

# MC study of 2km muon range detector for JHF neutrino experiment

A.V.Butkevich <sup>1</sup>

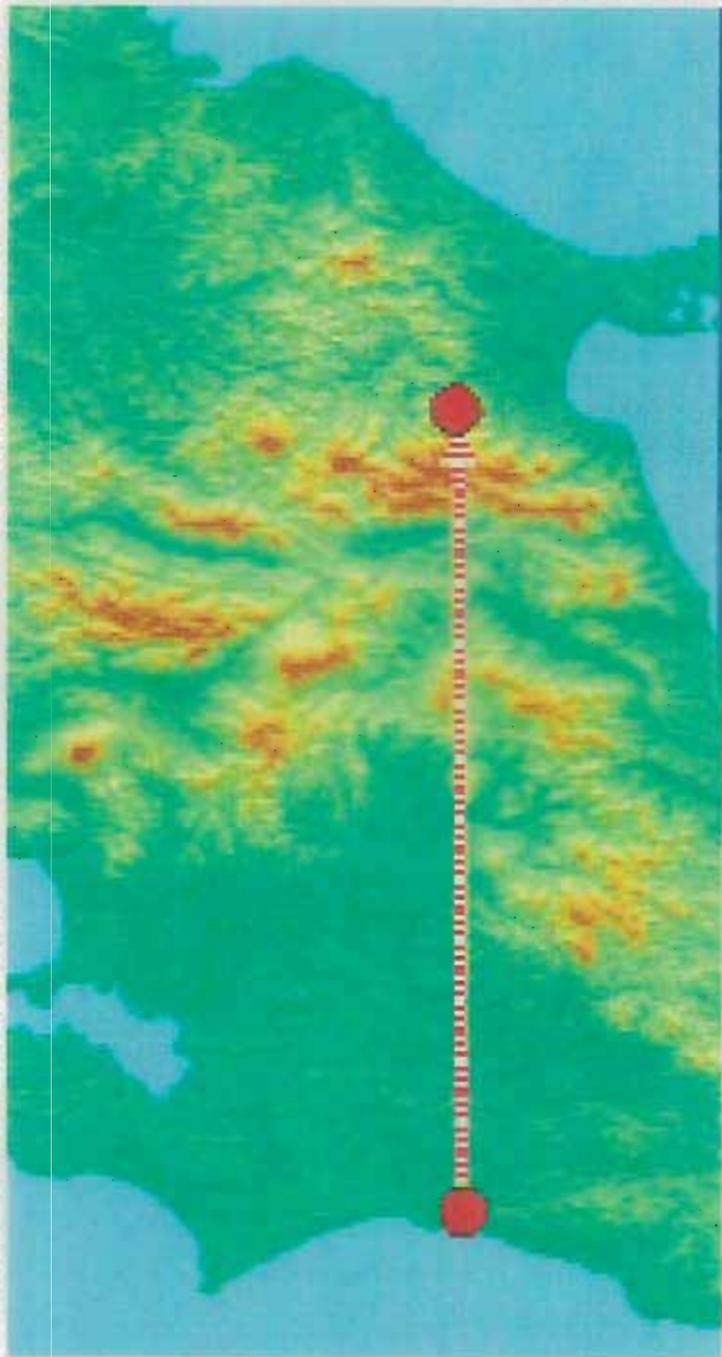
Institute for Nuclear Research of Russian Academy of Science  
Institute for Cosmic Ray Research, University of Tokyo

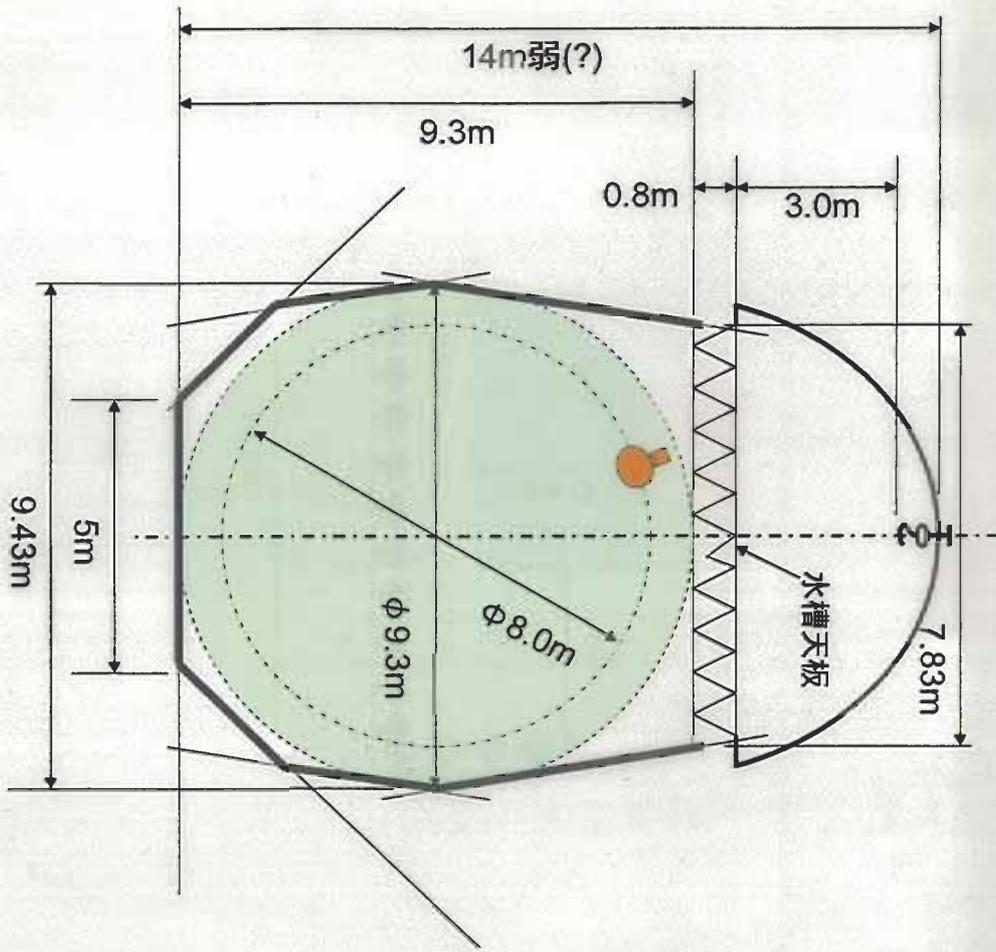
---

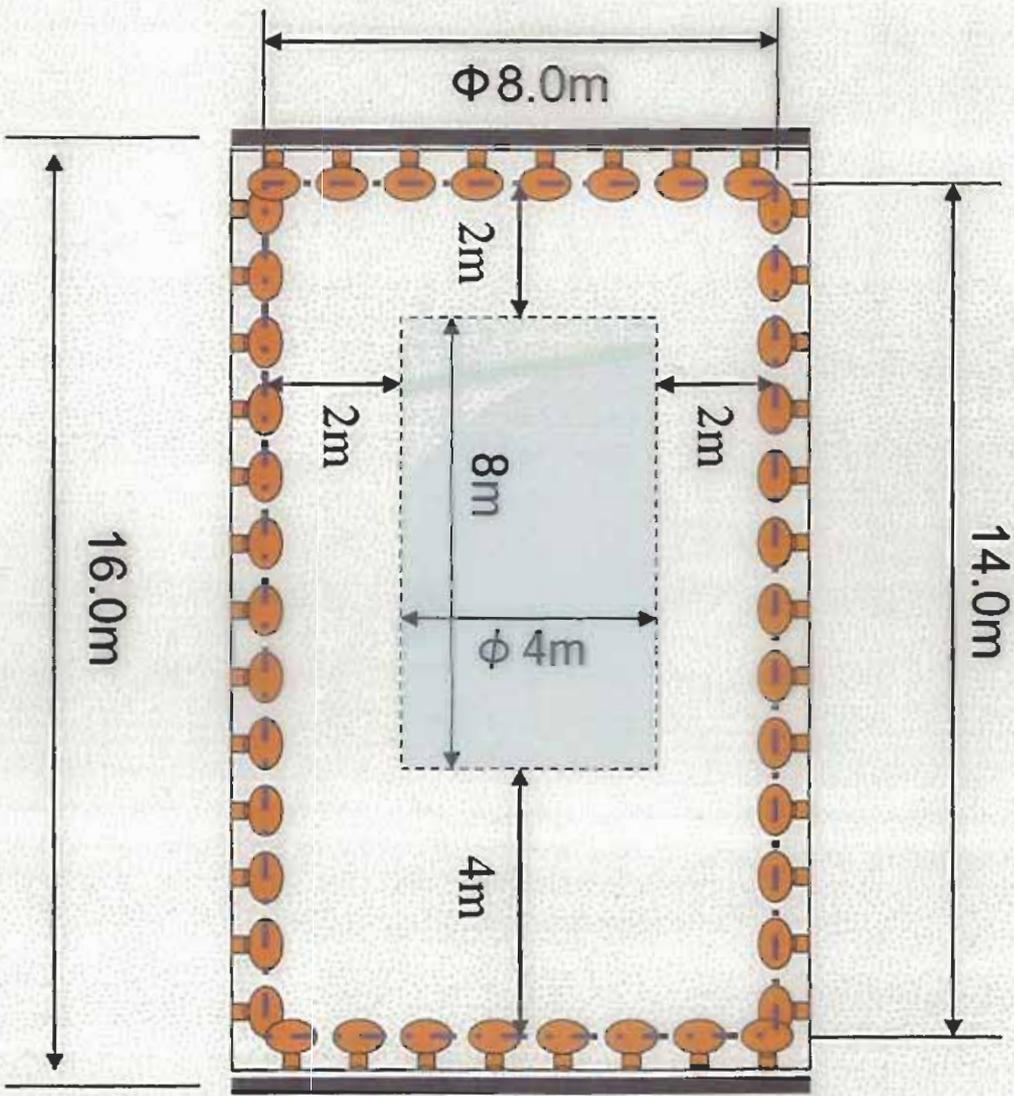
<sup>1</sup>ICRR, 25<sup>th</sup> December 2003

## Outline

- **Introduction: 2km detector**
- **MRD**
  - Physics goals
  - Design
  - Muon track reconstruction
  - Resolutions
- **Cherenkov detector + MRD**
  - Reconstarction efficiency
  - Resolutions
- **Concluding remarks**

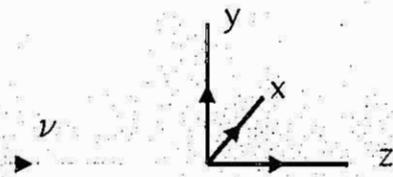
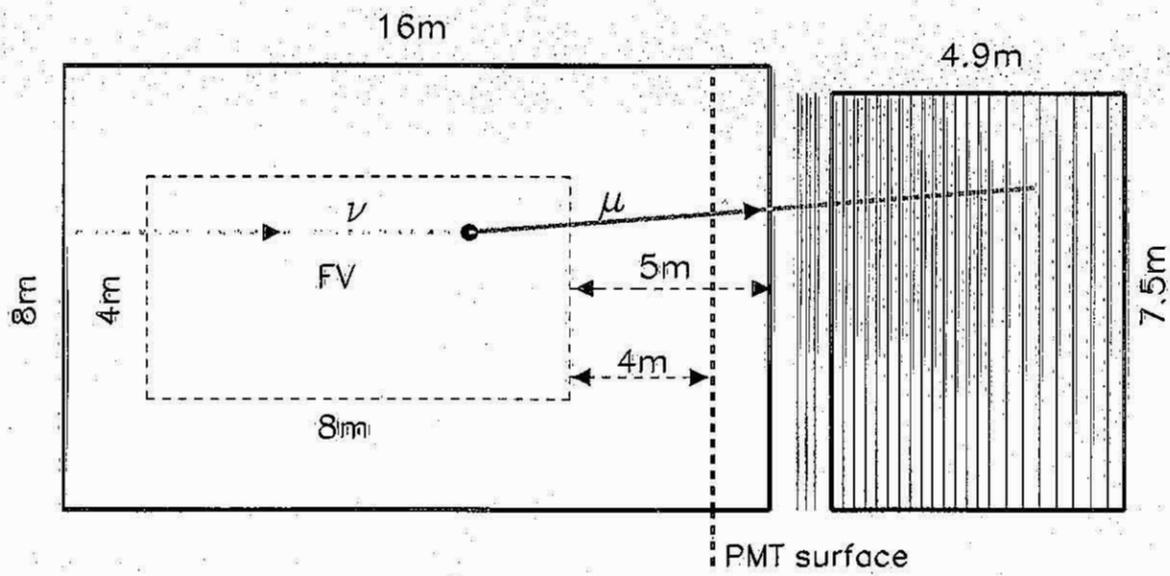






Cherenkov Detector

MRD

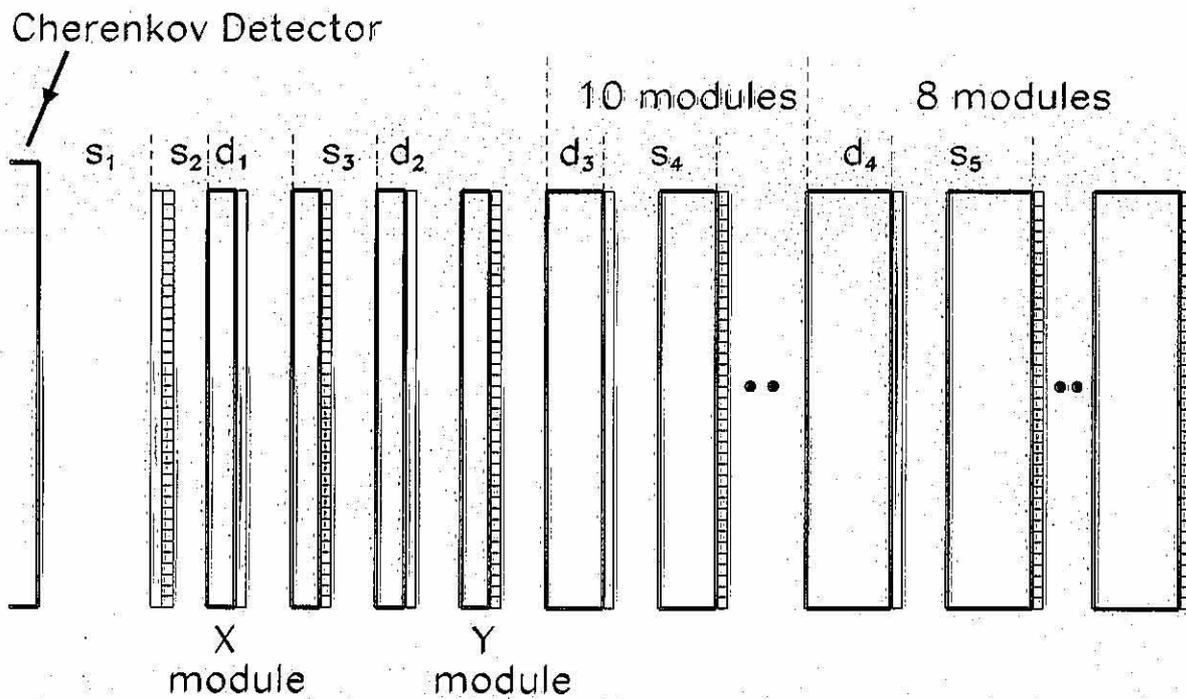


## Requirements of MRD

- Low energy threshold
- Good energy resolution for wide muon energy region
- High reconstruction efficiency
- Size can't be more than 5m length, 8m width and high

## MRD Side View

24 scintillator modules and 22 steel planes



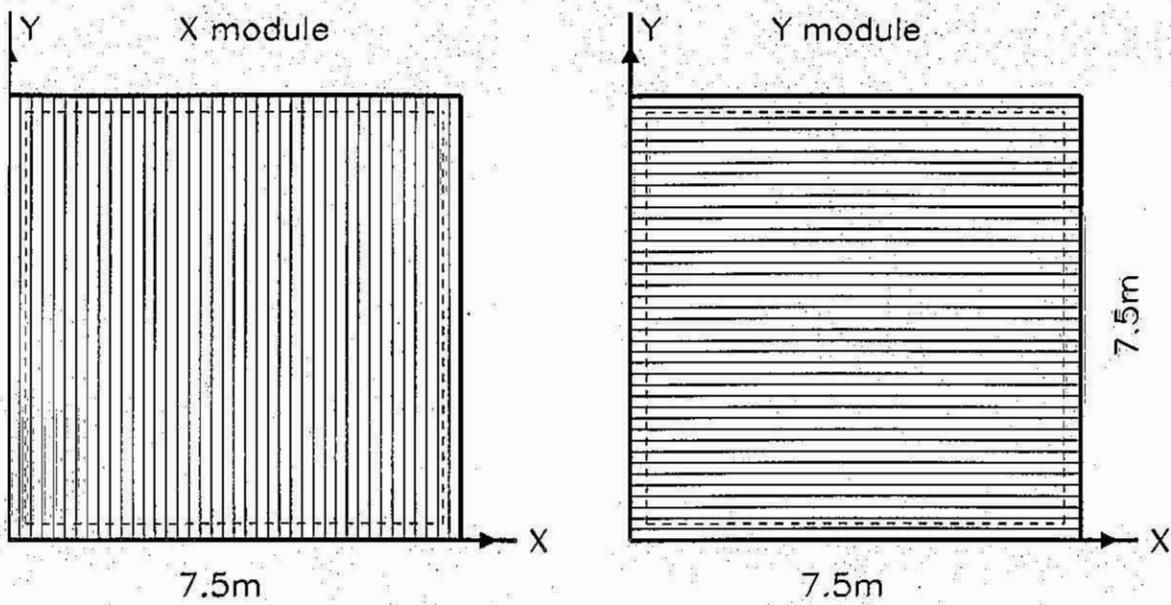
Thickness of steel planes

MRD2  $d_1=5\text{cm}$ ,  $d_2=5\text{cm}$ ,  $d_3=10\text{cm}$ ,  $d_4=20\text{cm}$ ,  $H=2224\text{g/cm}^2$

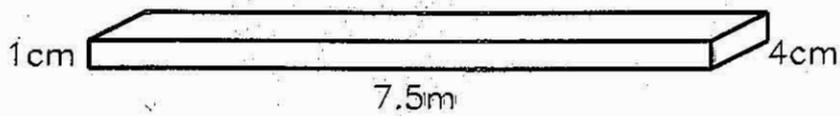
MRD3  $d_1=2.5\text{cm}$ ,  $d_2=5\text{cm}$ ,  $d_3=10\text{cm}$ ,  $d_4=20\text{cm}$ ,  $H=2185\text{g/cm}^2$

$s_1=74\text{cm}$ ,  $s_2=11\text{cm}$ ,  $s_3=19.1\text{cm}$ ,  $s_4=19.1\text{cm}$ ,  $s_5=29.1\text{cm}$

### Scintillator modules



187strips/module



## Muon tracking in MRD: linear fit

- A reconstructed track  $(x_i, y_i)$   
$$x_i = x_0 + b_x(z_i - z_0), \quad b_x = \left(\frac{P_x}{P_z}\right)_0$$
$$y_i = y_0 + b_y(z_i - z_0), \quad b_y = \left(\frac{P_y}{P_z}\right)_0$$
 $(x_0, y_0)$  - track origin  
 $(b_x, b_y)$  - slop of the track in xz and yz planes at its original

- Least squares method

$$\chi_x^2 = \sum_{i,j} (x_i^m - x_j) w_{ij} (x_j^m - x_j),$$

$$\chi_y^2 = \sum_{k,l} (y_k^m - y_l) w_{kl} (y_l^m - y_l)$$

$x^m, y^m$  - the measured value of track position

- Weight matrix

$$[w_{ij}] = [V_{ij}]^{-1}, \quad V = V_{er}(\epsilon) + V_{ms}(P)$$

$[V_{ij}]$  - covariance/error matrix

The uncertainties arise from limitations in the measurement accuracy of the detectors and from errors induced by multiple scattering.

- Measured errors are taken into account by a diagonal matrix  $[V_{ij}]_{er} = \epsilon^2 \delta_{ij}$   
 $\epsilon = r/\sqrt{12}$ ,  $r=4$  cm is strip's width.

- Multiple scattering matrix take into correlation between measurements

$$[V_{ms}]_{ij} = \langle P_i P_j \rangle.$$

$\langle P_N^2 \rangle$  - variation in displacement at plane  $N$

$\langle P_K P_N \rangle$  - covariance between the displacements at planes  $K$  and  $N$

- $\langle P_N^2 \rangle = \sum_{i=1}^N \theta_{0i}^2 \left( \frac{x_i^2}{3} + D(i, N)(x_i + D(i, N)) \right)$

$$\langle P_N P_K \rangle = \sum_{i=1}^K \theta_{0i}^2 \left( \frac{x_i^2}{3} + \frac{x_i}{2} (D(i, K) + D(i, N)) + D(i, K)D(i, N) \right),$$

$x_i$  and  $D(i, k)$  are width of steel and the distance from plane  $i$  to plane  $K$  along the path of muon

- Mean square scattering angle at the  $i$ -th plane

$$\theta_{0i} = \frac{\theta_0}{1 - \sum_{j=1}^i (dE/dX)x_j/p_0}$$

$dE/dX$  is the muon energy losses,  $p_0$  is an initial muon momentum in GeV

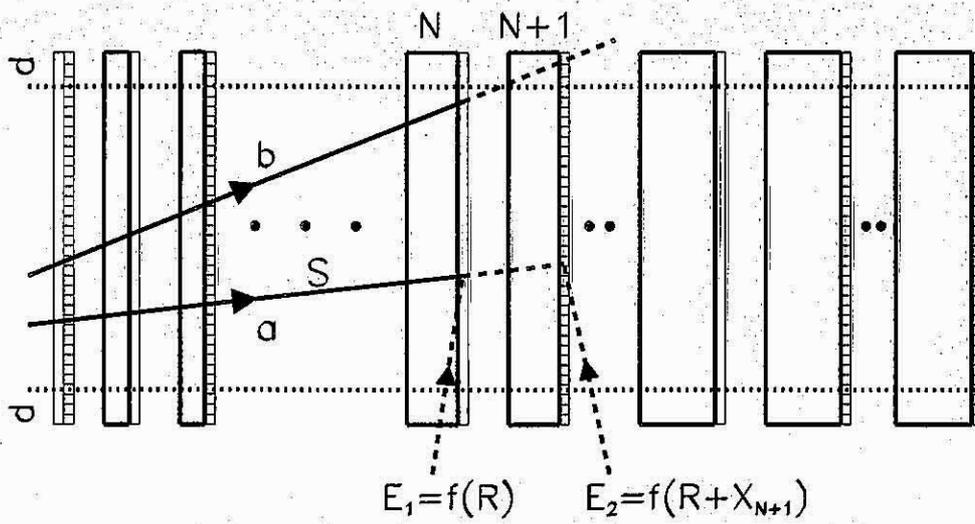
$$\theta_0 = \frac{0.0136}{p_0\beta} \left(\frac{x_i}{X_0}\right)^{1/2} [1 + 0.038 \ln \left(\frac{x_i}{X_0}\right)]$$

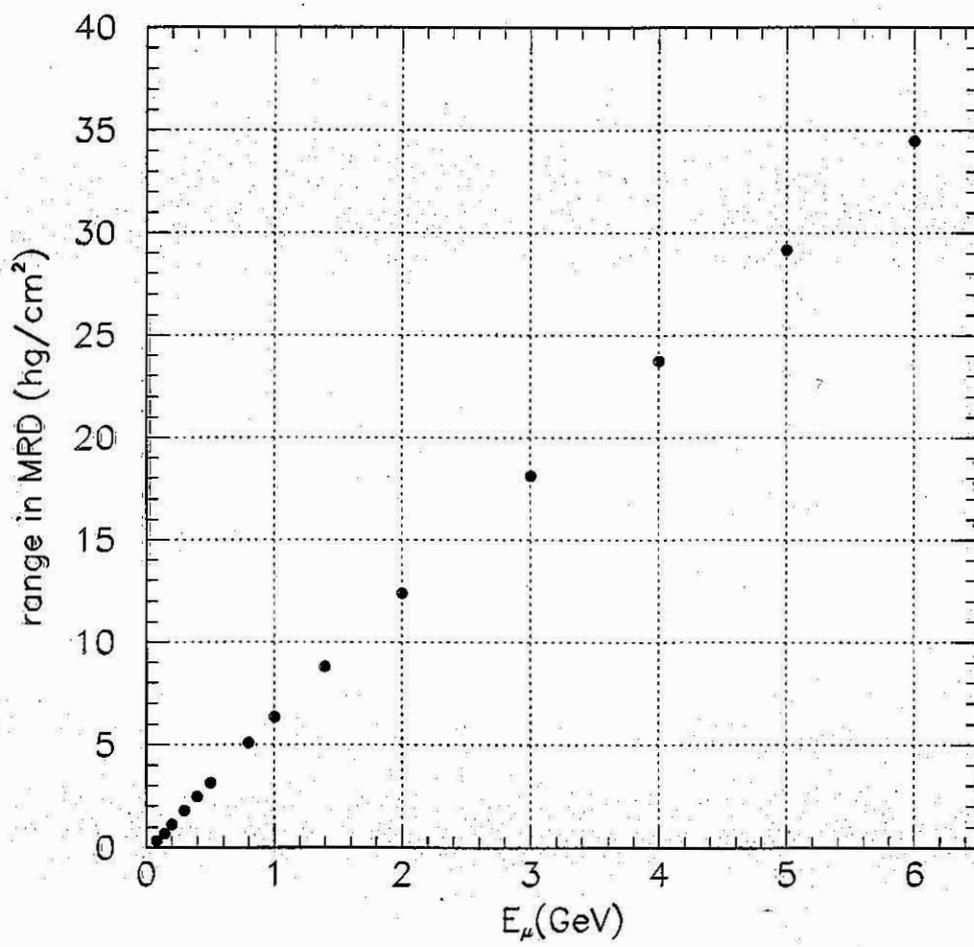
$X_0=1.76$  cm - radiation length of iron

## MRD: Muon track and energy reconstruction

- Number of the hit scintillator planes  $4 \leq N_h \leq 23$ , the track parameters  $x_0, y_0, b_x, b_y$  and track length  $S$  are calculated using  $V = V_{er}(\epsilon)$
- Energies  $E_1(S)$  and  $E_2(S + X_{N+1})$  are evaluated and the muon energy is estimated as  $E_{rec} = (E_1 + E_2)/2$
- Repeat with  $V = V_{er}(\epsilon) + V_{ms}(P)$ , using the reconstructed track direction and energy  $E_{rec}$  for calculation of  $V_{ms}$  matrix.
- Containment analyses requires a track to be contained in detector to make an energy measurement possible.

# MRD Side View





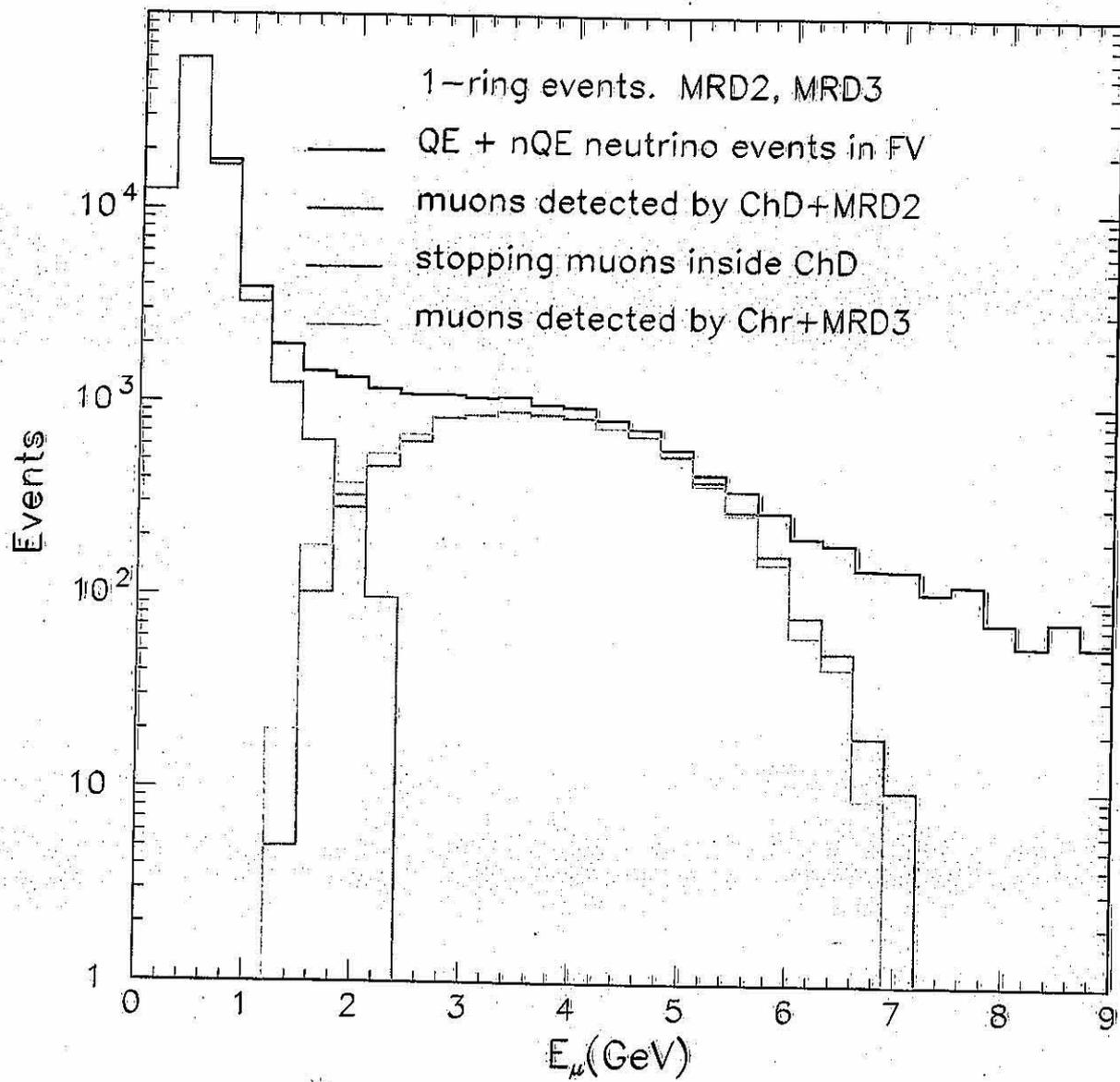
## NEUTRINO EVENTS

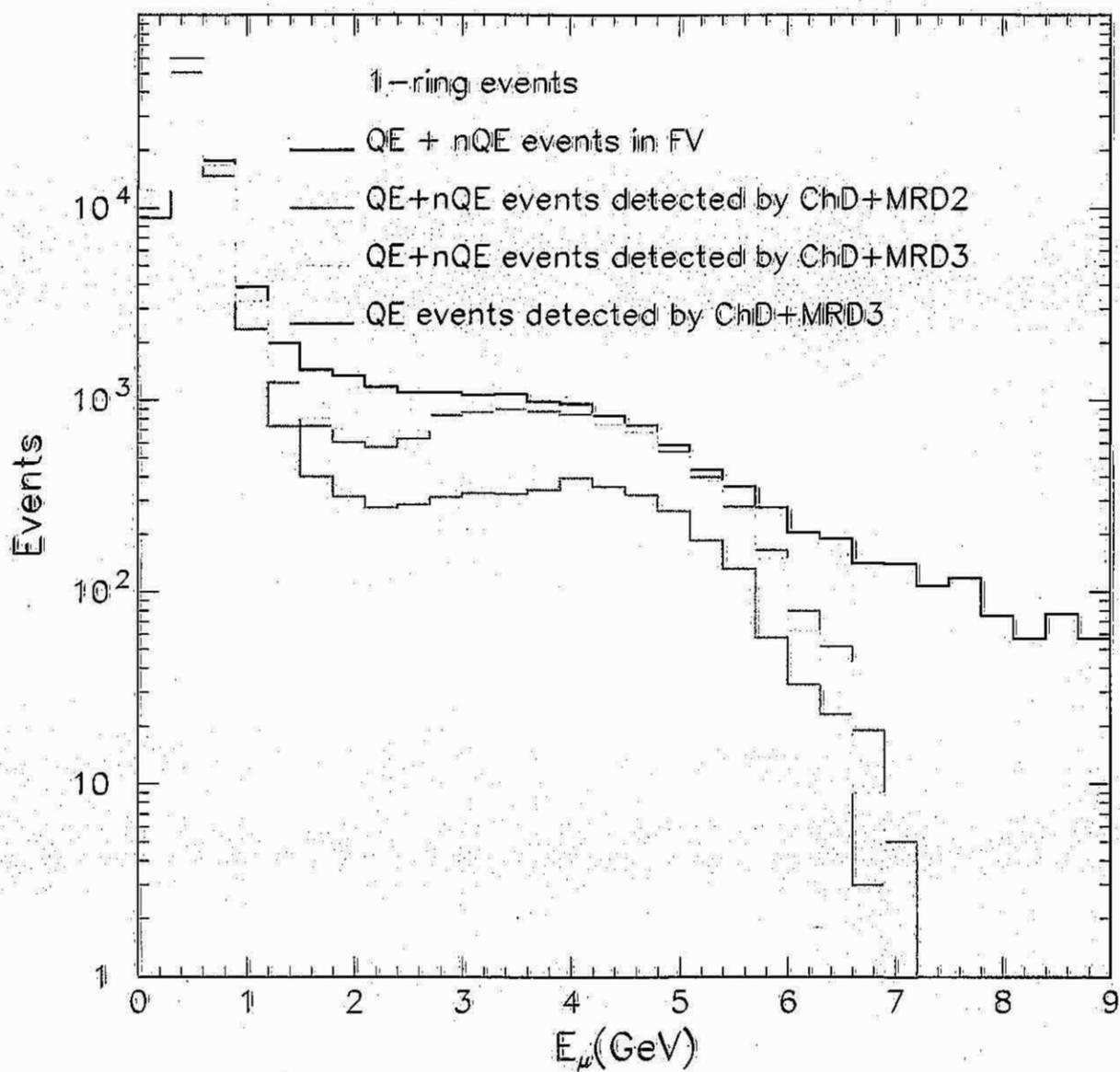
### FV of Cherenkov Detector

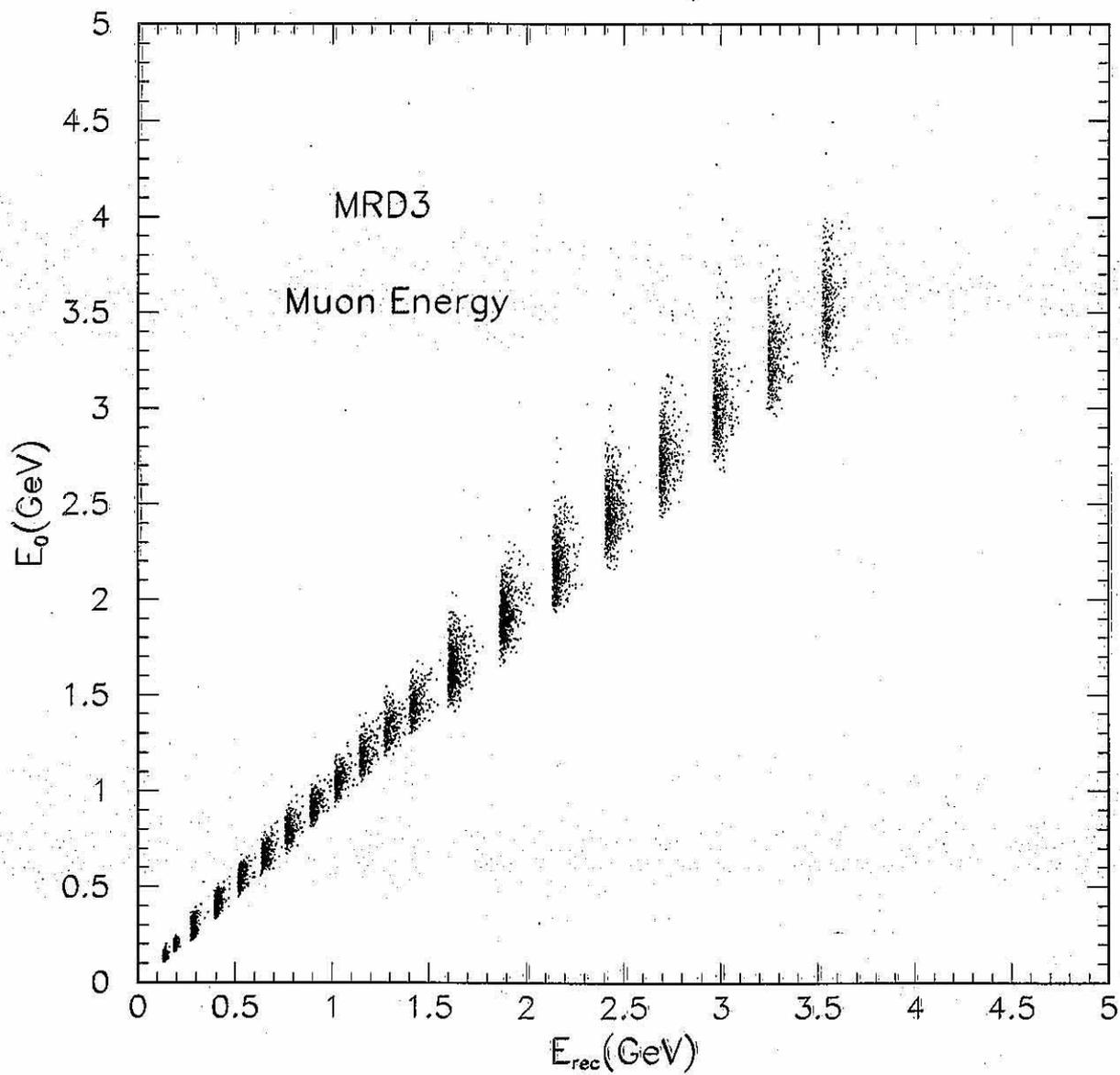
	1ring	total
QE	85172	116766
QE+nQE	110445	213791
QE/nQE	3.5	1.2

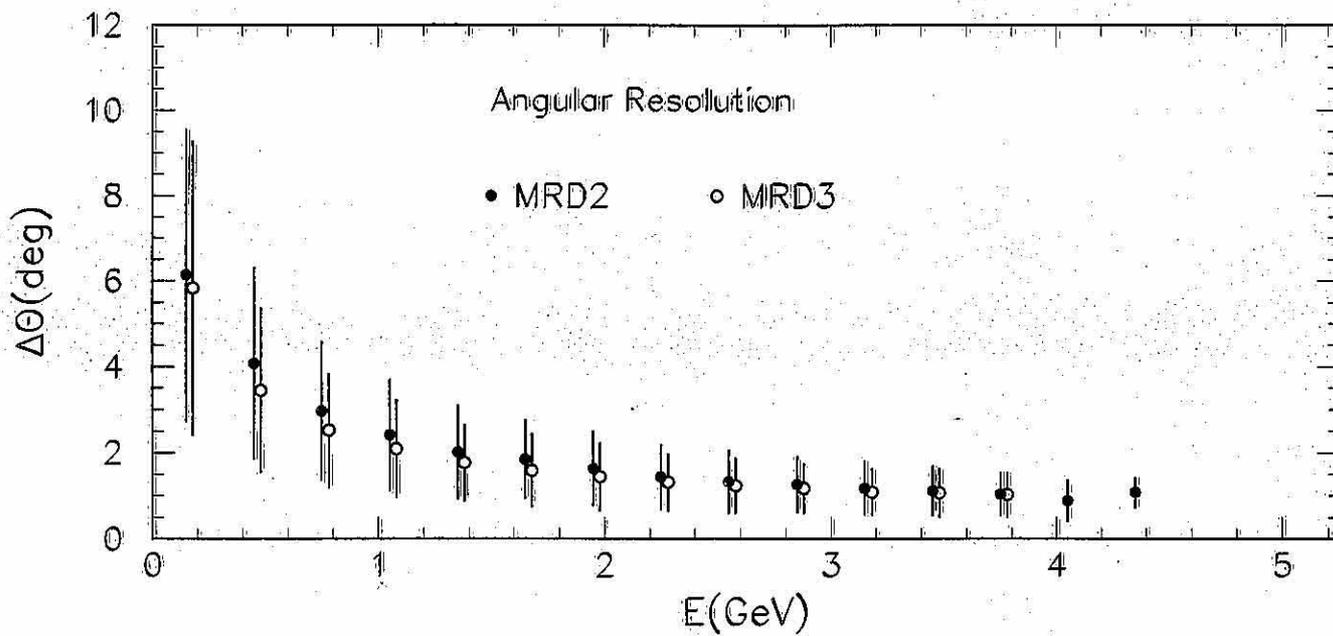
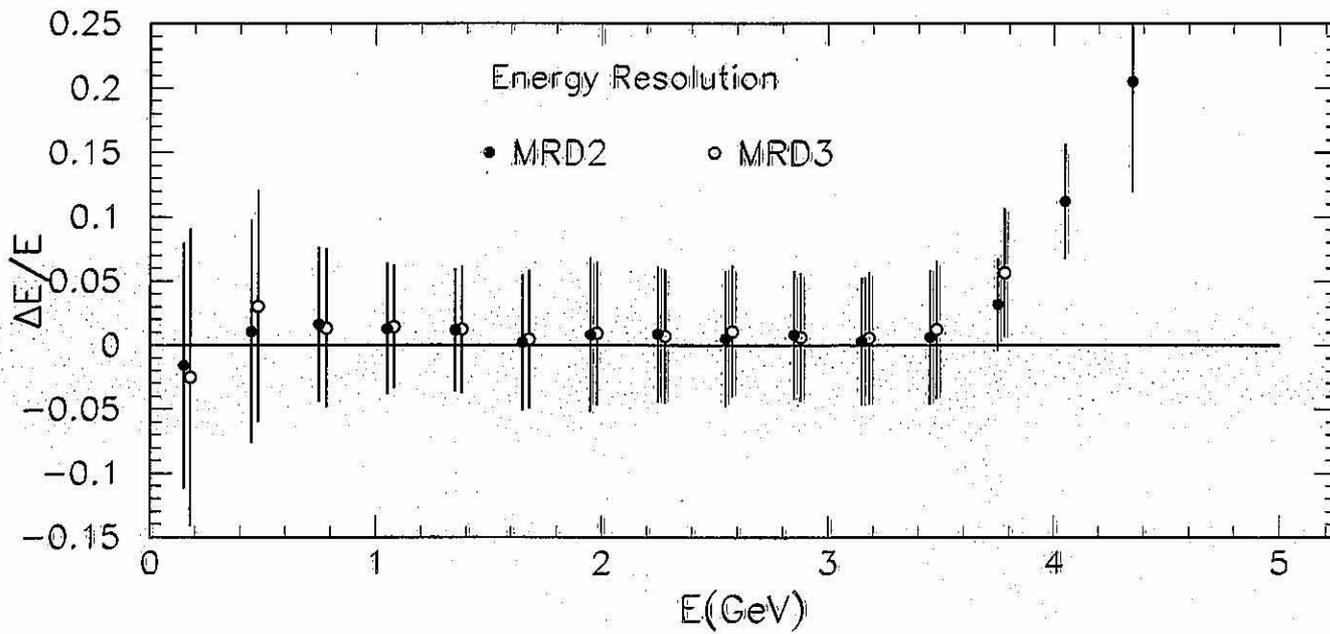
### Cherenkov Detector + MRD

	1ring	total
QE	3847	5475
QE+nQE	9014	27014
QE/nQE	0.75	0.25



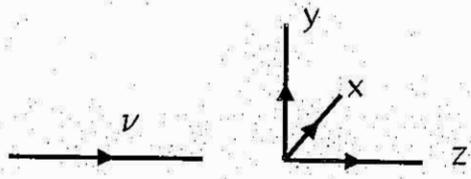
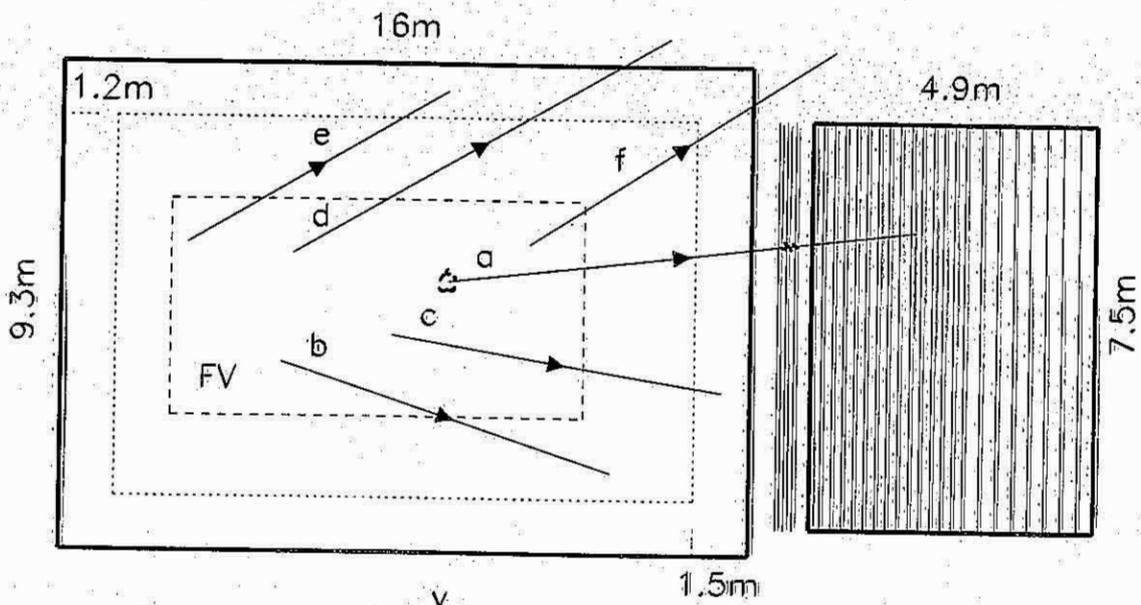


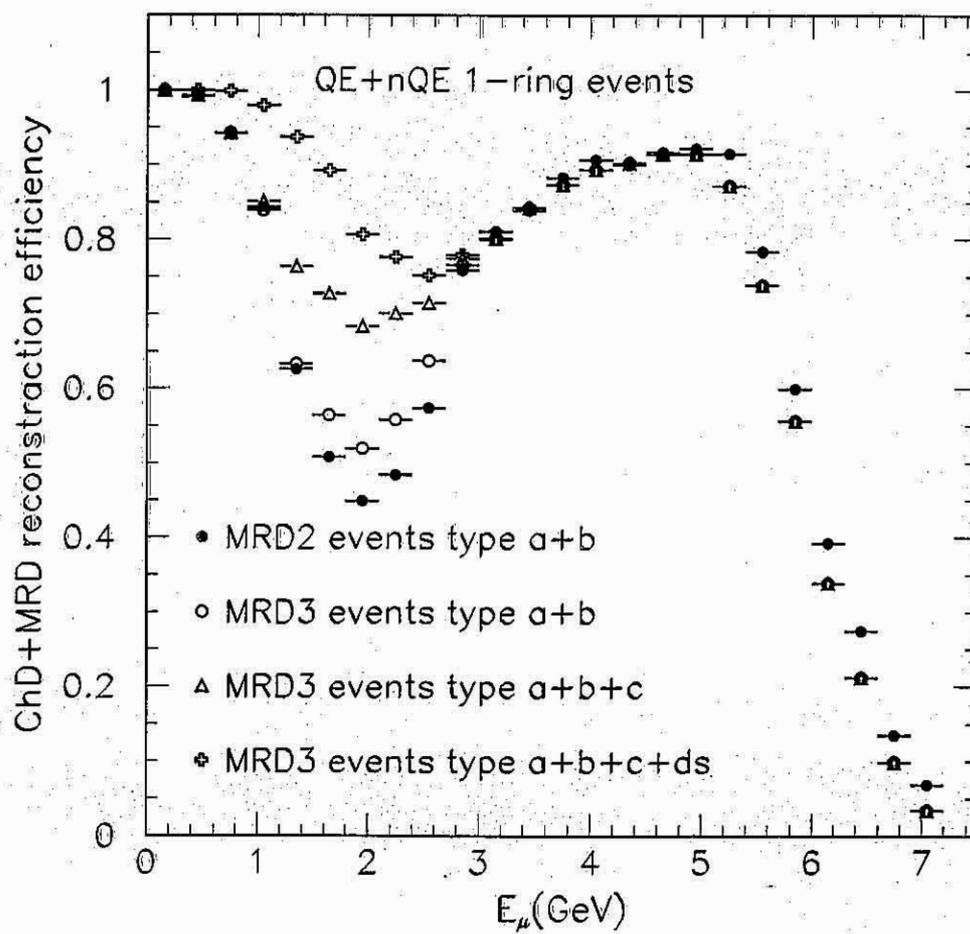




Cherenkov Detector

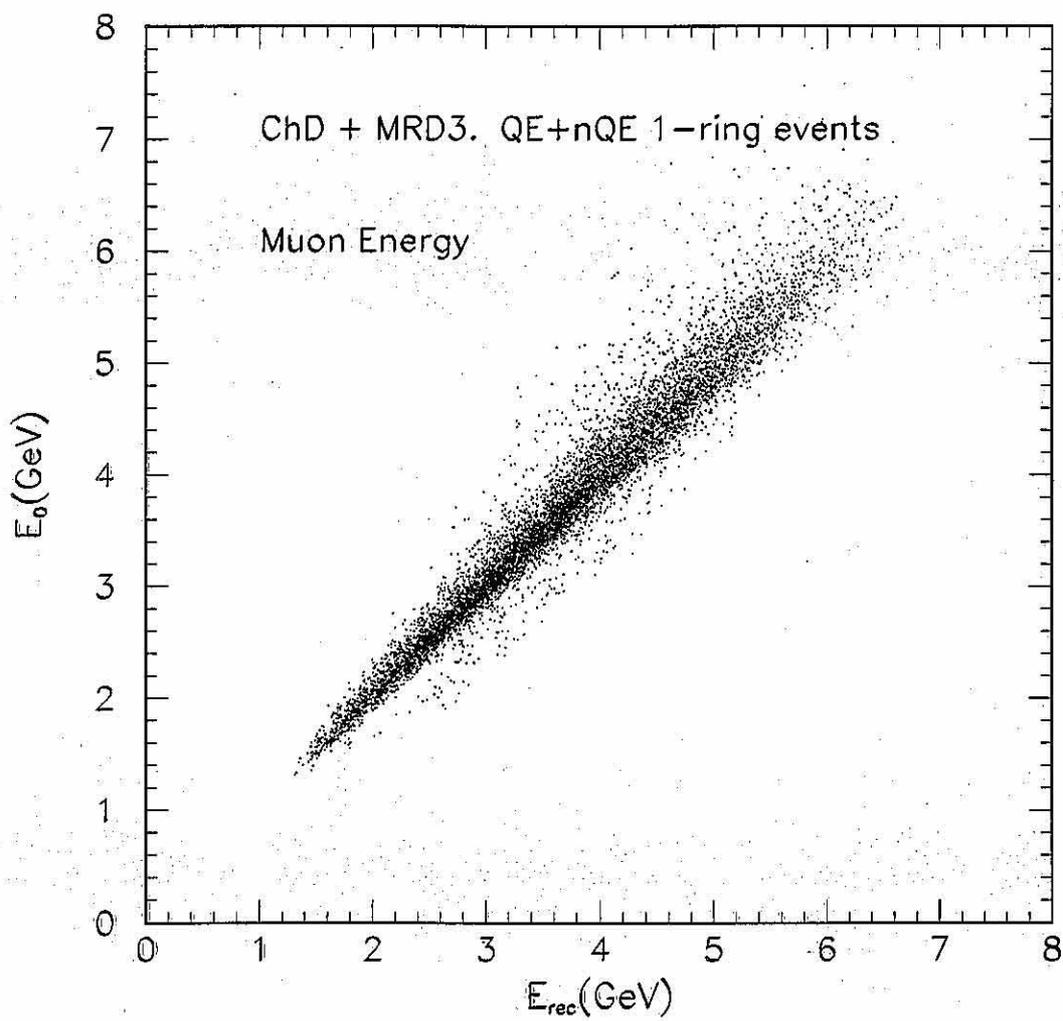
MRD

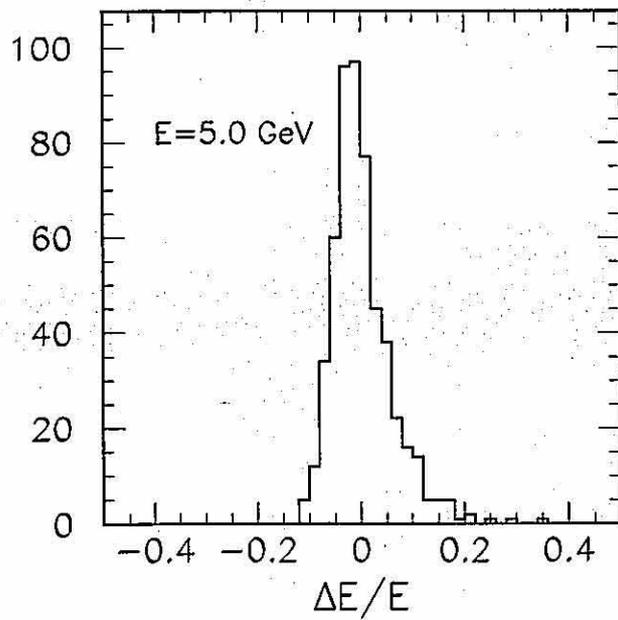
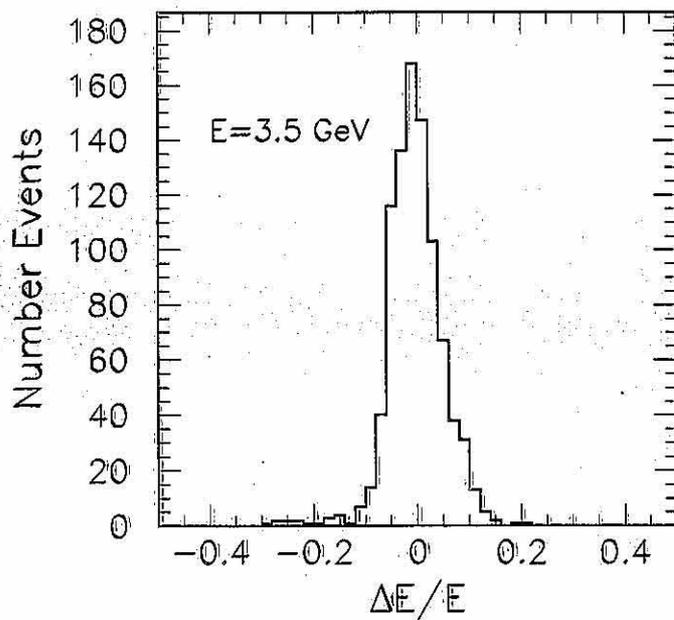
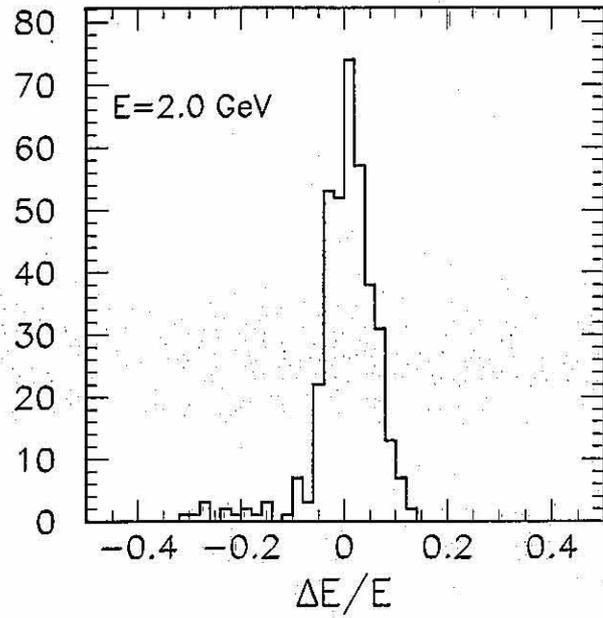
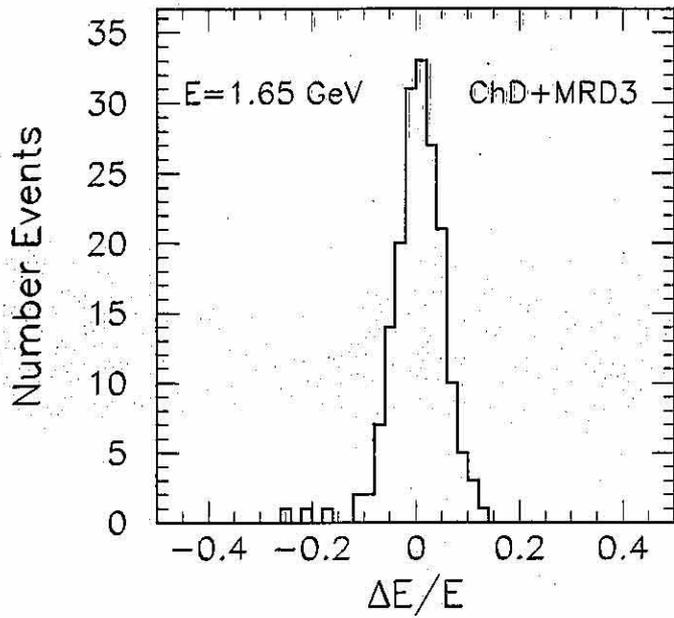


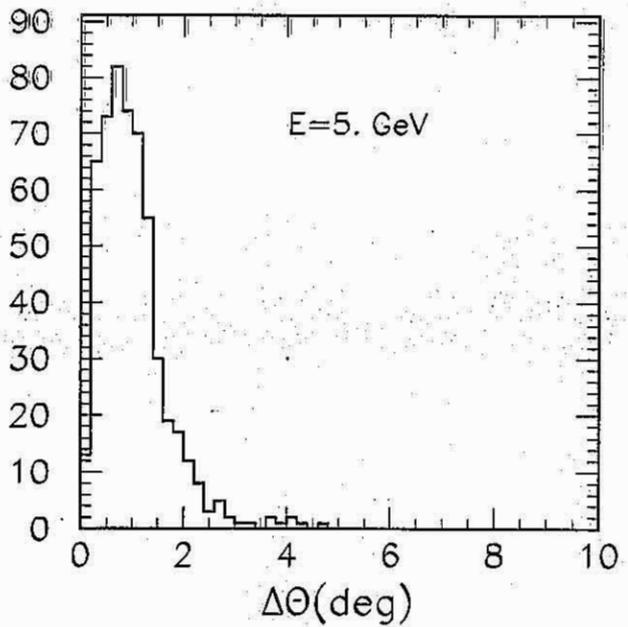
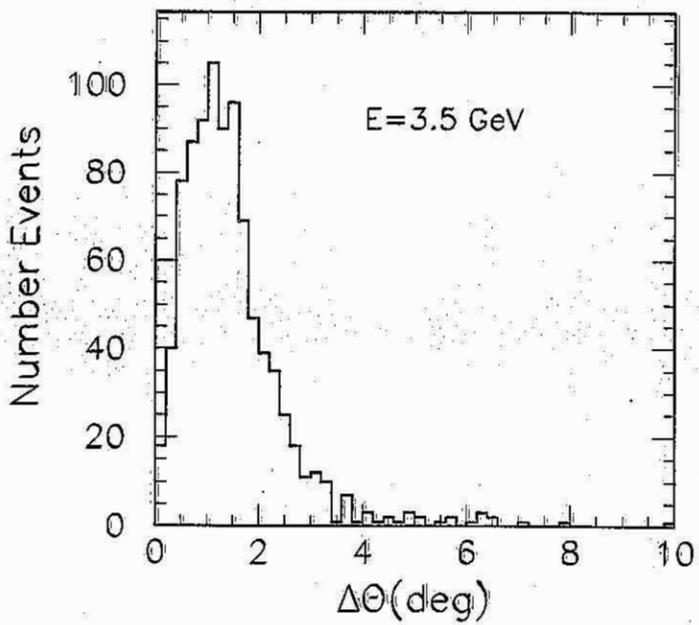
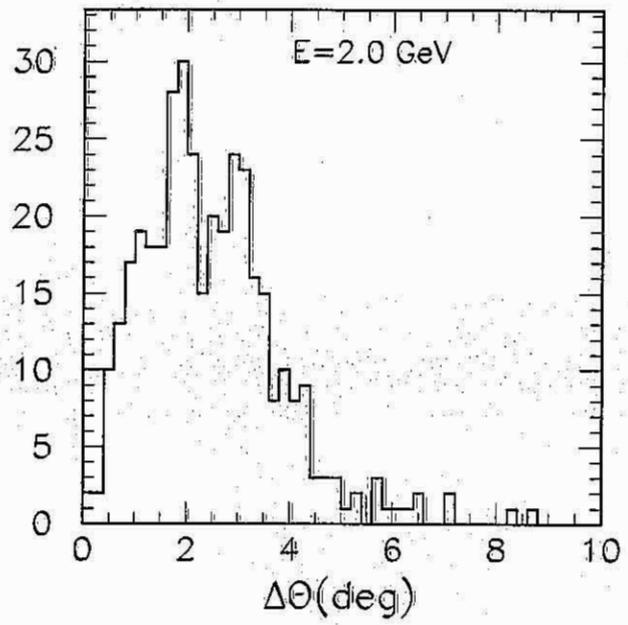
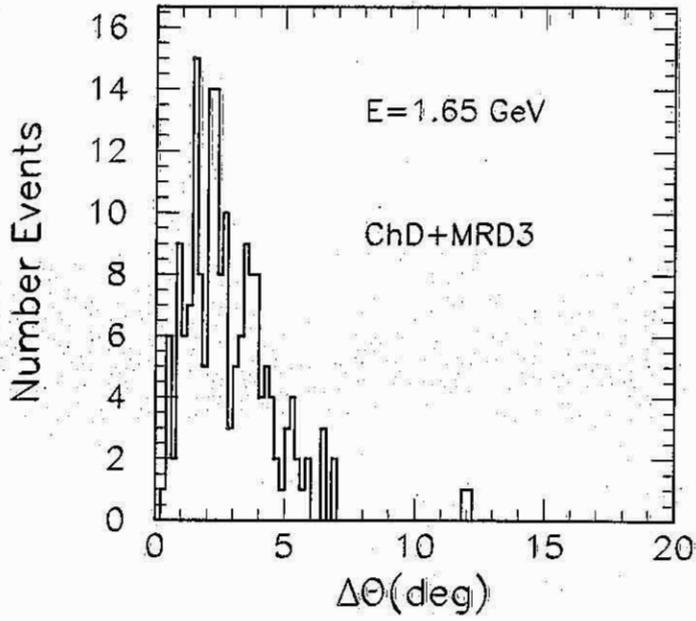


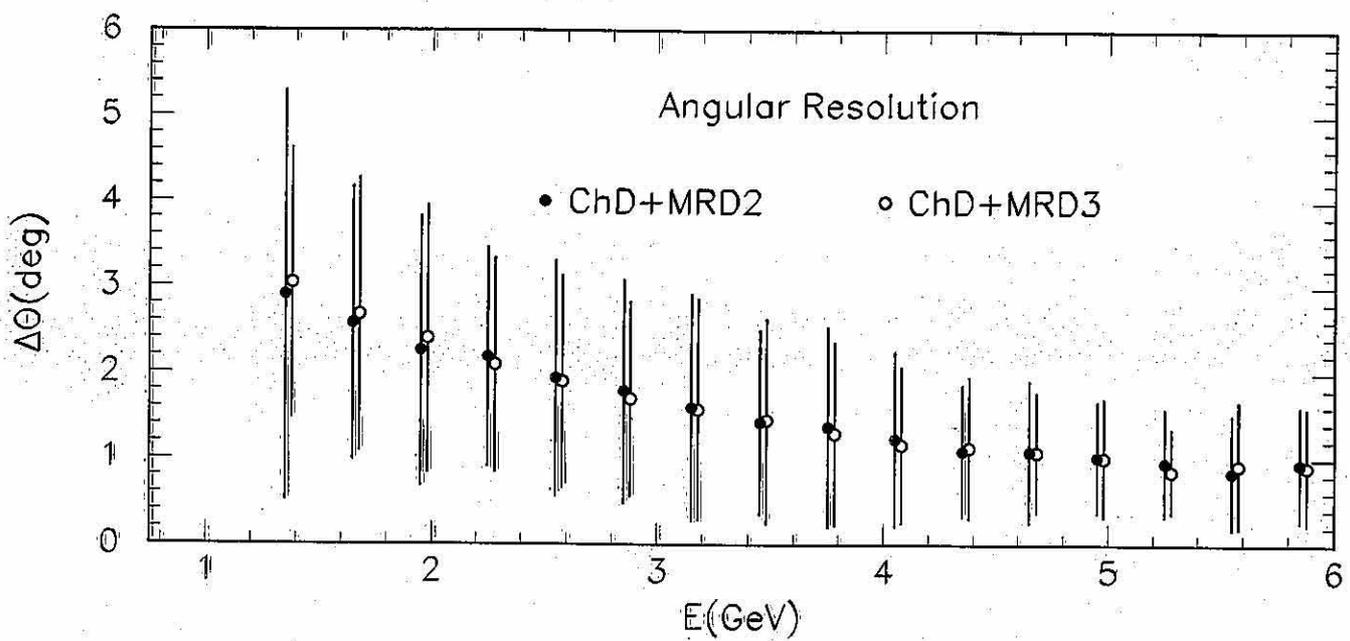
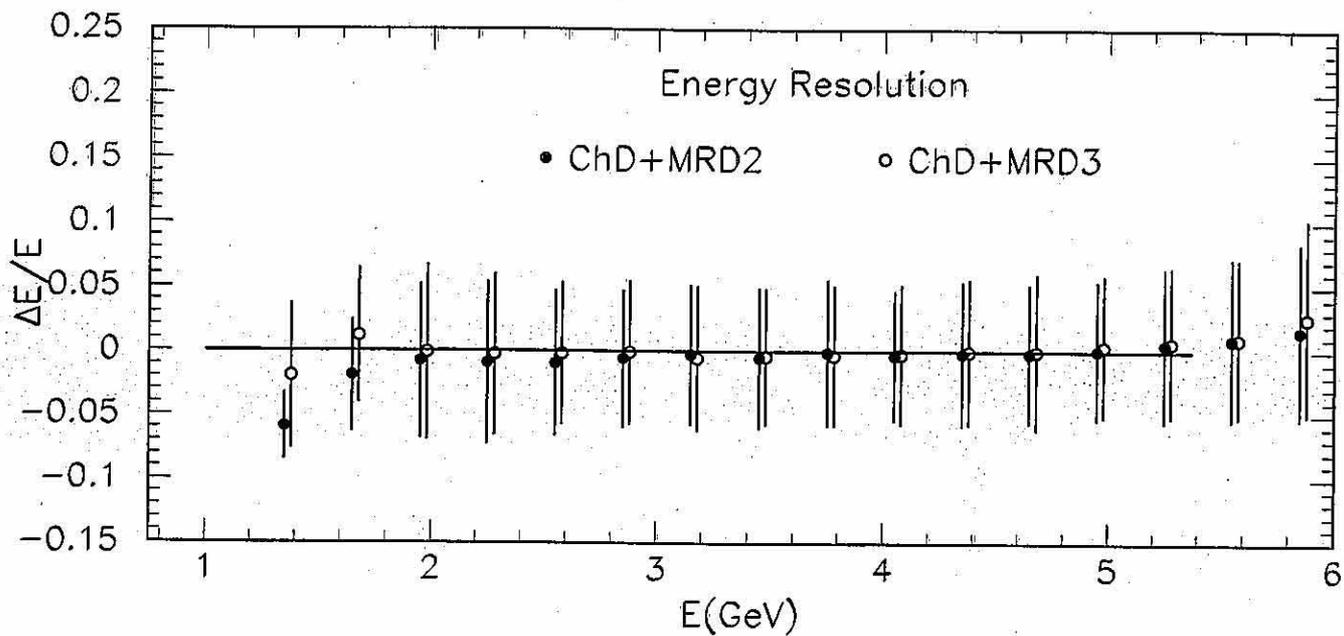
## ChD&MRD: Muon energy reconstruction method

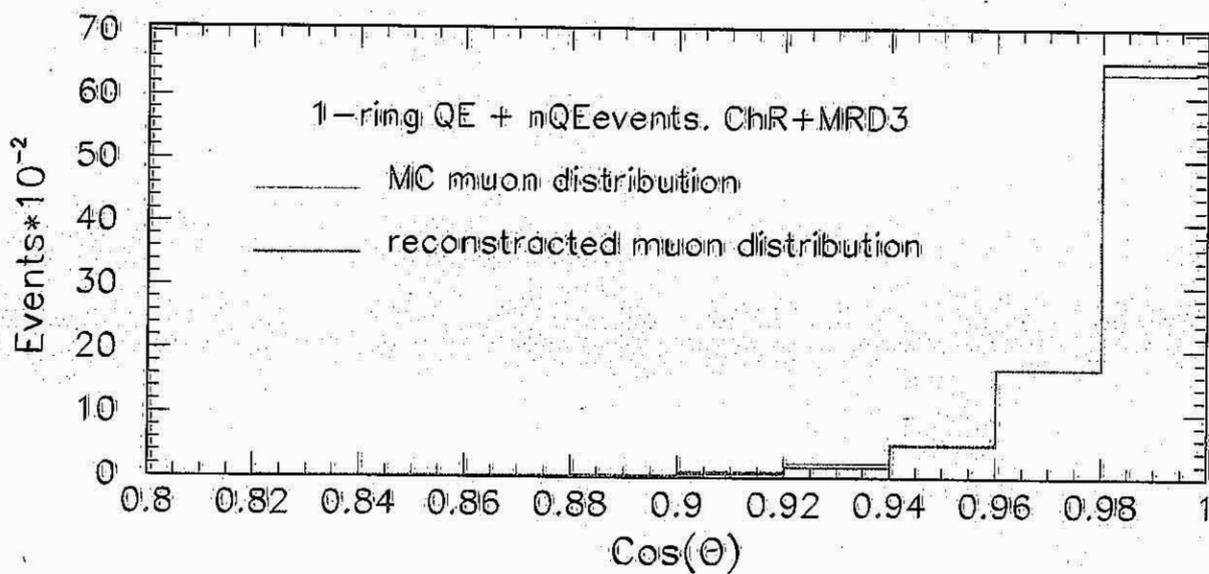
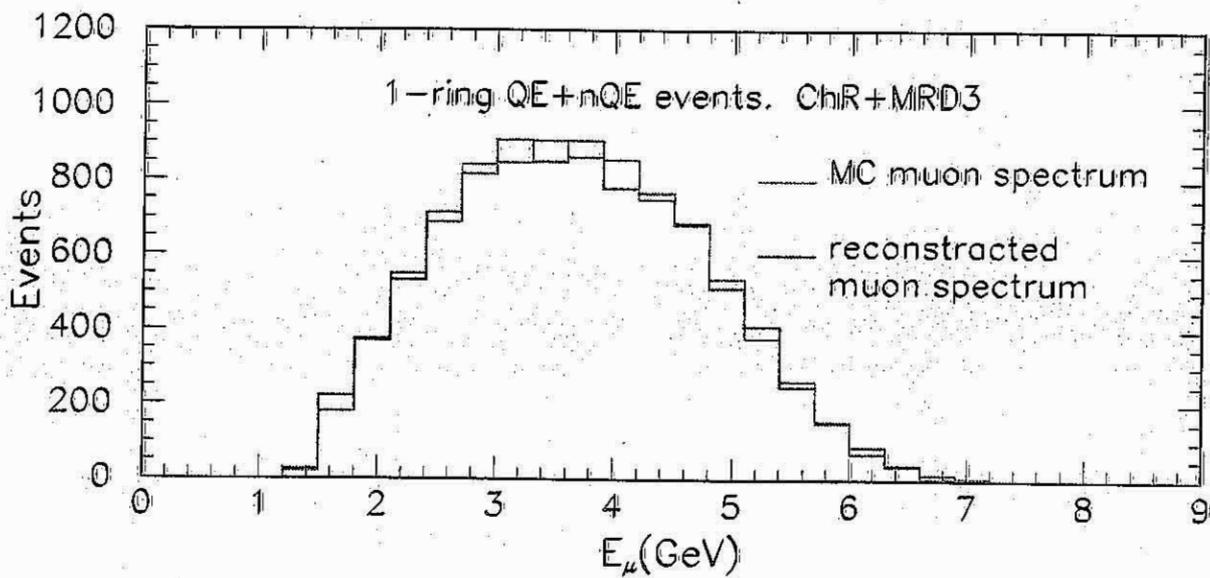
- Muon energy losed in MRD  $E_{MRD}$  is determined.
- Muon direction at the production is evaluated, using the reconstructed vertex position in the ChD and the coordinates of the hitted strips on stintillator planes between the ChD and MRD active regions
- Track length  $L_{ChD}$  in the ChD is calculated
- Energy losses in water  $dE/dX$  is fitted at the energy  $E_{MRD}$  and muon energy at the production is estimated as
$$E_i = E_{MRD} + (dE/dX)L_{ChD}.$$
- $dE/dX$  is fitted at energy  $E_f = \sqrt{E_i E_{MRD}}$  and the muon energy  $E_{i+1}$  is evaluated.
- Repeat until  $|E_{i+1} - E_i| / E_{i+1} \leq 0.01$











## Conclusion

### MRD

- Energy acceptance of MRD is from 0.12 GeV to 3.5 GeV with a resolution  $\sigma(\Delta E/E) = 5 \div 6\%$ .
- Track angular resolution  $\Delta\theta$  is about  $1^\circ \div 5^\circ$  and depends on muon energy.

### Cherenkov detector + MRD

- Energy acceptance is 1.2 ÷ 6.0 GeV with reconstruction efficiency  $\varepsilon \geq 0.5$  for events detected in active regions of the ChD and MRD.
- The efficiency will be  $\varepsilon \geq 0.75$  if an outer detector is used for detection of the stopping muons in the insensitive ChD region.
- Energy resolution  $\sigma(\Delta E/E) = 5 \div 6\%$
- Track angular resolution  $\Delta\theta$  decreases with energy from  $3^\circ$  up to  $0.7^\circ$

## Thanks!

First of all I'm thankful to Prof. T.Kajita, Prof. K.Kaneyuki and Dr. K.Okumura for useful discussions, their hearty hospitality and support during my stay in the RCCN of the Institute for Cosmic Ray Research.

I would like to thank ICCR Neutrino Center staff and students, for their friendly help and support.

I'm looking forward for fruitful collaborations in JPARC-nu experiment in the future