

# The Resolved Kennicutt-Schmidt Law in Nearby Galaxies

**Rieko Momose<sup>1,2,3</sup>**

Sachiko K. Okumura<sup>2,4</sup>, Jin Koda<sup>5</sup>, Robert C. Kennicutt, Jr<sup>6</sup>,  
Jennifer Donovan Meyer<sup>5</sup>, Daniela Calzetti<sup>7</sup>, Guilin Liu<sup>7,8</sup>, Fumi Egusa<sup>9</sup>,  
Nick Scoville<sup>10</sup>, Tsuyoshi Sawada<sup>2</sup>, Nario Kuno<sup>11</sup>

1. ICRR, University of Tokyo, 2. University of Tokyo, 3. Chile/NAOJ,  
4. Japan Woman's University, 5. Stony Brook University, 6. University of Cambridge,  
7. University of Massachusetts, 8. Johns Hopkins University, 9. JAXA, 10. Caltech,  
11. NRO/NAOJ

# CONTENTS

## 1. INTRODUCTION

1. The K-S Law
2. Physical background
3. Motivation

## 2. OBSERVATIONS & METHODS

1. Methodology
2. The K-S Law

## 3. RESULTS & DISCUSSIONS

1. From all galaxies
2. Among structures

## 4. SUMMARY

# ABSTRACT



not yet  
published



not yet  
published

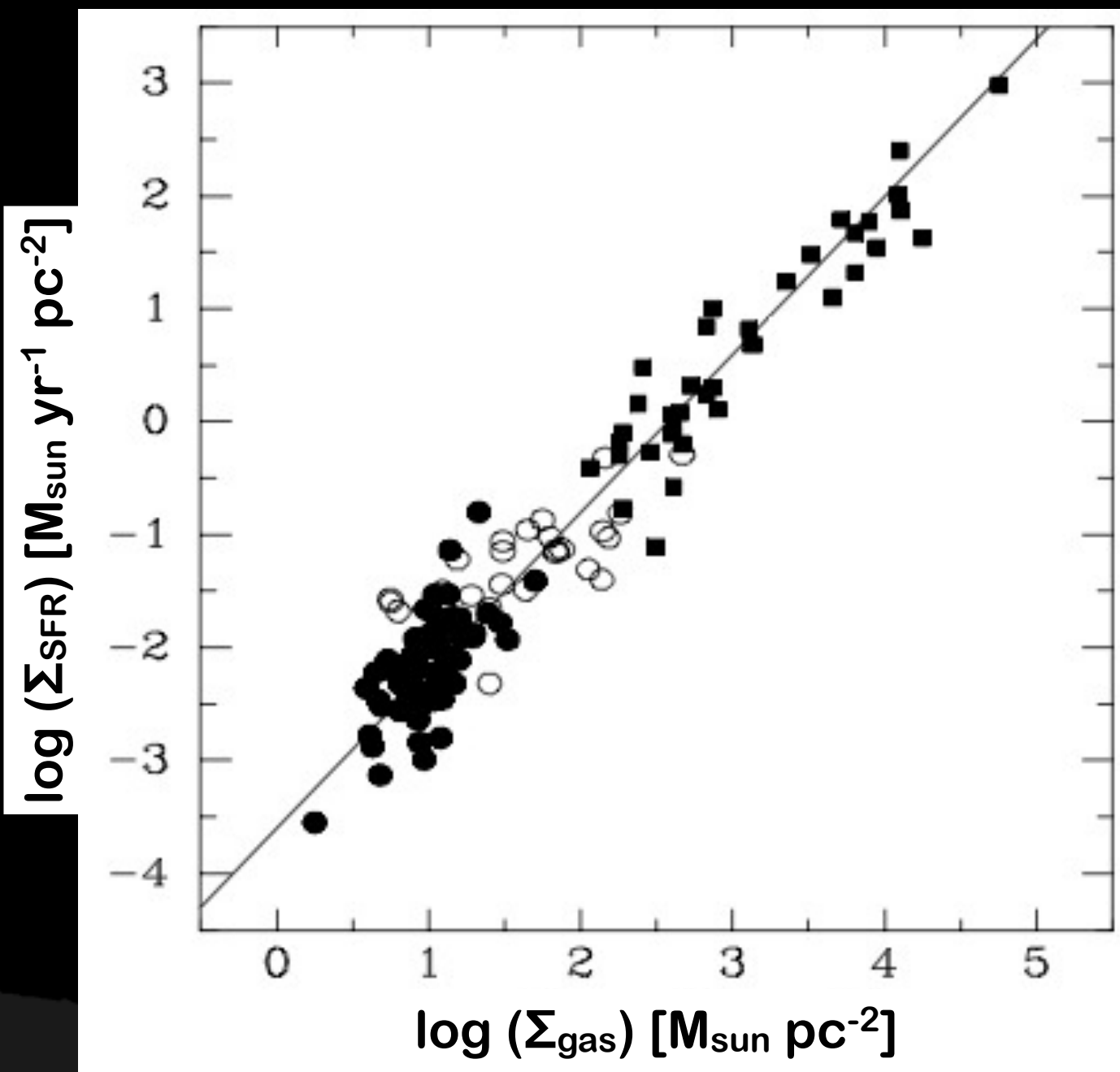
**We obtain super-linear  
slope of the K-S law on  
500 pc from 10 galaxies**

**Star formation mechanism  
indicated by the K-S law is little  
different among structures**

# 1. INTRODUCTION



# The Kennicutt-Schmidt Law



- Power law correlation between SFR and gas density (Schmidt 59)

$$\Sigma_{\text{SFR}} = A \Sigma_g^N,$$

- Observationally, the index has been estimated **among galaxies** (Kennicutt 98; following)
  - $N = 1.2 - 1.8$

※ Star Formation Rate (SFR)  
The amount of gas mass converted to stars per unit time [ $M_{\odot}/\text{yr}$ ]

# Physical Predictions From The K-S Law

- Star formation would be regulated by some physically motivated time scale

$$\Sigma_{\text{SFR}} = \epsilon_{\text{tphy}} \frac{\Sigma_{\text{gas}}}{t_{\text{phy}}},$$

$$\Sigma_{\text{SFR}} \propto \epsilon_{\text{tphy}} \Sigma_{\text{gas}}^{1-m} = \epsilon_{\text{tphy}} \Sigma_{\text{gas}}^N,$$

**N = 1**

- Constant star formation at fixed time scale

$$\text{SFR} = \frac{\rho_{\text{gas}}}{t_{\text{SF}}}.$$

$$\Sigma_{\text{SFR}} = \epsilon \Sigma_{\text{gas}},$$

**N > 1**

- Star formation at a free-fall time scale ( $t_{\text{ff}}$ )

$$t_{\text{ff}} = \sqrt{\frac{3\pi}{32G\rho_{\text{gas}}}}.$$

$$\Sigma_{\text{SFR}} = \epsilon_{\text{ff}} \frac{\Sigma_{\text{gas}}}{t_{\text{ff}}} \propto \Sigma_{\text{gas}}^{1.5},$$

**N = 1.5**

- Star formation regulated by collision time scale

$$t_{\text{cc}} = \frac{\lambda}{v} = \frac{hm_{\text{GMC}}}{\sqrt{2} \Sigma_{\text{gas}} a^2 \pi v} \quad \lambda_{\text{mfp}} = \frac{1}{\sqrt{2}ND},$$

$$\Sigma_{\text{SFR}} = \epsilon_{\text{cc}} \frac{\Sigma_{\text{gas}}}{t_{\text{cc}}} \propto \Sigma_{\text{gas}}^2.$$

**N = 2**

# Physical Predictions From The K-S Law

In order to understand the mechanism of star formation, estimating the index of the K-S law is important

**N = 1**

- Constant star formation at fixed time scale

$$\text{SFR} = \frac{\rho_{\text{gas}}}{t_{\text{SF}}}$$

$$\Sigma_{\text{SFR}} = \epsilon \Sigma_{\text{gas}},$$

**N > 1**

- Star formation at a free-fall time scale ( $t_{\text{ff}}$ )

$$t_{\text{ff}} = \sqrt{\frac{3\pi}{32G\rho_{\text{gas}}}}$$

$$\Sigma_{\text{SFR}} = \epsilon_{\text{ff}} \frac{\Sigma_{\text{gas}}}{t_{\text{ff}}} \propto \Sigma_{\text{gas}}^{1.5},$$

**N = 1.5**

- Star formation regulated by collision time scale

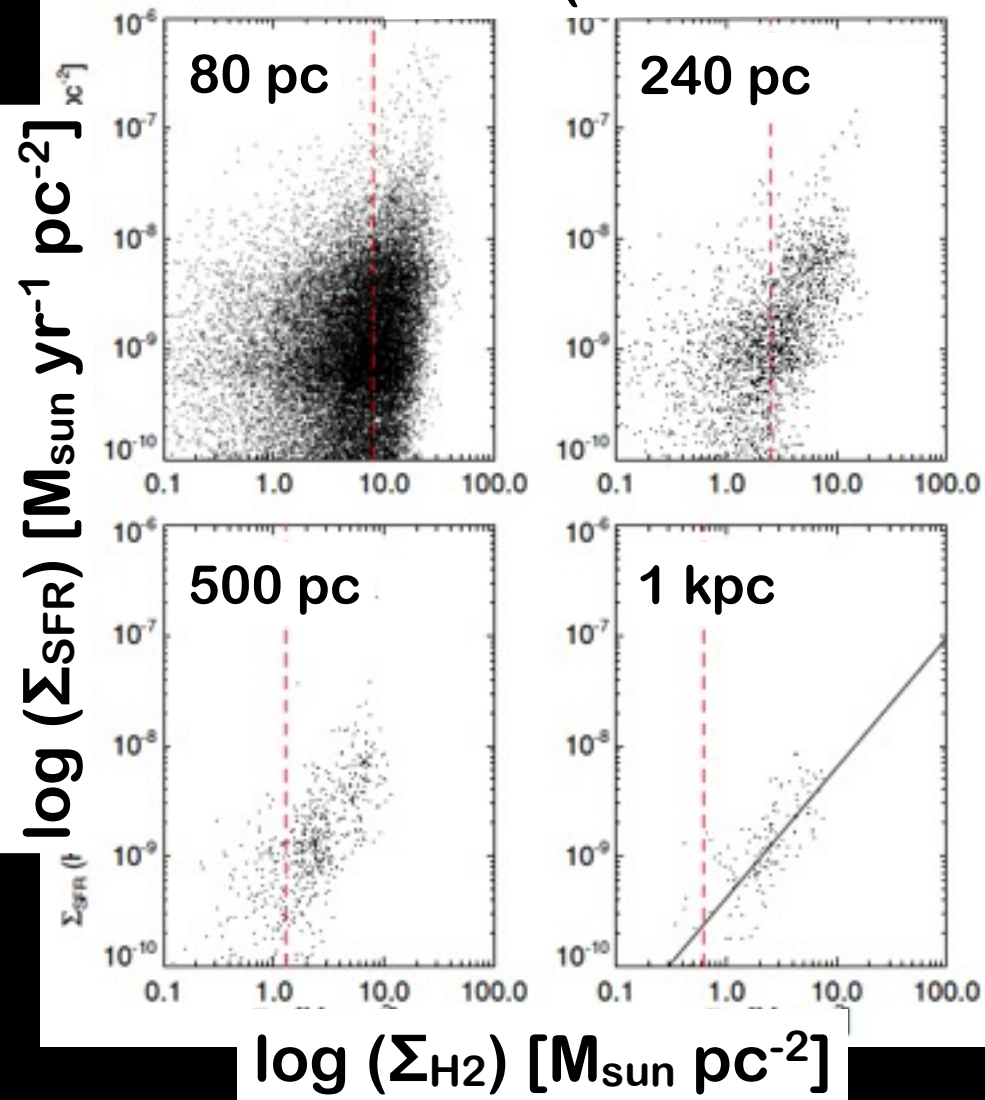
$$t_{\text{cc}} = \frac{\lambda}{v} = \frac{hm_{\text{GMC}}}{\sqrt{2} \Sigma_{\text{gas}} a^2 \pi v} \quad \lambda_{\text{mfp}} = \frac{1}{\sqrt{2}ND},$$

$$\Sigma_{\text{SFR}} = \epsilon_{\text{cc}} \frac{\Sigma_{\text{gas}}}{t_{\text{cc}}} \propto \Sigma_{\text{gas}}^2.$$

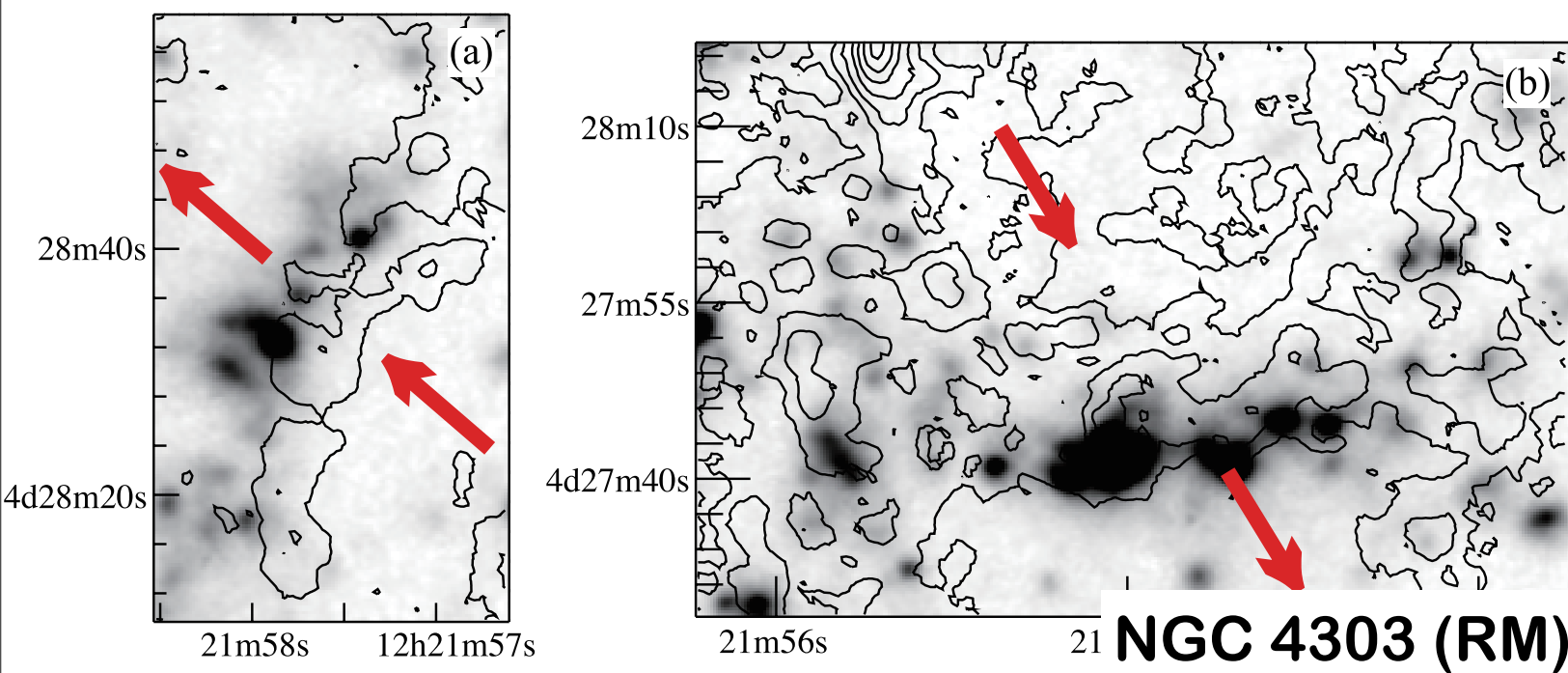
**N = 2**



# Recent Study of The K-S Law 1



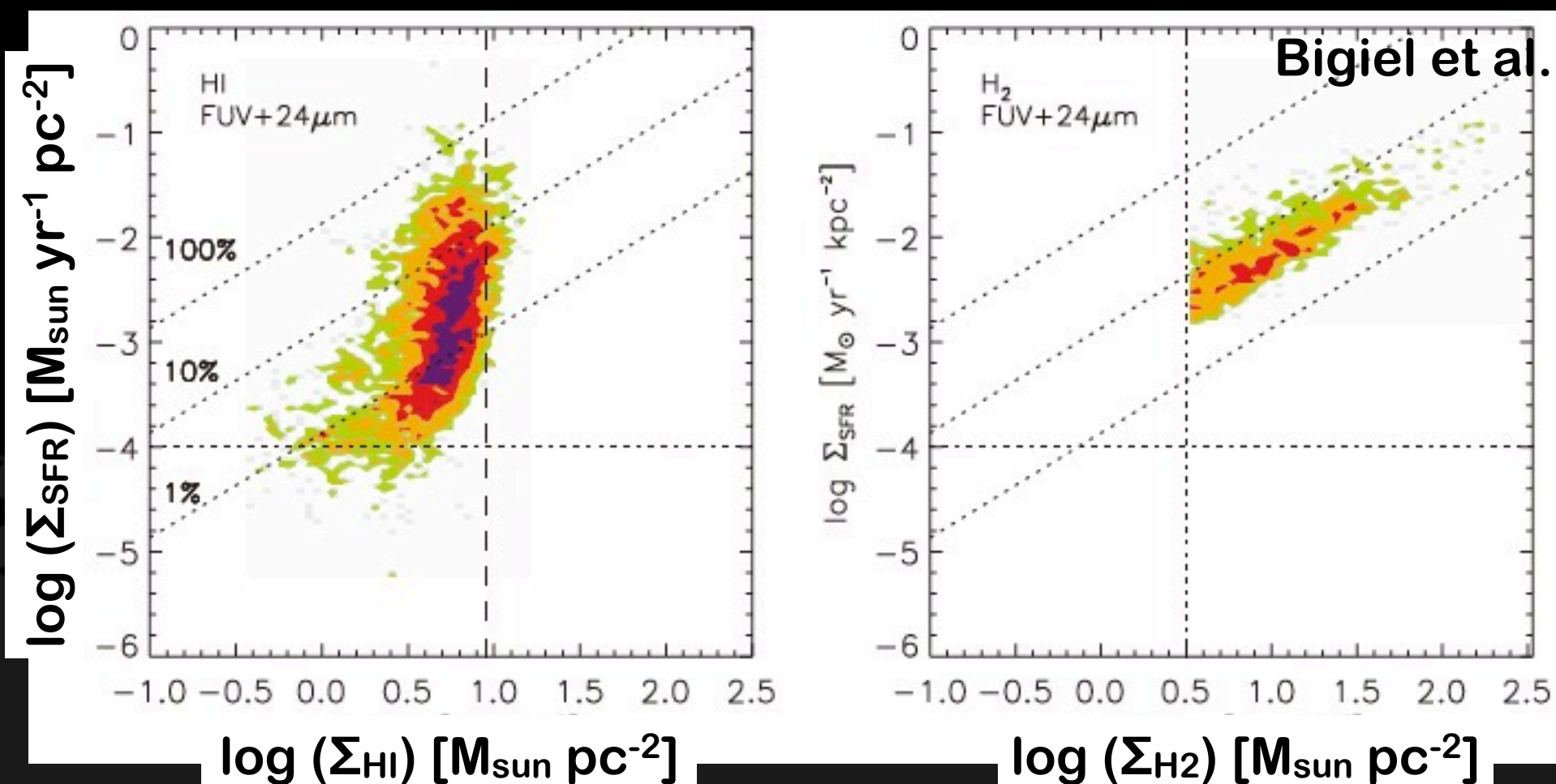
- The K-S law studies are shifting those of obtained **within a galaxy**
  - NGC 5194 (M 51):  $\sim 500$  pc (e.g. Kennicutt et al. 07; Blanc et al. 09)
  - M 33: 80-200 pc: poor--break down (Verley et al. 10; Onodera et al. 10)



The mechanism of triggering star formation operates at a scale smaller than **500 pc**.

# Recent Study of The K-S Law 2

- Statistical study of K-S law on sub-kpc scale (Bigiel et al. 08)
  - Molecular gas is dominant above  $10 M_{\odot} \text{ pc}^{-2}$
  - The correlation between  $\Sigma_{\text{SFR}}$  and  $\Sigma_{\text{H}_2}$  is  $N \sim 1$



# Remaining Problem And Motivations

## ✧ Resolved K-S law

- The K-S law traced by CO(J=1-0) breaks down on GMC scale (< 100 pc), but holds on 500 pc
- The index of  $N \sim 1$  becomes standard value of the K-S law
  - No study of the K-S law on sub-kpc scale using CO(J=1-0) line among several galaxies
  - Little study to estimate accurate SFR

## ✧ Motivation

- Studying the K-S law on 500 pc scale using CO(J=1-0) line as a tracer of the amount of molecular gas to 10 nearby galaxies
- We resolve galactic structures and study the K-S law to each structure

## 2. OBSERVATIONS & METHODS

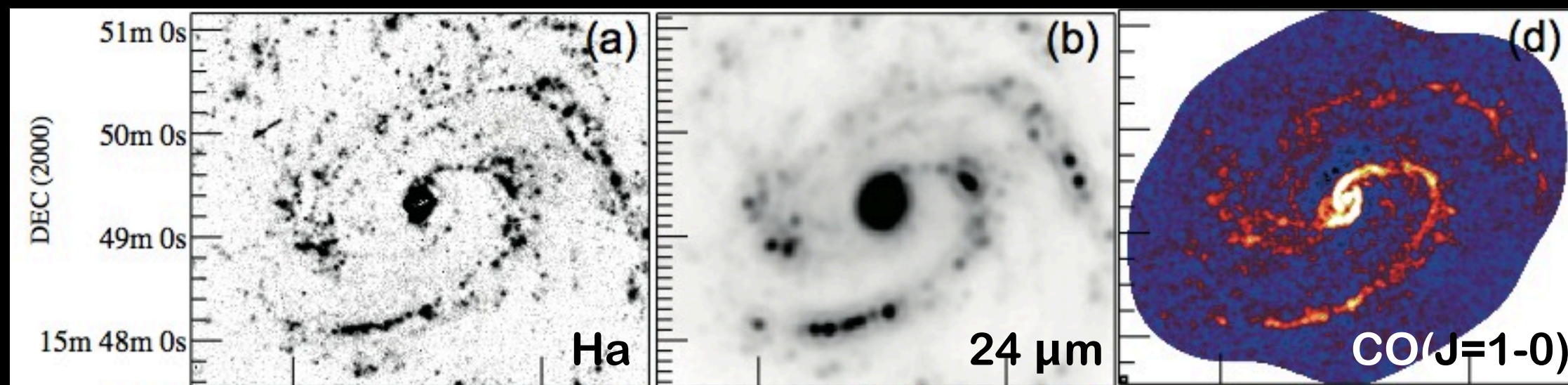
# Target of The Study

- Nearby disk galaxies which have relatively face-on viewing

NAME	Distance (Mpc)	$V_{\text{LSR}}$ (km/s)	Hubble Type	i (deg)	P.A. (deg)
NGC 3521	10.1	801	SABbc	64	163
NGC 3627	9.38	727	SABb	65	173
NGC 4254	16.5	2407	SAc	-30	24
NGC 4303	16.1	1556	SABbc	28	312
NGC 4321	14.32	1571	SABbc	32	30
NGC 4736	5.2	308	SAab	36	105
NGC 4826	7.48	408	SAab	-59	115
NGC 5055	7.8	484	SAbc	56	105
NGC 5194	7.62	463	SABbc *	20	163
NGC 6946	6.8	50	SABcd	32	53



# Methodology of The Study



- Re-gridding  $\Sigma_{\text{SFR}}$  and  $\Sigma_{\text{H}_2}$  data to 500 pc scale
  - $\Sigma_{\text{H}_2}$  : CO(J=1-0) data
  - $\Sigma_{\text{SFR}}$  : H $\alpha$  + 24 $\mu$ m data
- We compare pixel-by-pixel of both  $\Sigma_{\text{SFR}}$  and  $\Sigma_{\text{H}_2}$  data, and plot each data point in the K-S law plot
- Fitting the linear regression to the correlation

# Observations

- We observed CO(J=1-0) line by NRO45 and CARMA as a part of CANON survey
  - ✧ NGC 4303 by Momose et al. 10
  - ✧ NGC 5194 by Koda et al. 09

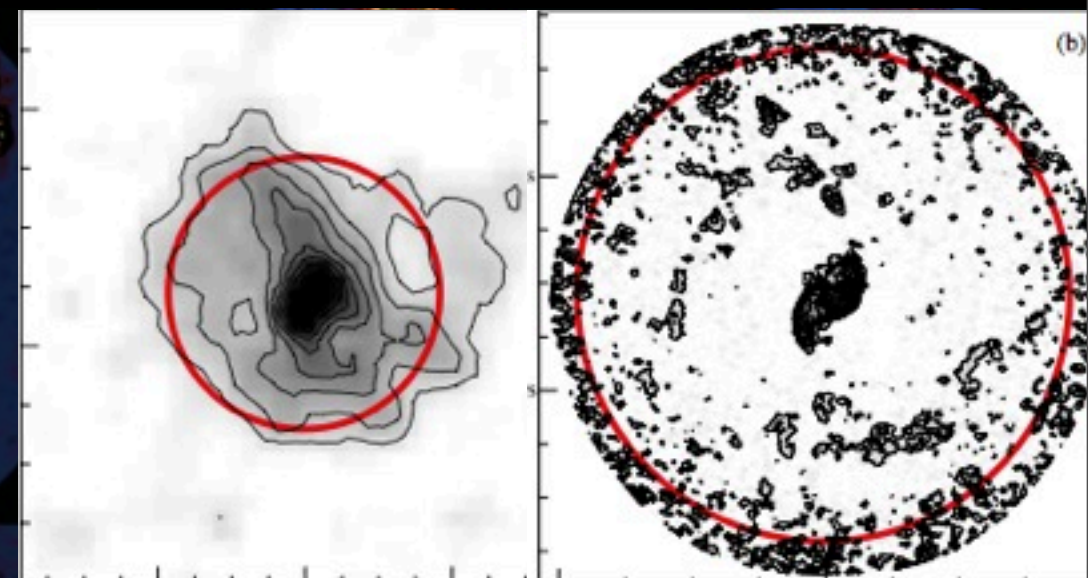


✧ CANON = CARMA and NOBEYAMA Nearby-galaxies CO survey



# Data Combining

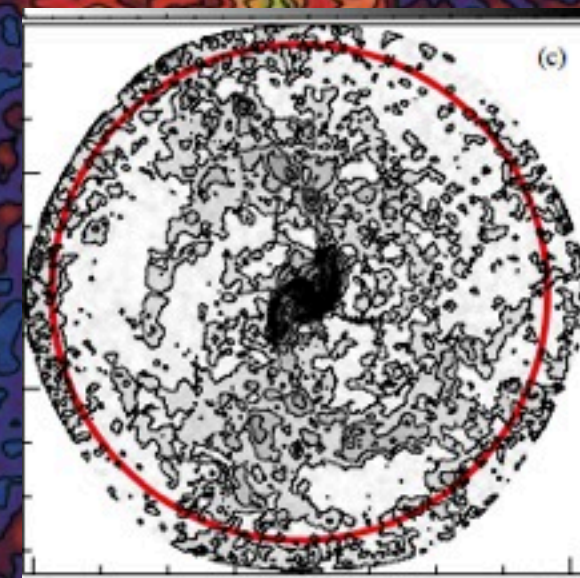
- We require the high spatial resolution data without missing flux
- Combining the data observed by single-dish and interferometer can recover the total flux with high spatial resolution



- Molecular gas surface density is estimated using the combined data

- $X_{\text{CO}} = 2.0 \times 10^{20} [\text{cm}^{-2} (\text{K km s}^{-1})^{-1}]$

$$\Sigma_{\text{H}_2} [M_{\odot} \text{ pc}^{-2}] = 4.8 \times \cos(i) \left( \frac{I_{\text{CO}}}{\text{K km s}^{-1}} \right) \left( \frac{X_{\text{CO}}}{\text{cm}^{-2} (\text{K km s}^{-1})^{-1}} \right)$$





# Data Combining

- We require the high spatial resolution data without missing flux
- Combining the data observed by single-dish and interferometer can recover the total flux with high spatial resolution
- Molecular gas surface density is estimated using the combined data
- $X_{\text{CO}} = 2.0 \times 10^{20} [\text{cm}^{-2} (\text{K km s}^{-1})^{-1}]$

$$\Sigma_{\text{H}_2} [M_{\odot} \text{ pc}^{-2}] = 4.8 \times \cos(i) \left( \frac{I_{\text{CO}}}{\text{K km s}^{-1}} \right) \left( \frac{X_{\text{CO}}}{\text{cm}^{-2} (\text{K km s}^{-1})^{-1}} \right)$$



# Estimating SFR

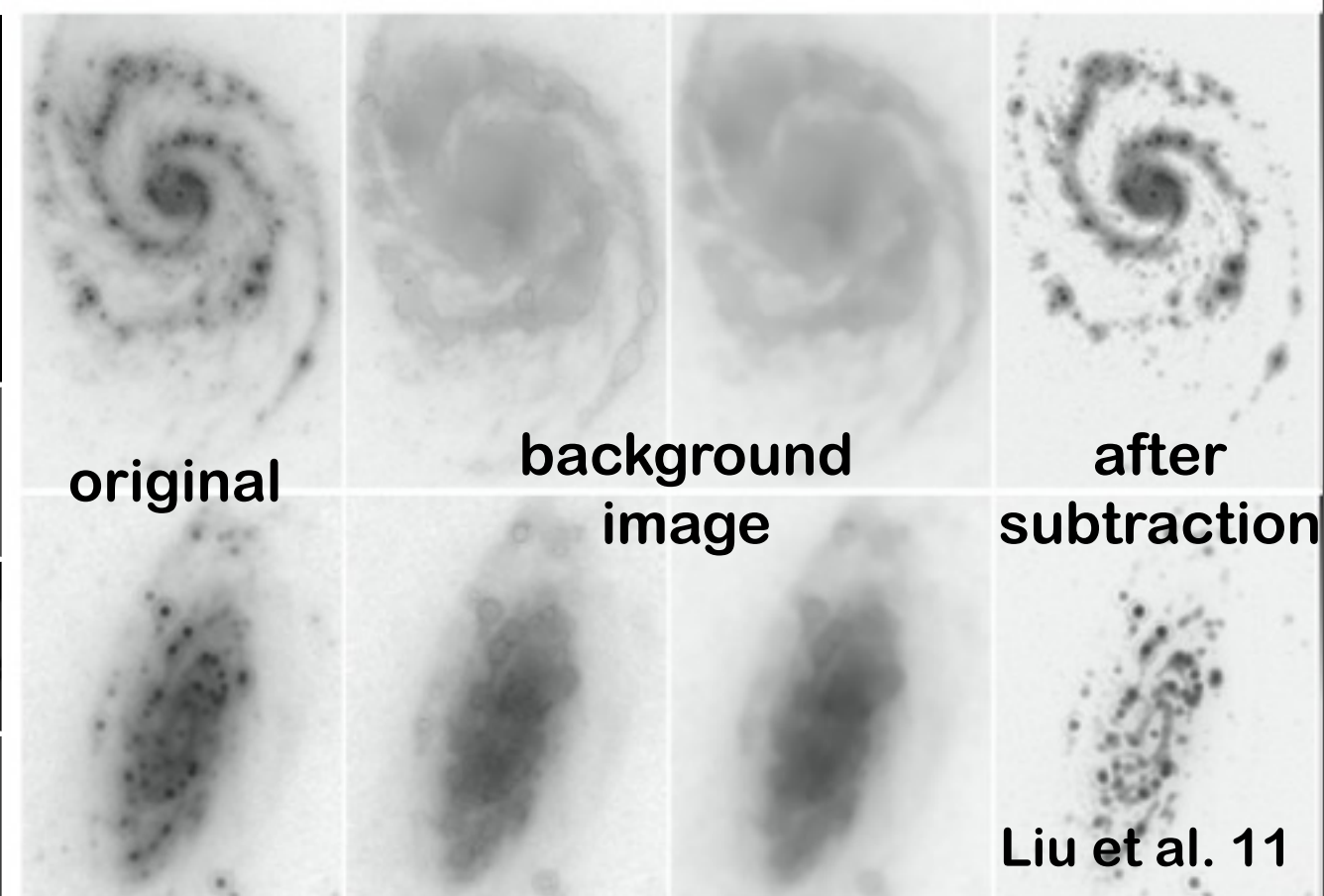
- We can estimate SFR counting emission from young stars and assuming IMF
- Ha and 24um images are used to estimate SFR (Kennicutt 98; Calzetti et al. 07)

$$\begin{aligned} \text{SFR } [M_{\odot} \text{yr}^{-1}] &= 7.9 \times 10^{-42} L(\text{H}\alpha)_{\text{corr}} [\text{erg s}^{-1}] \\ &= 7.9 \times 10^{-42} \{L(\text{H}\alpha)_{\text{obs}} + (0.031 \pm 0.006)L(24\mu\text{m})\} [\text{erg s}^{-1}] \end{aligned}$$

- We carried out diffuse ionized gas (DIG) subtractions

## Diffuse Ionized Gas

	local	not young stars	fraction
Ha	leak emission	ionized by hot old stars	30-50 % of total $L_{\text{Ha}}$
24um	---	dust heated by old stars	20-80 % of total $f_{24\text{um}}$



# 3. RESULTS & DISCUSSIONS

## 3.1 GLOBAL K-S LAW

## 3.2 STRUCTURAL K-S LAW



3.1

# GLOBAL K-S LAW FROM ALL GALAXIES



# The K-S Law From All 10 galaxies



not yet  
published

We obtain super-linear slope of the K-S law on 500 pc

- We obtain the index of
  - $N = 1.75 \pm 0.09$  (DIG subtracted) on 500 pc
  - $N = 1.26 \pm 0.03$  on 750 pc



# Discussion 1

## Physical Predictions

- Star formation would be regulated by some physically motivated time scale

$$\Sigma_{\text{SFR}} = \epsilon_{\text{tphy}} \frac{\Sigma_{\text{gas}}}{t_{\text{phy}}},$$

$$\Sigma_{\text{SFR}} \propto \epsilon_{\text{tphy}} \Sigma_{\text{gas}}^{1-m} = \epsilon_{\text{tphy}} \Sigma_{\text{gas}}^N,$$

**N = 1**

- Constant star formation at fixed time scale

$$\text{SFR} = \frac{\rho_{\text{gas}}}{t_{\text{SF}}}.$$

$$\Sigma_{\text{SFR}} = \epsilon \Sigma_{\text{gas}},$$

**N > 1**

~100 pc

- Star formation at a free-fall time scale ( $t_{\text{ff}}$ )

$$t_{\text{ff}} = \sqrt{\frac{3\pi}{32G\rho_{\text{gas}}}}.$$

$$\Sigma_{\text{SFR}} = \epsilon_{\text{ff}} \frac{\Sigma_{\text{gas}}}{t_{\text{ff}}} \propto \Sigma_{\text{gas}}^{1.5},$$

**N = 1.5**

- Star formation regulated by collision time scale

$$t_{\text{cc}} = \frac{\lambda}{v} = \frac{hm_{\text{GMC}}}{\sqrt{2} \Sigma_{\text{gas}} a^2 \pi v} \quad \lambda_{\text{mfp}} = \frac{1}{\sqrt{2}ND},$$

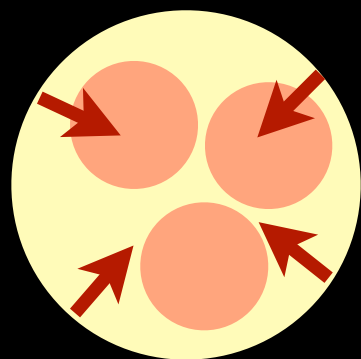
$$\Sigma_{\text{SFR}} = \epsilon_{\text{cc}} \frac{\Sigma_{\text{gas}}}{t_{\text{cc}}} \propto \Sigma_{\text{gas}}^2.$$

**N = 2**

# Discussion 1

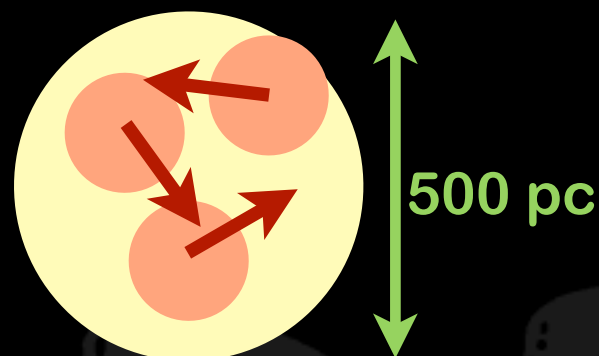
## Physical Predictions on 500 pc

$N = 1.5$



- Within a 500 pc, there may be some GMCs to form stars
- $N > 1$
- Existence of molecular gas is not enough to form clouds which will bear stars

$N = 2$

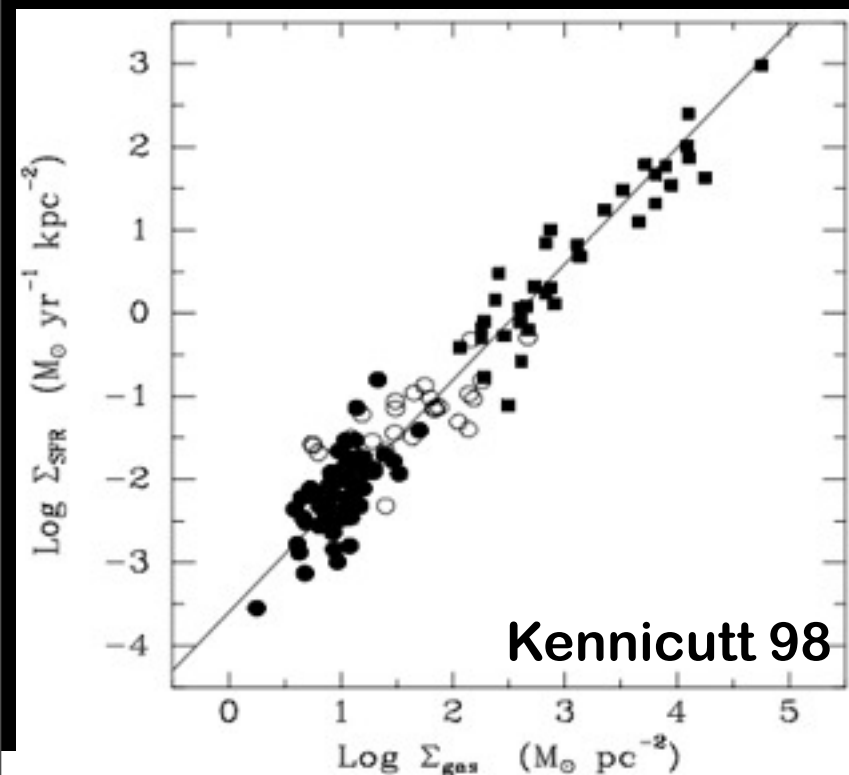


- Self-gravity of molecular gas
- Cloud-cloud collision
- Mixture of both

Any mechanism is required for triggering star formation on 500 pc scale

# Discussion 2

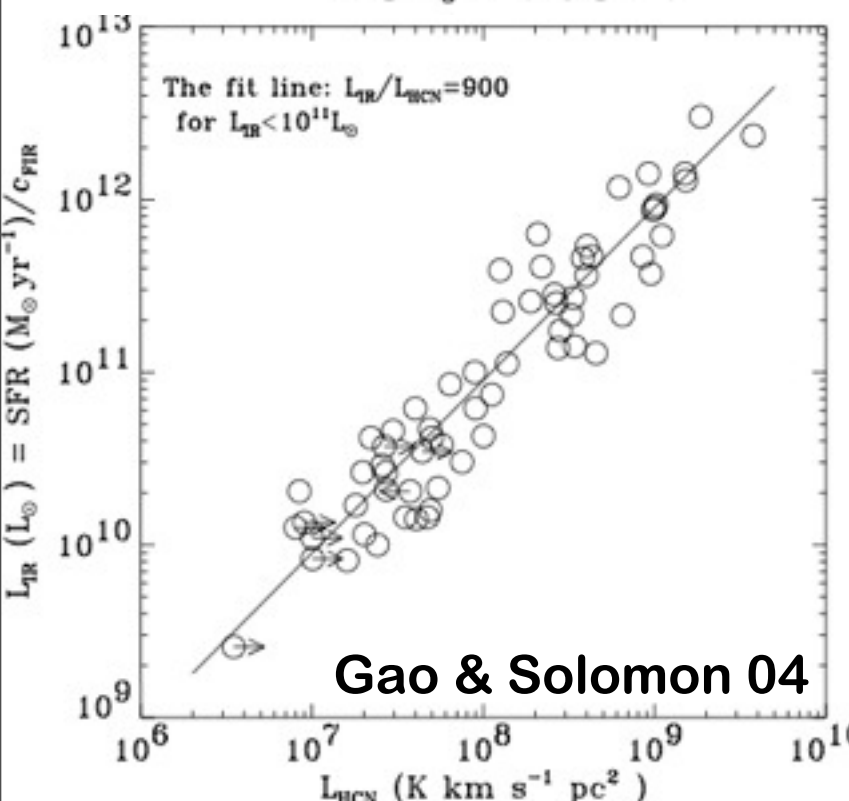
## Gas Tracers and The Index



The relation between the amount of molecular gas and SFR (e.g. Kennicutt 07; Liu et al. 11; **This study**)

-  **$N > 1$**

\* The relation between the molecular gas which can form stars and SFR



The relation between dense gas and SFR (e.g. Gao & Solomon 04; Komugi et al. 07; Iono et al. 09)

-  **$N \sim 1$**

\* The relation between star-forming cores and SFR

Line	$n_{\text{crit}}$ [cm <sup>-3</sup> ]
CO (J=1-0)	$3 \times 10^2$
HCN (J=1-0)	$\sim 10^{5-6}$
CO (J=3-2)	$\sim 10^4$
CO (J=2-1)	$\sim 10^3$

CO(J=1-0) is a best tracer to understand SF mechanism

# Summary 1

- We estimate the K-S law index by estimating accurate SFR and molecular gas surface density from CO(J=1-0)
- The K-S law on 500 pc shows the super-linear slope with  $N = 1.75$ .
- Some trigger mechanisms, such as self-gravity of molecular gas or cloud-cloud collision or mixture of them are necessary for forming stars.
- The correlation between SFR and amount of the molecular gas traced by CO(J=1-0) can indicate the mechanism of star formation rather than the correlation seen by dense gas tracers

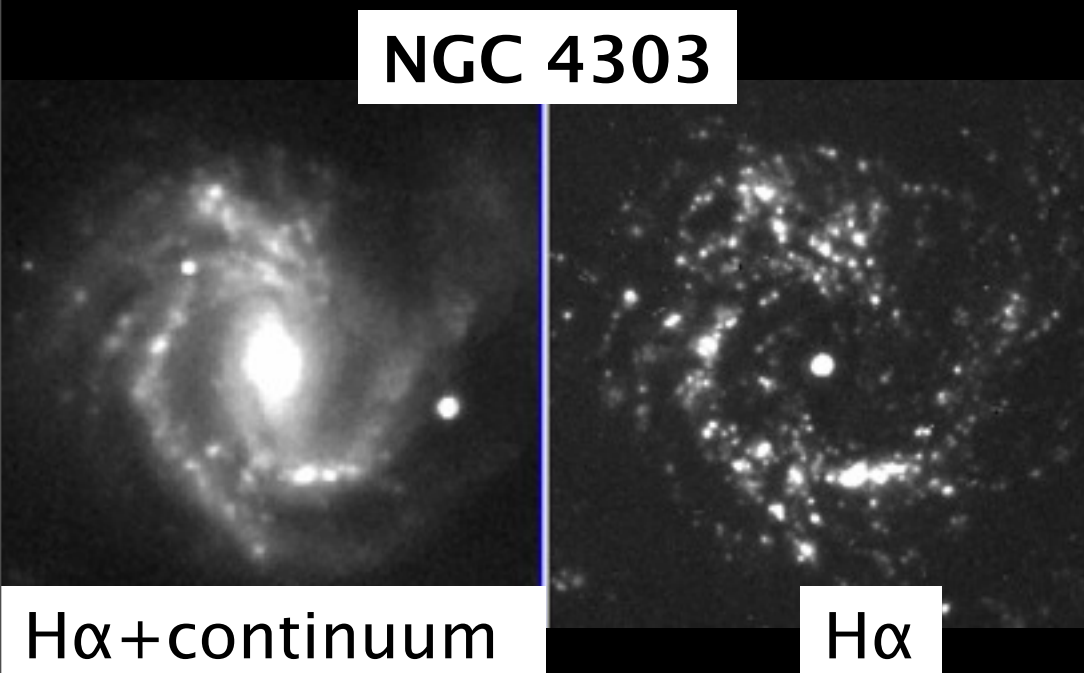
# 3.2

## STRUCTURAL K-S LAW WITHIN A GALAXY



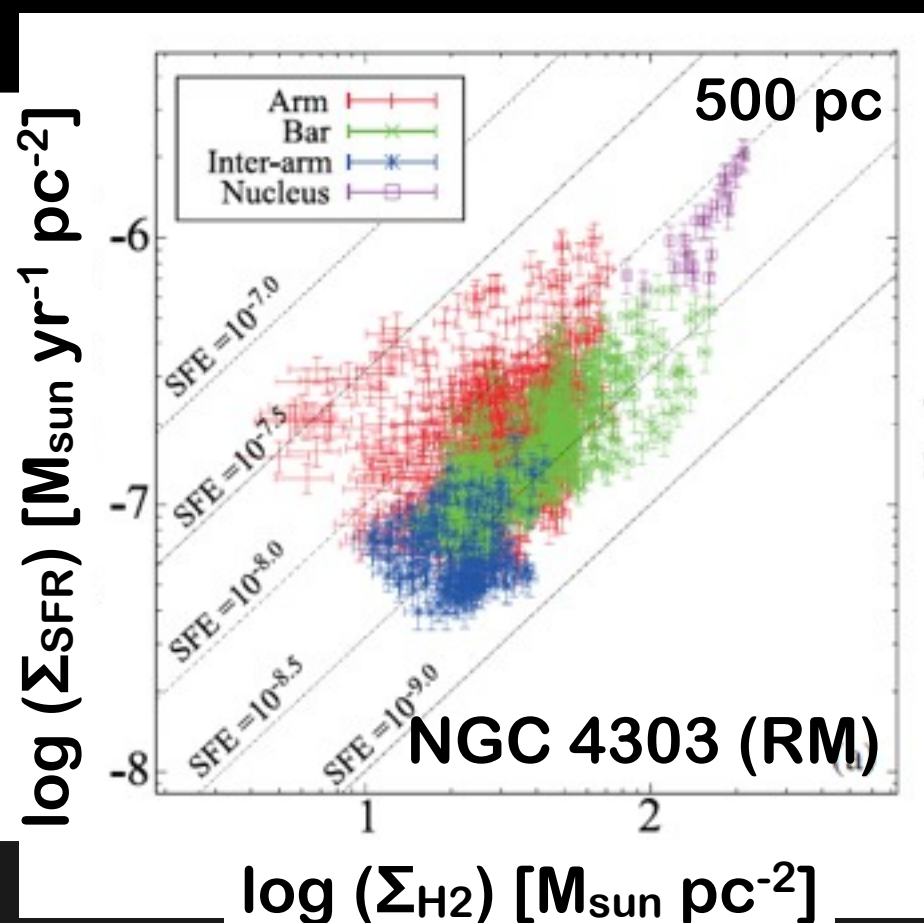


# Star Formation Within A Galaxy



We confirmed the difference of SFE depending on galactic structures (Momose et al. 10)

- Active star formation in the spiral arms
- Star formation is suppressed in the bar
- Galactic dynamics (e.g. galactic shock, shear in the bar) would regulate star formation (e.g. Downes et al. 96; Sheth et al. 02, 04 )



※ Star Formation Efficiency (SFE)  
The ratio of SFR to gas mass [/yr]

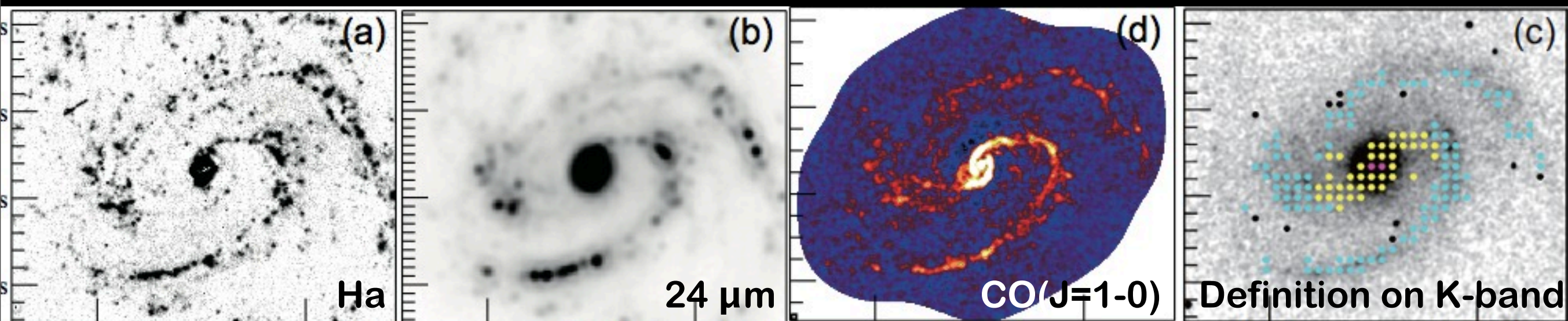
# Definition of Galactic Structures

- We divide all regions of a galaxy to the nucleus, spiral arms, bar and other regions where is including inter-arm region

Nucleus: central 500 pc around the dynamical center

Bar: estimate using IRAF

Spiral Arms/Other: estimated using GALFIT



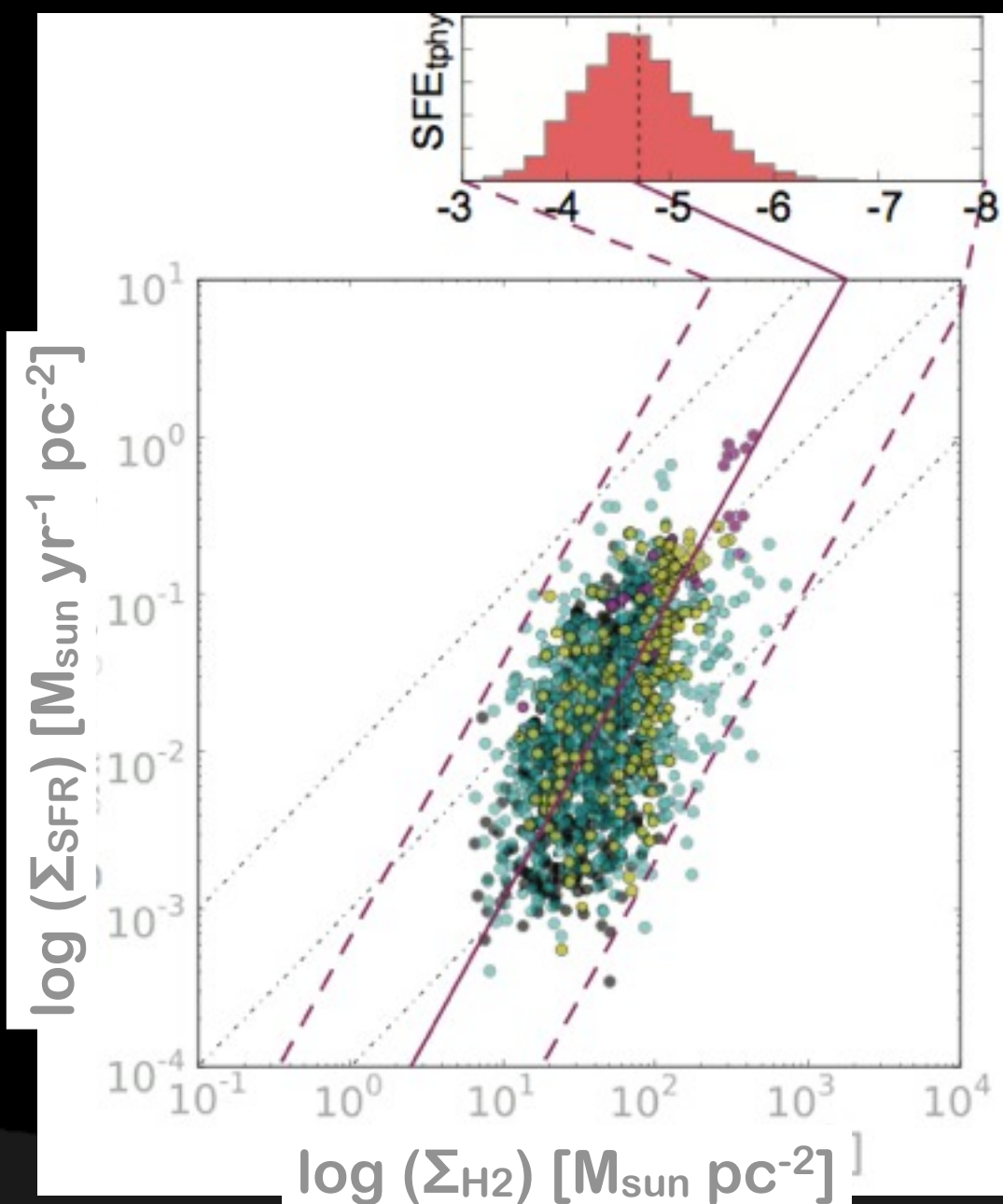
# Star Formation in Each Structure



not yet  
published



# The Variation of Star Formation Activities



- We define  $\text{SFE}_{\text{tphy}}$  as following:

- \* Efficiency toward the averaging star formation mechanism given area

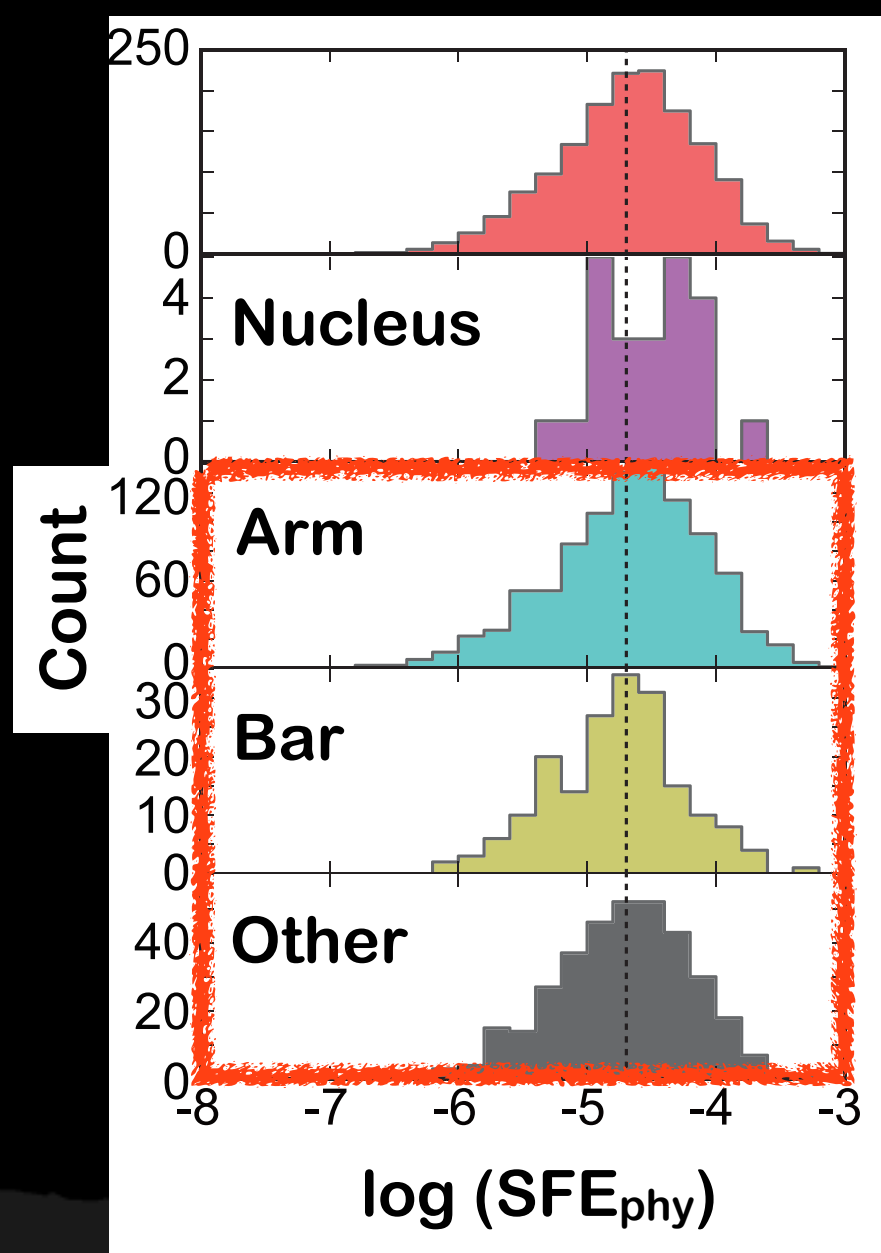
$$\Sigma_{\text{SFR}} = \epsilon_{\text{tphy}} \frac{\Sigma_{\text{gas}}}{t_{\text{phy}}},$$

$$\Sigma_{\text{SFR}} \propto \epsilon_{\text{tphy}} \Sigma_{\text{gas}}^{1-m} = \epsilon_{\text{tphy}} \Sigma_{\text{gas}}^N,$$

$$\text{SFE}_{\text{tphy}} = \frac{\Sigma_{\text{SFR}} [M_{\odot} \text{ yr}^{-1} \text{ kpc}^{-2}]}{\Sigma_{\text{H}_2}^N [M_{\odot} \text{ pc}^{-2}]^N}.$$

$$N = 1.75$$

# The Variation of Star Formation Activities



- We compare peak and distribution of the  $\text{SFE}_{\text{tphys}}$ 
  - Scatter: the nucleus is smaller than rest of structures
  - Peak: little different
- The  $\text{SFE}_{\text{tphys}}$  are little different among structures from all galaxies
  - Averaging mechanism of star formation can be the same
  - The nucleus may be different than rest of structures

# Summary 2

- We compare the K-S law for each structure of a galaxy -- the nucleus, bar, spiral arms and other region
- Averaging structures from all sample galaxy show the little difference of star formation mechanism
- Star formation mechanism may be different from other structures of a galaxy



# 4. SUMMARY



# Summary

## Aim & Method

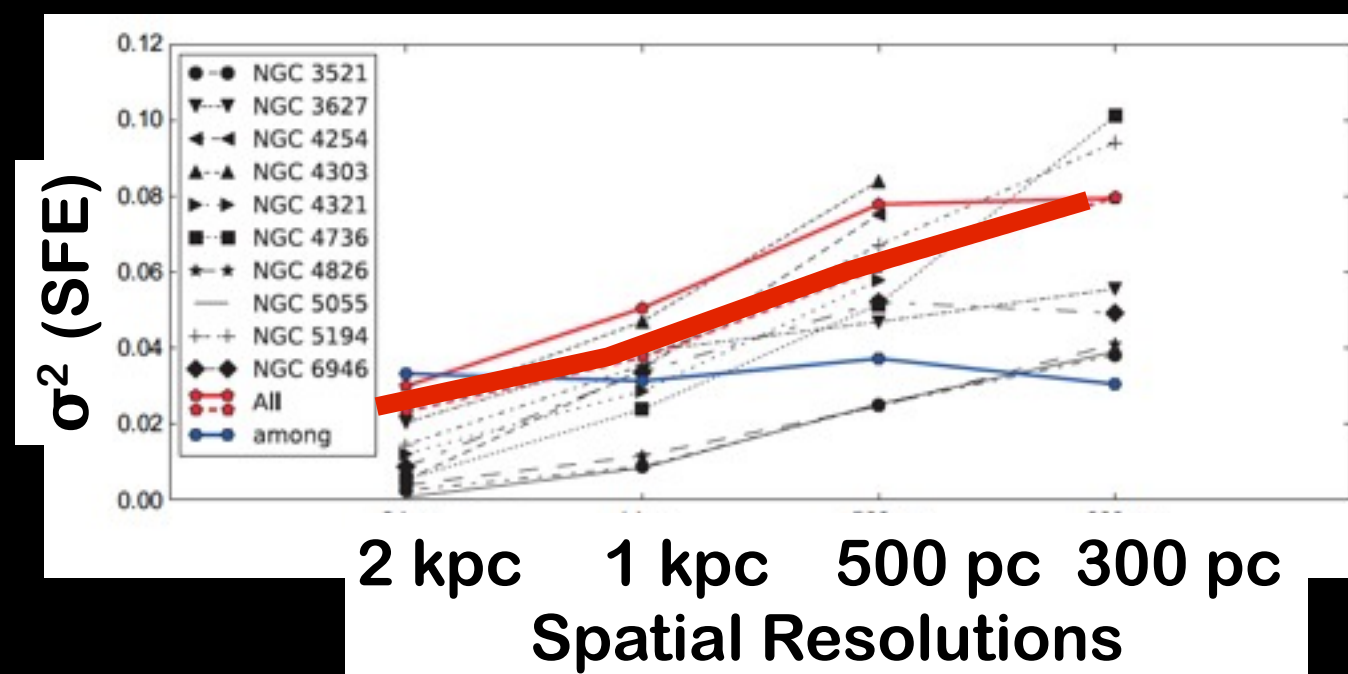
- We study the K-S law on 500 pc scale among nearby 10 galaxies
- In order to estimate the amount of molecular gas and SFR accurately,
  - we combined CO(J=1-0) data observed by CARMA and NRO45, respectively
  - we subtracted DIG from Ha and 24um images

## Result & Discussion

- We obtained the super-linear slope of the K-S law ( $N = 1.75$ )
  - Star formation can be induced by some trigger mechanism
  - The correlation between SFR and bulk of molecular gas traced by CO(J=1-0) can show the mechanism of star formation
- Self-gravity of molecular gas is a main mechanism for star formation in each structure of a galaxy

# Future Works

- Bulk gas vs SFR
- Global K-S law  $\rightarrow N > 1$  ( $\sim 1.5$ )
- **Resolved K-S law  $\rightarrow N > 1$**
- GMC scale  $\rightarrow$  break-down



Which scale causes break-down of the K-S law?

