

ステライルニュートリノ探索の実 験的レビュー

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I cannot cover all experiments (e.g.; T2K-ND280 searches; arXiv:1410.8811 [hep-ex]) since there are a lot of on-going efforts.

Also $\nu_\mu \rightarrow \nu_\mu$ disappearance was discussed in Yasuda-san's talk and will be discussed later for more details.
(by E.Richard)

Physics (sterile neutrino search)

- Anomalies, which cannot be explained by standard neutrino oscillations for 15 years are shown;

Experiments	Neutrino source	signal	significance	E(MeV), L(m)
LSND	μ Decay-At-Rest	$\bar{\nu}_\mu \rightarrow \bar{\nu}_e$	3.8σ	40, 30
MiniBooNE	π Decay-In-Flight	$\nu_\mu \rightarrow \nu_e$	3.4σ	800, 600
		$\bar{\nu}_\mu \rightarrow \bar{\nu}_e$	2.8σ	
		combined	3.8σ	
Ga (calibration)	e capture	$\nu_e \rightarrow \nu_x$	2.7σ	<3, 10
Reactors	Beta decay	$\bar{\nu}_e \rightarrow \bar{\nu}_x$	3.0σ	3, 10-100

- Excess or deficit does really exist?
- The new oscillation between active and inactive (sterile) neutrinos? ($\Delta m^2 \sim 1.0 \text{ eV}^2$)?
- One of the hottest topics in the neutrino community;
 - About quarter of the Neutrino2014 talks mentioned sterile ν_s
 - P5 (US) endorsed the short baseline neutrino experiment for sterile
 - A lot of experiments are on-going (especially, using reactors).

Neutrino oscillations with $\Delta m^2 \sim 1 \text{eV}^2$ region

$$\begin{bmatrix} \nu_e \\ \nu_\mu \\ \nu_\tau \\ \nu_s \\ \bullet \end{bmatrix} = \begin{bmatrix} U_{e1} & U_{e2} & U_{e3} & U_{e4} & \bullet \\ U_{\mu1} & U_{\mu2} & U_{\mu3} & U_{\mu4} & \bullet \\ U_{\tau1} & U_{\tau2} & U_{\tau3} & U_{\tau4} & \bullet \\ U_{s1} & U_{s2} & U_{s3} & U_{s4} & \bullet \\ \bullet & \bullet & \bullet & \bullet & \bullet \end{bmatrix} \begin{bmatrix} \nu_1 \\ \nu_2 \\ \nu_3 \\ \nu_4 \\ \bullet \end{bmatrix}$$



Matrix elements, which are considered in 3x3 mixing framework.

$$\sum_{j=1,3} U_{ej}^* U_{\mu j} = -U_{e4}^* U_{\mu 4}$$

Small mixture with active ν 's $U_{e4}, U_{\mu 4} \sim 0.1$ $U_{s4} \sim 1$ $m_4 \sim 1 \text{eV} \gg m_{1,2,3}$

$$P_{e\mu} = -4 \sum_{i=1,3} (U_{e4}^* U_{\mu 4} U_{ei} U_{\mu i}^*) \sin^2 \frac{(m_4^2 - m_i^2)L}{4E_\nu} \sim 4 |U_{e4}|^2 |U_{\mu 4}|^2 \sin^2 \frac{\Delta m_4^2}{4} \frac{L}{E}$$

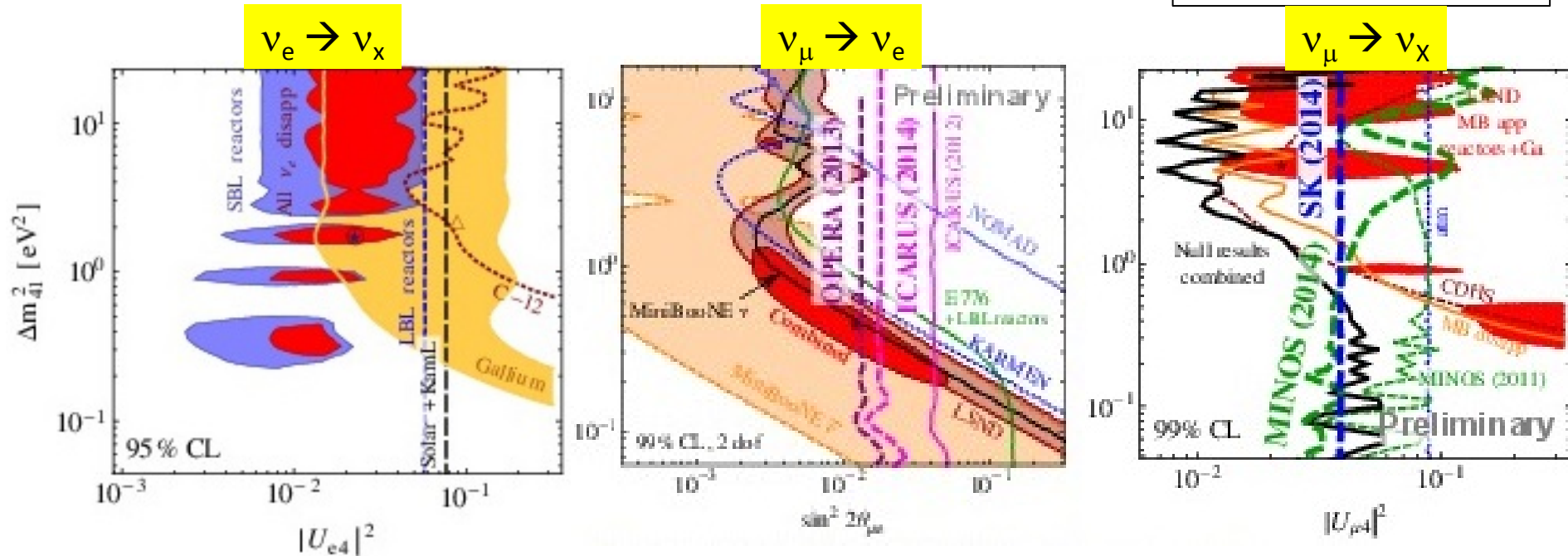
$$P_{es} = -4 \sum_{i=1,3} (U_{e4}^* U_{s4} U_{ei} U_{si}^*) \sin^2 \frac{(m_4^2 - m_i^2)L}{4E_\nu} \sim 4 |U_{e4}|^2 |U_{s4}|^2 \sin^2 \frac{\Delta m_4^2}{4} \frac{L}{E}$$

$$P(\nu_\mu \rightarrow \nu_e) = \sin^2 2\theta \cdot \sin^2 \left(\frac{1.27 \cdot \Delta m^2 \cdot L}{E_\nu} \right)$$

(3+1) model

Status as of 2014

J. Kopp's
talk at Neutrino2014



- ICARUS / OPERA experiments had new results in 2014 for appearance channel.
- SK / MINOS had new results in ν_μ disappearance channel.
- Daya-Bay had latest results on ν_e disappearance in the last summer

Discrepancy between ν_μ and others?

- Due to issue on theoretical model (3+1) ?
- **Confirming or refuting the anomalies with various E/L is first thing to do for experimentalists.**

典型的ニュートリノ源・検出器

- ニュートリノ源

- (反)電子 ν 消失

- 原子炉
 - 線源

- $\nu_\mu \rightarrow \nu_e$ 出現

- 電磁ホーンを使った ν ビーム
 - 加速器ダンプ ν
 - μ 貯蔵リングによるビーム

- $\nu_\mu \rightarrow \nu_\mu$ 消失

- 電磁ホーンを使った ν ビーム
 - 大気 ν

- 検出器

- 液体シンチレータ

- 反電子 ν 検出に優れた性能
 - 原子炉・線源・ダンプ ν が ν 源の実験に良く使われる。

- 水・氷・ミネラルオイルチェレンコフ

- コストが安い。PID等も確立。
 - ホーンを使った ν 、大気 ν 等。

- 液体アルゴンTPC

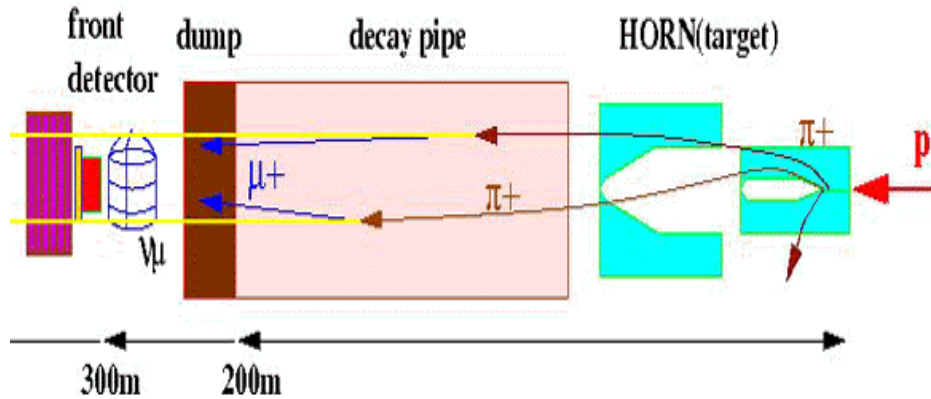
- π^0 /電子の区別に優れている
 - ホーンを使った ν で使用。

- 鉄とシンチ(またはチェンバー)

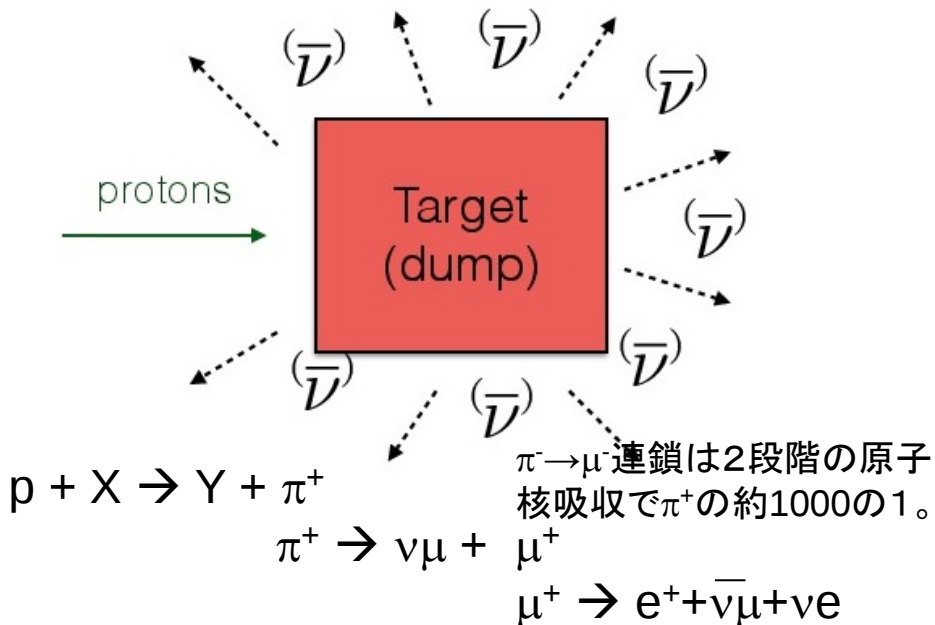
- 磁場をかけていれば粒子の電荷符号が分かる。
 - ホーンを使った ν 、 μ 貯蔵リングによるビーム

ニュートリノ源

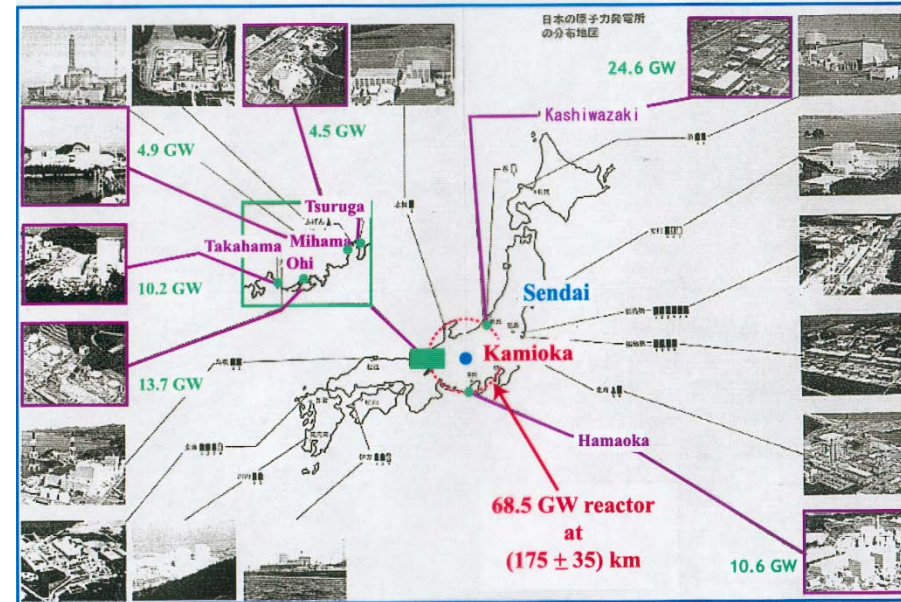
ホーンで作る ν (ν_μ を π DIFから作る)



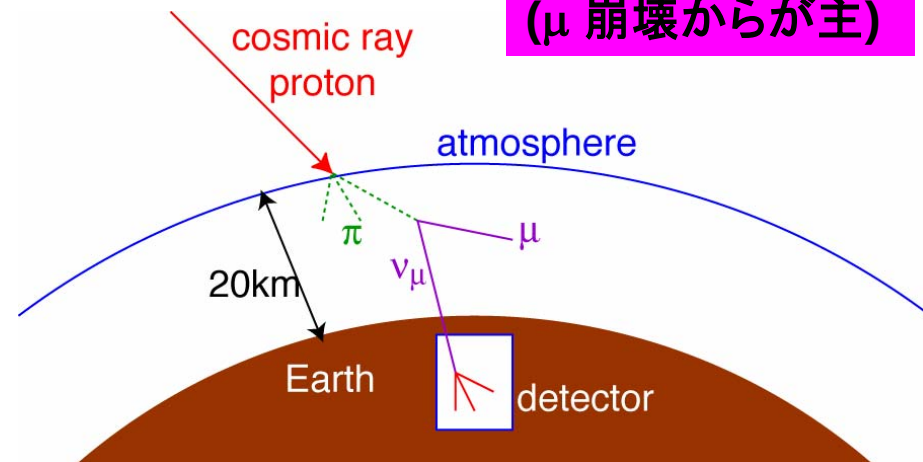
ダンプで作る ν (ν をDARから作る)



原子炉 ($\bar{\nu}_e$ 核分裂から)



大気 ν ($\nu_e + \nu_\mu$) (μ 崩壊からが主)

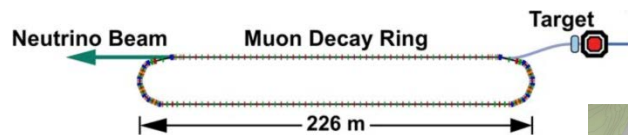


nuSTORM (neutrinos from STORed Muons)

A low-energy muon storage ($P_\mu=3.8$ GeV/c) ring based on existing technology to:

- Address the large Δm^2 neutrino oscillations
- Provide beams for precision ν_e and ν_μ cross section measurements
- Provide an accelerator technology test bed (ν -Factory & μ -Collider)
- Provide a neutrino Detector Test Facility
 - Proposal: [arXiv:1308.6822](https://arxiv.org/abs/1308.6822)
 - Project Definition Report: [arXiv:1309.1389](https://arxiv.org/abs/1309.1389)

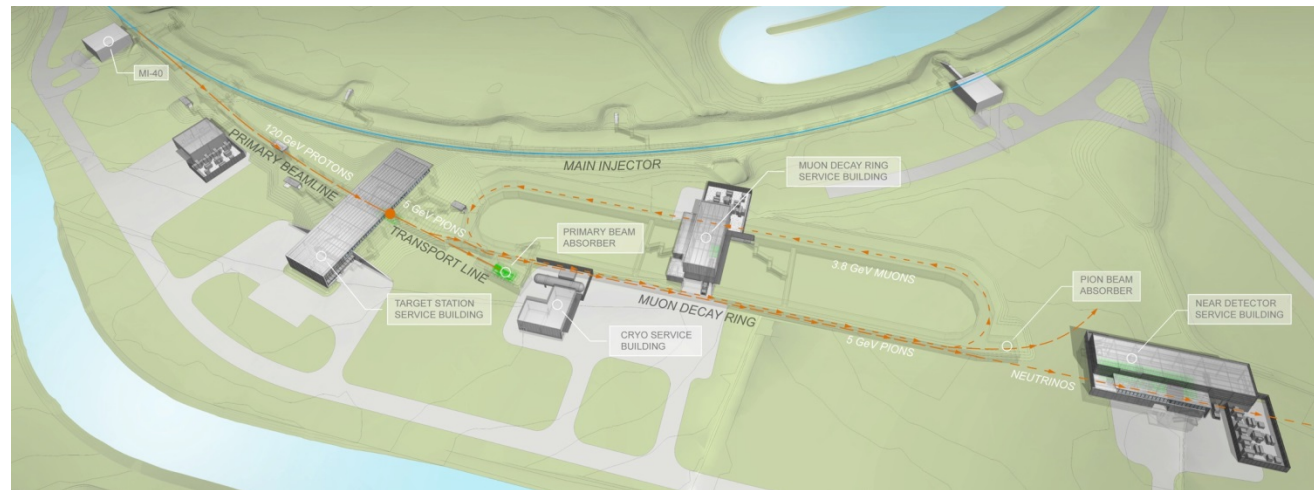
Technologically limited Schedule: 5-7 years



The nuSTORM facility

Table 1: Decay ring specifications

Parameter	Specification	Unit
Central momentum P_μ	3.8	GeV/c
Momentum acceptance	$\pm 10\%$	
Circumference	480	m
Straight length	185	m
Arc length	50	m
Arc cell	DBA	
Ring Tunes (ν_x, ν_y)	9.72, 7.87	
Number of dipoles	16	
Number of quadrupoles	128	
Number of sextupoles	12	



注； 昨年P5でエンドースを受けられなかったため、FNALで行う可能性はほぼなくなり、CERNでの可能性を残すのみとなった。

Liquid scintillator detector

Many experiments for sterile neutrino search use liquid scintillator.

→ Explaining the detection principle at first

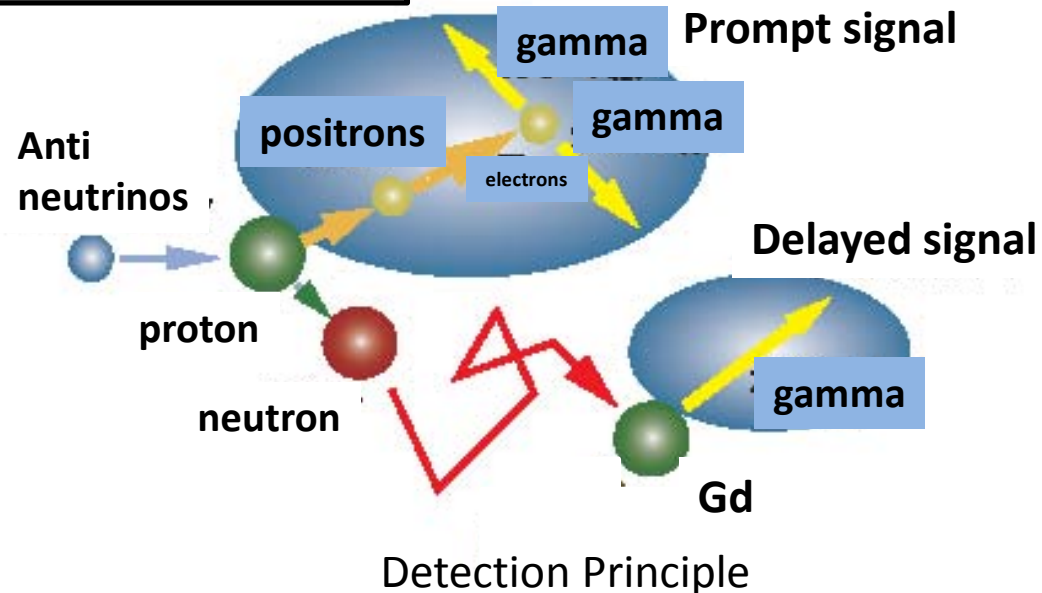
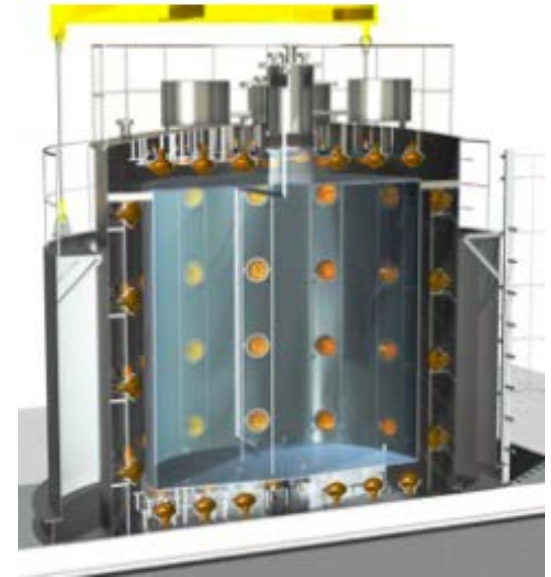
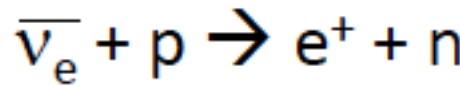
Identify neutrinos with detecting e⁺ and gammas from n capture on Gd (H).

=>Can reduce accidental BKG
(Gd~8MeV γs, capture time
~ several tens μs; Gd case).

Prompt $E_{\text{vis}} + 0.8\text{MeV} \sim E_{\nu}$;

- < a few MeV (for beta source)
- 2~ 3MeV (for reactor)
- ~ 40MeV (for muon DAR)

Inverse Beta Decay
(coincidence)



Severest BKG is fast neutrons from cosmic ray or reactor or acc. Beam.

→ Creating prompt + delayed correlated BKG.

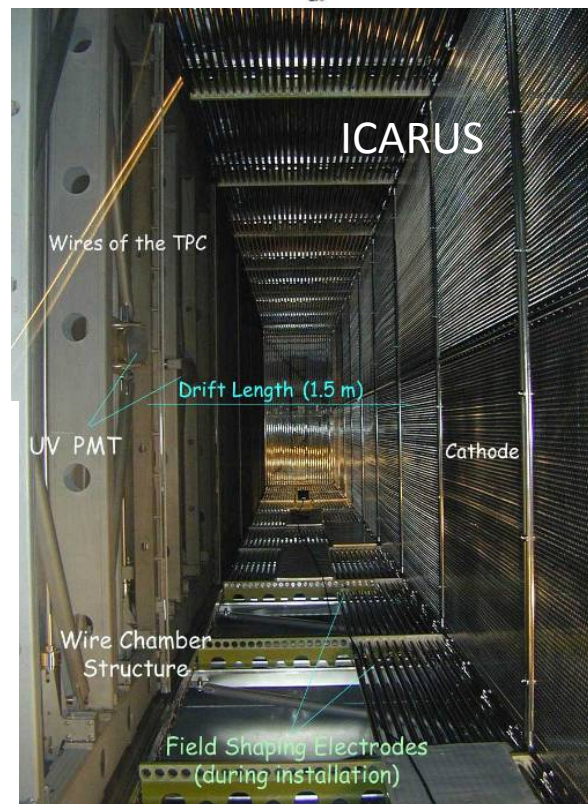
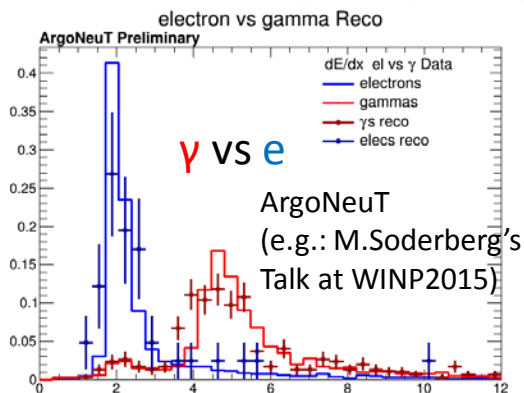
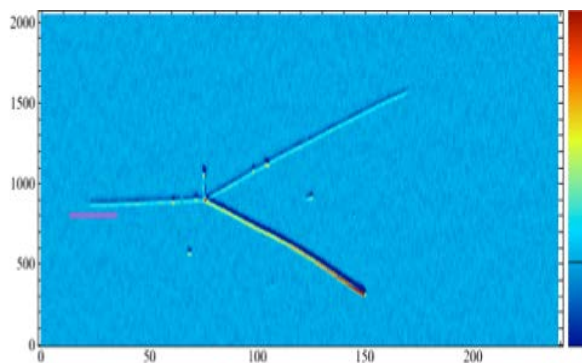
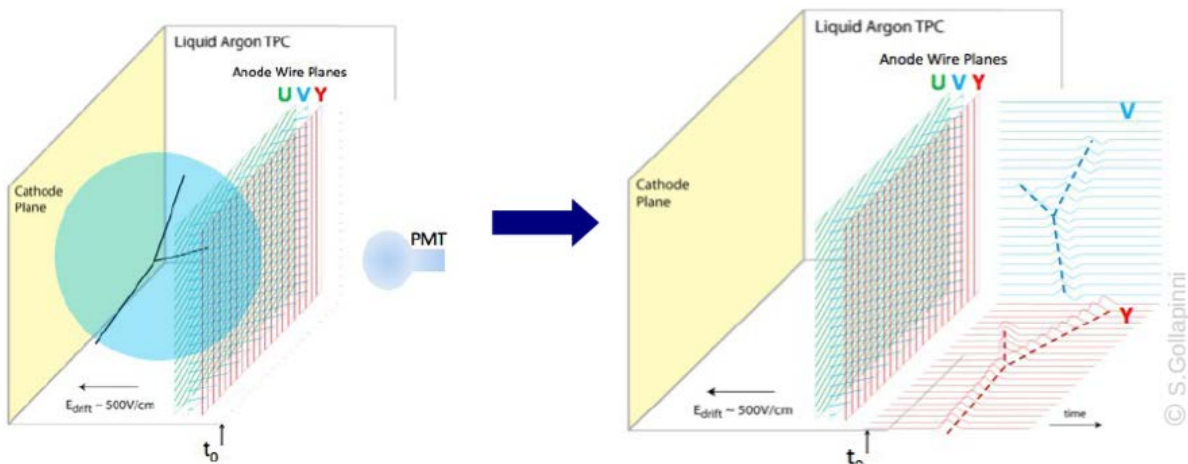
電離電子信号

- ~ 50000 電離電子/cm (MIP)
- $\sim 500\text{V/cm}$ の電場により $\sim \text{mm}/\mu\text{s}$ でドリフト
- 高純度の液体アルゴンを使い、ICARUSではほぼ1.5mのドリフト長で減衰8%程度 ($< 0.02\text{ppb}$)
- 液体中では増幅無し
- ドリフト電子のディフュージョンは小さい
- TPCとして三次元飛跡再構成

電子と π^0 の識別が得意。

- 全ての荷電粒子トラックが分かるので、 $\gamma \rightarrow e+e$ -事象が分かる。(vertex付近から2電子分の電荷)
- 電子/ γ 識別が得意と言った方が正確かも。

LAr TPC



Recent status

- Sterile neutrino is **one of the most interesting topics** in the neutrino community.
 - A quarter of the presentations in Neutrino 2014 conference mentioned the sterile neutrino.
 - P5 (Particle Panel) endorsed the short baseline neutrino experiments to search for sterile neutrinos.

Presentation and Discussion of P5 Report by S.Ritz (on 22-May-2014)

- **Recommendation 15: Select and perform in the short term a set of small-scale short-baseline experiments that can conclusively address experimental hints of physics beyond the three-neutrino paradigm.**

Based on the P5 report, neutrino physics during intermediate span (5-10years) were discussed at BNL Feb 4-6 (WINP 2015)

<https://indico.bnl.gov/conferenceDisplay.py?confId=918>

Very interesting topics are discussed. Please see if you are interested.

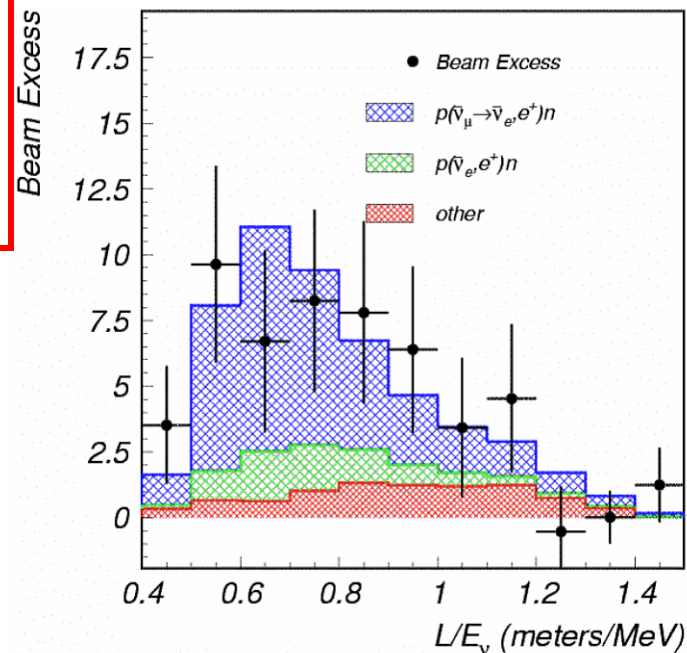
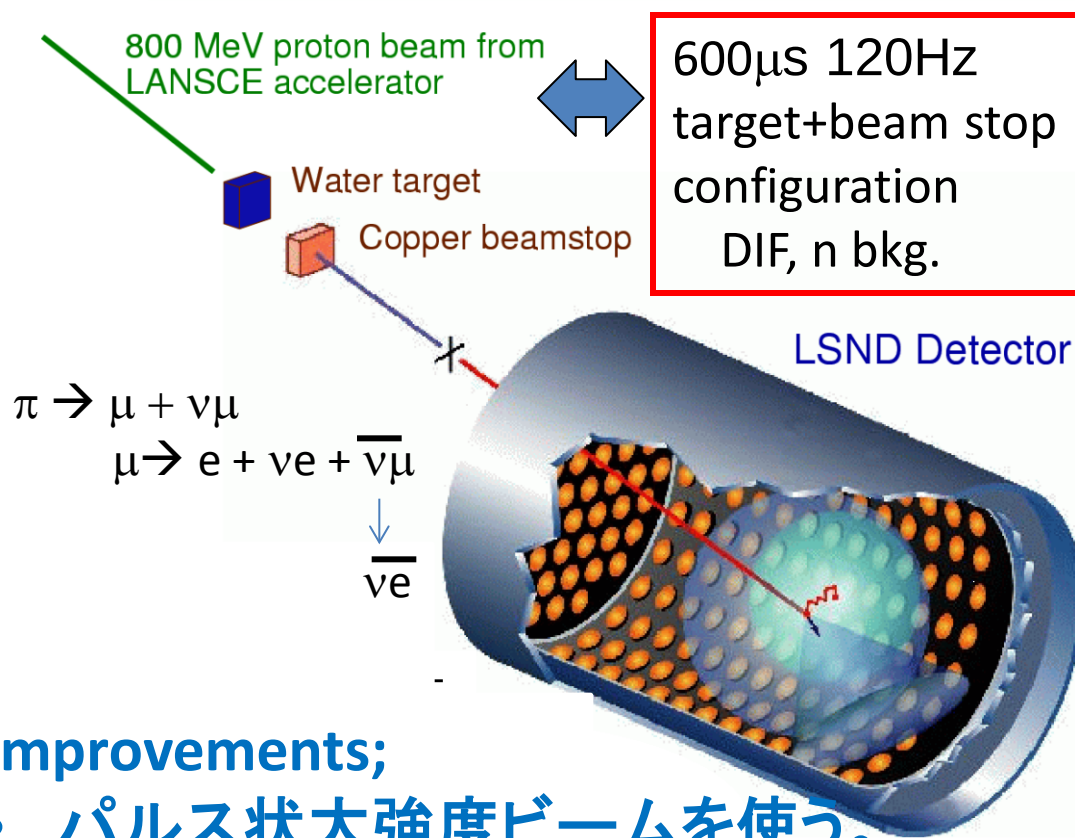
最近の傾向・動向

- 振動の議論をする前に、事象超過や消失が本当なのかどうか(隠れた)系統誤差を極限まで減らして結論を出す。
 - ビームの不定性
 - 検出器の不定性
 - 背景事象の不定性
- LSND, MiniBooNE, Ga, Reactor とともに、振動であれば見えるべき振動パターンを見ていない。→ 見るべし
 - 距離の関数として
 - エネルギーの関数として
 - L/E の関数として
- $(3+1)$ モデルが正しいのであれば、 $\nu_\mu \rightarrow \nu_\mu$ 消失も $\nu_e \rightarrow \nu_e$ 消失と同じ確率で起こるはず。
 - SK、MINOS+、ICECUBE
 - 今のところ、兆候なし。

Appearance

LSND $\bar{\nu}_\mu \rightarrow \bar{\nu}_e$ Signal

1998



Saw an excess of:
 $87.9 \pm 22.4 \pm 6.0$ events.

With an oscillation probability of
 $(0.264 \pm 0.067 \pm 0.045)\%$.

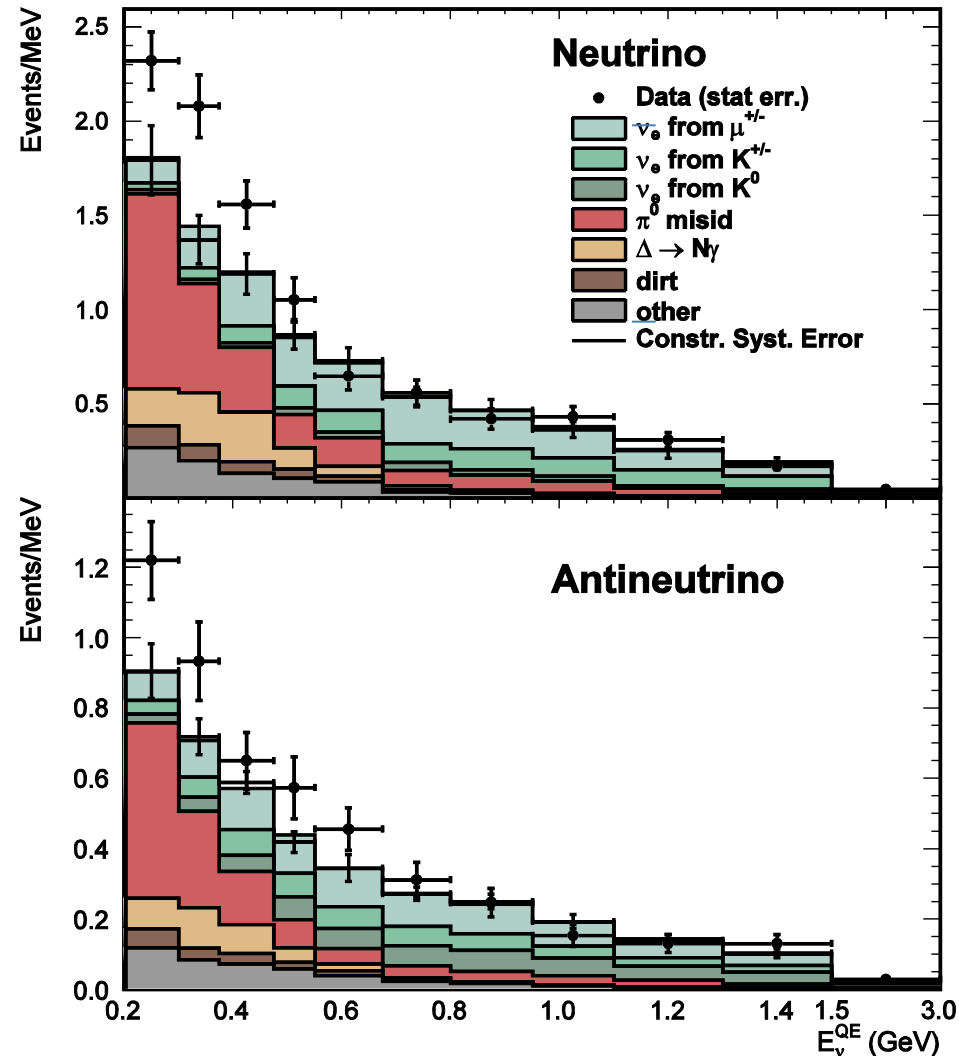
3.8 σ evidence for oscillation.

π^-, μ^- absorbed before decay into ν 's
 there should not be $\bar{\nu}_e$ at the level of 7×10^{-4}

Signal : $\bar{\nu}_e p \rightarrow e^+ n$ $np \rightarrow d \gamma(2.2\text{MeV})$

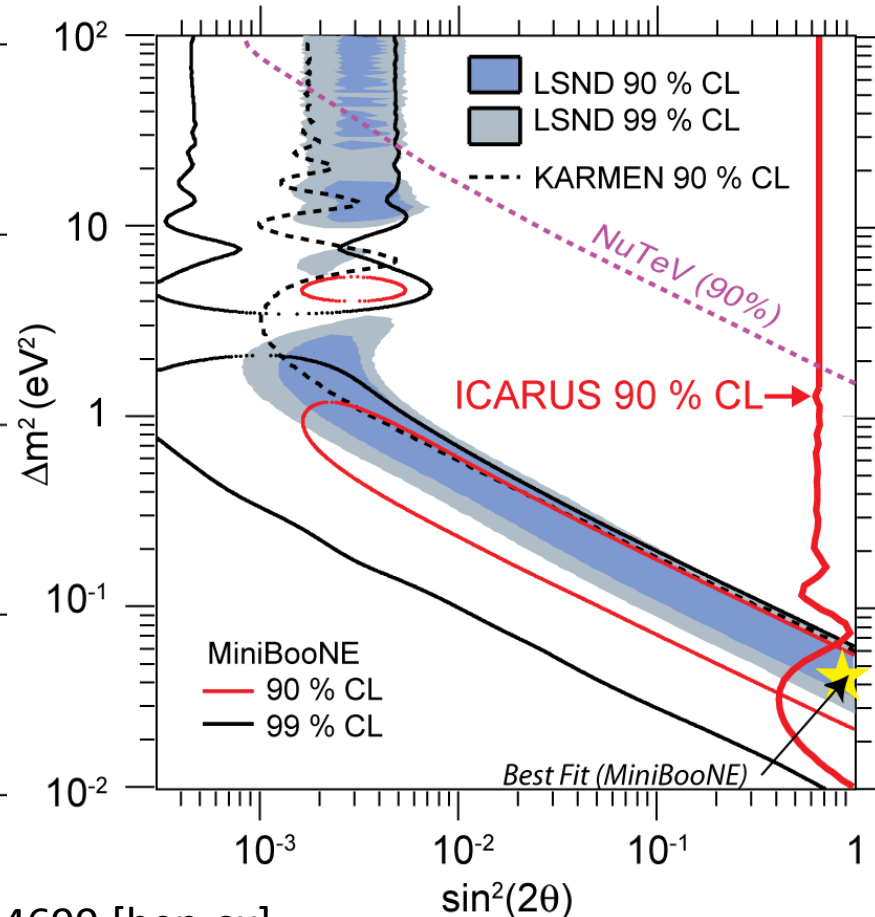
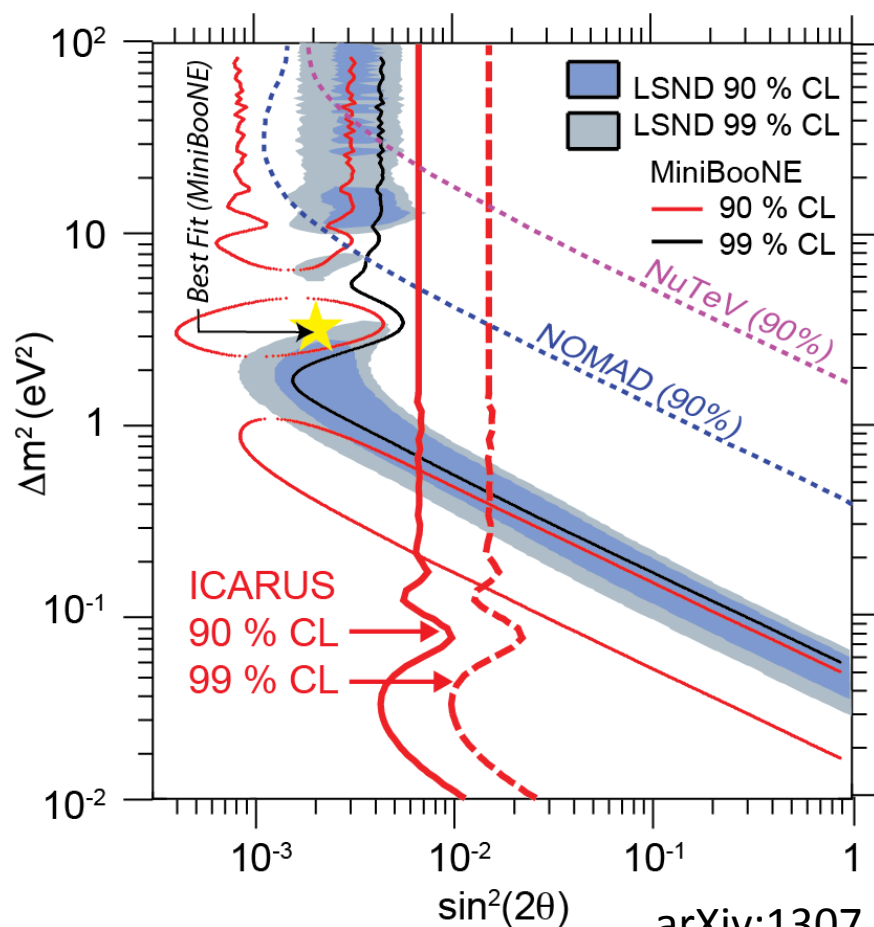
MiniBooNEからのimprovements

- 弱点とimprovements
 - 背景事象の絶対量が多い。
→ MiniBooNEはミネラルオイルチェレンコフ。 π^0 からの 2γ を電子と間違ってしまうことも多い。右図のうち、液体アルゴンTPCを使えば、赤部の π^0 misIDがなくなる。
 - そもそも、こんなに多い背景事象をMCのみに頼って予言するのは難しい。
→ 2つ以上の検出器を設置して、nearとfarの検出器でのエネルギー分布を比較する。
 - E_ν 再構成; 原子核効果影響



Status of excess of $\bar{\nu}_\mu \rightarrow \bar{\nu}_e$ or $\nu_\mu \rightarrow \nu_e$

Neutrino Antineutrino



arXiv:1307.4699 [hep-ex].

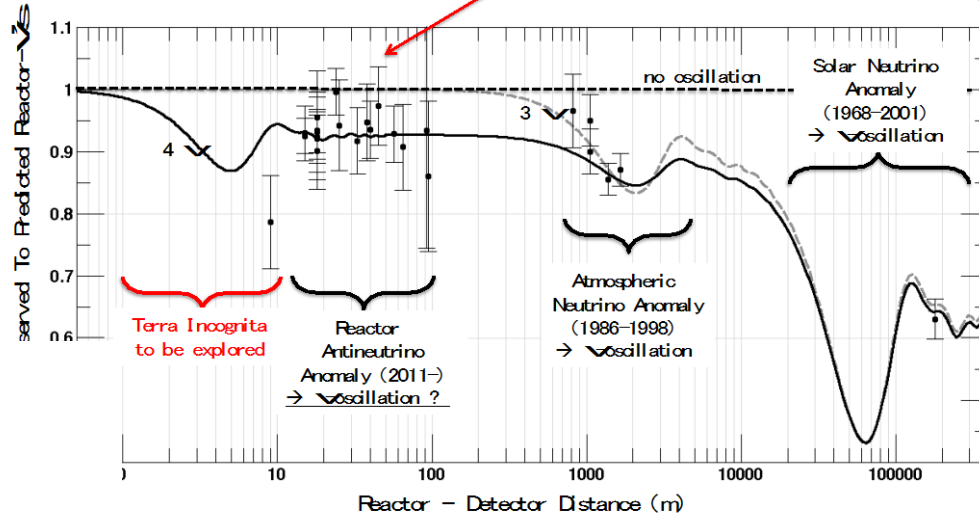
- LSND and MiniBooNE see the excess of the events.
- 3 generation model cannot explain oscillation with $\Delta m^2 > \sim 1.0 \text{ eV}^2$ region.
- Z measurements conclude 3 active $\nu \rightarrow$ sterile
- 15年間解決しない、この混沌とした状況にどう終止符を打つか、実験屋の腕の見せどころ。

ν_e disappearance in reactor and β -source

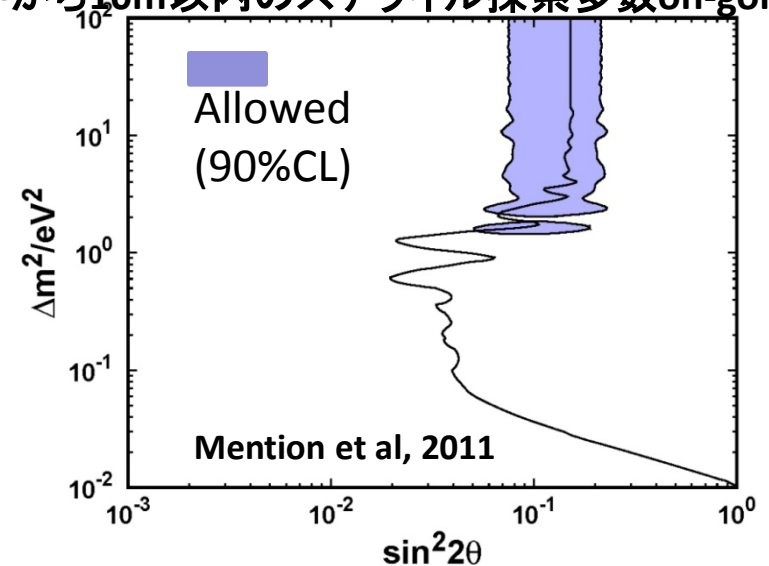
Chris Polly, Thierry Lasserre
NEUTRINO2012

The Reactor Antineutrino Anomaly

- Observed/predicted averaged event ratio: $R=0.927 \pm 0.023$ (3.0 σ)



- 原子炉からの反電子 ν のフラックス精密調査中
- 炉から10m以内のステライル探索多数on-going

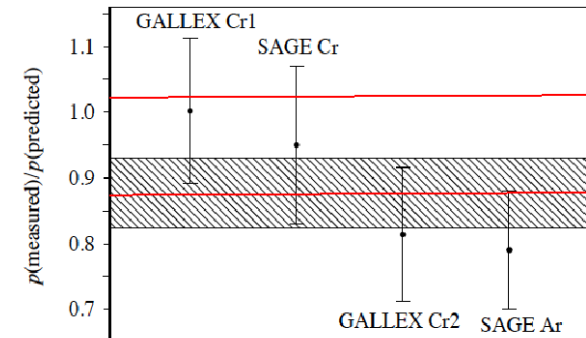
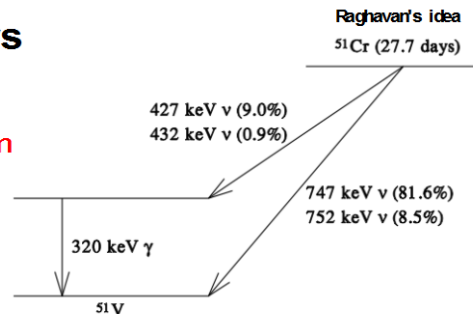


The Gallium Neutrino Anomaly

- Test of solar neutrino detectors
GALLEX and SAGE (ν_e 's)

- $E \approx \text{MeV}$, Baseline range $\approx$ few m

- 4 calibration runs
 $\approx 1 \text{ MQ EC } \nu_e$ emitters



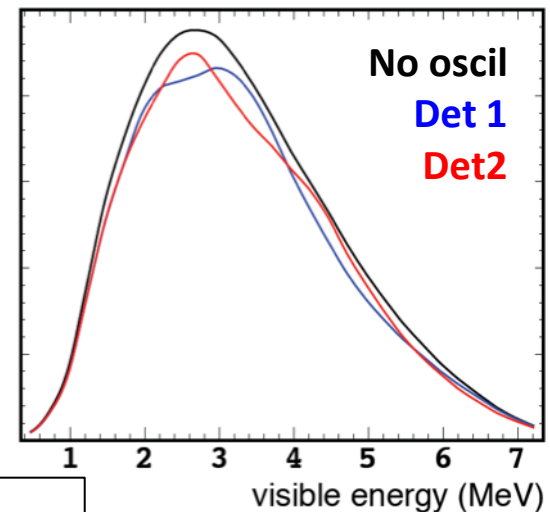
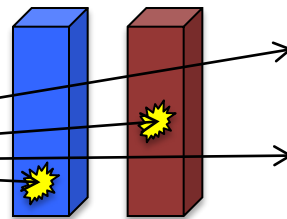
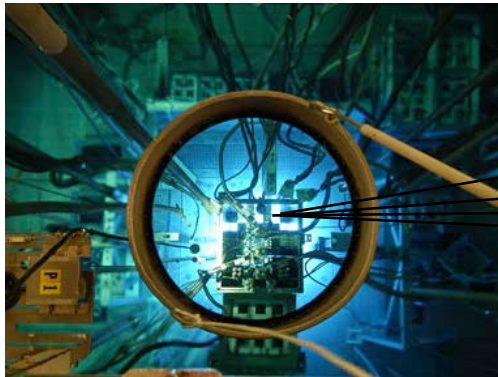
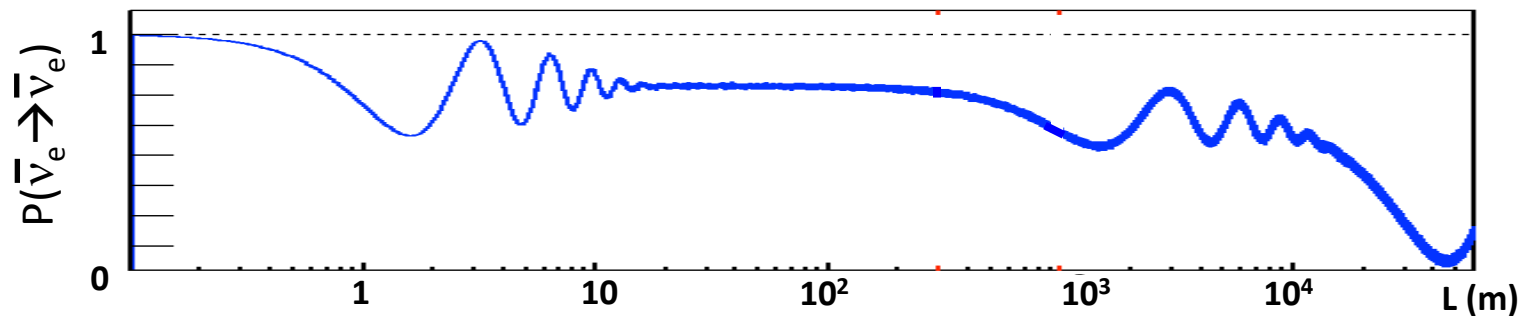
$\bar{\nu}_e \rightarrow \bar{\nu}_e$ disappearance

Many talks will be presented later, so
I just show the prospects briefly.

Testing the New Oscillation Hypothesis

Direct test of a new oscillation pattern in E & L

$$\Delta m^2 > 0.1, \sin^2 2\theta > 0.05 \rightarrow L_{\text{osc}} = [1-10] \text{ m}$$



- Relative shape distortion in identical detector modules
- Complemented by rate info.

David Lhuillier's
talk at Neutrino2014

Candidate detectors

Picking up typical experiments.
Not covering all (sorry!)

Requirements

- Underground location;
- Large mass, ultra-pure detectors;
- Capability to measure E and L (oscillometry);

Existing large liquid scintillator experiments (Borexino, KamLAND, Daya-Bay)
Future large liquid scintillator experiments (SNO+, LENS, JUNO)
Future experiments based on other techniques (RICOCHET, BEST)



Experiments with radioactive sources

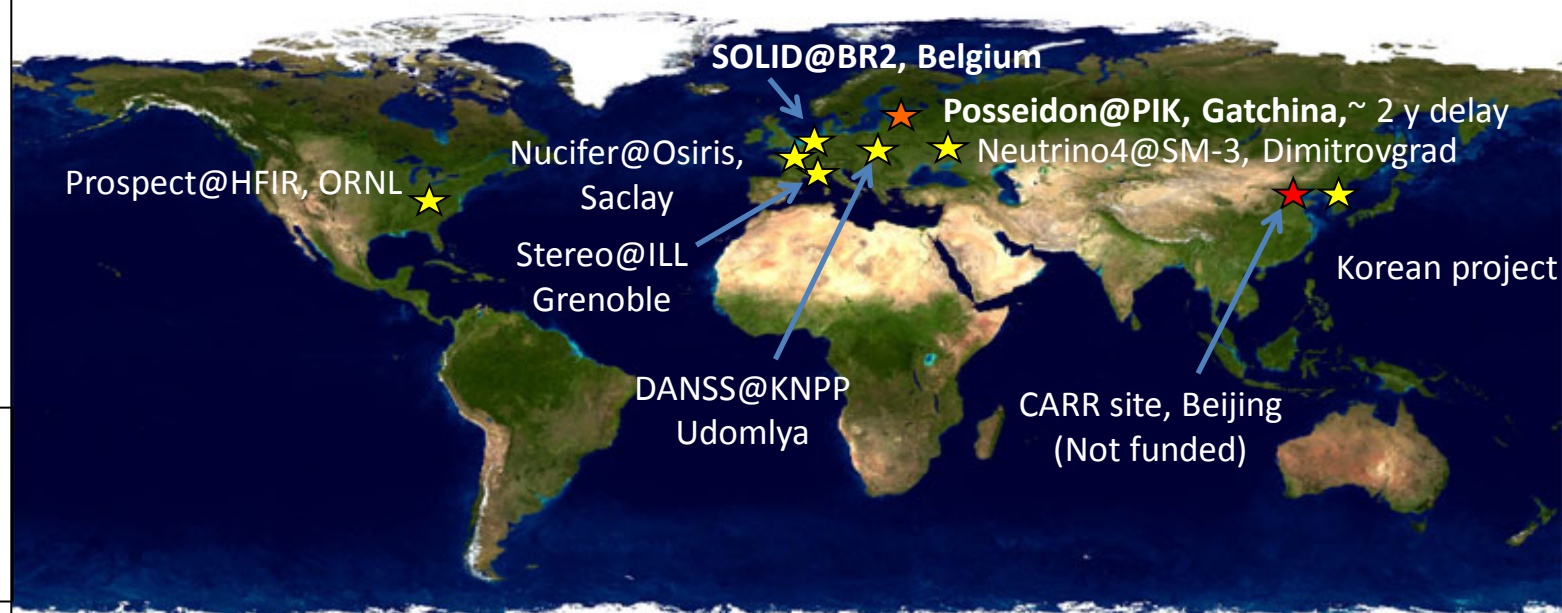
Barbara Caccianiga's talk at Neutrino 2014

Many interesting and exciting results are coming in near future. (2015~)

Many talks will be also coming

Experiments with reactors

David Lhuillier's talk at Neutrino2014



Barbara Caccianiga’s talk at Neutrino 2014

- Need **~10MCi** level radio active source
 ^{51}Cr , ^{144}Ce - ^{144}Pr are considered.
Typically $E_\nu < \text{a few MeV}$
- Need to transport the source to detector

Technique	Detector	Sources	Reaction	Activity	Reference
Large Liquid scintillator detectors	SOX (Borexino)	^{51}Cr	$\nu + e \rightarrow \nu + e$	10MCi	<i>JHEP08(2013)038</i> ,
		^{144}Ce - ^{144}Pr	$\nu + p \rightarrow e^+ + n$	100kCi	<i>Phys. Rev. Lett. 107, 201801 (2011)</i>
	KamLAND	^8Li (ISODAR)	$\bar{\nu} + p \rightarrow e^+ + n$	8.2×10^{14} v/sec	<i>arXiv:1205.4419</i> , <i>arXiv:1310.3857</i>
		^{144}Ce (CeLAND)	$\bar{\nu} + p \rightarrow e^+ + n$	100kCi	<i>arXiv:1312.0896</i>
	Daya-Bay	^{144}Ce - ^{144}Pr	$\bar{\nu} + p \rightarrow e^+ + n$	500kCi	<i>arXiv:1109.6036</i>
	LENS	^{51}Cr	$\nu + ^{115}\text{In} \rightarrow ^{115}\text{Sn}^* + e$	10MCi	<i>Phys.Rev.D75 093006(2007)</i>
	JUNO	^8Li (ISODAR)	$\bar{\nu} + p \rightarrow e^+ + n$	8.2×10^{14} v/sec	<i>arXiv:1310.3857</i>
Radiochemical	BEST	^{51}Cr	$\nu + ^{70}\text{Ga} \rightarrow ^{71}\text{Ge} + e$	3MCi	<i>arXiv:1204.5379</i>
Bolometers	Richochet	^{37}Ar	$\nu + N \rightarrow \nu + N$	5MCi	<i>Phys. Rev. D85, 013009, (2012)</i>

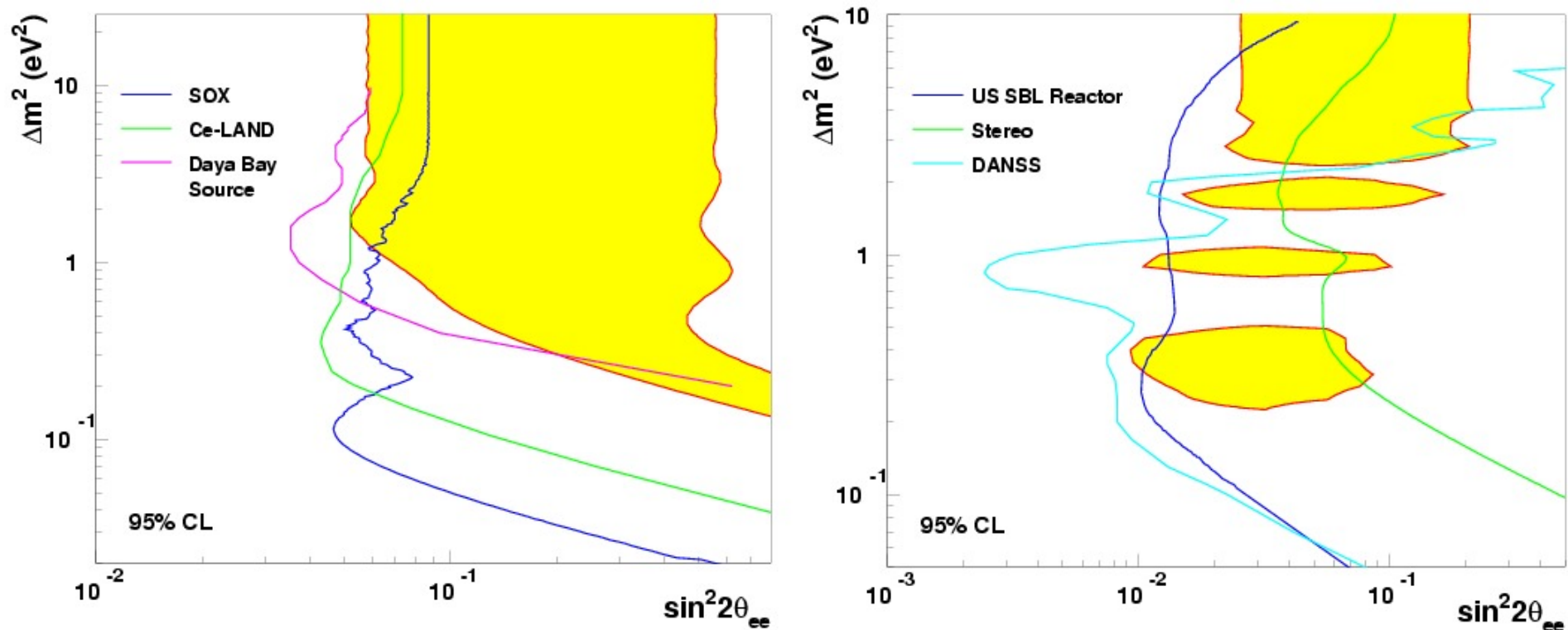
	P_{th} (MW)	M_{target} (tons)	L (m)	Depth (m.w.e.)
On-going Nucifer (FRA)	70	0.8	7	13

- Need PID to reject fast neutrons from reactors. PSD (Pulse Shape Discrimination) helps.
- Need sufficient shield.
- Need to see “Oscillation pattern”

Updated table from David Lhuillier’s talk at Neutrino 2014

	Poseidon (RU)	100	~ 3	5-8	~ 15
Start in 2015	Stéréo (FRA)	57	1.75	8.8-11.2	18
Start in 2015	Neutrino 4 (RU)	100	1.5	6-12	~ 10
Start in 2015	Hanaro (KO)	30-2800	~ 1	6	few
Final detector In 2015	DANSS (RU)	3000	0.9	9.7-12.2	50
White paper was submitted	Prospect (USA)	85	1 & 10	7-18	few
Aim to operate In 2015	NuLat (USA)	20-1500	~ 1	2.5-8	2-10
Start in 2016	SoLid (UK)	45-80	2.9	6-8	10

Sensitivities (arXiv; 1310.4340)



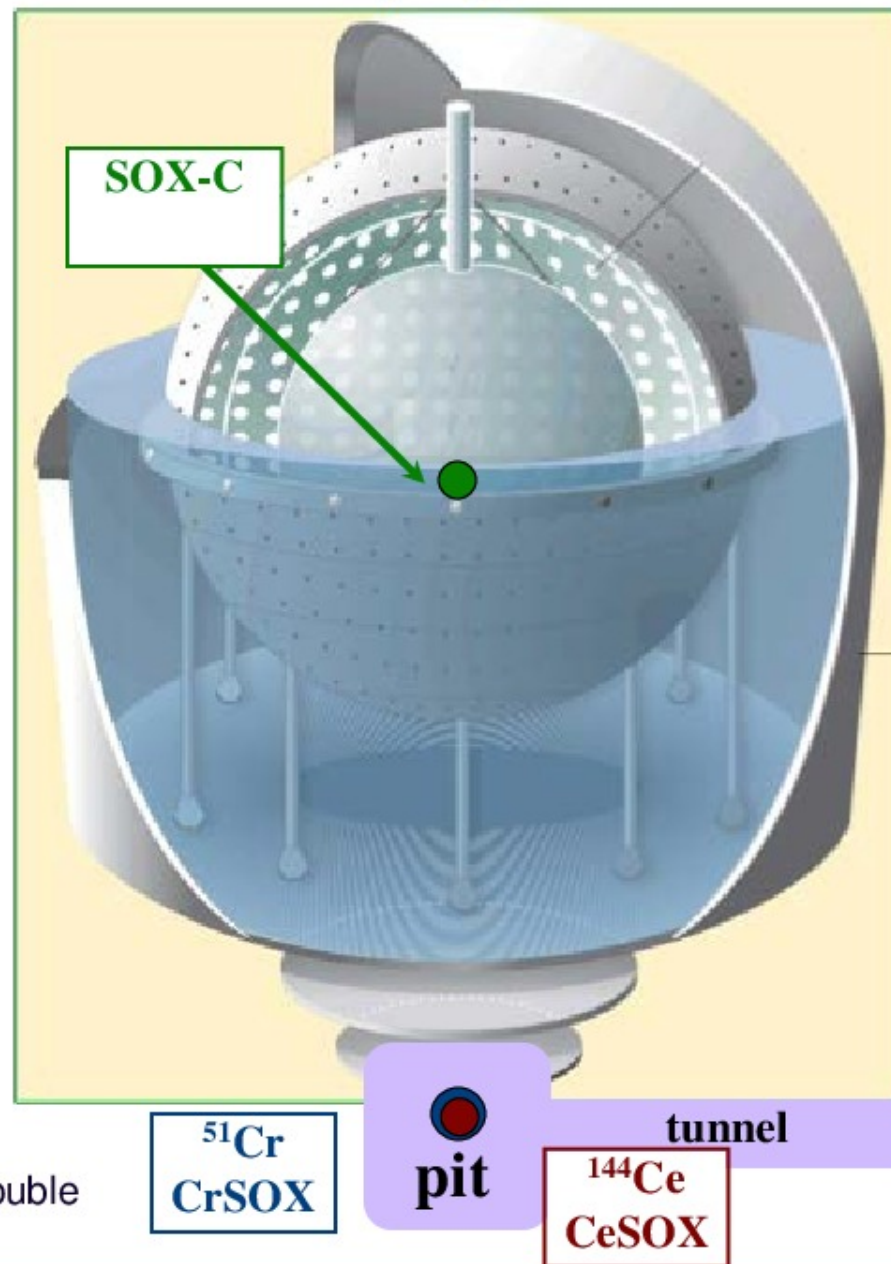
They can achieve the sensitivity with short time scales.
(typically < a few years for 95% C.L.)



SOX: the potential full ^{144}Ce and ^{51}Cr program

Gioacchino Ranucci's talk @ WINP 2014

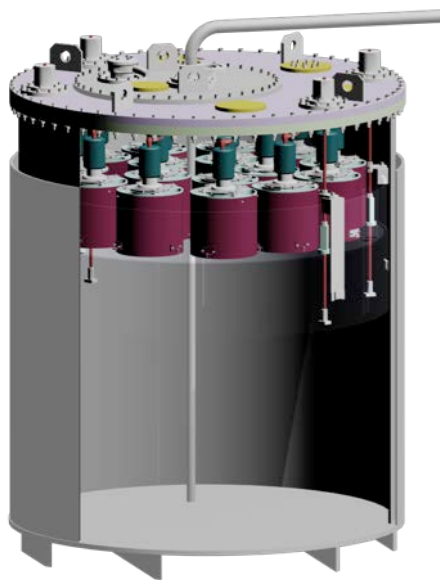
- **Mission:** test the existence of low L/E ν_e and/or $\bar{\nu}_e$ anomalies by placing well known artificial sources close to or inside **Borexino**
- **CrSOX**
 - ^{51}Cr source in pit beneath detector
 - 8.25 m from center
 - Activity 10 M Ci (or two exposures at 5 M Ci)
- **CeSOX First to be accomplished**
 - ^{144}Ce - ^{144}Pr source in in pit beneath detector
 - 8.25 m from center
 - Activity 100-150 kCi
 - **Planned start-up date middle 2016**
- **SOX-C**
 - ^{144}Ce - ^{144}Pr or ^{51}Cr source in the center
 - Would imply a full refurbishment of the detector (suited also for a possible future Double Beta Decay exp. with Xenon)



Nucifer @ OSIRIS

既にon-goingなNUCIFERの話をする

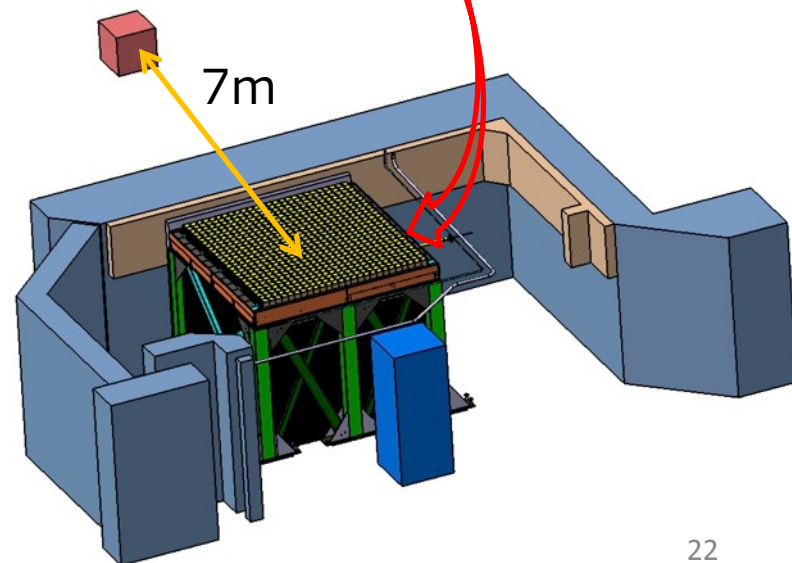
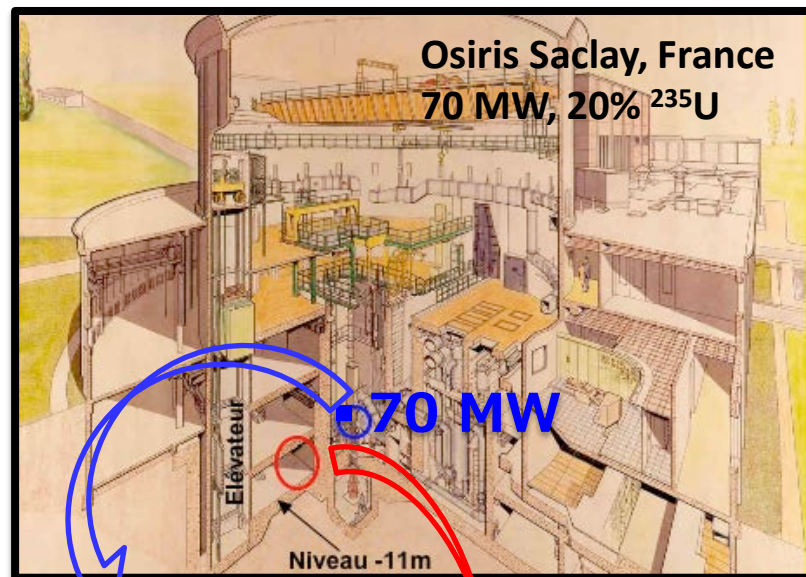
Simple design for reactor monitoring studies



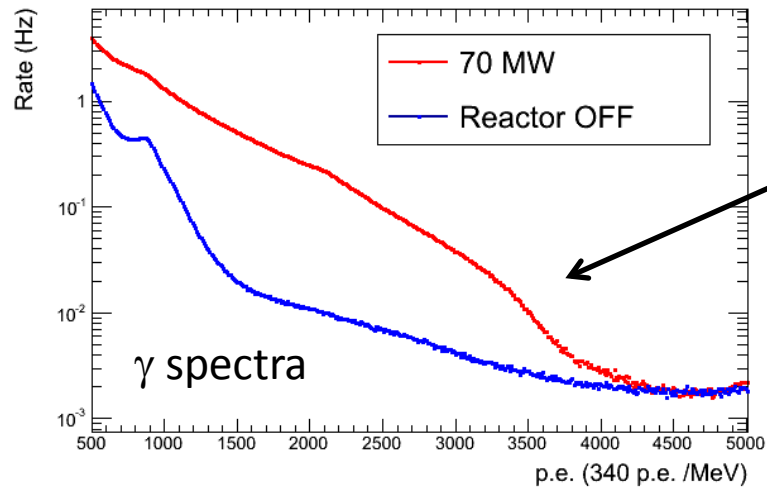
- 850 l Gd-loaded LS
- Compact core (60x60x60 cm³)
- Short baseline: 7m
- 10 mwe overburden
- ~300 detected ν /day expected

→ Some sensitivity to Sterile ν

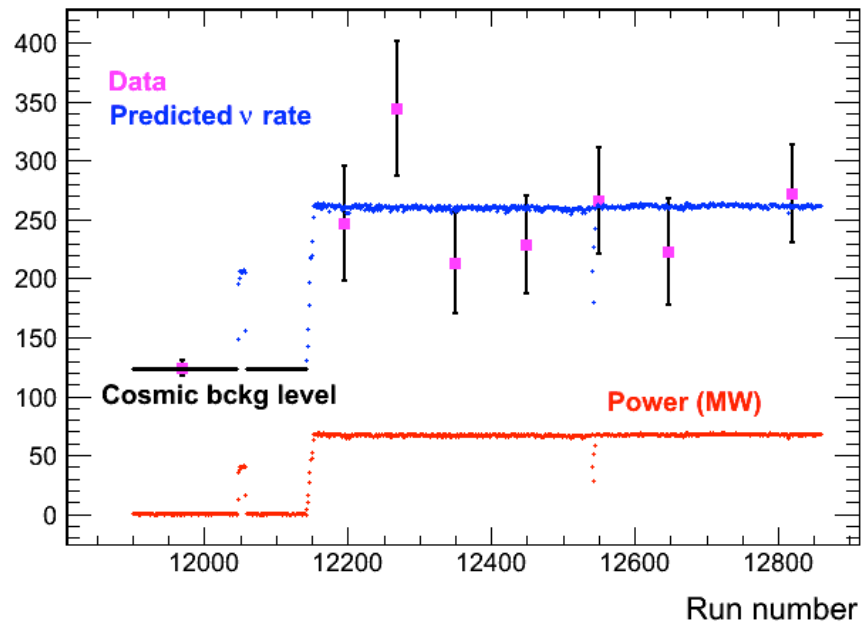
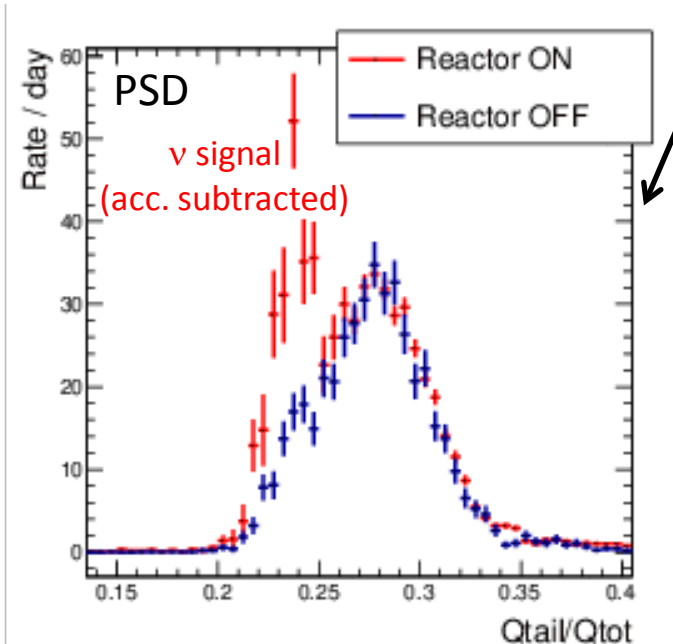
But challenging reactor background



Nucifer @ OSIRIS



- Need further γ attenuation to reach $\nu/\text{Acc} \sim 1$. New data taking next week with extra 4 cm of lead on reactor side.
- Rejection of cosmic background with PSD
No reactor induced fast n @ 7m from core.



$\nu_{\mu} \rightarrow \nu_e$ appearance

Accelerator based neutrino
experiments.

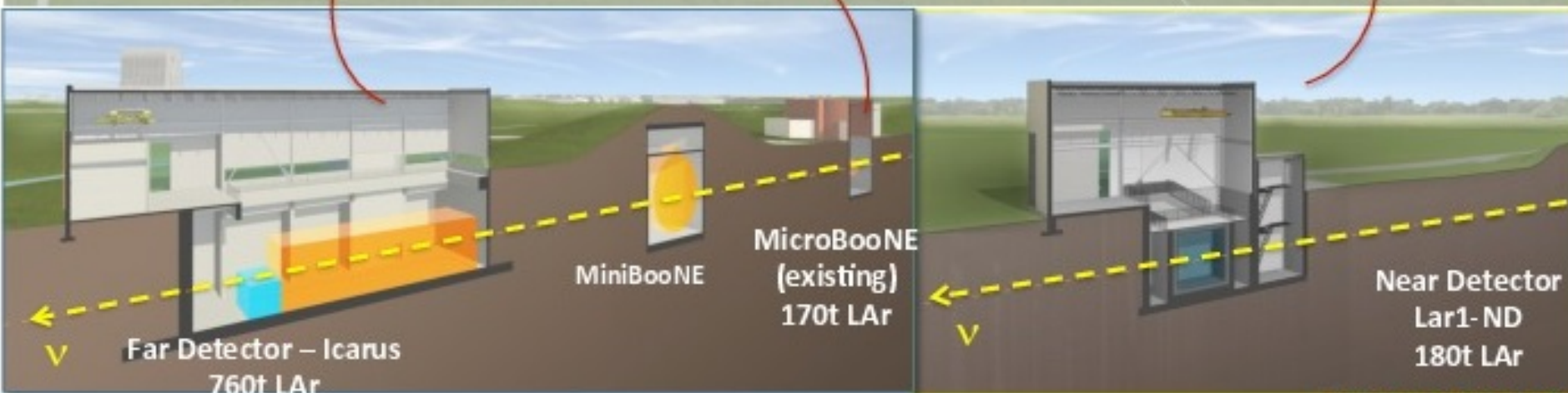
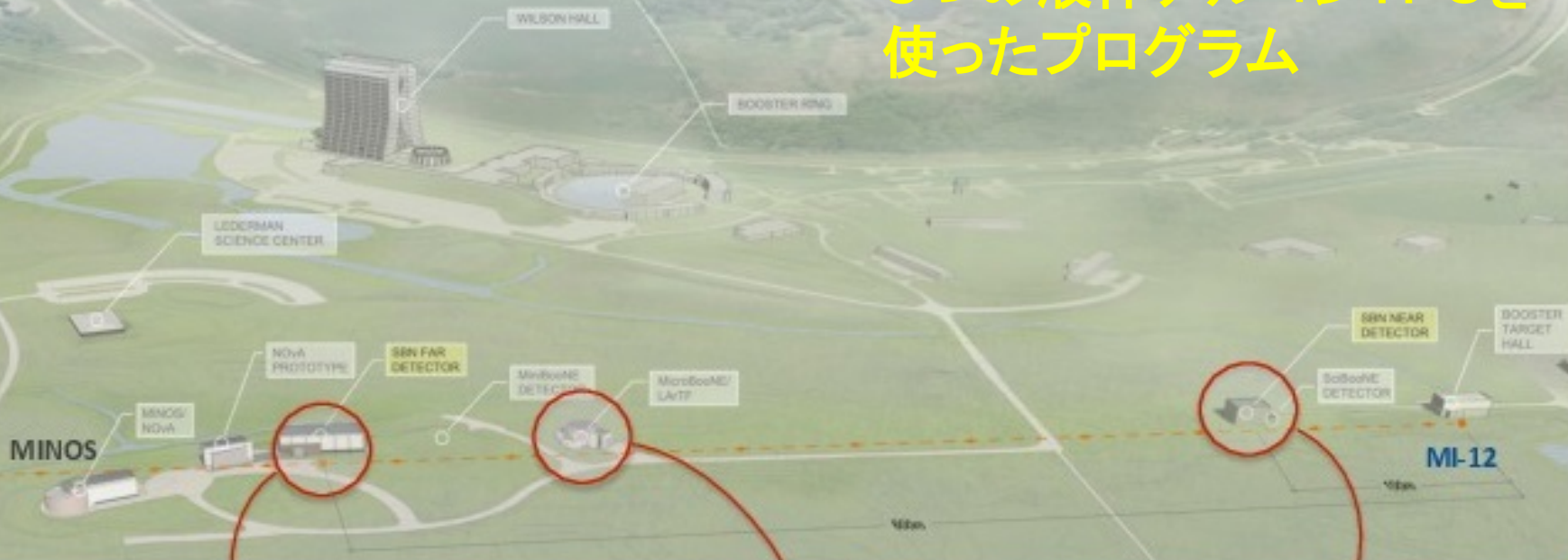
$\nu_\mu \rightarrow \nu_e$ appearance

- LSND / MiniBooNE have significant excess -> to be checked
 - μ Decay-At-Rest source ; J-PARC E56 @MLF and OscSNS (white paper was submitted)
 - Better pulsed beam than LSND ($\mu^+ \rightarrow e^+ + \nu_e + \bar{\nu}_\mu$; $\bar{\nu}_\mu \rightarrow \bar{\nu}_e$)
 - Better liquid scintillator detector than KARMEN / LSND (PID / Gd loaded)
 - Direct test for LSND anomaly (same detector material + μ DAR)
 - Conventional horn focused beam;
 - LAr; MicroBooNE, ICARUS, LAr1-ND -> triple LAr@FNAL ($\pi^+ \rightarrow \mu^+ + \nu_\mu$; $\nu_\mu \rightarrow \nu_e$ / reversed horn polarity)
 - Better detector (BKG rejection, e ID, ν energy reconstruction) than MiniBooNE $\rightarrow$ reduced #BKG and systematics on BKG
 - Triple detectors $\rightarrow$ flux uncertainty can be reduced.
 - same beam condition as the MiniBooNE
 - WC; ν PRISM 50m tall x 10-12m diameter WC @1km (T2K beamline)
 - IsoDAR/DAEdALUS ; 10mA level very powerful cyclotron.
 - New type of beam + Fe+scintillator detector; nuSTORM
 - Using neutrinos from STORed Muons. ($\mu^+ \rightarrow e^+ + \nu_e + \bar{\nu}_\mu$)

SBN Program Layout

Peter Wilson's talk @ WINP workshop

3つの液体アルゴンTPCを使ったプログラム

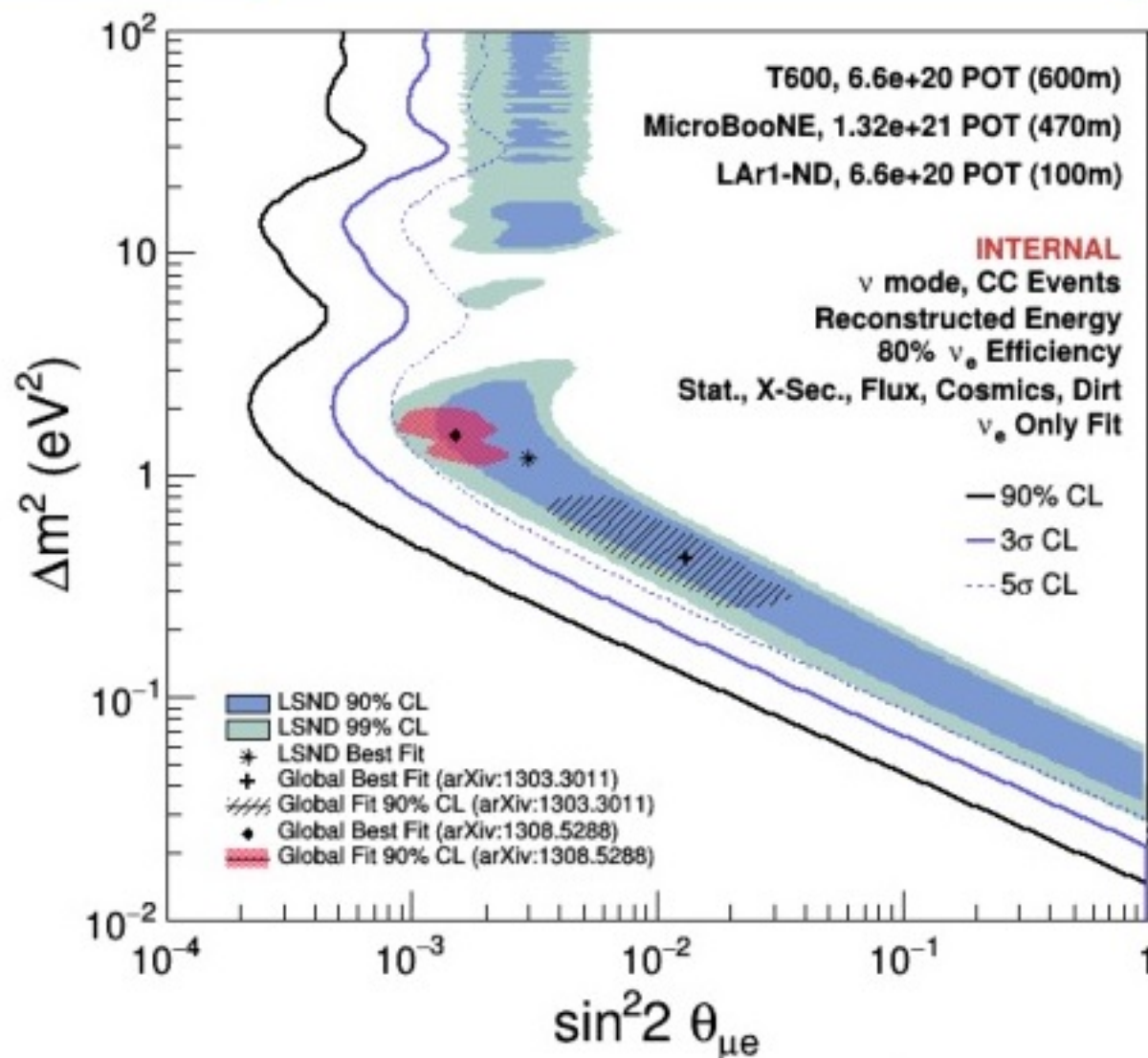
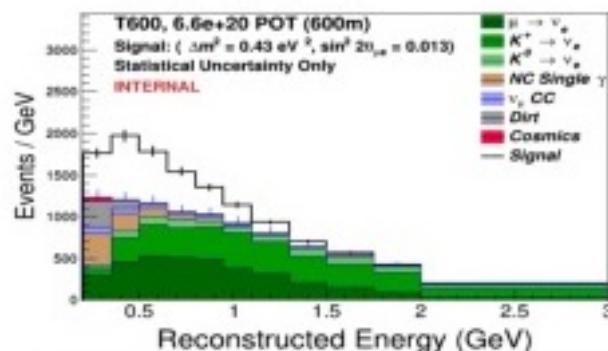
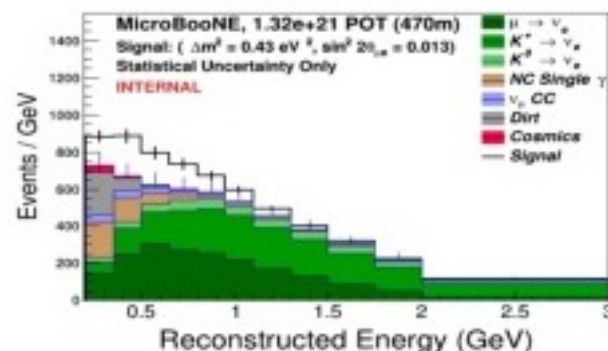
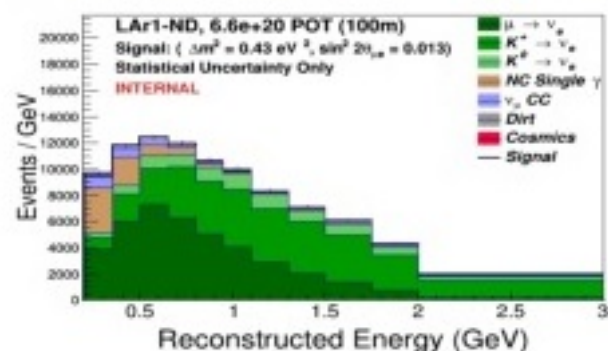


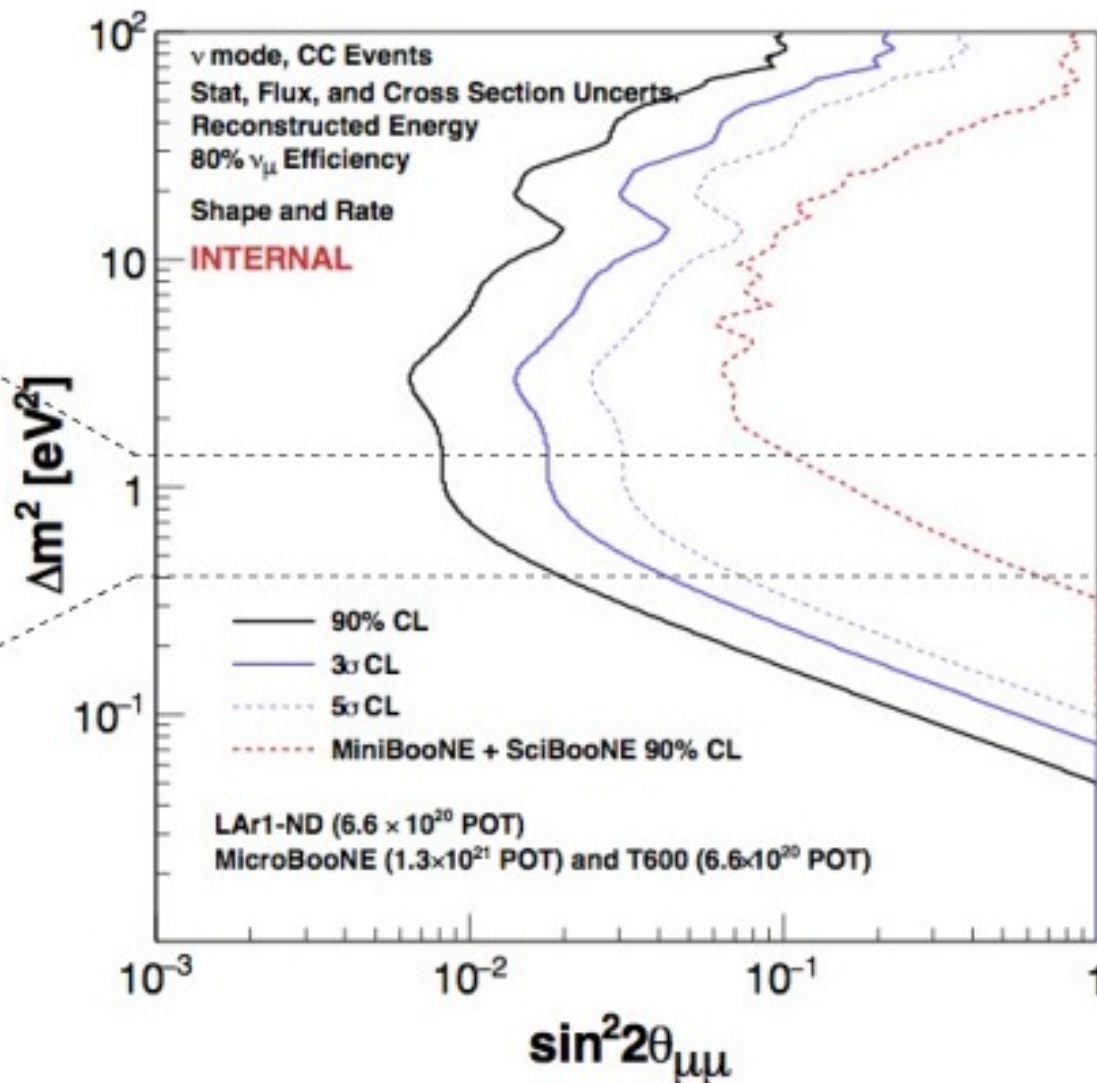
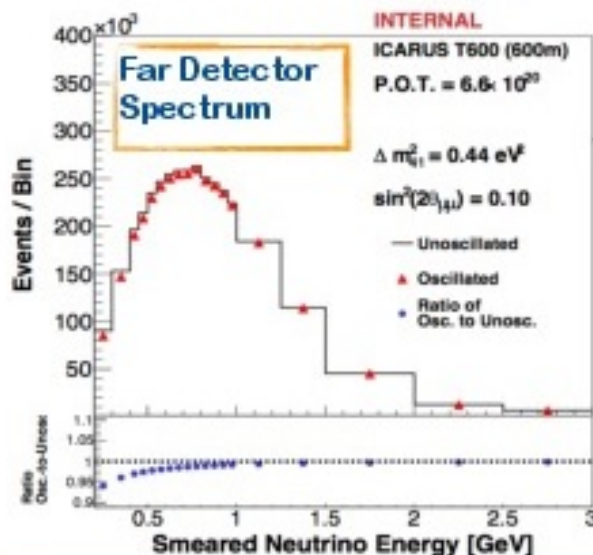
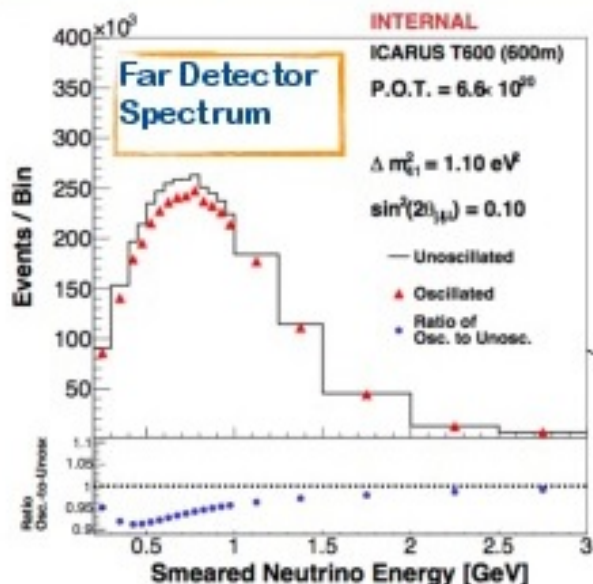
Fermilab

液体中の電離電子ドリフト速度は遅い(m/ms) →
地上でやるので、宇宙線が問題になるかも。

SBN ν_e Appearance Sensitivity

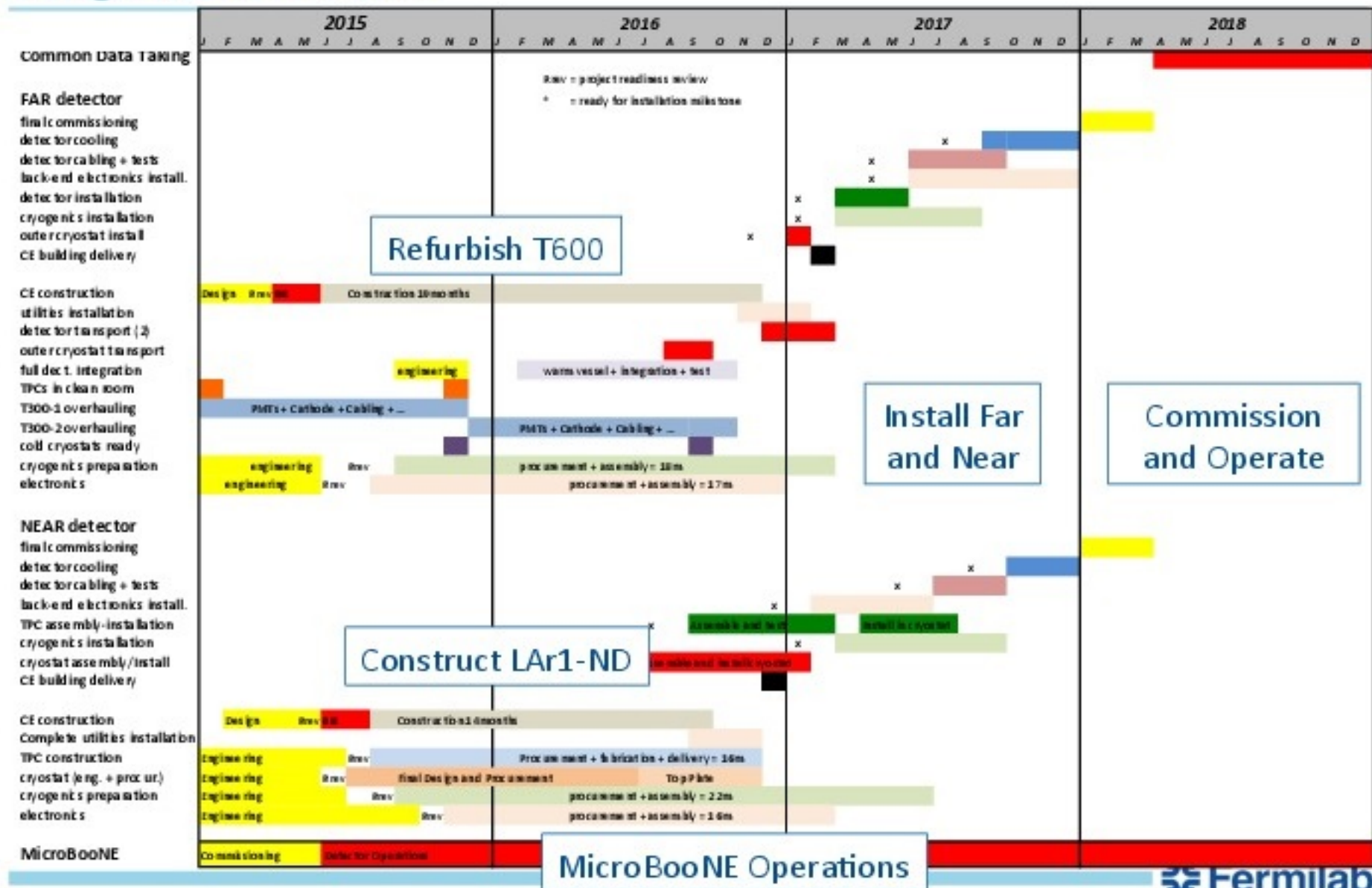
Peter Wilson's talk @ WINP workshop





Program Schedule

Peter Wilson's talk @ WINP workshop



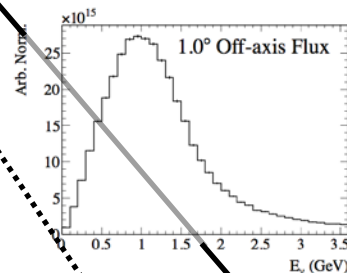
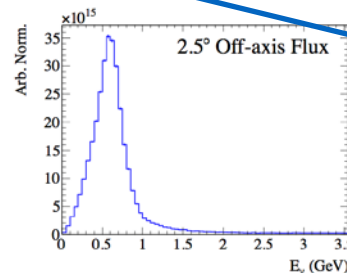
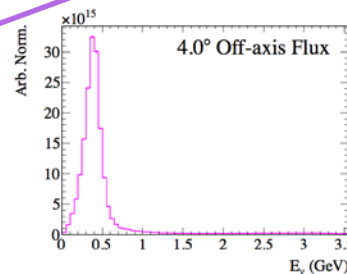
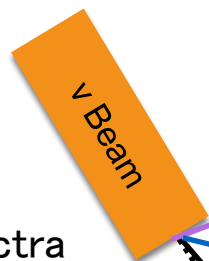
NuPRISM

water-Cherenkov
detector at 1 km in the
J-PARC beam

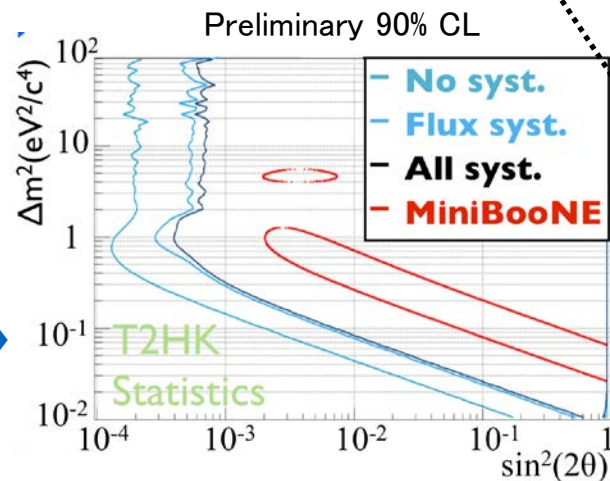
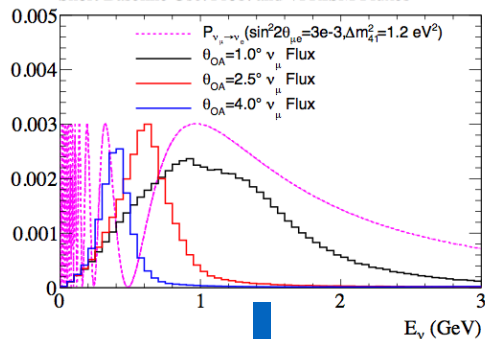
Range of off-axis angles
= Range of neutrino spectra

0.5% ν_e contamination near peak
= Short-baseline ν_e appearance
measurement

Unique capability to see oscillation
pattern in observed off-axis angle as
well as reconstructed energy



Short Baseline Osc. Prob. and νPRISM Fluxes



J-PARC E56 Sterile ν search @MLF

<http://research.kek.jp/group/mlfnu/>

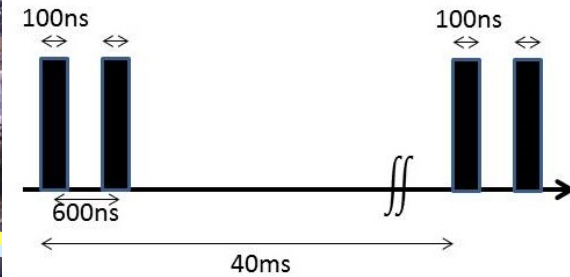
J-PARC Facility
(KEK/JAEA)

South to North

400MeV

3 GeV RCS

Neutrino Beams
(to Kamioka)



25Hz 300kW now &
will be 1MW

Materials and Life
Experimental Facility

30GeV MR

Hadron hall

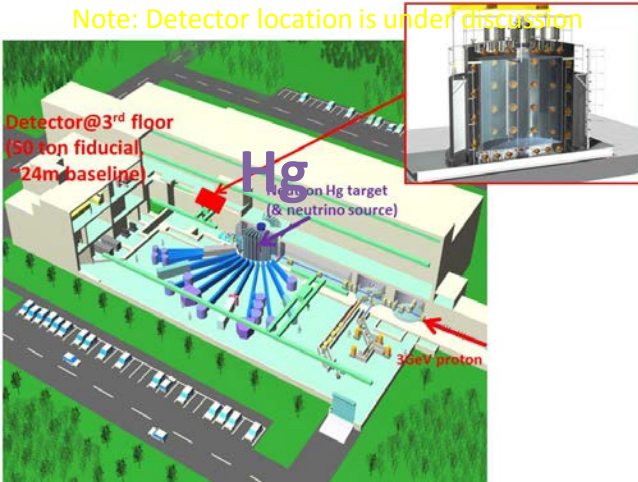
— CY2007 Beams
— JFY2008 Beams
— JFY2009 Beams

Bird's eye photo in January of 2008

Sterile neutrino search @MLF (proposal in 2013)

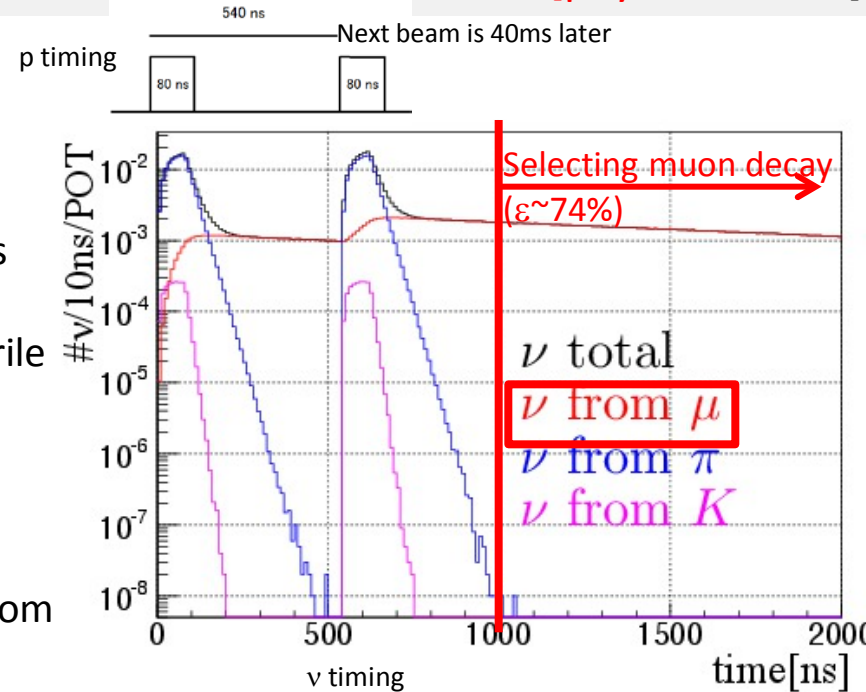
M. Harada *et al*, [arXiv:1310.1437](https://arxiv.org/abs/1310.1437) [physics.ins-det]

Note: Detector location is under discussion

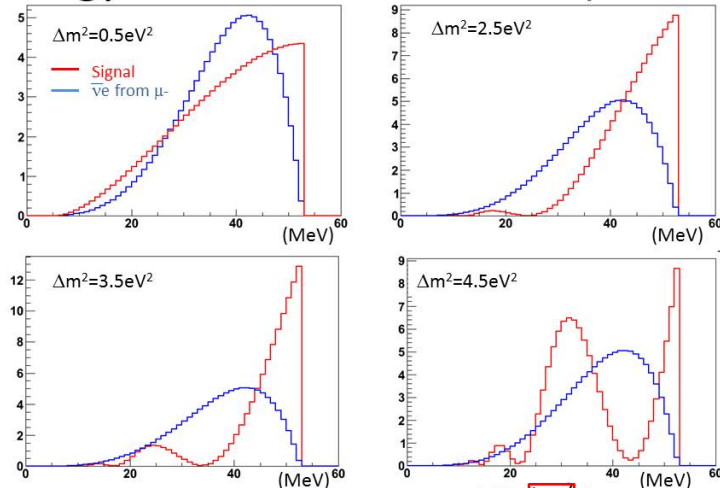


J-PARC E56

- confirms or refutes the neutrino oscillation with sterile neutrino ($\bar{\nu}_\mu \rightarrow \bar{\nu}_e$)
- uses ultra-pure neutrinos from stopping μ^+
- separates signals from BKG by measuring energy distortion

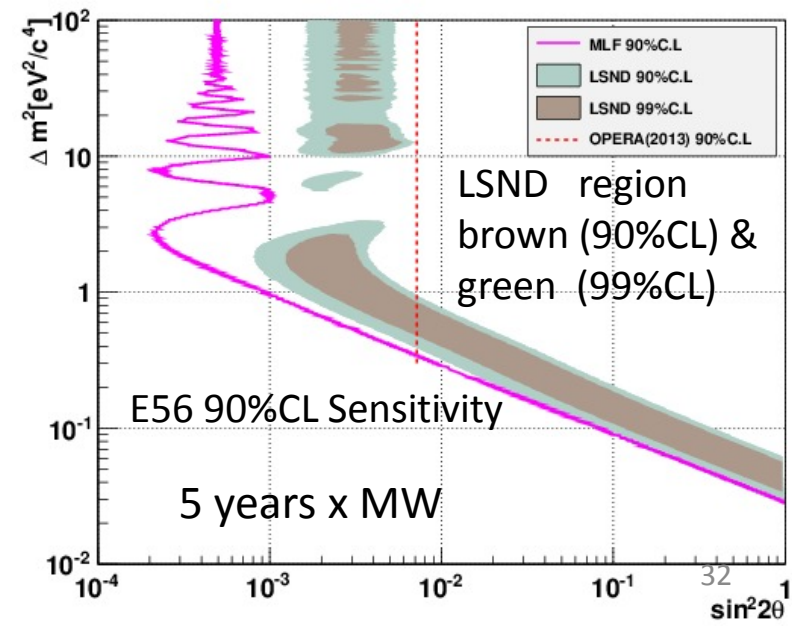


Energy distribution of events (L=24m)



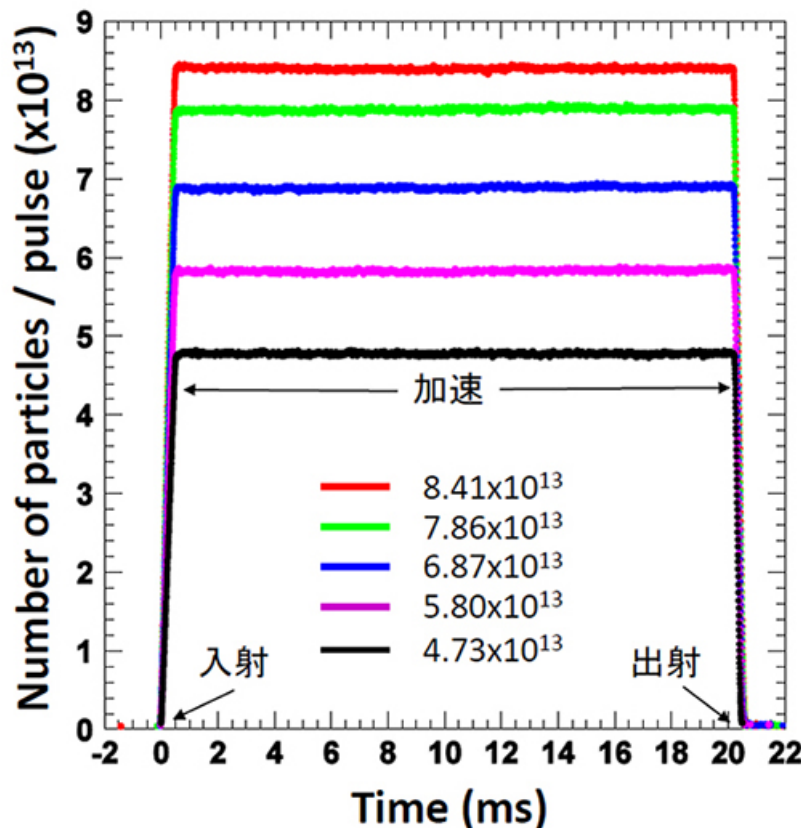
$$P(\nu_\mu \rightarrow \nu_e) = \sin^2 2\theta \cdot \sin^2\left(\frac{1.27 \cdot \Delta m^2 \cdot L}{E_\nu}\right)$$

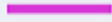
- Energy is smeared by 15%/sqrt(E) (detector E resolution)



Prospects; RCS/MLF beam

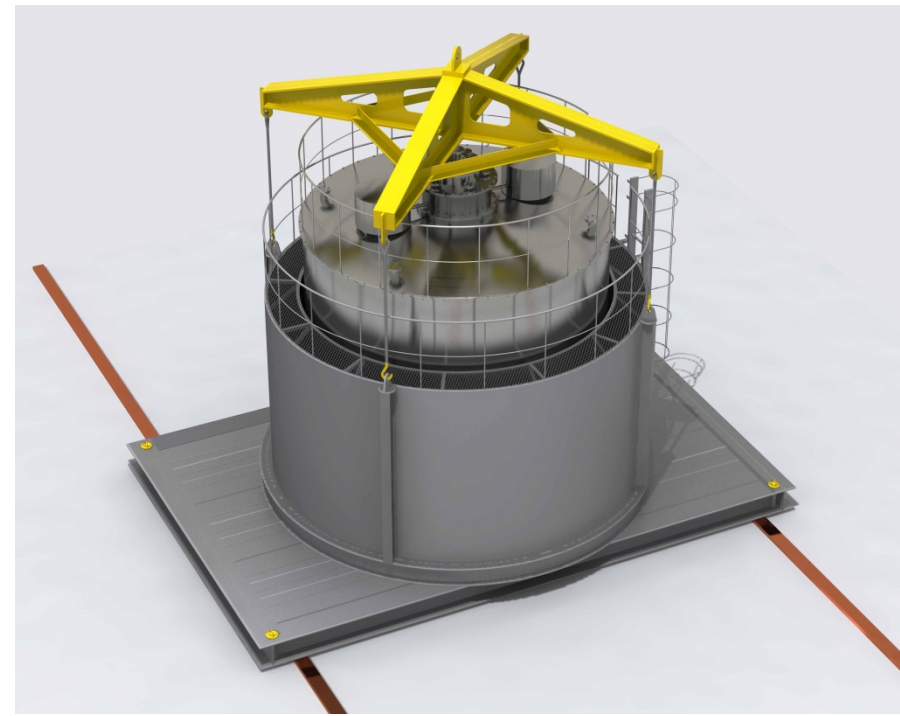
- Current beam power is 300kW.
- 1MW trial during the very short period was succeeded. (下図) <http://j-parc.jp/ja/topics/2015/Pulse150206.html>
- The beam power will be slowly increased. Next immediate step is 300 → 400kW.



	1パルス当たりの粒子数	25Hz運転時のパワー換算値
	8.41×10^{13}	1010 kW
	7.86×10^{13}	944 kW
	6.87×10^{13}	825 kW
	5.80×10^{13}	696 kW
	4.73×10^{13}	568 kW

Detector

- 50 tons Gd-loaded liquid scintillator.
- Well established technique.
- Strength of tank / endurance for earthquakes were calculated.
- Some of Double Chooz / Daya-Bay collaborators joined E56.
- Will be located at MLF 3rd floor : the maintenance area
→ need to avoid interference

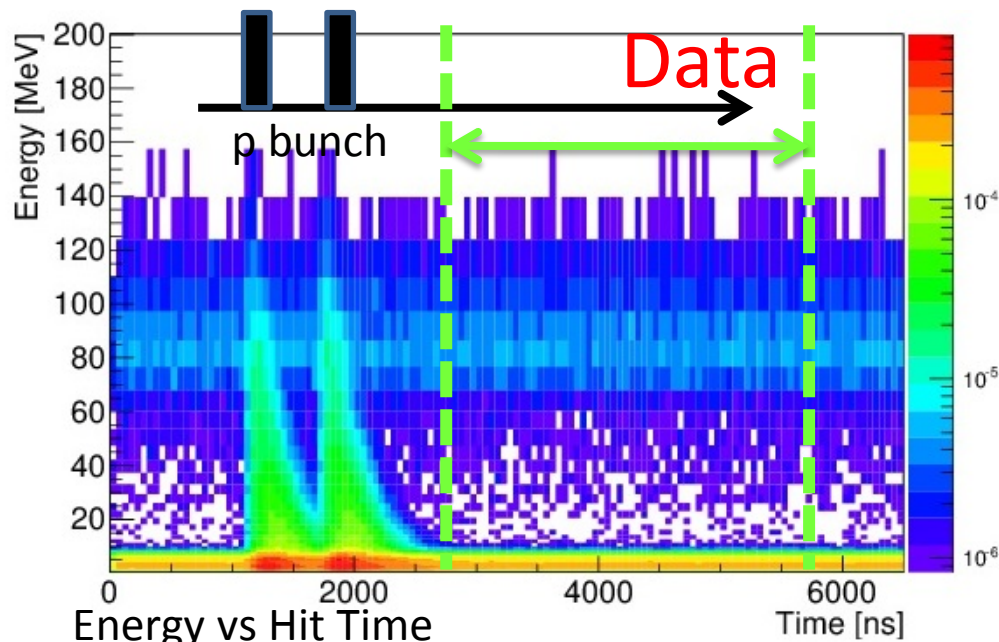
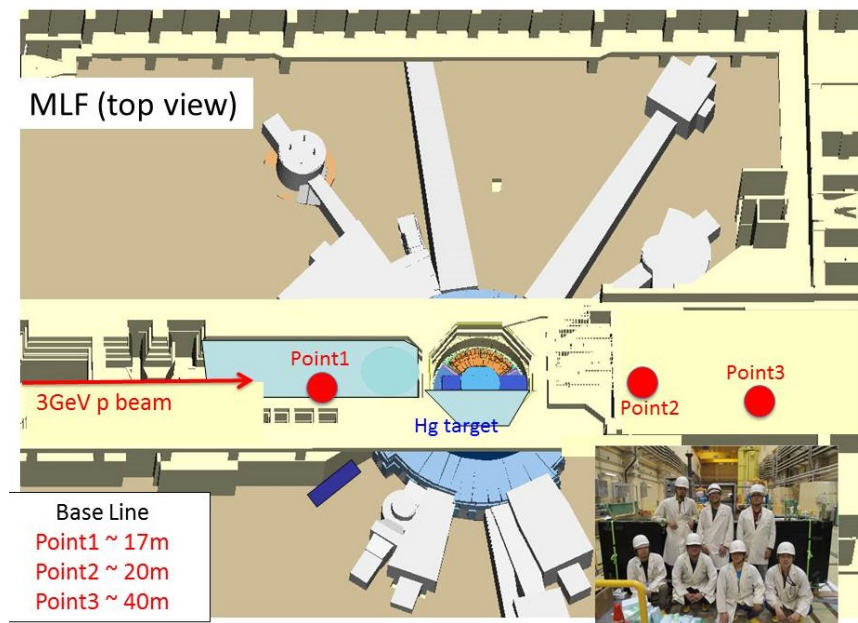


	Time from beam	Energy
Prompt signal	$1 < T_p < 10 \mu s$	$20 < E < 60 \text{ MeV}$
Delayed signal	$T_p < T_d < 100 \mu s$	$6 < E < 12 \text{ MeV}$

J-PARC E56; Achievements so far

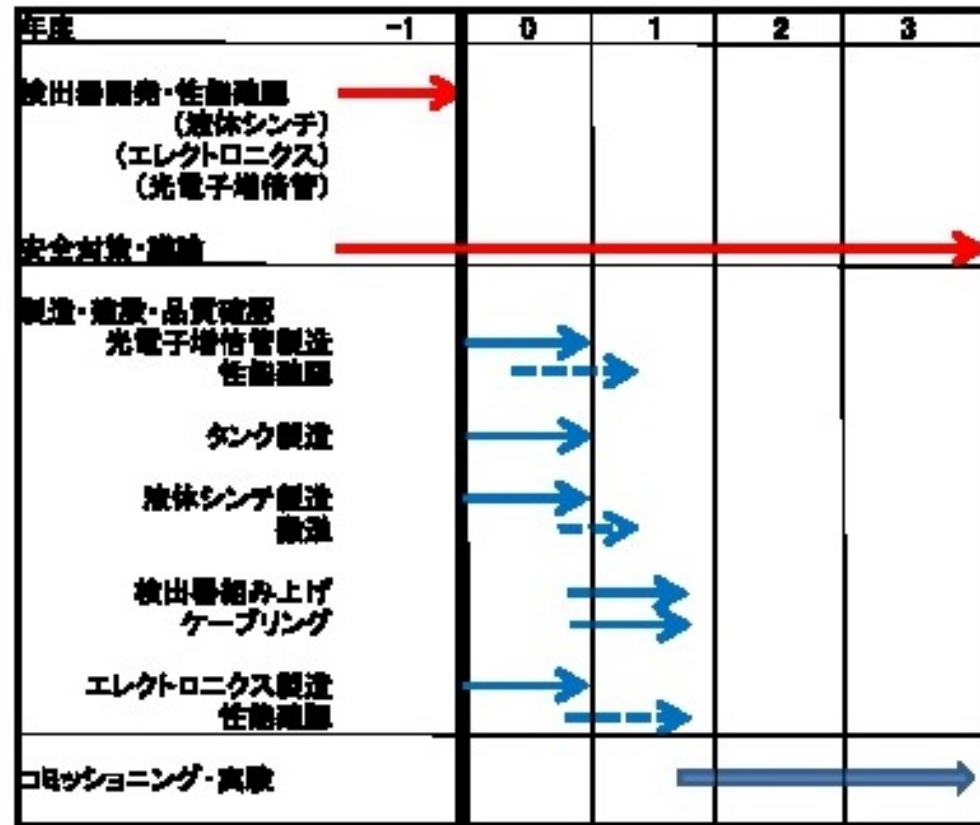
- 2013 Feb-May; A background measurement on the 1st floor
- 2013 Sep; A proposal was submitted to the J-PARC PAC
 - The PAC recommended to measure the background at the detector candidate site (3rd floor of MLF)
- 2014 Apr-Jul; We measured the BKG rate at 3rd floor. (MLF 2014BU1301 test experiment).
-> Small beam/cosmic BKGs to perform E56 experiment.
[arXiv:1502.02255 \[physics.ins-det\]](https://arxiv.org/abs/1502.02255)
- 2014-Dec; A status report were submitted to J-PARC PAC. We **obtained the stage-1** approval from the PAC.
http://j-parc.jp/researcher/Hadron/en/pac_1412/J-PARC-PAC-19-final_141205.pdf

2014BU1301 experiment (@MLF 3rd floor)



J-PARC E56 Timescale

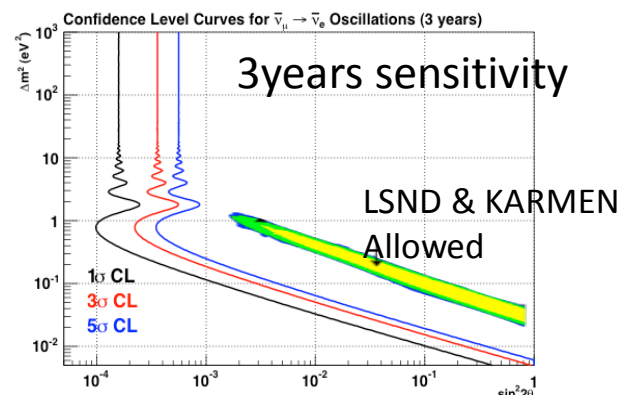
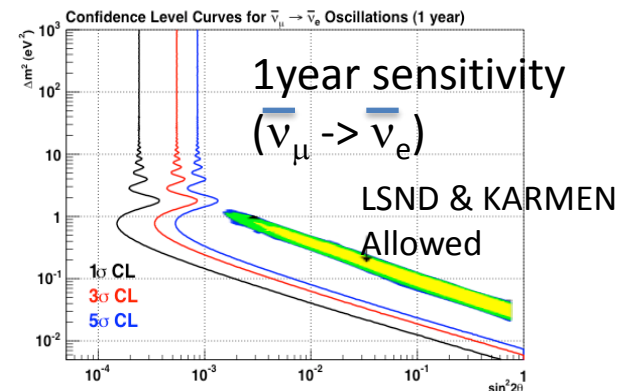
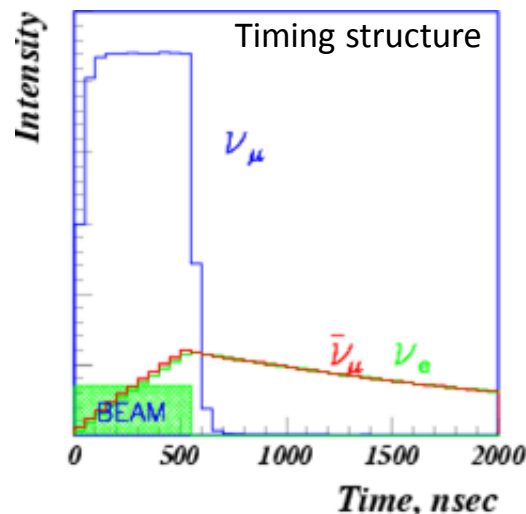
- 2015;
 - 液体シンチのR&D・性能確認(含ビームテスト)
 - エレクトロニクスR&D
 - タンク等のデザインの最終化.
 - 追加のマイナーな背景事象測定。
 - MLFと安全面・検出器のメンテナンスに関する議論
- 検出器建設から約1.5年で実験をスタート。



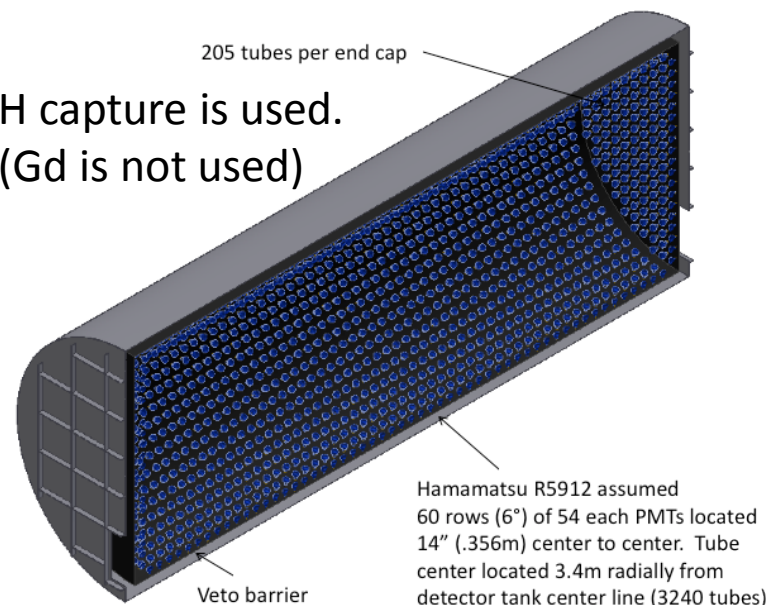
- 新しいビームライン／検出器建屋(穴)が必要ないので、非常にコスト効率が高い。
検出器1基 ～2億円
- また、上記に加え、検出器も良く理解されたものを使うことで、世界に先駆けて結果を出せる可能性が高い。

OscSNS (arXiv:1307.7097)

- Use Spallation Neutron Source (SNS) at ORNL
- ~1GeV protons on Hg target (1.4MW)
- Free source of neutrinos
- Well understood flux of neutrinos

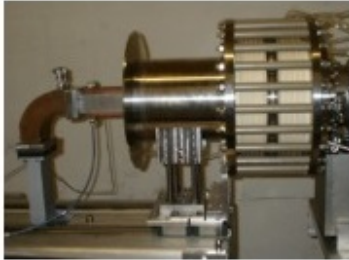


H capture is used.
(Gd is not used)



IsoDAR / DAE δ ALUS

Ion source

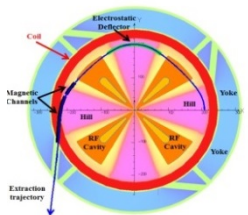


Superconducting
ring cyclotron
(DAE δ ALUS)

J.Spitz's talk @ WINP2015

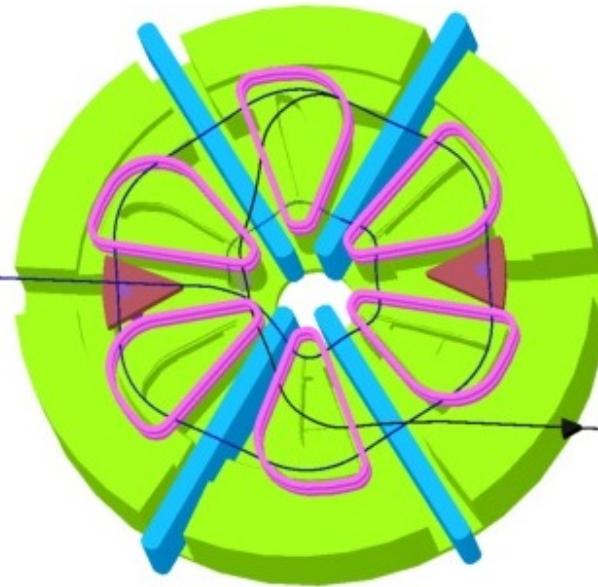
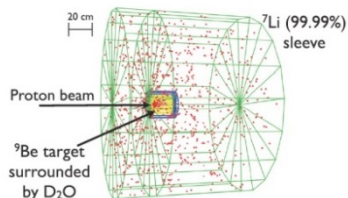


Injector cyclotron
(IsoDAR)

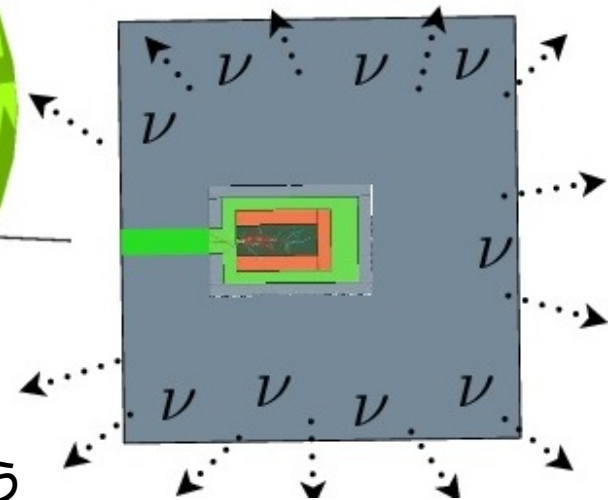


Accelerating H_2^+ ,
60MeV/n @ 5ma

Neutrons capture
on ^7Li to make ^8Li .



Target/dump



- 強力なCyclotronビームを使う
- DAE δ ALUSの最終的なビームは700MeVで5MWくらい。
- IBDが起こる巨大検出器に向かって打ち込む。(加速器を距離を変えて、3つくらい作る)

mode	experiments	ν source	Detector	Pros	comments
$(\bar{\nu}_e) \rightarrow (\bar{\nu}_x)$	Nucifer, STEREO, etc	Reactor	Gd-LS	Ton-scale detector, variation of L	BKG from reactor/CR
	PROSPECT	Reactor	Li-LS	PSD for delayed n signal (double PSD)	R&D
	SOX, etc	Radio active source (Cr,Ce-Pr)	Gd-LS	Fastest results(?) / ν and $\bar{\nu}$ sources	MCi level source / transportation
$(\bar{\nu}_\mu) \rightarrow (\bar{\nu}_e)$	J-PARC E56/Osc SNS	μ DAR	Gd-LS / LS	Well-establish det. / small BKG / Ev rec	MLF maintenance / new hole (SNS)
	IsoDAR/DAE δ ALUS	μ DAR	LS or Gd-H ₂ O	Very strong beam	R&D
	Triple LAr	π DIF	LAr-TPC	Beam / MicroBooNE / ICARUS are available. e/π^0 sep	CR on ground Erec(?) LAr1-ND plan
	nuPRISM	π DIF	WC	Beam is available / different Ev	Ev rec
$(\bar{\nu}_\mu) \rightarrow (\bar{\nu}_\mu)$	SK/MINOS+/T2 K-ND/ICECUBE	π DIF	WC/Fe-Scinti	On-going	---

Summary

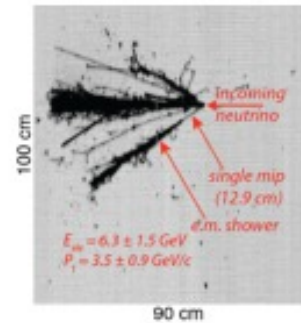
- LSND、MiniBooNE、Ga較正、原子炉で見られる事象超過と事象欠損が本当にあるかどうかは、系統誤差が小さい実験で早く決着をつけるべき問題。
 - 現在、出現・消失モードに関する多数の実験がon-going または予定されている。
 - 線源・原子炉と液体シンチを用いた ν_e 消失
 - ダンプ ν と液体シンチを用いた反 ν_e 出現
 - 電磁ホーン ν とLAr/鉄シンチ/WCを用いた反 ν_e 出現、 ν_μ 消失。
 - 後、5-10年の間に全てのモードでだいたいの結果が出そろおう。
 - 非常にエキサイティングな時期
 - 小回りや自分のアイディアが活かされる実験も多い。
- 是非、みなさんの知恵を実験に！！！！

backup

Conventional beam + LAr TPC

- LAr TPC can see all of charged particle tracks.
- Clear electron, π^0 and single γ PID can be performed.

FIGURES



ICARUS arXiv:1307.4699 [hep-ex].

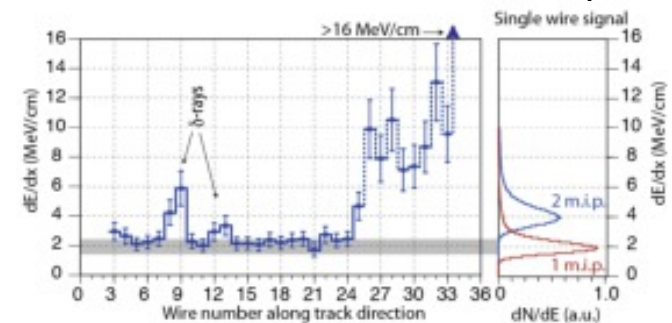
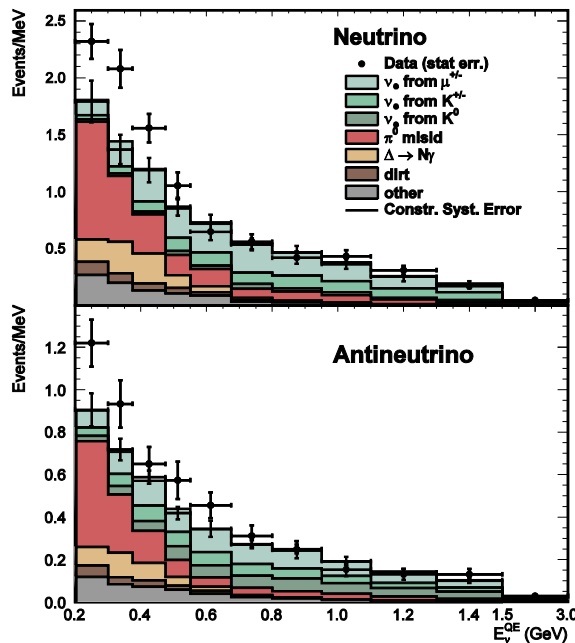


Figure 1. Experimental pictures of the first of the two events with a clear electron signature found in the additional sample of 904 neutrino interactions. The evolution of the actual dE/dx from a single track to an e.m. shower for the electron shower is shown along the individual wires. The event has a total energy of ~ 27 GeV and an electron of 6.3 ± 1.5 GeV with a transverse momentum of 3.5 ± 0.9 GeV/c.

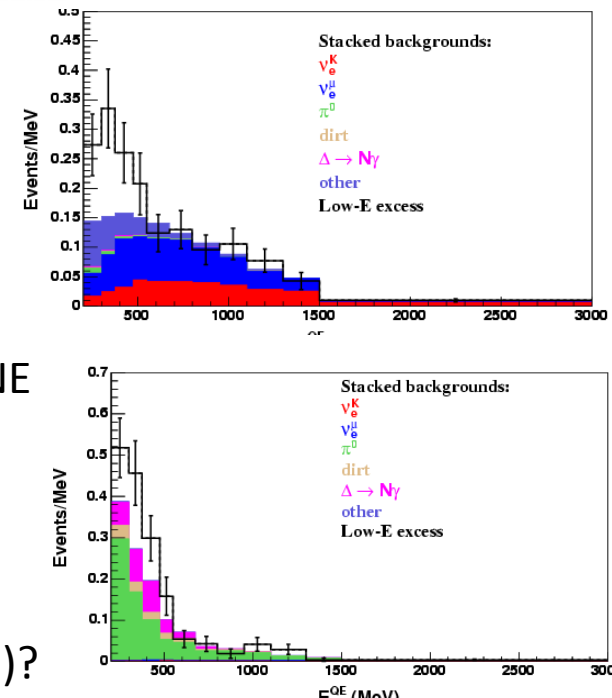


MiniBooNE
Excess
PRL 110.161801,2013.

Electron excess
(oscillation)?

MicroBooNE

or γ excess
(uncertainty of
 ν cross section)?



CMB and Neff

- They first calculate the whole radiation energy density from the thermal fluctuation . (ρ_{rel})
- Subtracting the energy density of the photons from the whole energy density. ($\rho_{\text{rel}} - \rho_{\text{photon}}$)
- Then the Neff is calculated by dividing the energy density by the neutrino energy density of 1 generation

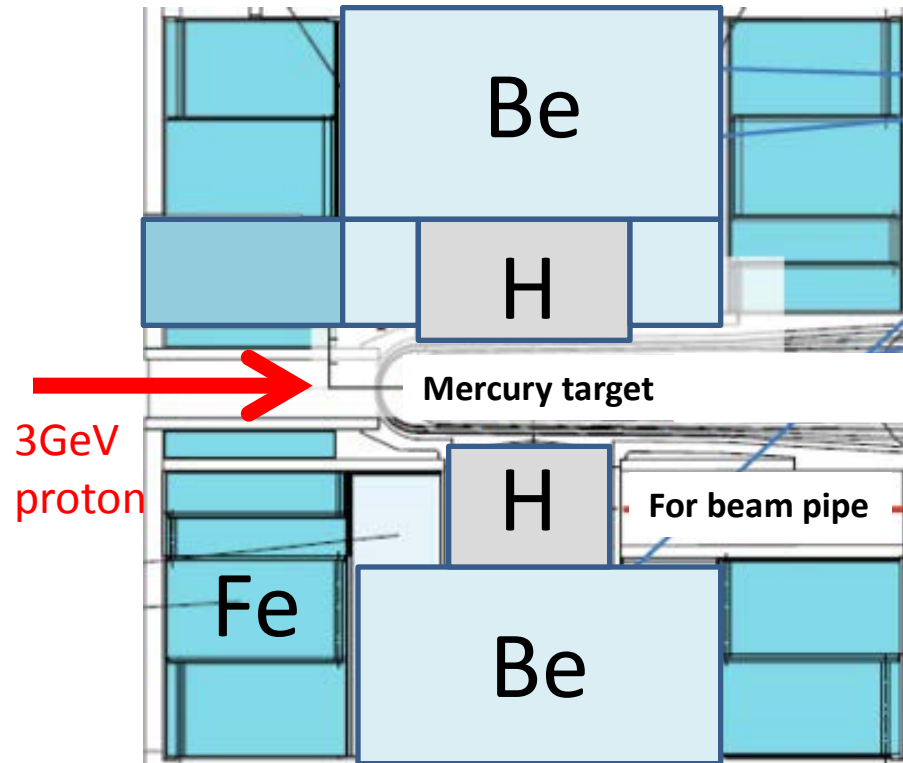
$$N_{\text{eff}} = (\rho_{\text{rel}} - \rho_{\text{photon}}) / \rho_{\text{nu,thermal}}$$

- Here we assume neutrinos are in thermal equilibrium to calculate the energy density of the neutrinos for one generation.
- If neutrinos are in thermal equilibrium, the energy density of neutrinos are related to that of photon with the equation;

$$(\rho_{\text{nu}} = 0.22 * N_{\text{eff}} * \rho_{\text{photon}})$$

- Sterile neutrinos may or may not be in thermal equilibrium via neutrino oscillations (depending on the model), so Neff could not prove the number of generations of sterile neutrinos. (E.g.; see <http://arxiv.org/pdf/1303.5368v2.pdf> for further details)

MLF mercury target and Intrinsic $\bar{\nu}_e$ BKG estimation



	Target	π^- absorb	μ^- capture	suppression	$\times \pi^-/\pi^+$
LSND	H2O	96%	88%	5×10^{-3}	$\times 0.13$
J-PARC E56	Hg(+Fe+Be)	99%	$\sim 80\%$	1.7×10^{-3}	$\times 1.$

We will assume $\sim 1.7 \times 10^{-3}$ Intrinsic background hereafter.