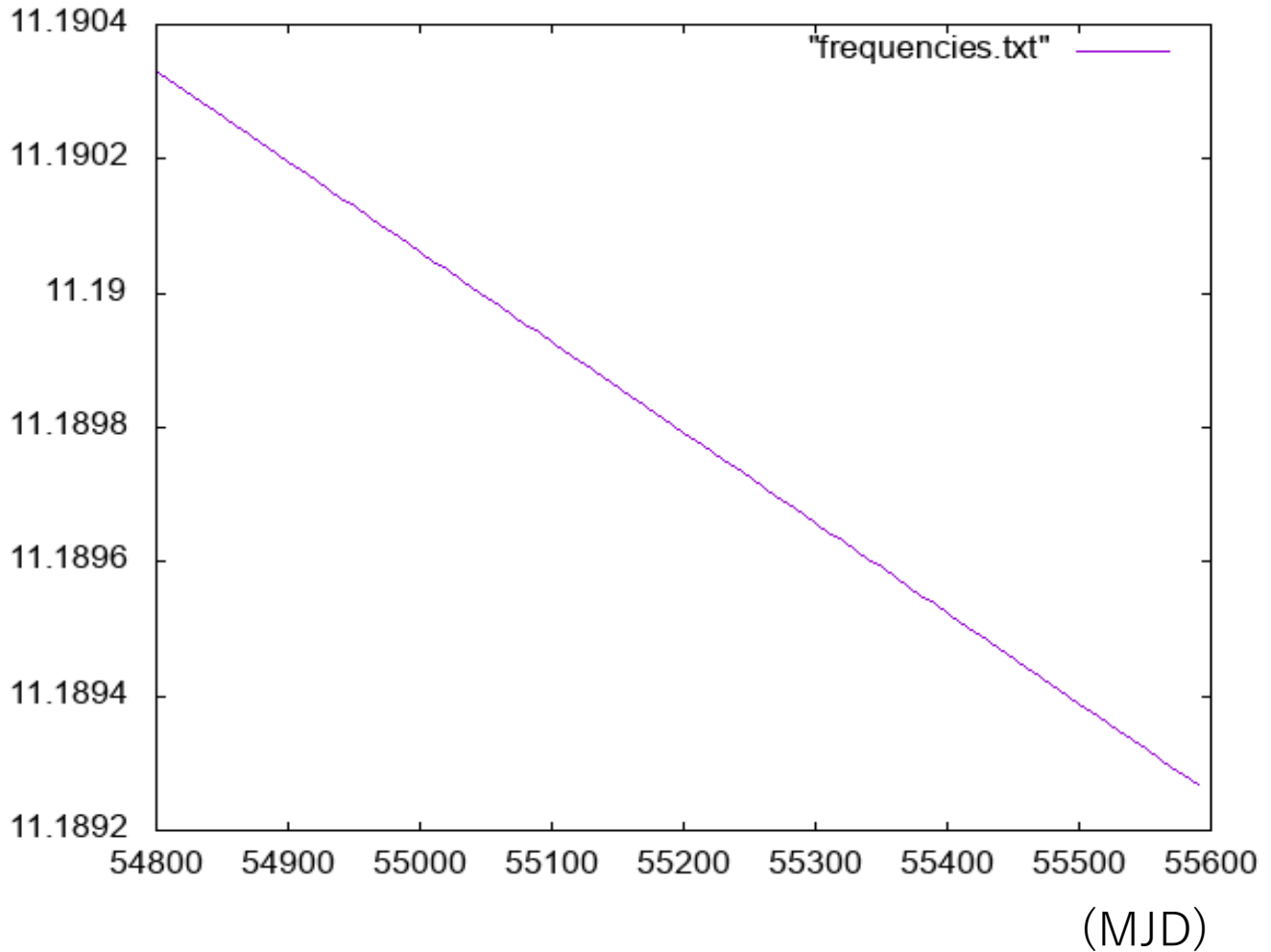


Frequency is decreasing...

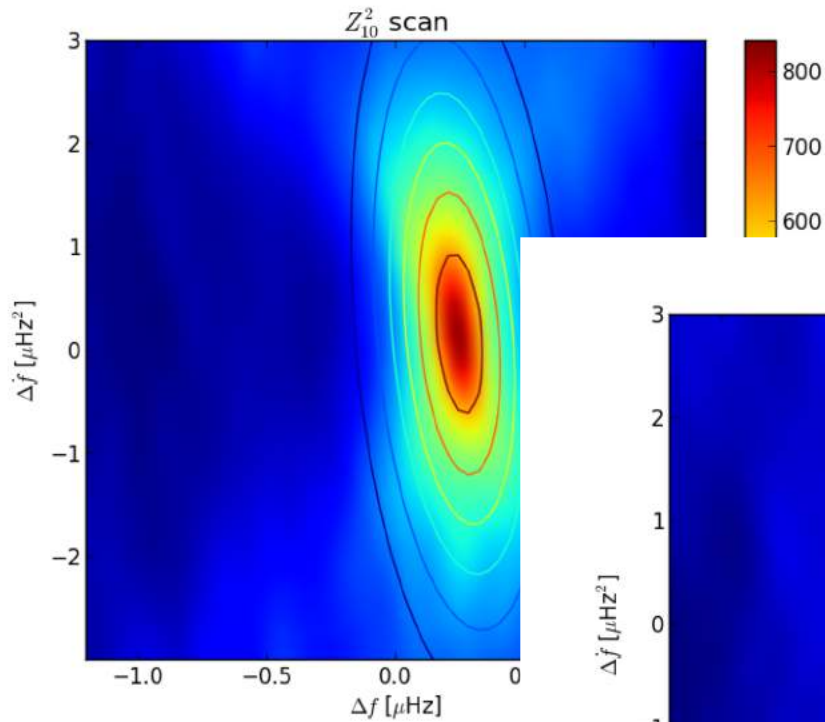
Vela

2008 Nov 30
- 2011 Feb 8

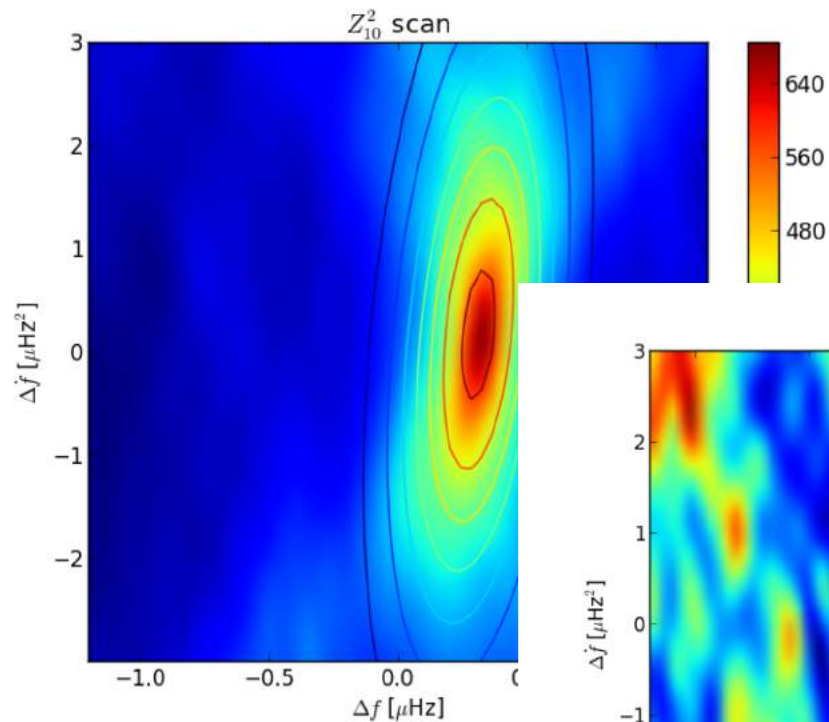
Frequency
[Hz]



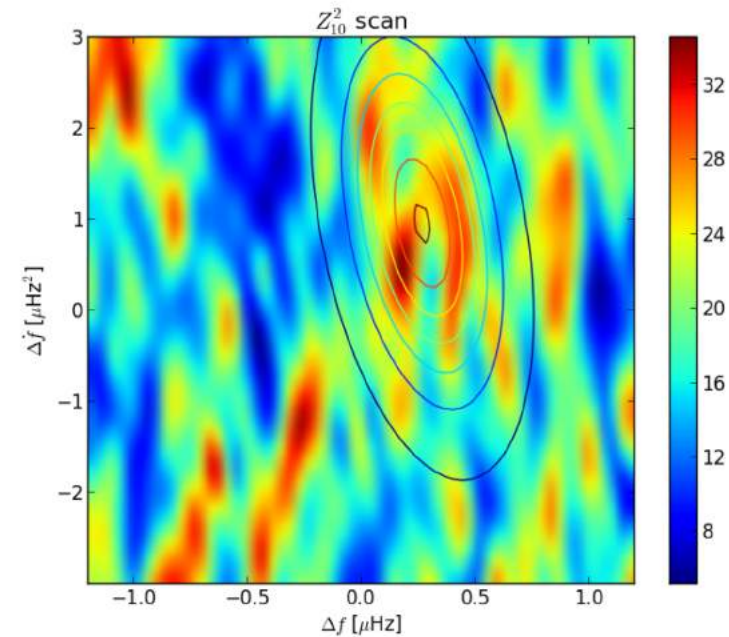
Frequency ...changed ?



MJD 58051~60



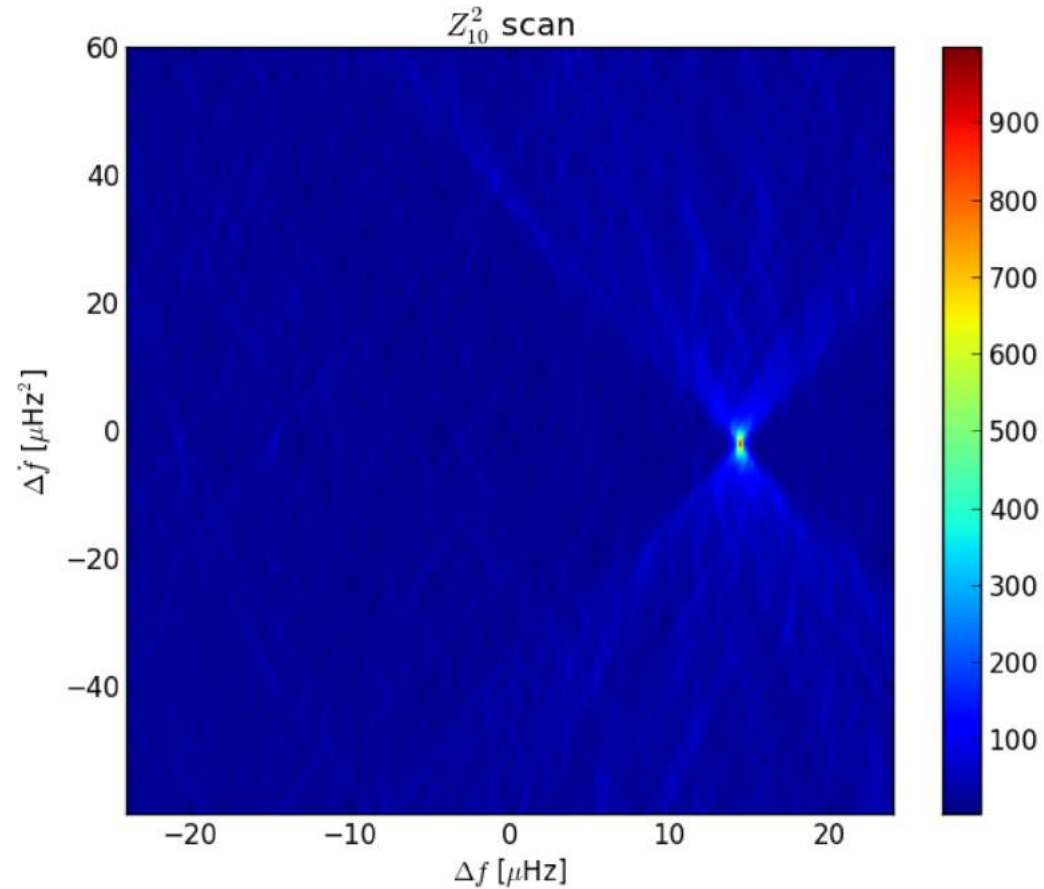
MJD 58061~70



MJD 58071~ 80

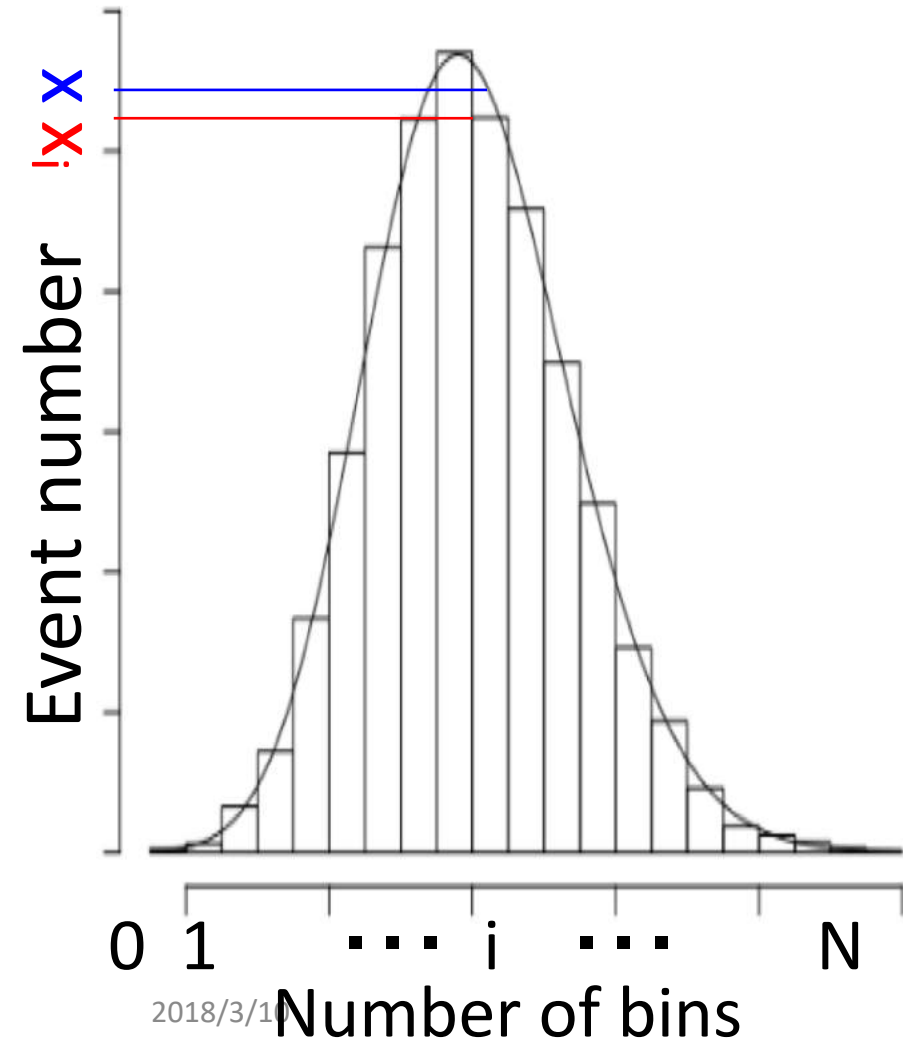
Frequency ...changed ?

MJD 58071~ 80



4. Discussion

Chi-Square test



Evaluation of deviation between experimental data and model

$$\chi^2 = \sum \frac{1}{N-1} \frac{(x_i - X)^2}{(\sqrt{N})^2}$$

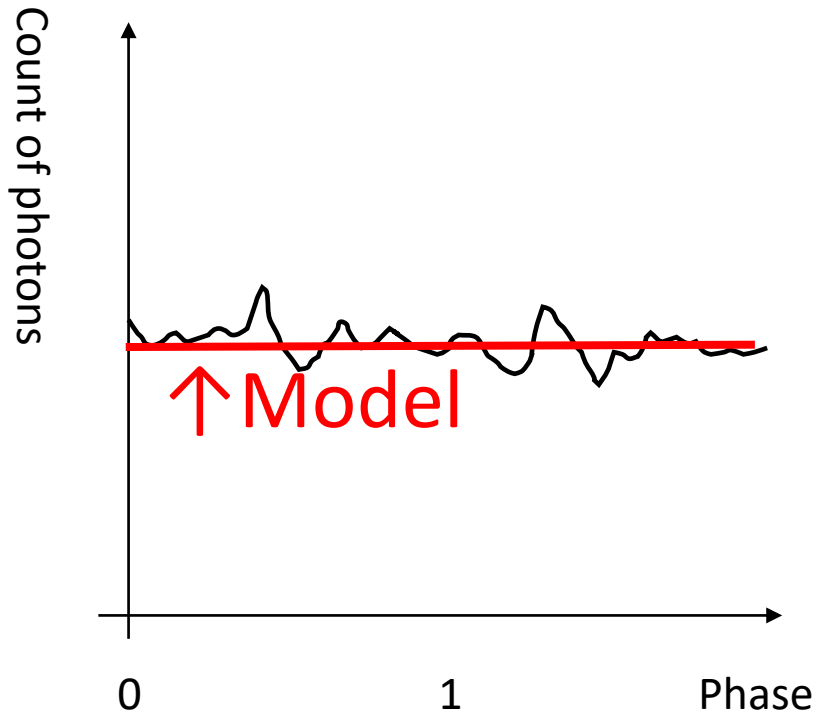
N: Number of bins

x_i : Counts of photons
(experimental data)

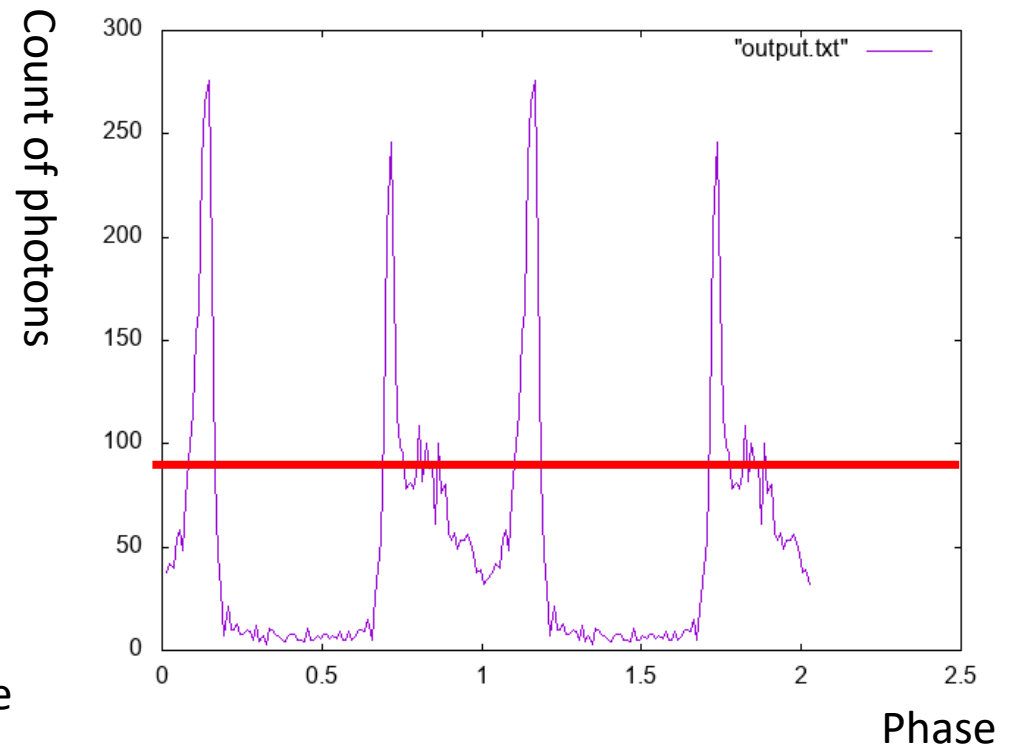
X: Counts of photons
(model data)

Statistical test of pulsation

$$\chi^2 \sim 1$$



$$\chi^2 \gg 1$$

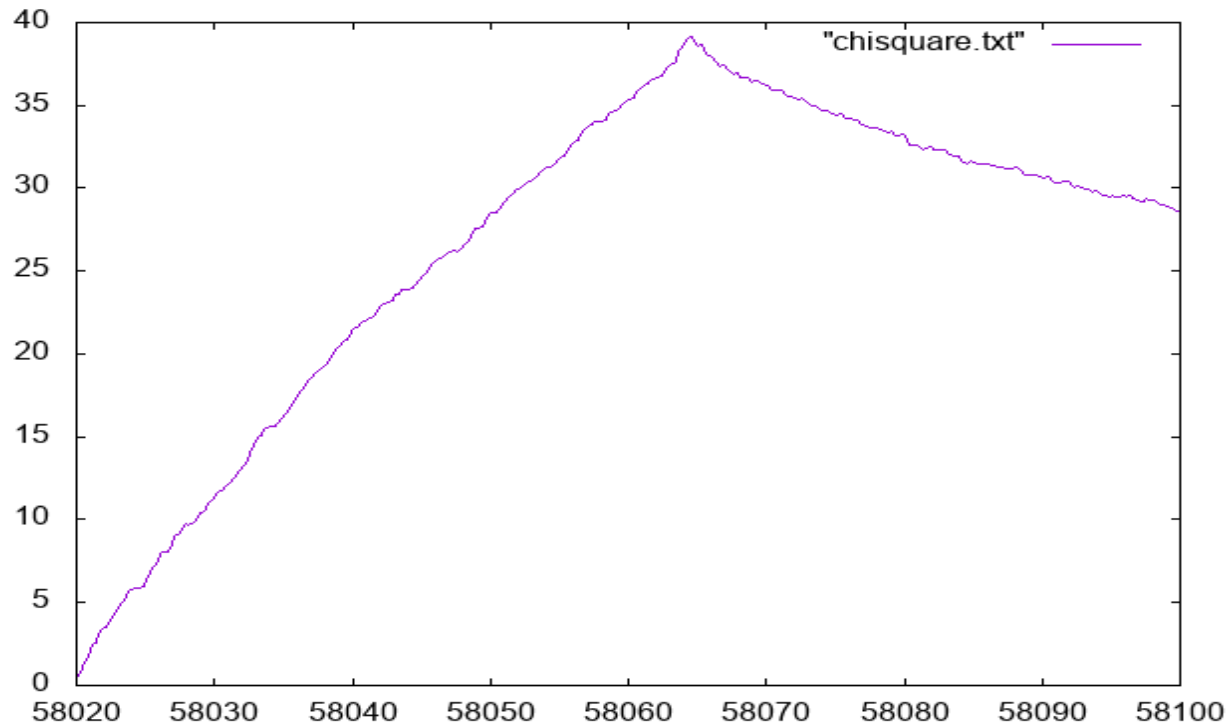


How to fit

We tried to know when the event happened with chi square statistics.

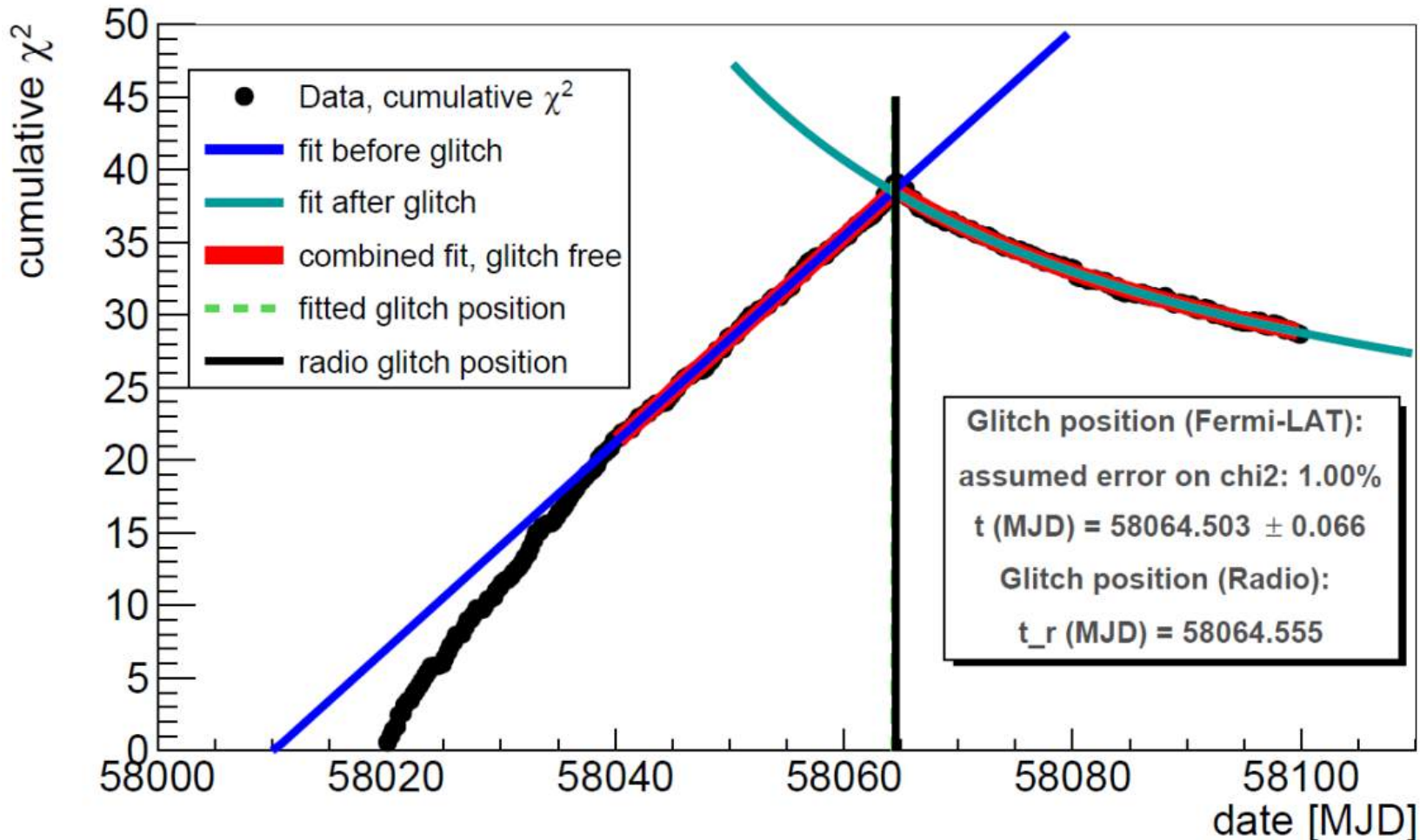
We discovered the edge in graph.

→ It is possible that at this time, the event occurred.



Decide when the event happened

Crab Pulsar Glitch



We investigated the edge's time by changing two line.

→ The event happened at “ 58064.503 ± 0.066 ”.

Compare with other wave area

We checked

“The Jodrell Bank Observatory with the 42-ft Lovell Telescope(1400 MHz)” .

Reference:

Shaw+2017(Atel#10939)

→The event happened at
58064.555(3)(MJD)
(58064.503(66))

⇒**The two data are consistent within a statistical error.**



https://en.wikipedia.org/wiki/Lovell_Telescope#/media/File:Lovell_Telescope_5.jpg

What was the event?

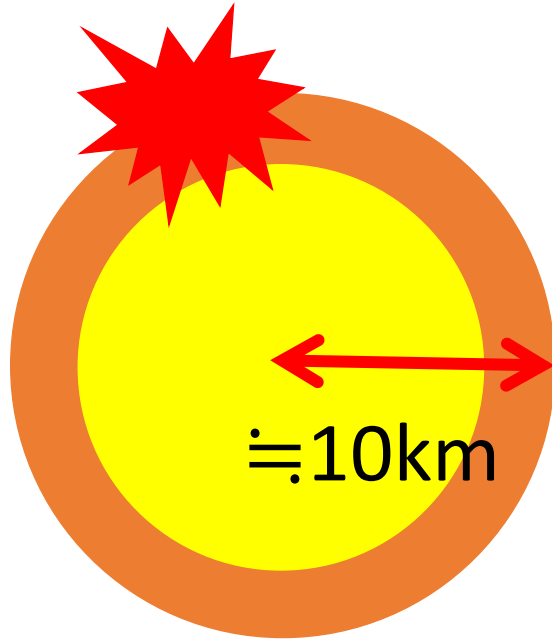
We discovered the two things.

1. The gamma ray from the Crab pulsar changed and the time is consistent in two wave areas.
2. The pulsar's frequency increased after the event.

⇒ We expected that the event was “GLITCH”.

What is happened on the CRAB?

STARQUAKE!!!!



- We think ...

Starquake

on the CRAB.

- If we assume that the preservation of angular momentum is valid...

$$\frac{\Delta R}{R} \sim \frac{\Delta \nu}{\nu}$$

$\Delta R = 5 \sim 10 \text{mm}$ Sedimentation

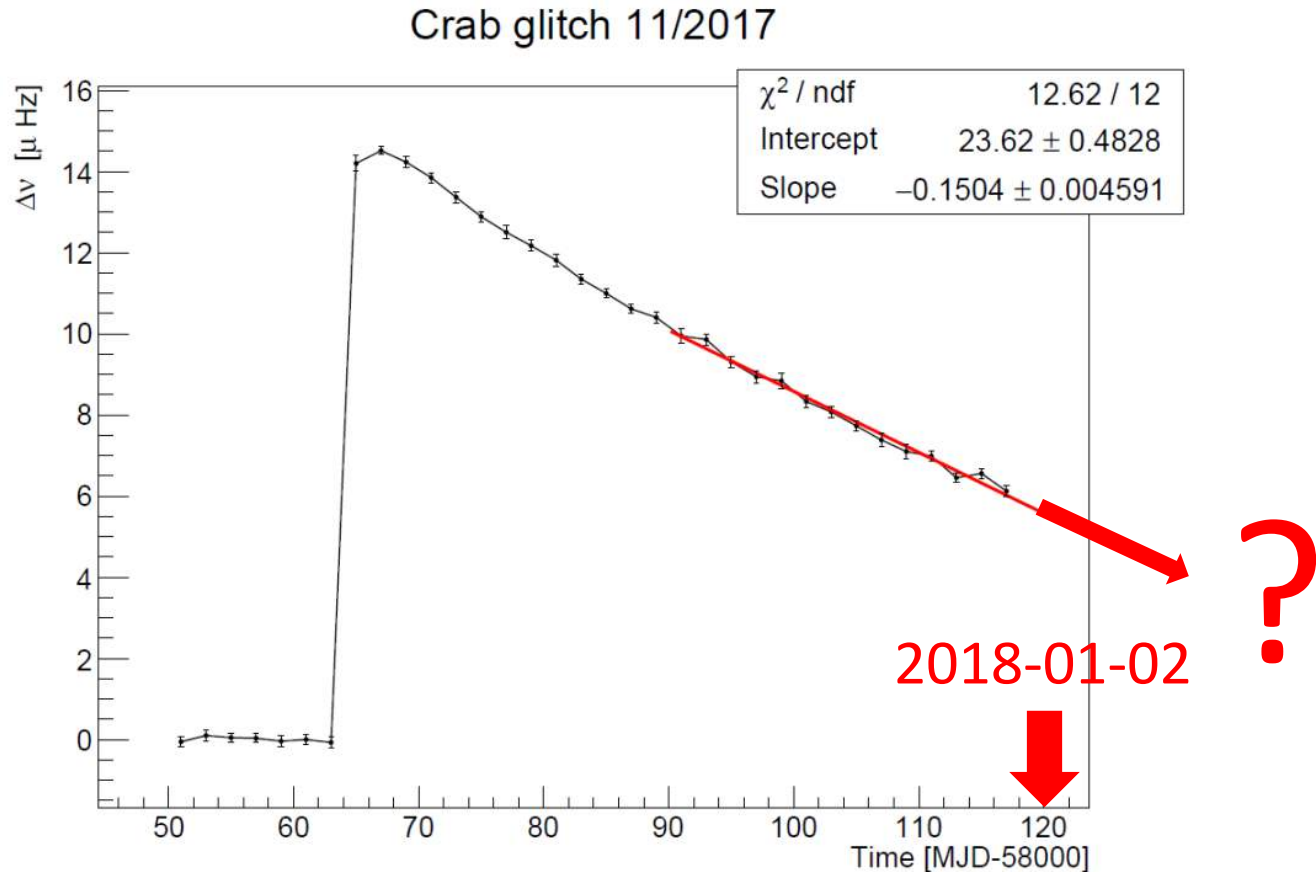
5. Summary

How much did the frequency change?

We investigated how much the frequency changed.

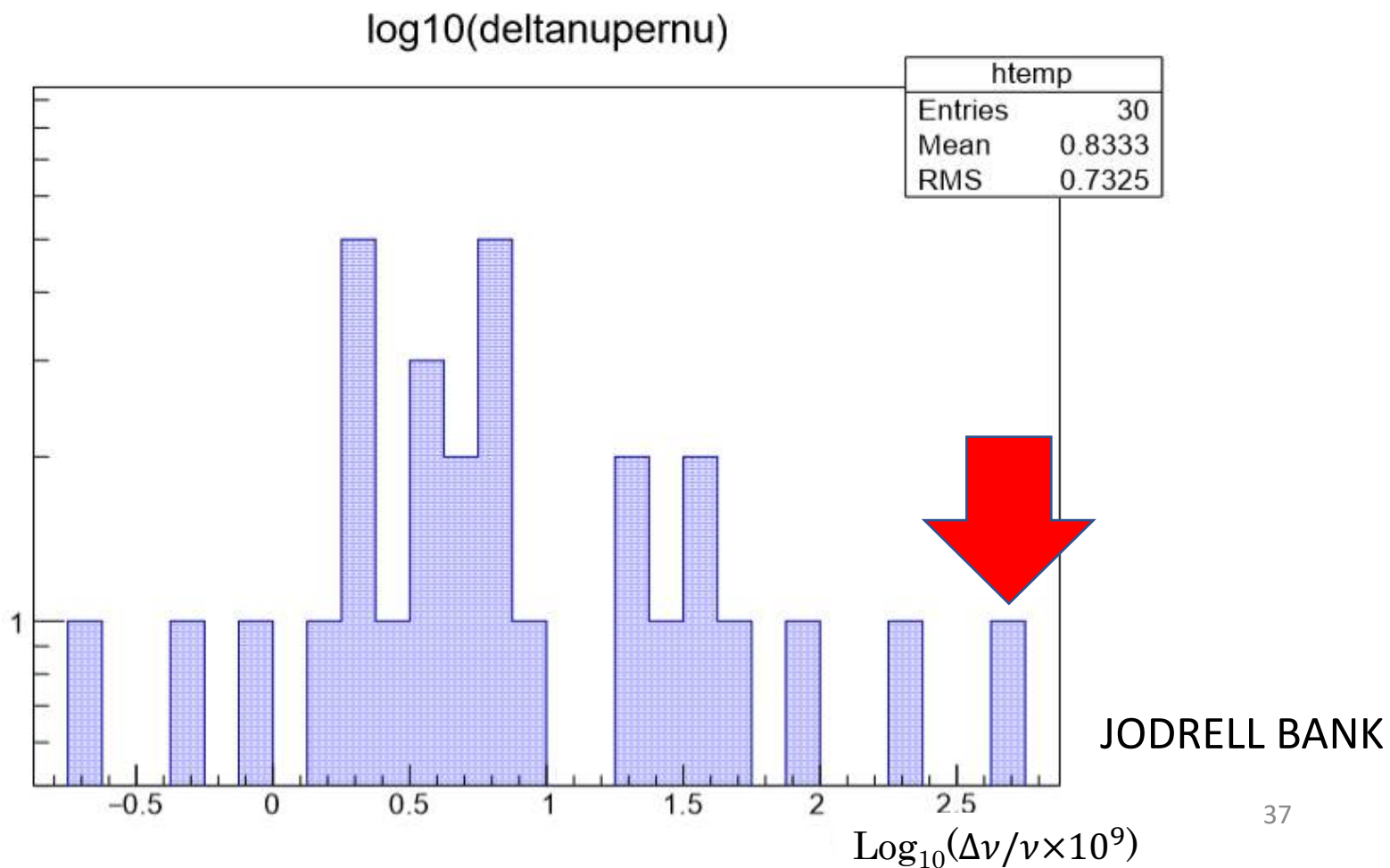
→The frequency increased by $14\mu\text{Hz}$.

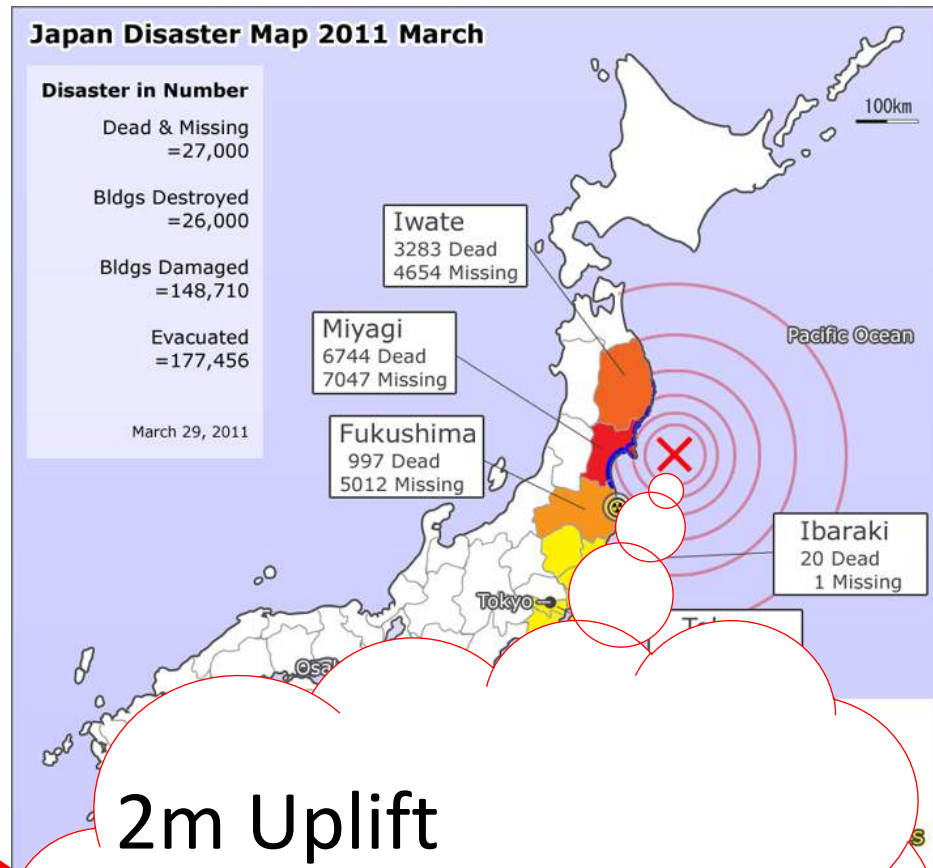
Question? : We need more study !!



We made a histogram of $\Delta\nu/\nu$.

It shows that the Glitch we discovered is the biggest one.





0.01m Sedimentation
 Radius of neutron star
 $\cong 10^4\text{m}$
 Magnitude = 23.5Mw

2m Uplift
 Earth radius $\cong 10^6\text{m}$
 Magnitude = 9.0Mw



Thank you for your
attention

Back up

delta_nu
v.s. MJD

When the Glitch happen?

Cumulative
chi2 v.s.
MJD

Relaxation...???

Glitch recovery
(exponential
decay in nu-
MJD plot)