



# IceCube 2009 status report in light of the ICRR Inter-university program

Shigeru Yoshida

Chiba University

<http://www.ppl.phys.chiba-u.jp>

# IceCube status

IceCube Lab

50 m

IceTop

80 Strings each with  
2 IceTop Cherenkov Detector Tanks  
2 Optical Sensors per tank  
320 Optical Sensors

2004 Project Start	1 Hole
2009 Current Status	59 Holes
2011 Projected Completion	86 Holes

IceCube In-Ice Array

86 Strings, 60 Sensors  
5160 Optical Sensors

AMANDA-II Array  
(Precursor to IceCube)

Deep Core

6 Strings - Optimized for low energies  
360 Optical Sensors



Eiffel Tower  
324 m

1450 m

2450 m

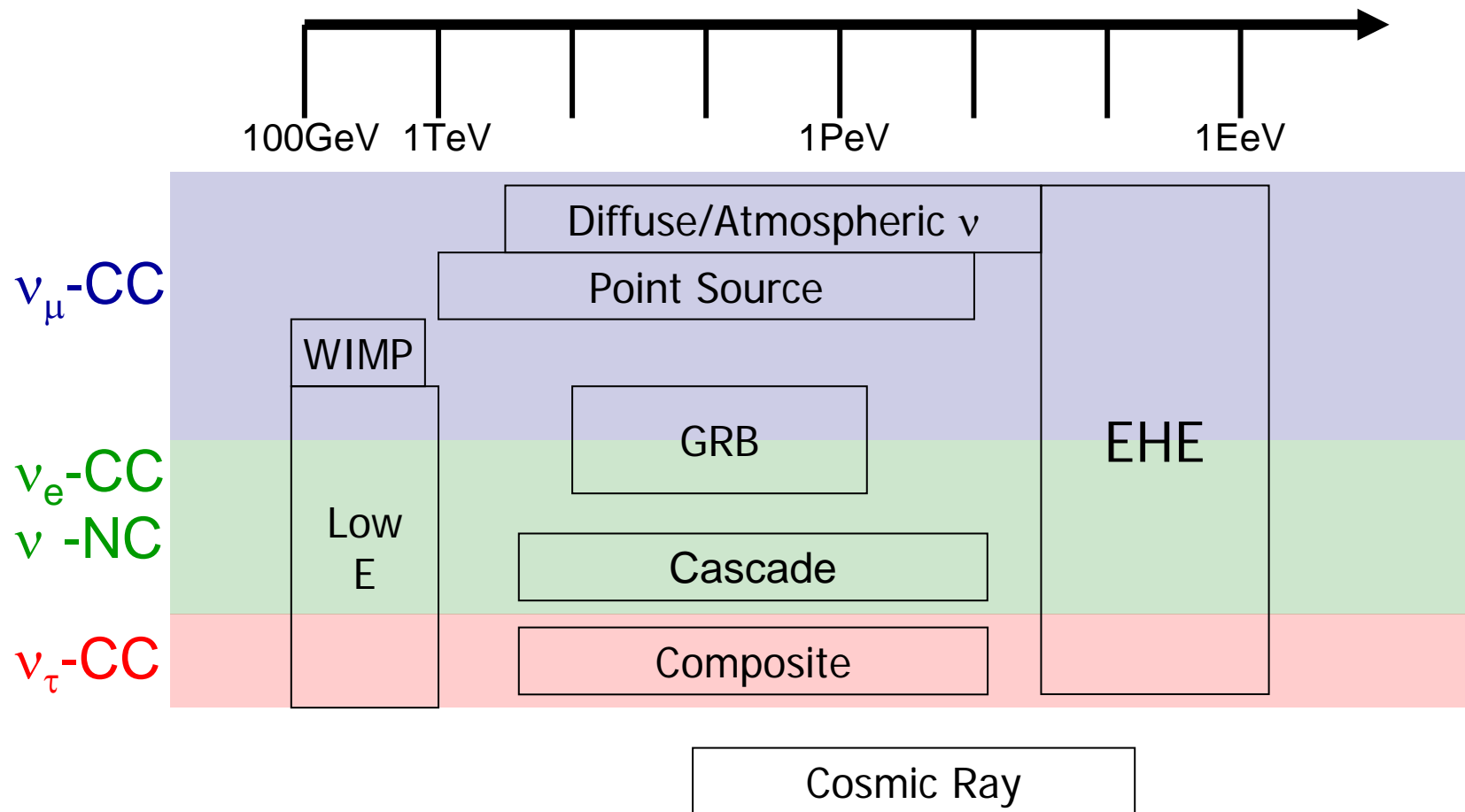
2820 m

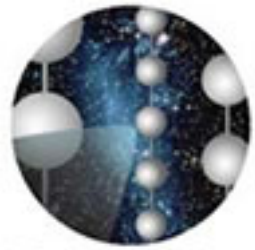
Bedrock

- Total of 59 strings and 118 IceTop tanks → over two thirds complete!
- Completion with 86 strings: January 2011
- Detector is taking data during construction phase.

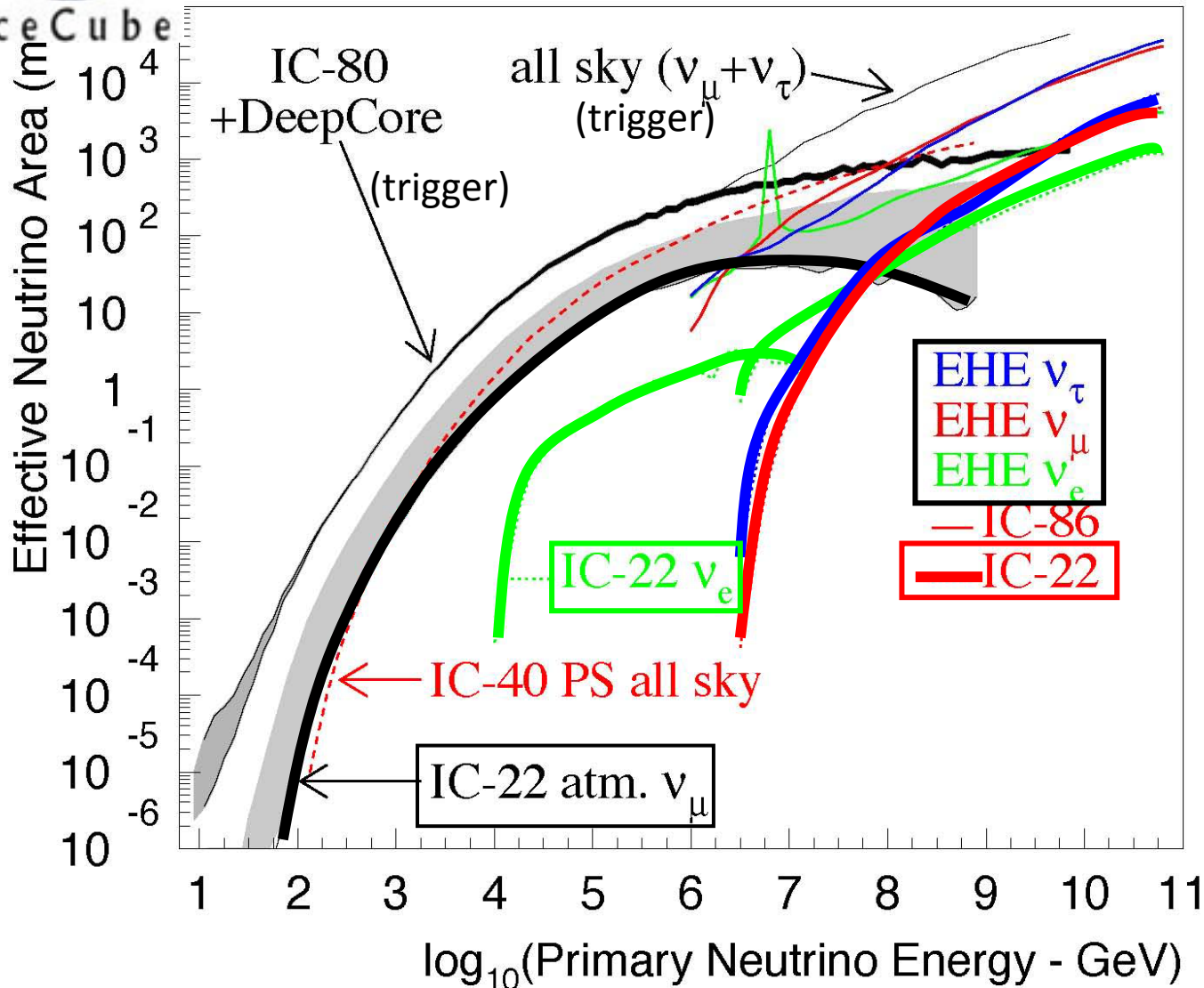


# IceCube working groups

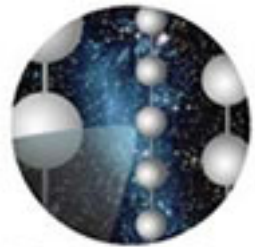




# Effective Area

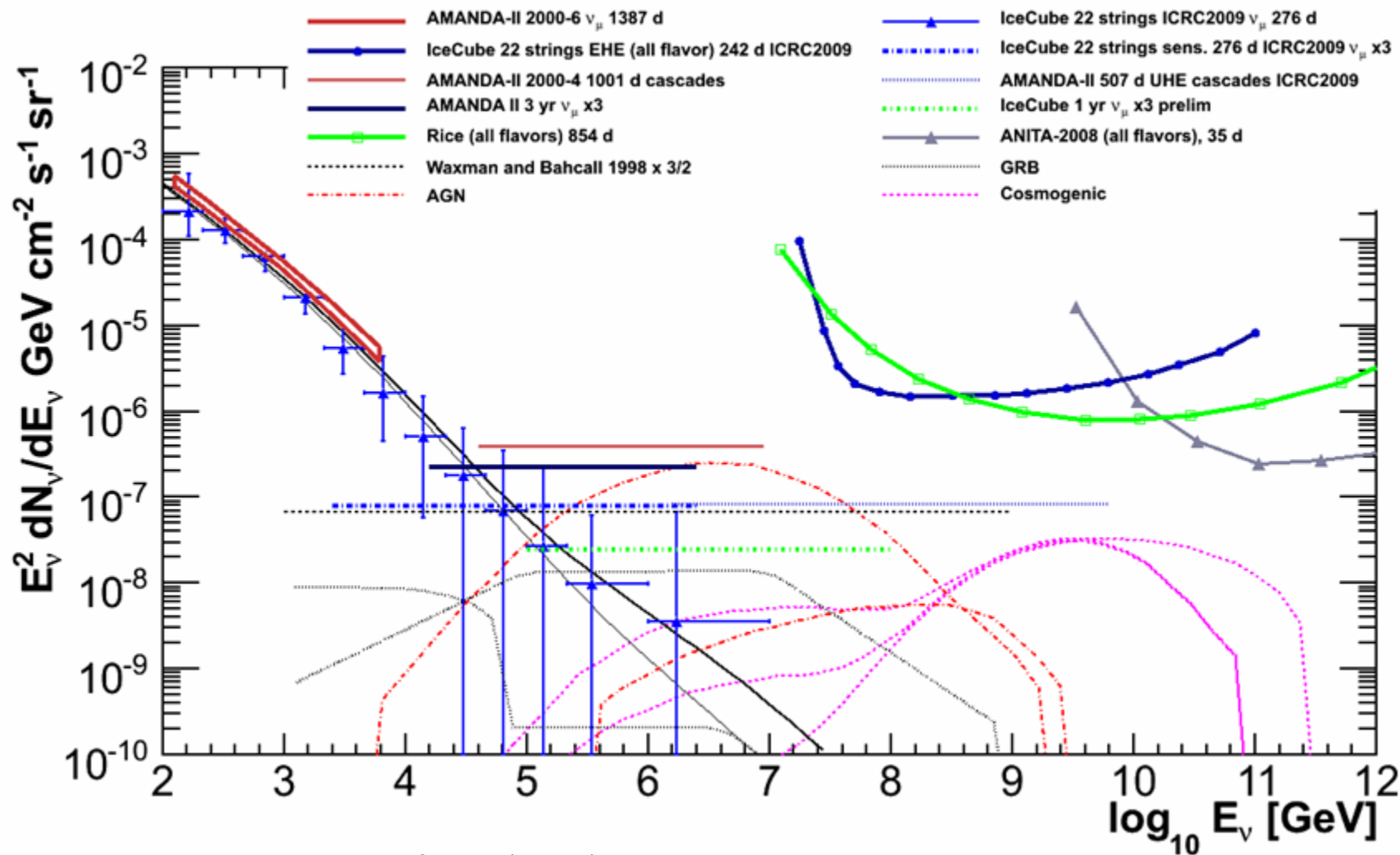


Effective area for ν<sub>μ</sub>  
 Strong rise with energy:  
 -  $\sigma \propto E_\nu$   
 - Increase of muon range with energy up to PeV



IceCube

# And $\nu$ flux limits



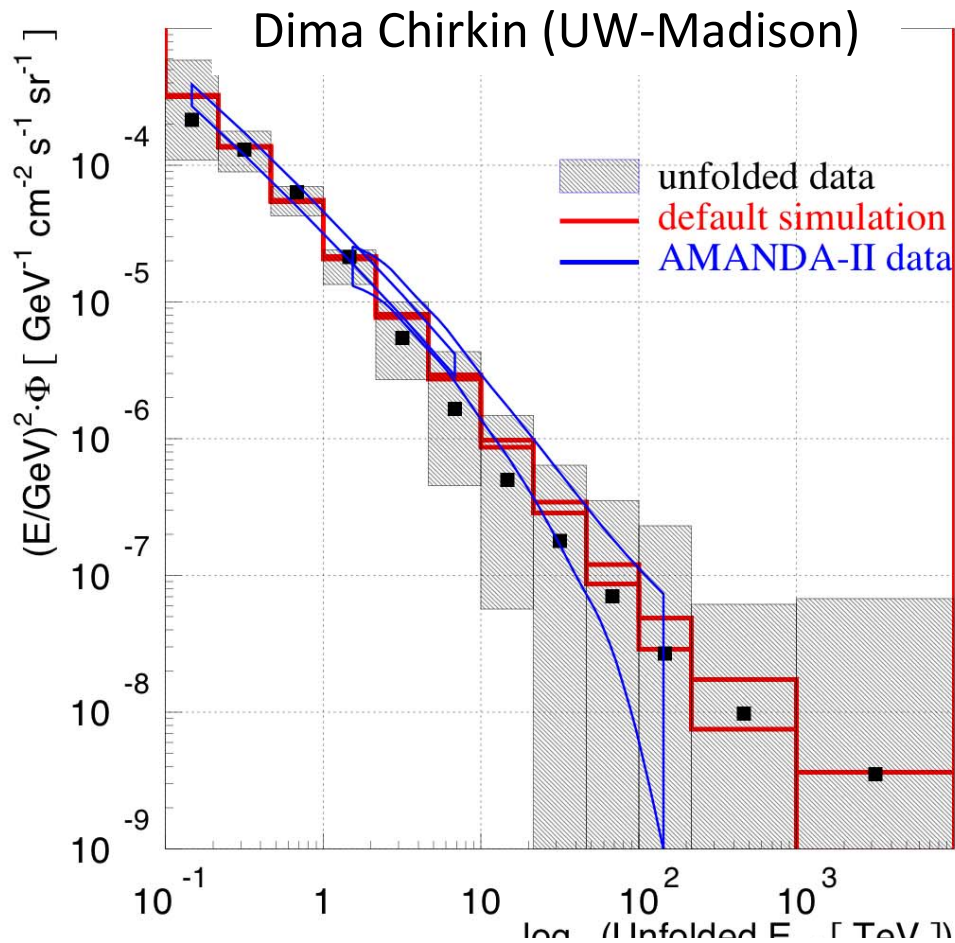
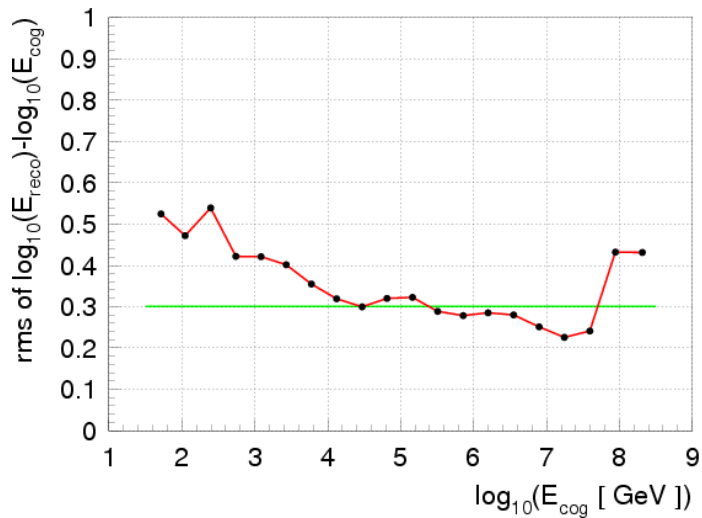
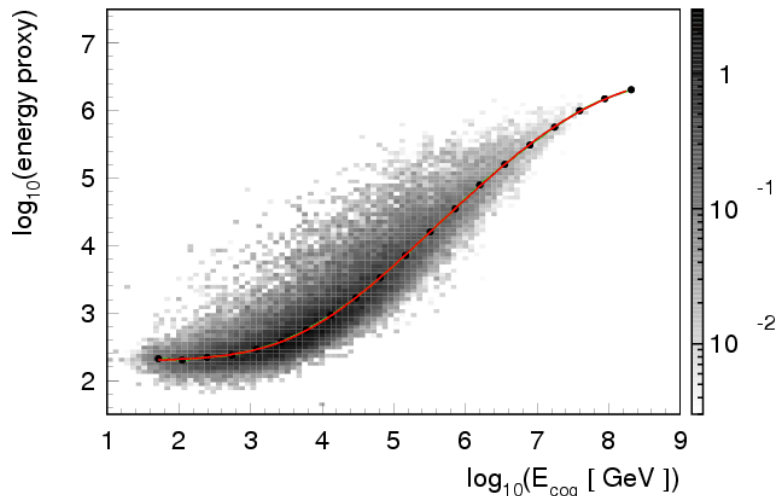
Models and limits are shown as all flavor (1:1:1).



IceCube

# Atmospheric $\nu$ : muon

Muon energy resolution:  $\sim 0.3$  in  $\log(E)$



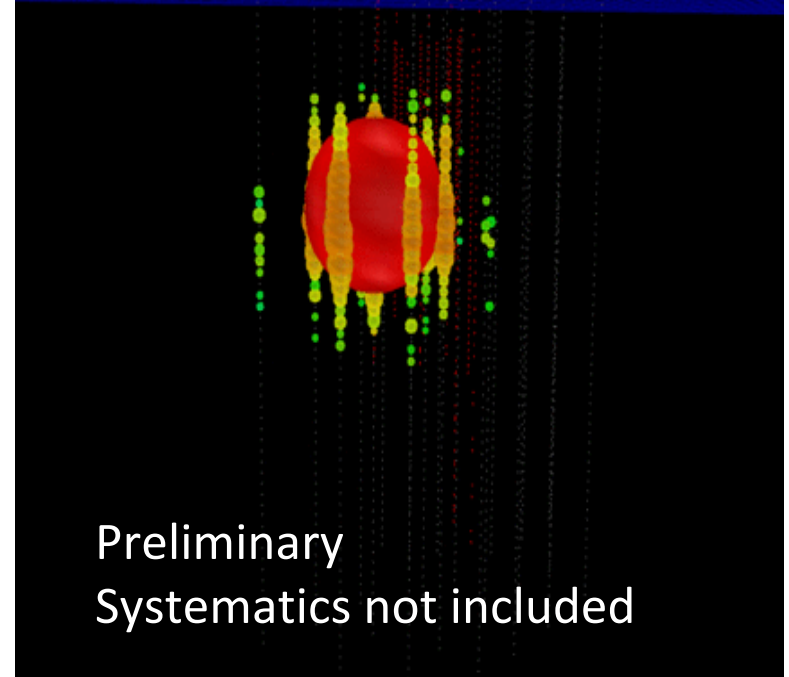
- IceCube 22 string analysis
- 4492 neutrino events at high purity ( $>95\%$ )

# Atmospheric $\nu$ ? : cascade



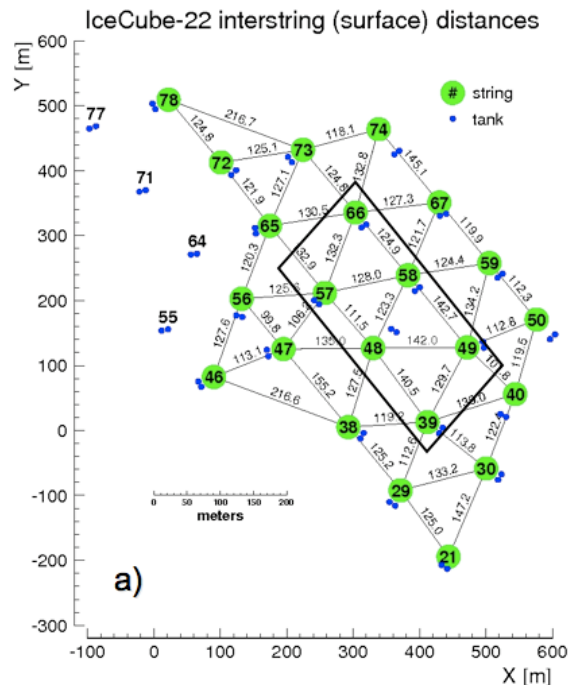
- Cascade searches look for electron-, tau-neutrinos and neutral current interactions.
- Challenge: bremsstrahlung events from cosmic ray muons
- Apply veto techniques (first hits must all be inside defined volume) and test cascade fit quality parameters.
- Approaching the level where atmospheric cascade events are expected.

cascade-like event in IC22  
 $E = \sim 130 \text{ TeV}$

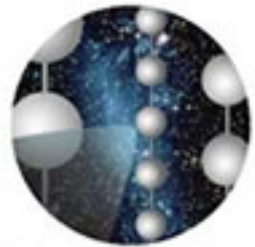


Preliminary  
Systematics not included

Michalangelo d'Agostino (UC-Berkley)  
Joanna Kiryluk (LBNL)



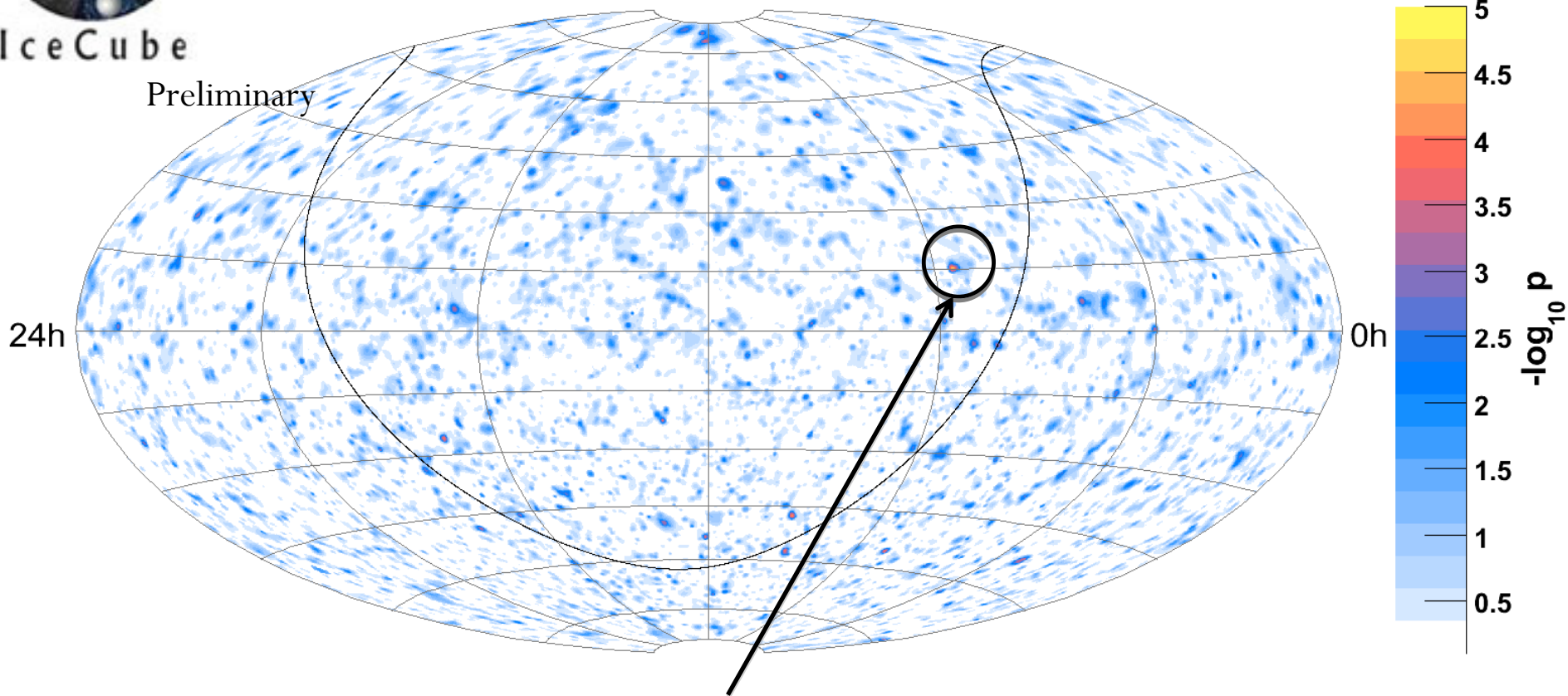
IC22



IceCube

# Search for point sources - 40-string(6month) all-sky results

Preliminary



Hottest location in the all-sky search is:  
r.a.=114.95° , dec.=15.35°

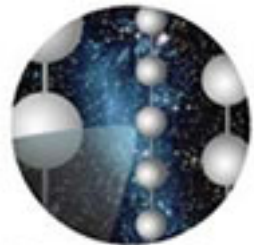
**No excess found!**

**⇒all-sky p-value is 61%, not significant**

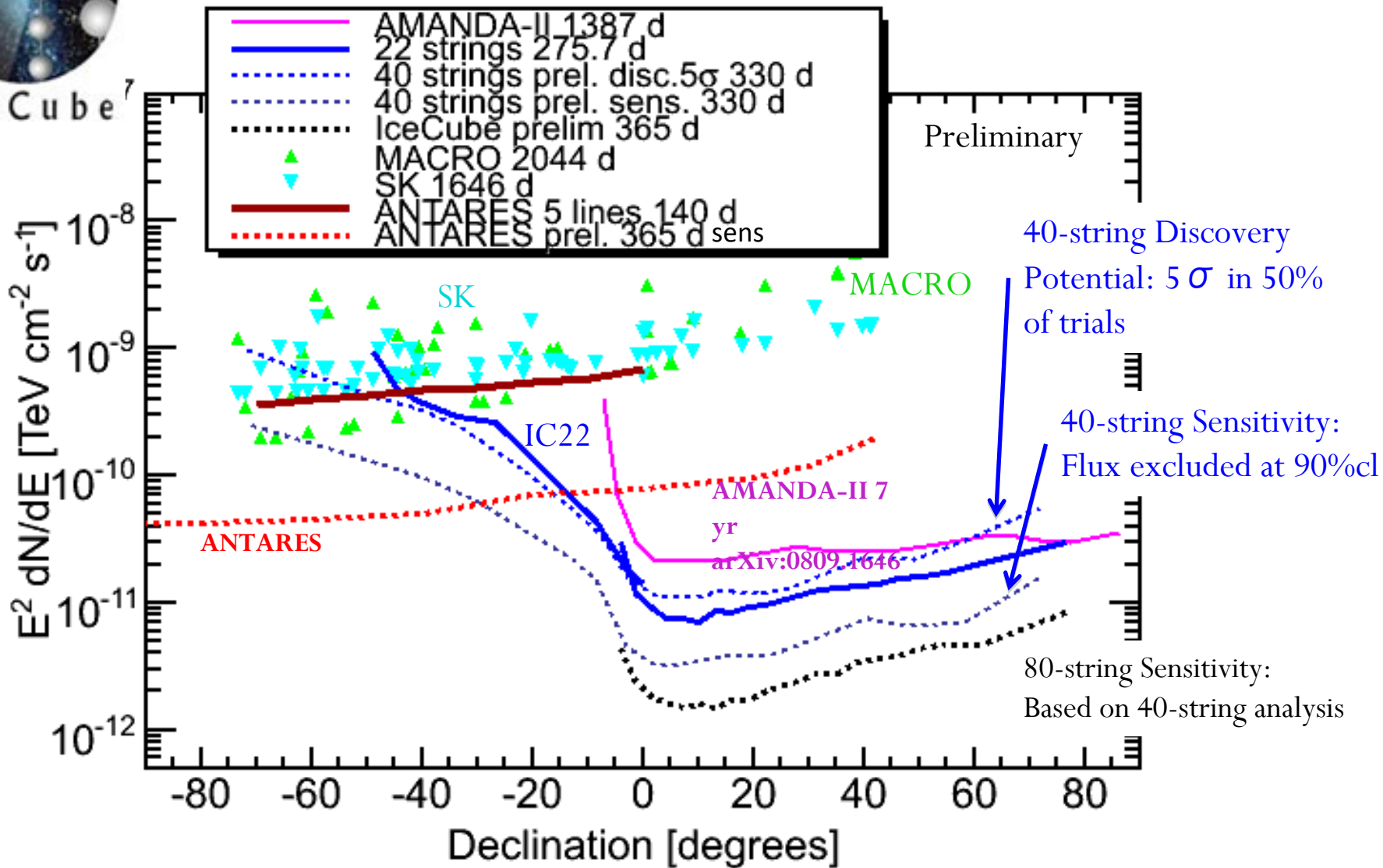
Pre-trial  $-\log_{10}(\text{p-value}) = 4.43$   
Best-fit # of source events = 7.1  
Best-fit spectral index = 2.1

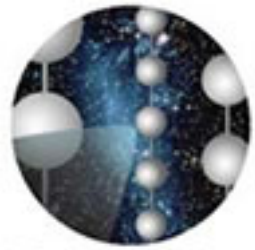
Jon Dumm (UW-Madison)





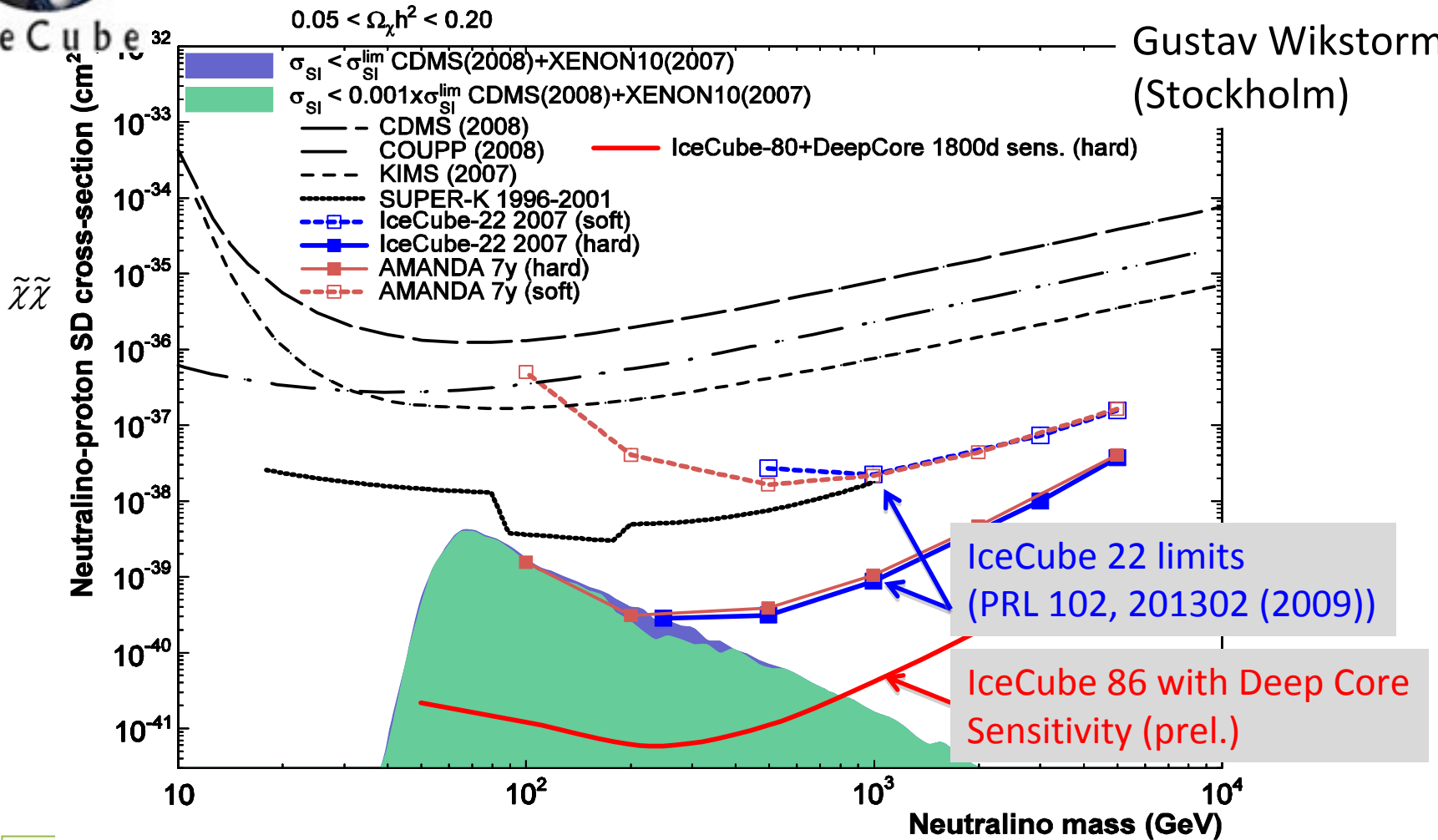
# $E^{-2}$ Sensitivities, limits vs zenith angle





# Dark Matter search: neutrinos from WIMP annihilation in the sun

Gustav Wikstorm  
(Stockholm)

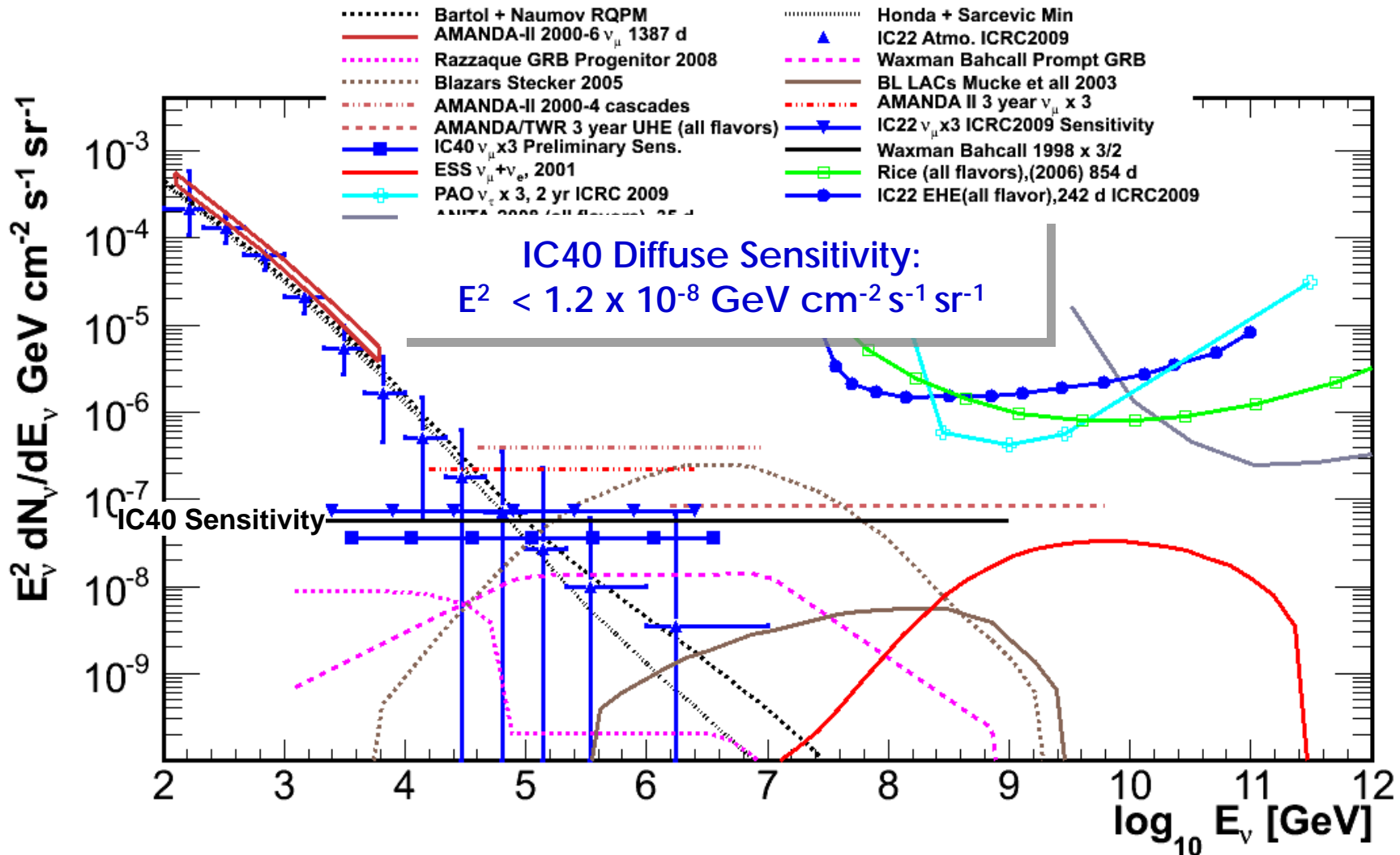


→ Deep core enhancement under construction will greatly enhance sensitivity.



# Diffuse $\nu$ flux

Kotoyo Hoshina/Sean Grullon  
(UW-Madison)



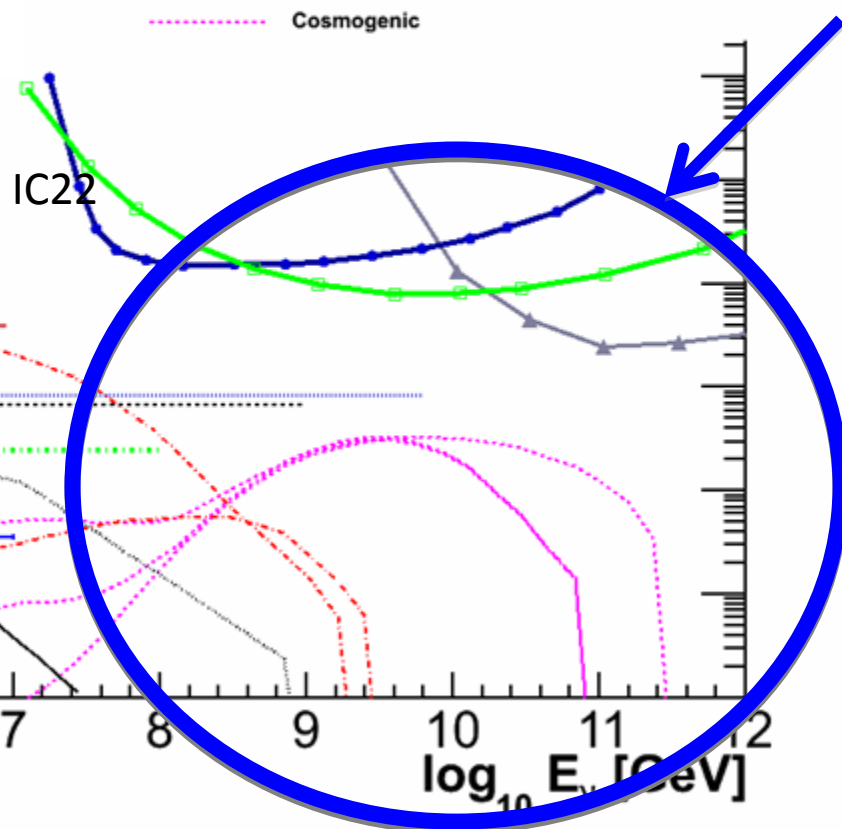
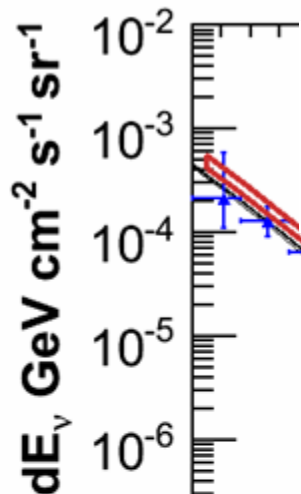


# EHE (EeV and higher)

**Astrophysical sources:  
100 PeV to 10 EeV  
AGN, Cosmogenic neutrinos  
(GZK)**

— AMANDA-II 2000-6  $\nu_\mu$  1387 d  
— IceCube 22 strings EHE (all flavor) 242 d ICRC2009

Aya Ishihara, Keiichi Mase  
Shigeru Yoshida (Chiba)



Event rates for

Flux: Engel, Seckel, Stanev, 2001)

(Factor of 10 higher still allowed by current limits, including IceCube)

- IceCube-22strings, through going, 240 days:  $\sim 0.1$  events/yr
- IC86, total:  $\sim 0.5$  event/yr
- $10 \times 10 \text{ km}^2$  radio array:  $\sim 10$  events/yr



# 2007 Extremely High $E_\nu$ search Data Chain

Level1 :  $\text{NDOM} \geq 80$  with excluding bad runs

"EHE" filter data  $\sim 1.5 \text{ Hz}$  242 days

Level2 :  $\text{NPE} \geq 10^4$

6516 events

Level3 :  $\cos(\text{Zenith}) < 0.8$

2011 events

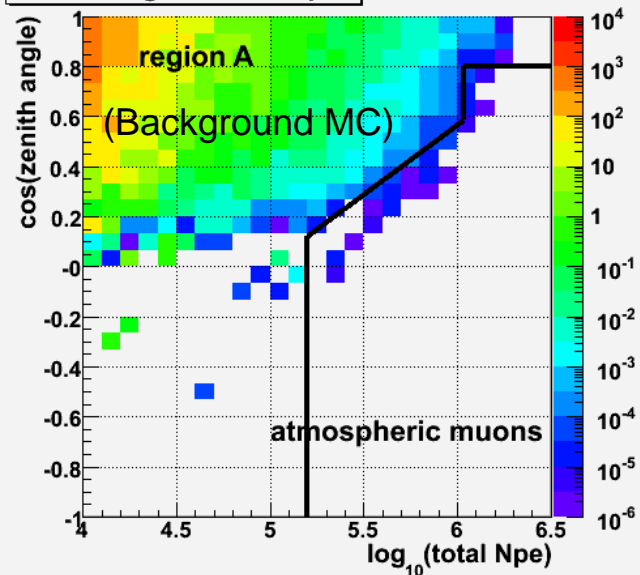
Level4 : cuts on NPE and Zenith

0 events

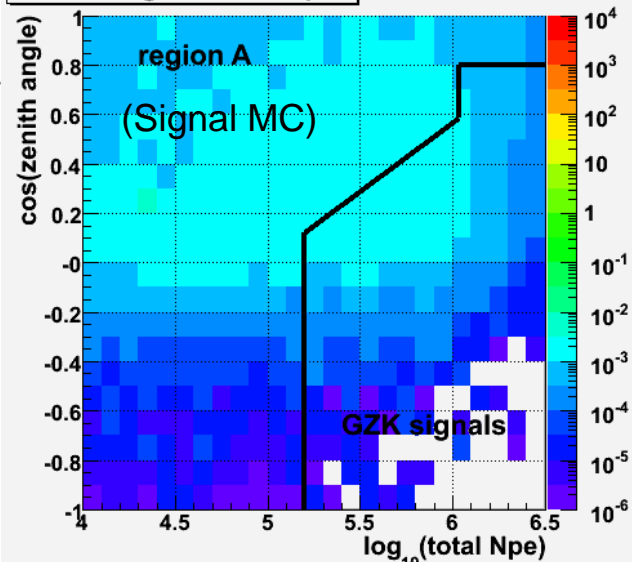
region A:  $-250 < \text{CoGZ} < -50$  m and  $\text{CoGZ} > 50$  m

region B:  $\text{CoGZ} < -250$  m and  $-50 < \text{CoGZ} < 50$  m

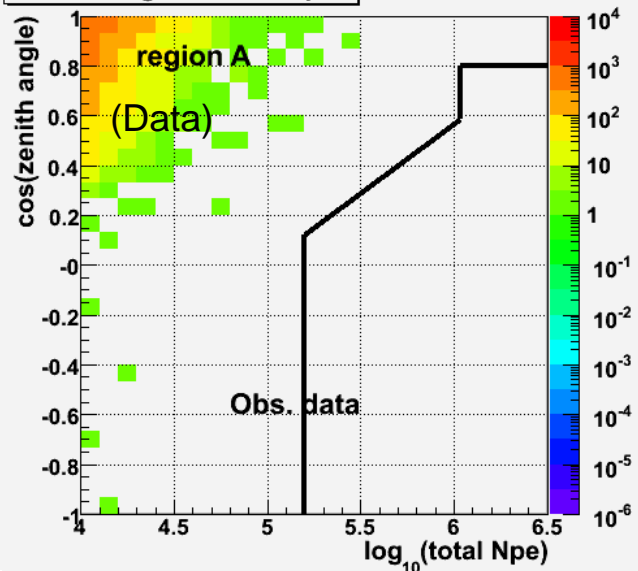
Zenith angle Vs total Npe



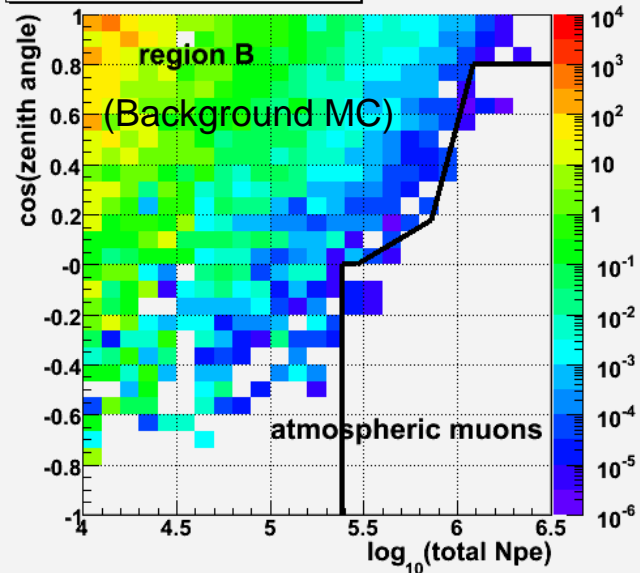
Zenith angle Vs total Npe



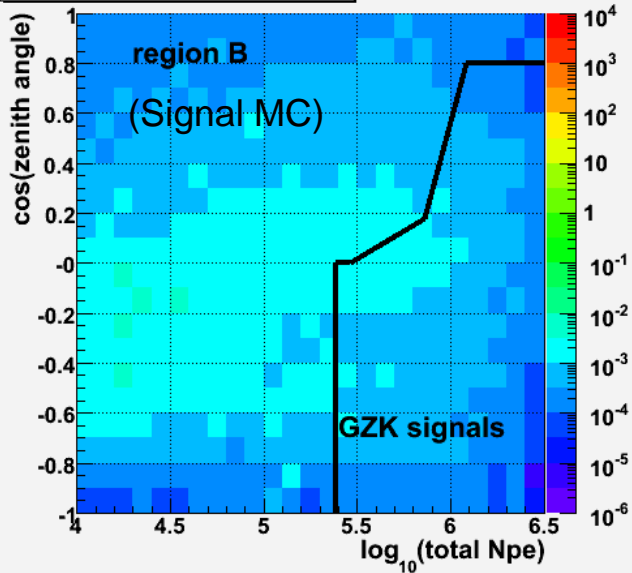
Zenith angle Vs total Npe



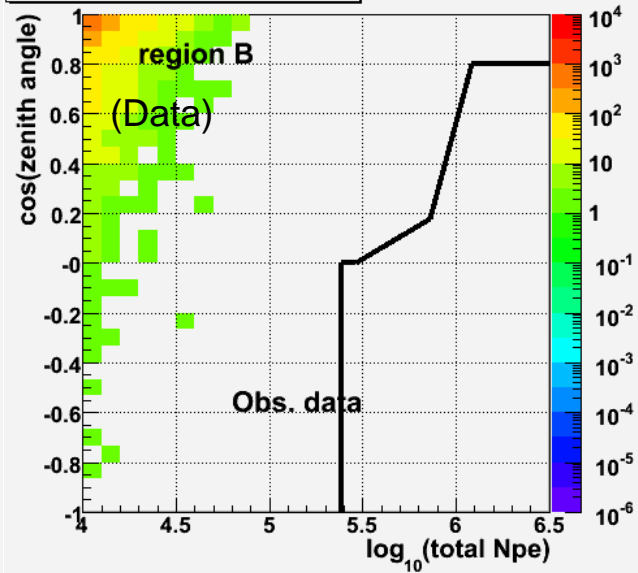
Zenith angle Vs total Npe



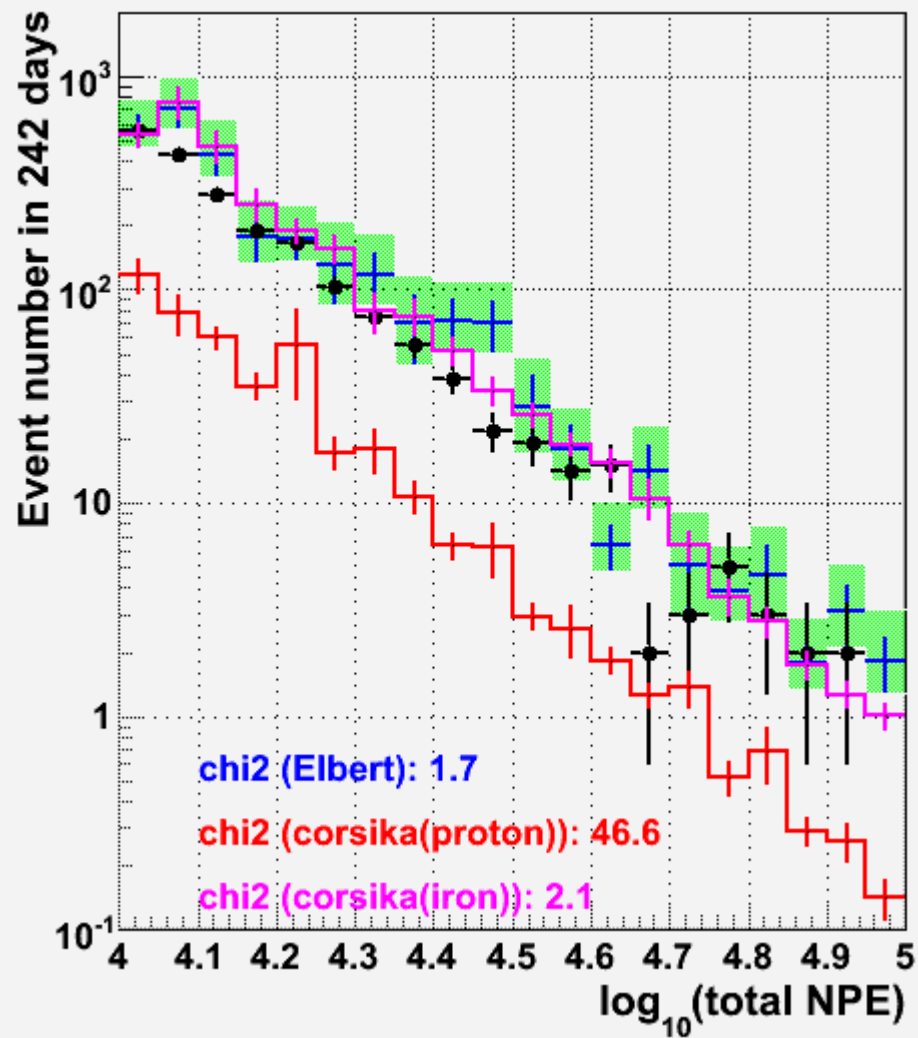
Zenith angle Vs total Npe



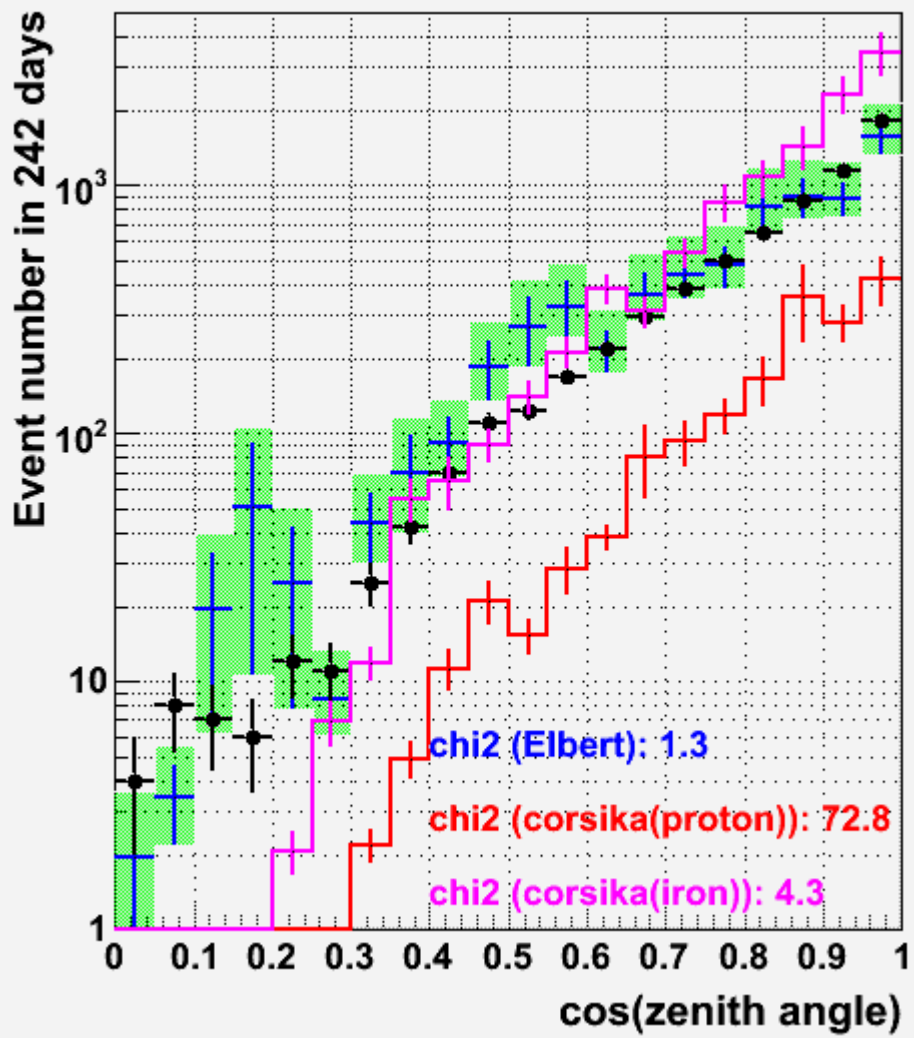
Zenith angle Vs total Npe



Total NPE distribution

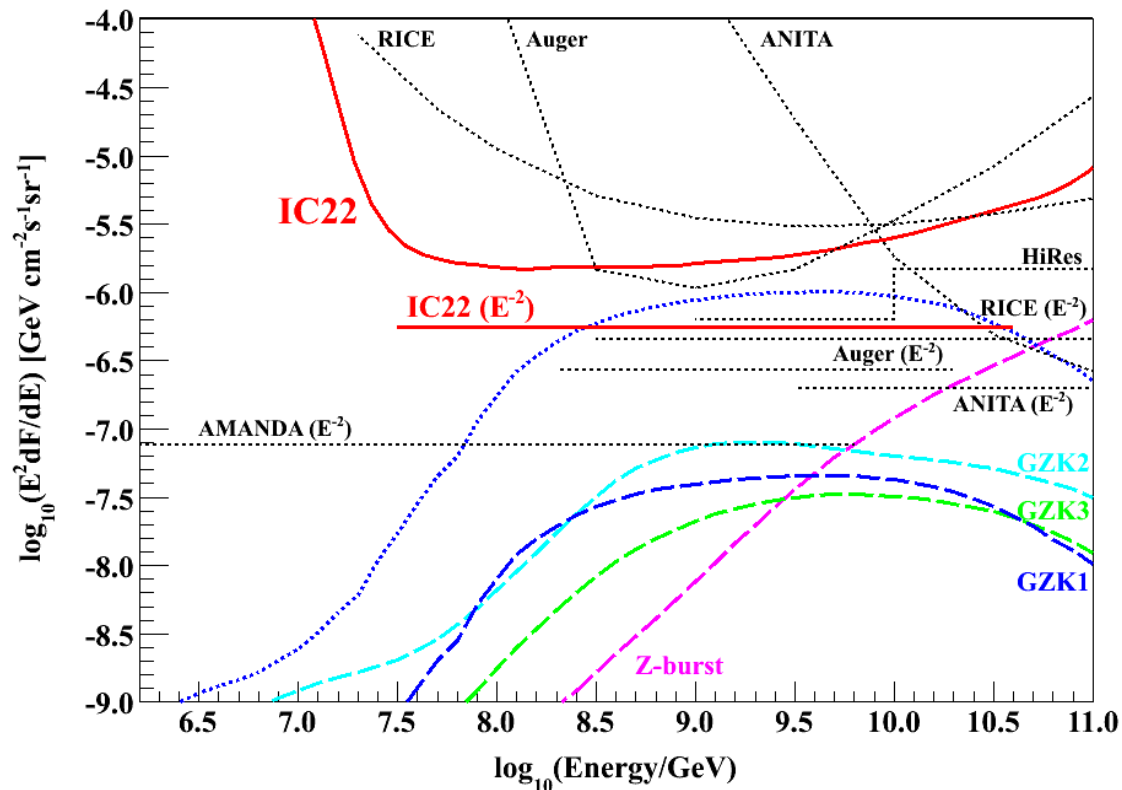
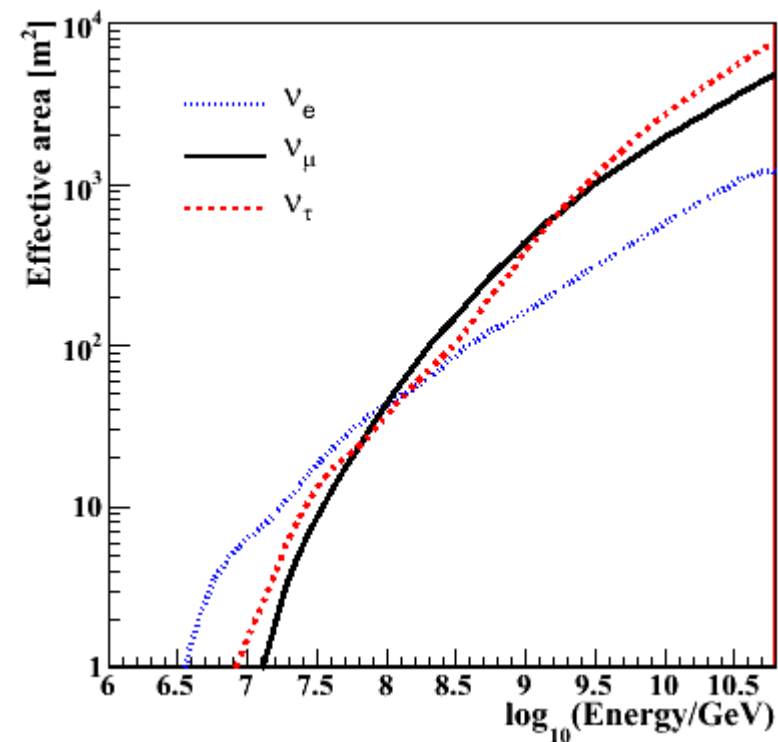


Zenith angle distribution



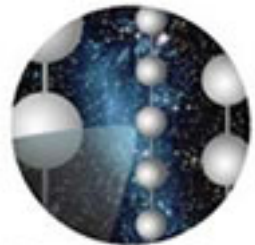


# Extremely-High Energy $\nu$ limits with 241 days observation in 2007



Aya Ishihara, Keiichi Mase  
Shigeru Yoshida (Chiba)

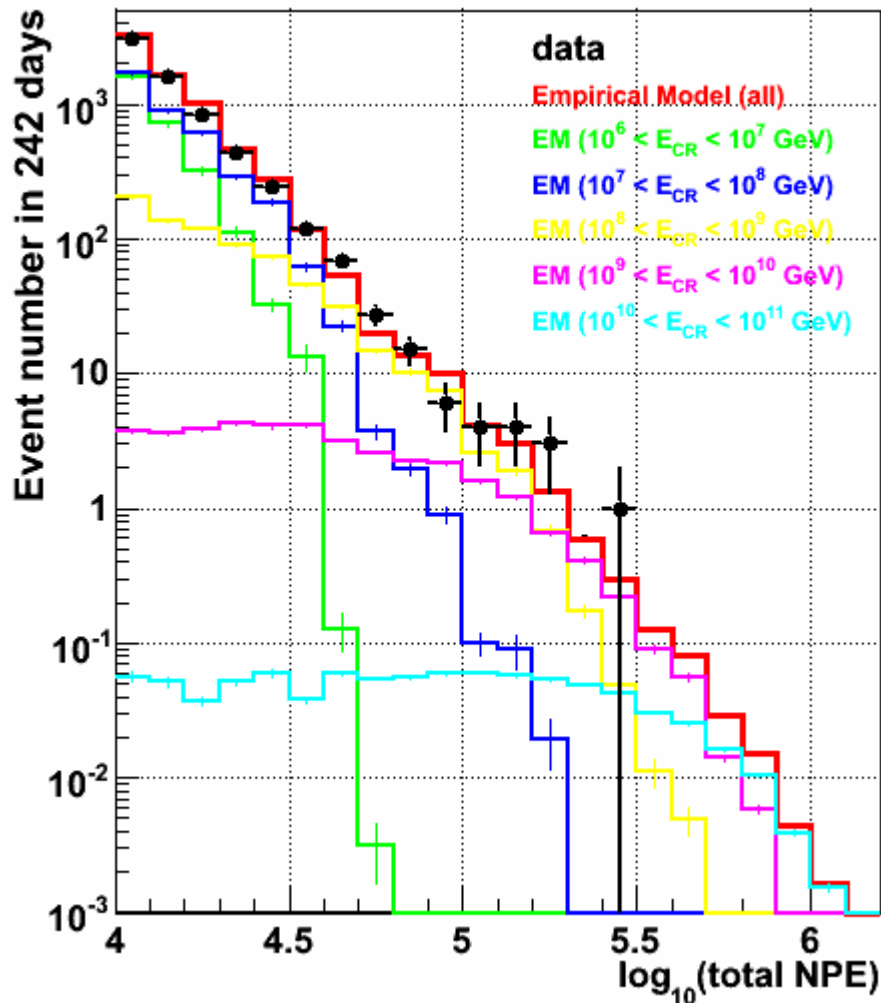




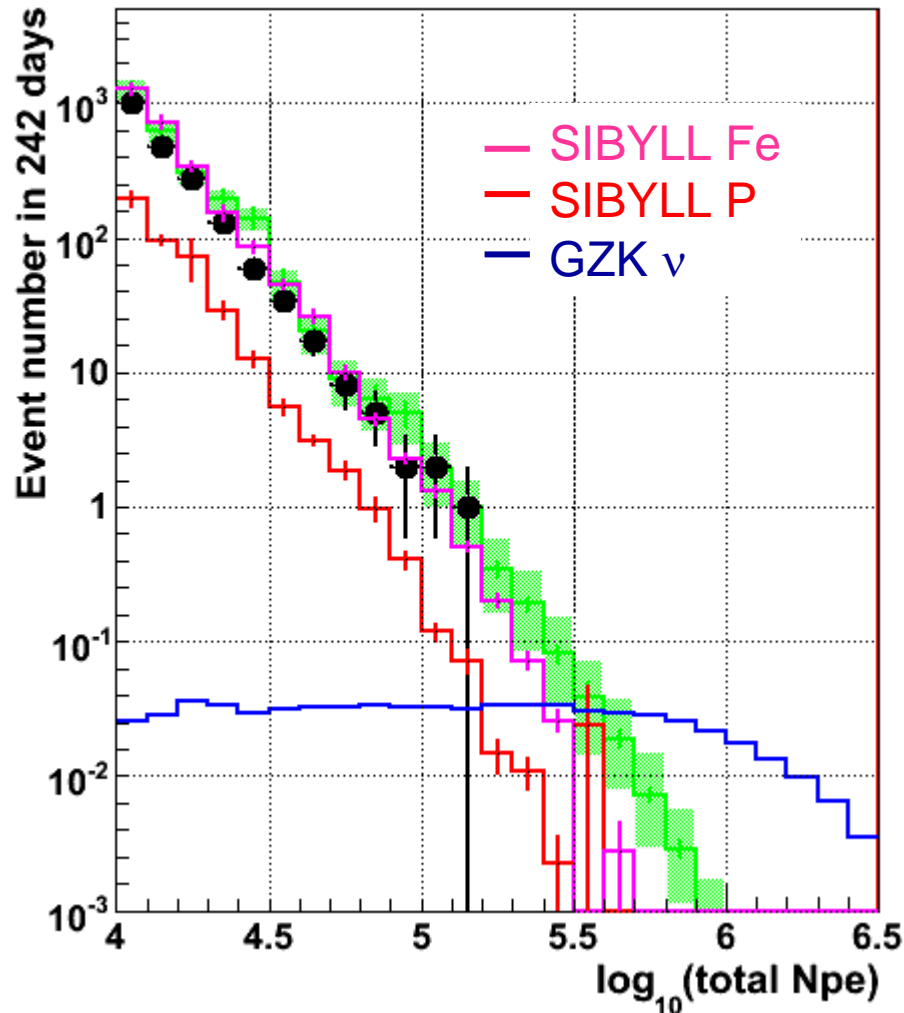
# By products – UHECR composition?

IceCube

Total NPE distribution



Total Npe distribution



Aya Ishihara, Keiichi Mase (Chiba)

# 2008-2009 Data Status

Aya Ishihara (Chiba)

## 2008/4/17~2009/5/20: 40String operation

– livetime ~370days (>93%)

month	June08	July08	Aug08	Sept08	Oct08	Nov08	Dec08	Jan09	Feb09
live-time[day]	28.59 (95.3%)	28.86 (93.1%)	29.85 (96.3%)	26.92 (89.7%)	28.11 (90.1%)	28.95 (96.5%)	28.59 (92.2%)	26.83 (86.5%)	25.01 (89%)

- Trigger Rate (8 Optical Module)
- High Energy Trigger Rate
  - Bright Event Condition:  $NPE > 630 \rightarrow 1.23\text{Hz}$
  - High Energy Condition:  $NPE > 10000 \rightarrow 45\text{evts/day}$

Strings	Year	Livetime	$\mu$ rate	$\nu$ rate	HE rate
IC9	2006	137days	80Hz	1.7/day	4.3/day
IC22	2007	250days	550Hz	28/day	27/day
IC40	2008	370days	1000Hz	110/days	45/days
IC59	2009	~365days	1500Hz	150/days	??/days

## 2009/5/20 ~ : 59String operation

- Trigger Rate (8 Optical Module)
- High Energy Trigger Rate
  - Bright Event Condition:  $NPE > 630 \rightarrow 2.29\text{Hz}$

We'll have **one full IceCube year of data** by the end of 59 string operation !

2009/Nov-2010/Feb: (59+18=77strings) 2010/Nov-2011/Feb: Complete (86 Strings)

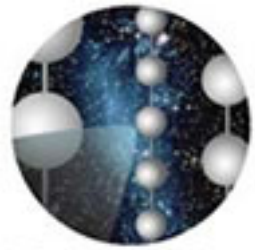


# Systematic Error Budget (EHE analysis)

TABLE III: List of the statistical and systematic errors. The signal rate is estimated by assuming the high evolution flux  $(m, Z_{\max}) = (4, 4)$  in Ref [6].

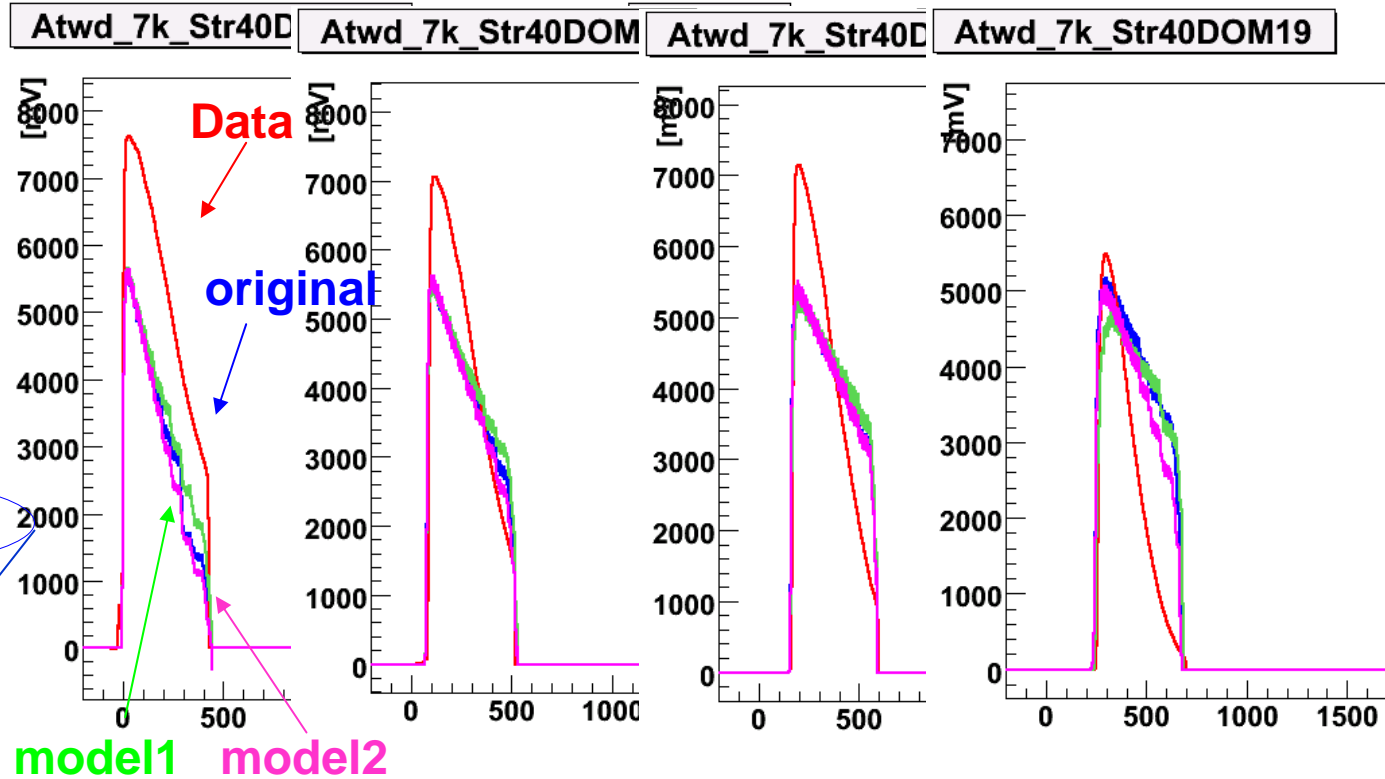
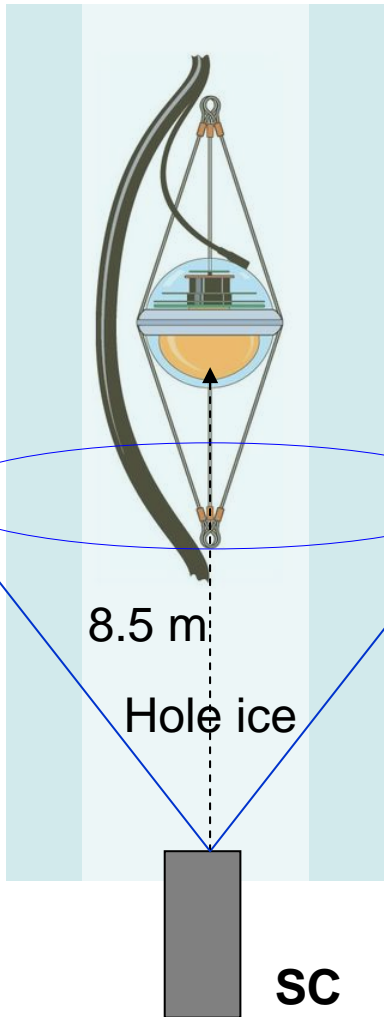
Error source	backgrounds	signals (GZK)
	rate	rate
Statistical error	$\pm 22\%$	$\pm 0.6\%$
Detector sensitivity	-	$\pm 8\%$
Yearly variation	$\pm 17\%$	-
Empirical model	+99 -59%	-
Hadronic int. model	$\pm 4\%$	-
NPE yield	-	-32%
Neutrino cross-section	-	$\pm 9\%$
Photo-nuclear interaction	-	+10%
LPM effect	-	$\pm 1\%$
Total	$\pm 22\%$ (stat.)	$\pm 0.6\%$ (stat.)
	+101 -62% (sys.)	+16 -34% (sys.)

# The biggest disagreement



IceCube

Waveforms of the detector just above the SC, **Data** > **MC**



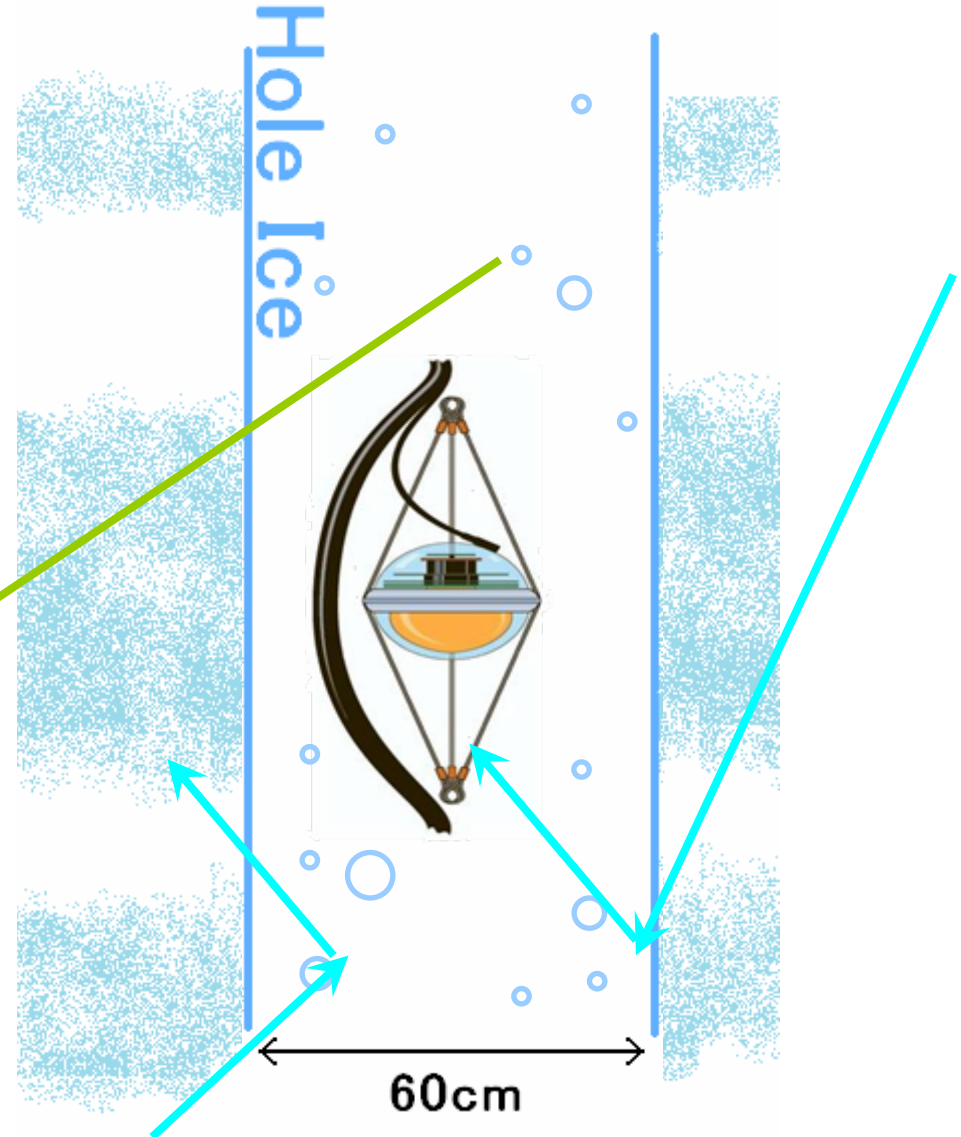
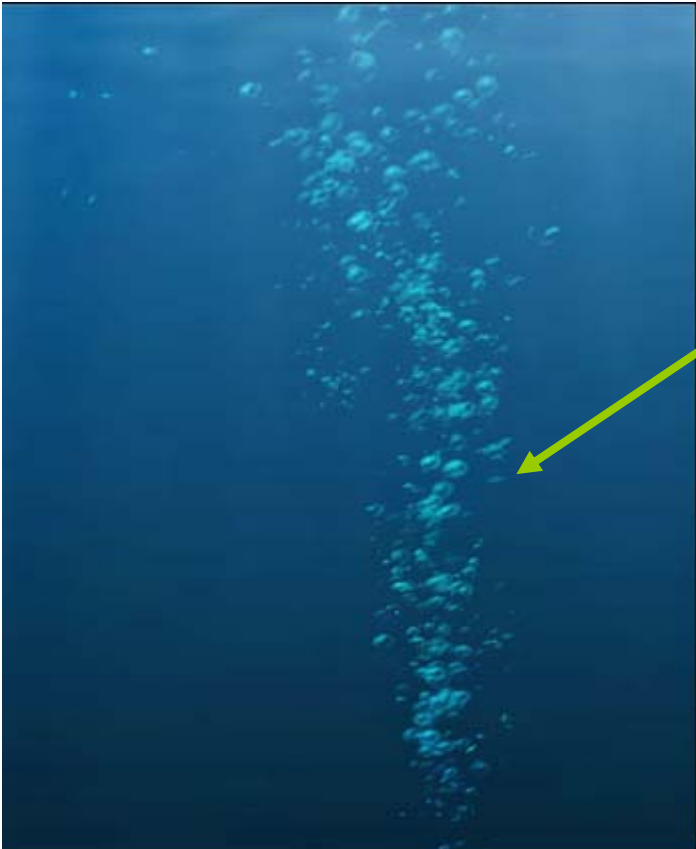
The disagreement of the peak suggests that there are photons coming directly from the SC in the real.

# Mission



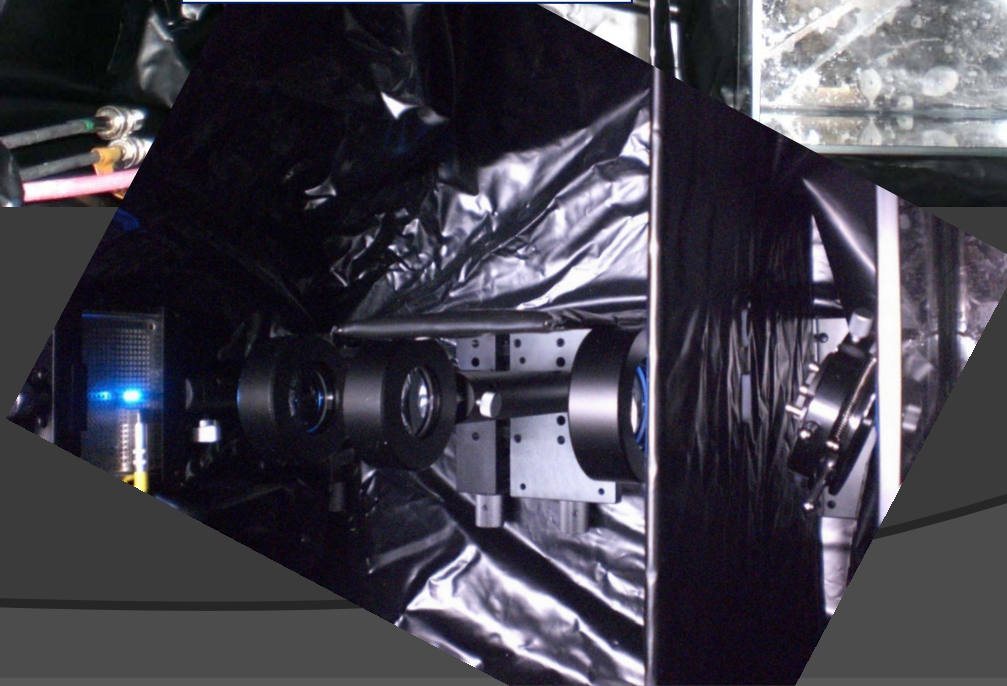
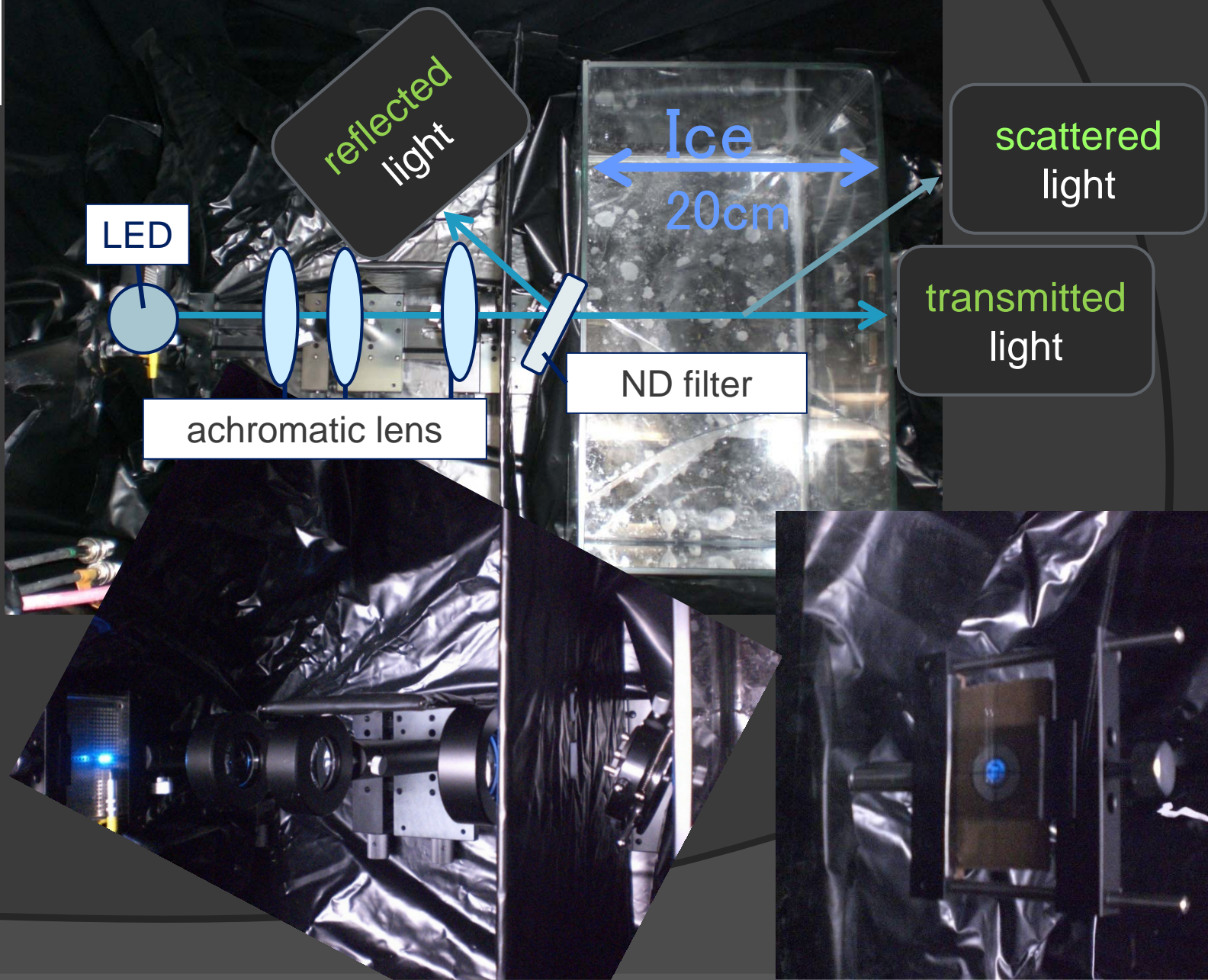
IceCube

Count the number of  
“Bubbles”  
in a path in hole-ice



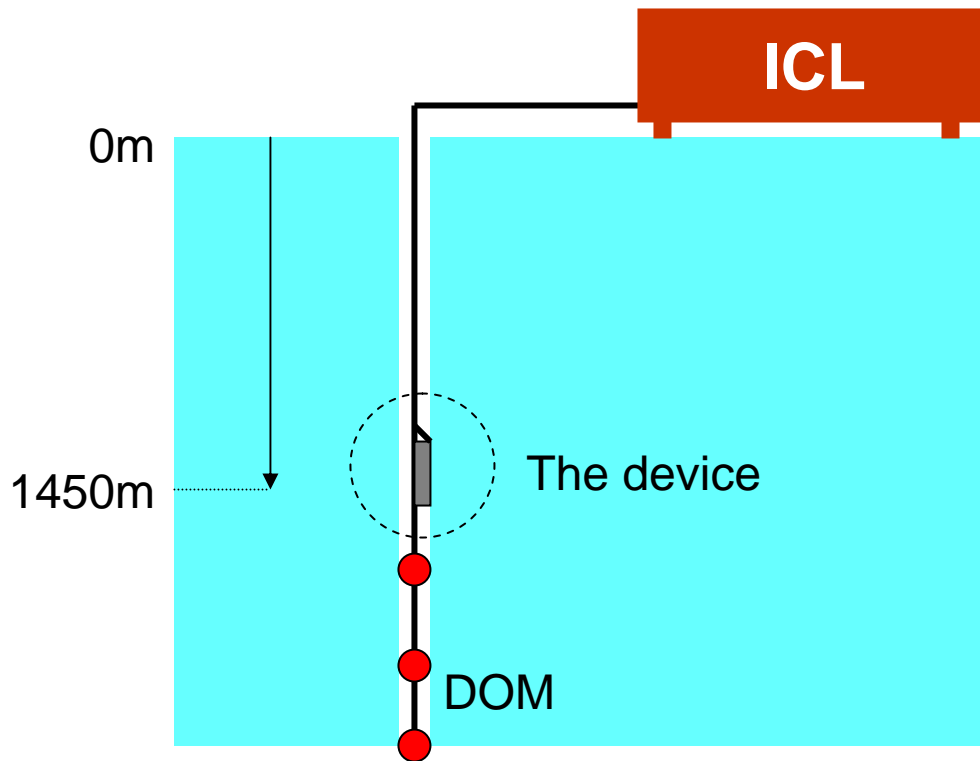


# The system overview

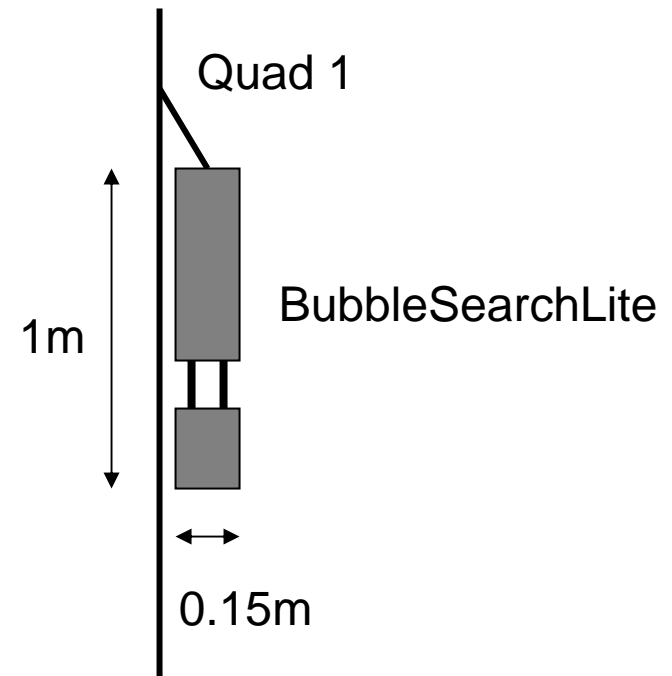




# BubbleSearchLite Deployment Configuration



IceCube main cable





# The BubbleSearchLite

