

Optical Follow-up Observations for IceCube-170922A

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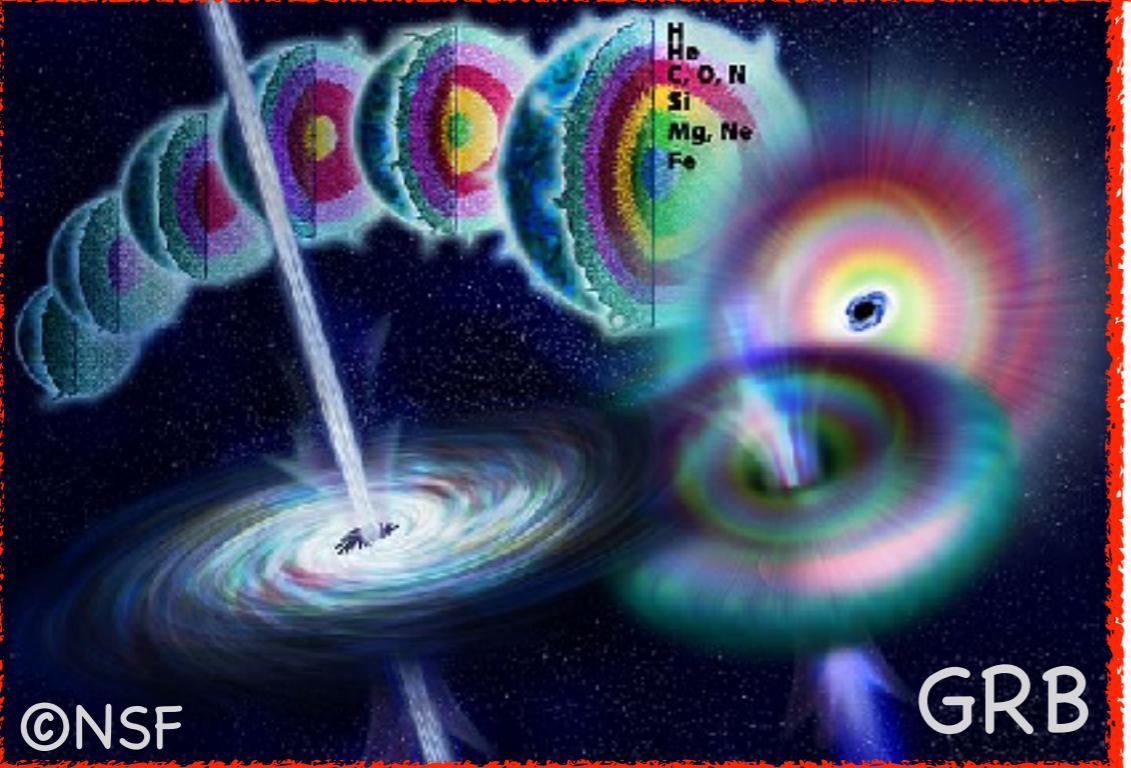
- IceCube neutrino source candidates
- optical observations in general
- our optical follow-up observation strategy for IceCube neutrino sources
- optical observations for IceCube-170922A
 - imaging observations
 - spectroscopic observations
 - polarization observations
- Summary

Origin of high-energy (TeV) neutrinos (high-energy cosmic ray)

transient (variable)

©U. Tokyo

(peculiar) supernova



©NSF

GRB

©NASA/Fermi

blazar: AGN relativistic jet



©NASA

starburst galaxy

optical/NIR telescopes

↑ aperture [m]

30

3x 30m telescopes in 2020s?

~10x 8-10m telescopes (largest now)

10

P200

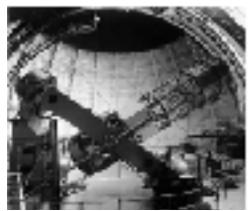


CFHT

Keck



Subaru



Hooker

P48



HST

PS1



E-ELT



TMT



GMT



LSST

now

year

©NSF
JTF Multi-messenger astronomy prompted by the neutrino alert IceCube-170922A

1900

1950

1990

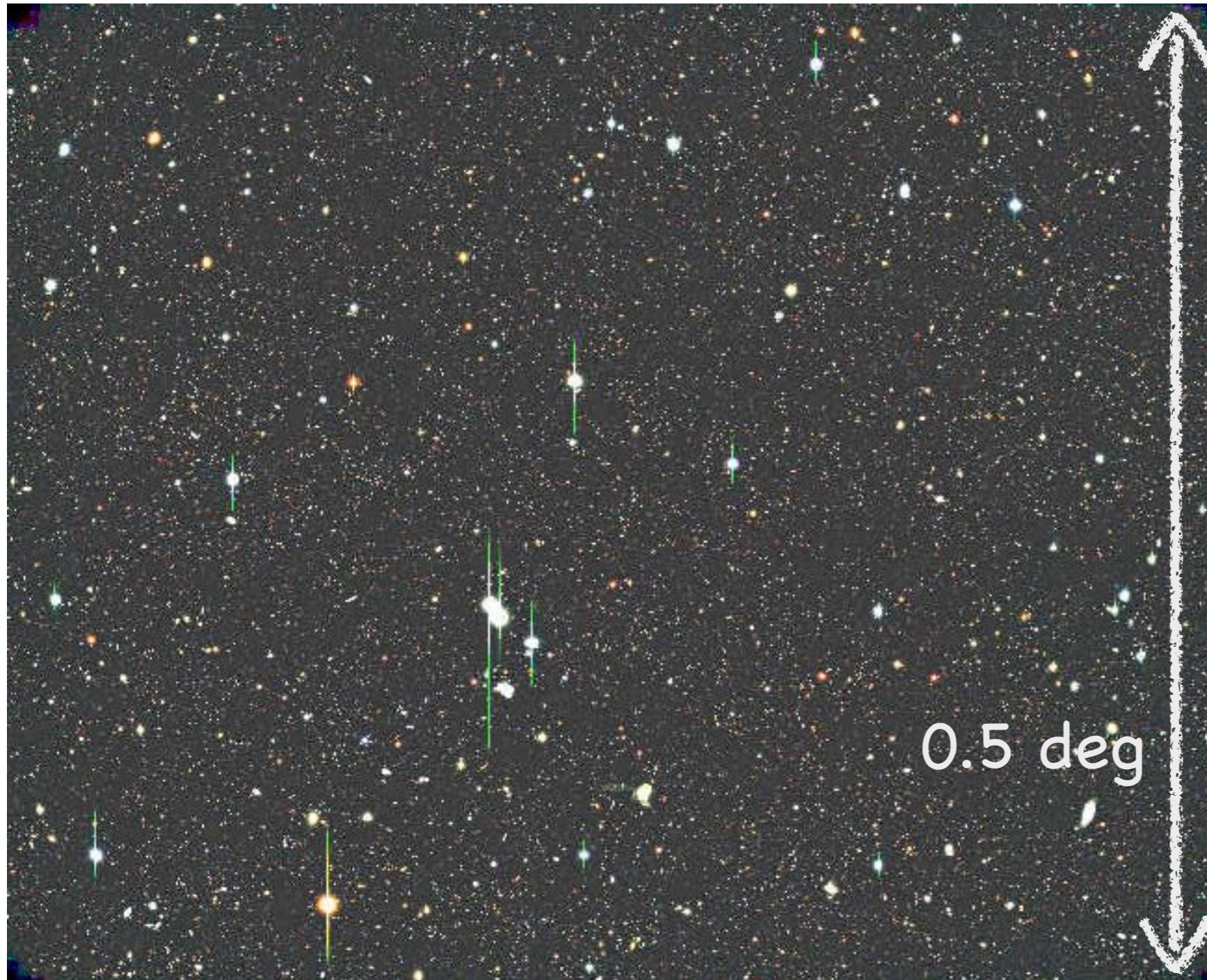
2000

2010

2020
2018/09/20

objects in optical wavelengths

- non-transient: $O(10^{5-6}) \text{ deg}^{-2}$ (Furusawa+2008)
- transient: $O(10^2) \text{ deg}^{-2}$ (TM+2008)



3-color image taken with
Subaru (8.2m)
Suprime-Cam

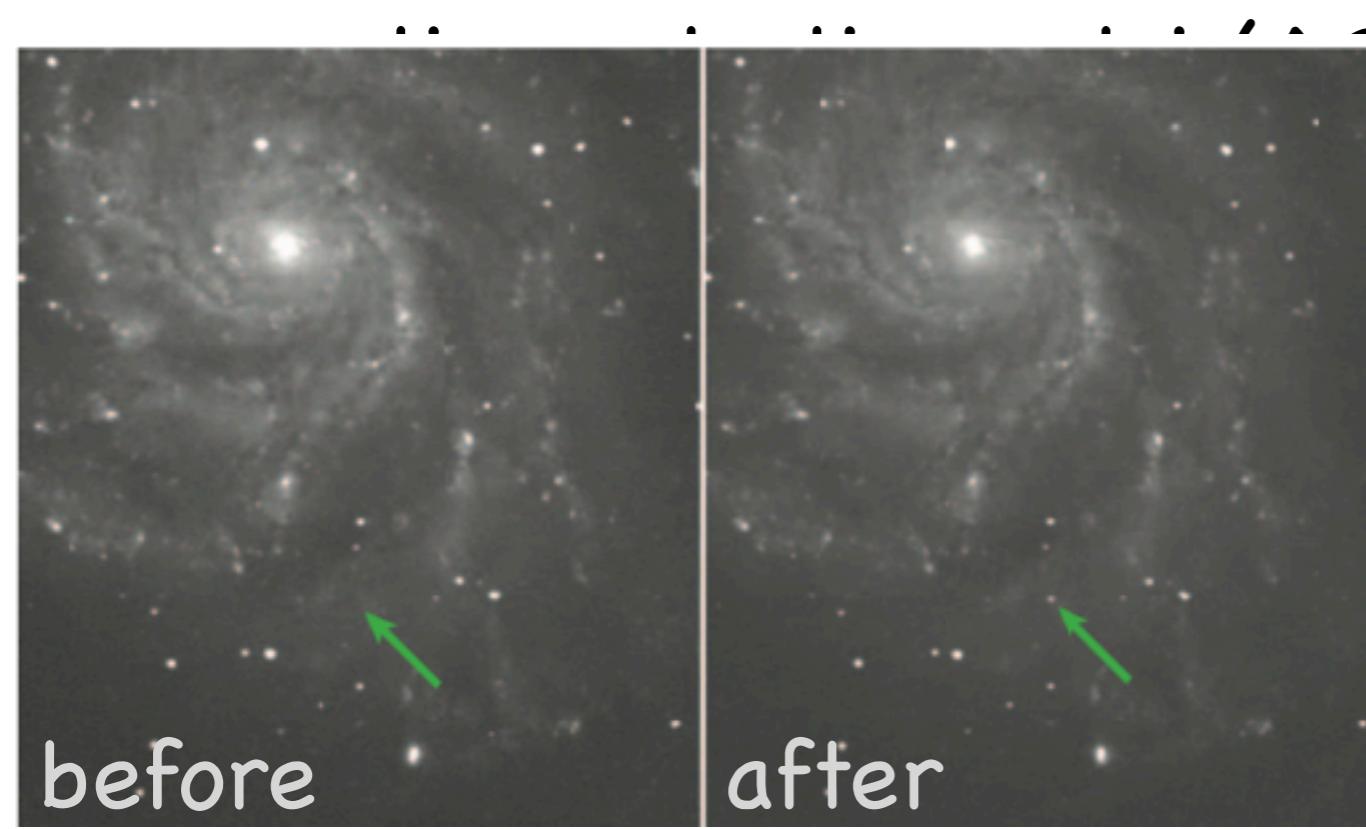
field-of-view: 0.25 deg^2
 $\Leftrightarrow 40,000 \text{ deg}^2$ (all-sky)

objects in optical wavelengths

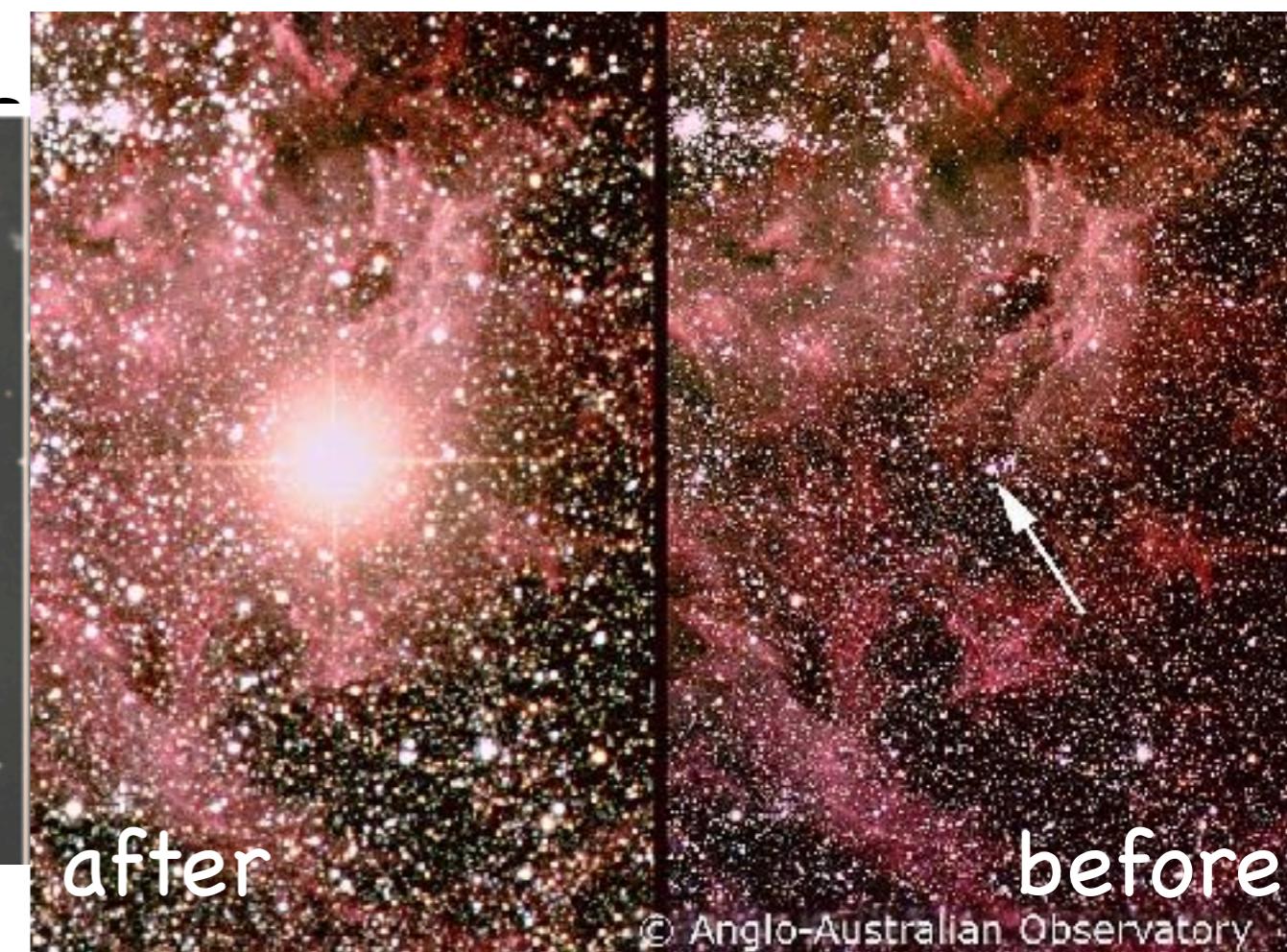
- non-transient: $O(10^{5-6}) \text{ deg}^{-2}$ (Furusawa+2008)
- transient: $O(10^2) \text{ deg}^{-2}$ (TM+2008)
 - variable star
 - [high proper motion star]
 - [asteroid]
 - supernova (SN)
 - active galactic nuclei (AGN)
 - gamma-ray burst (GRB)
 - fast radio burst (FRB): Parkes, ...
 - Gravitational Wave source (kilonova)
 - neutrino source

objects in optical wavelengths

- ☐ non-transient: $O(10^{5-6})$ deg $^{-2}$ (Furusawa+2008)
 - ☐ transient: $O(10^2)$ deg $^{-2}$ (TM+2008)
 - ☐ variable star
 - ☐ [high proper motion star]
 - ☐ [asteroid]
 - ☐ supernova (SN)



Nugent+2011, SN 2011fe@M101

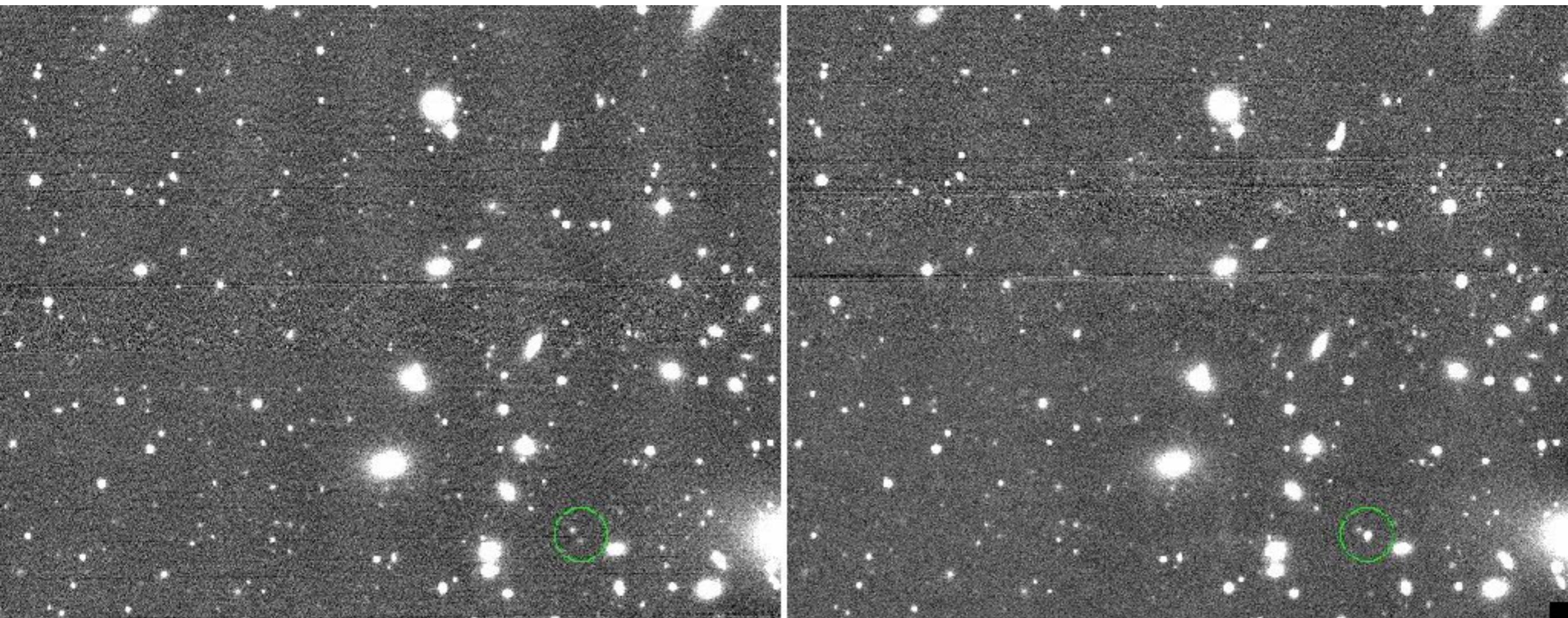


© Anglo-Australian Observatory

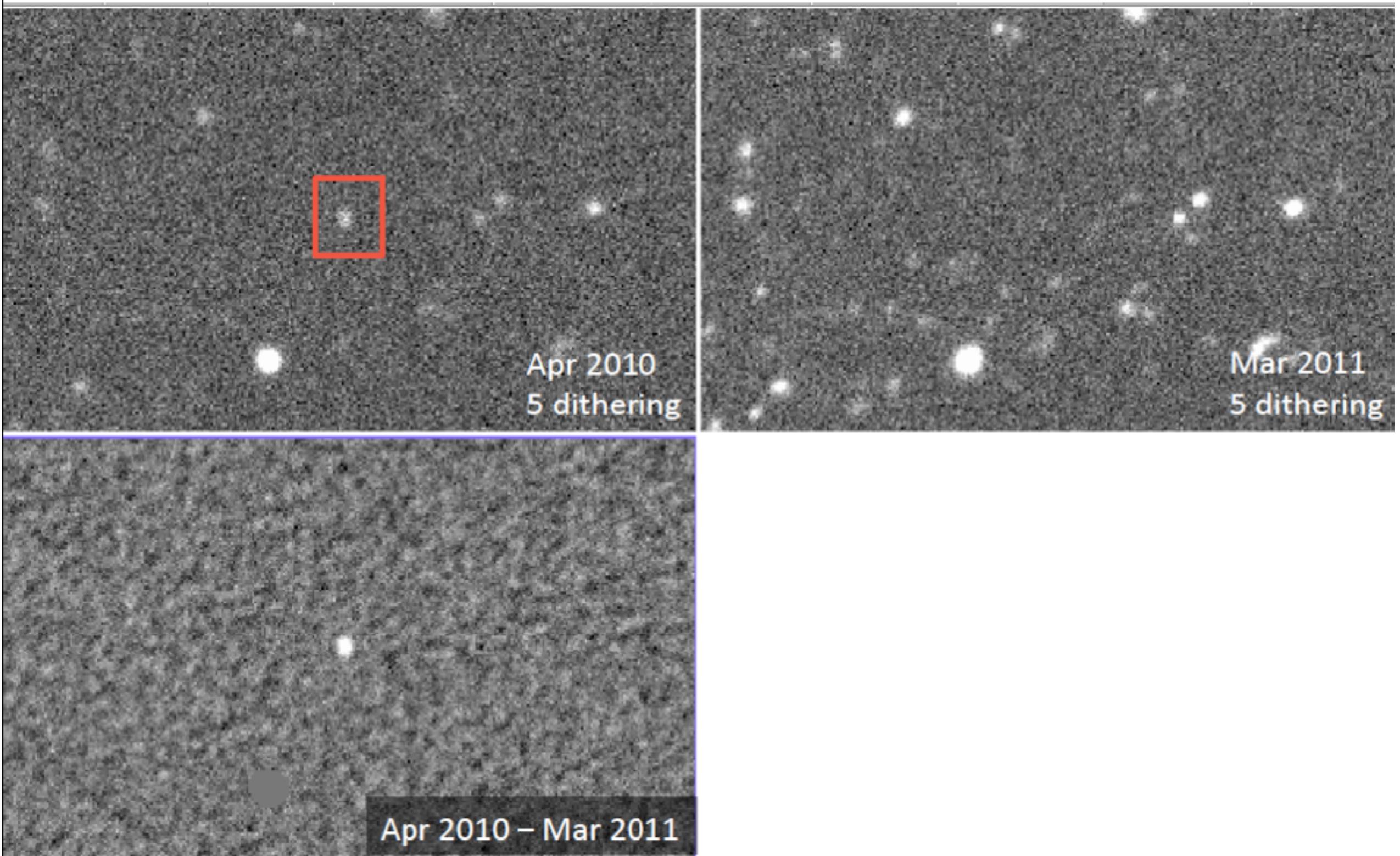
!!! Let's find a new supernova !!!

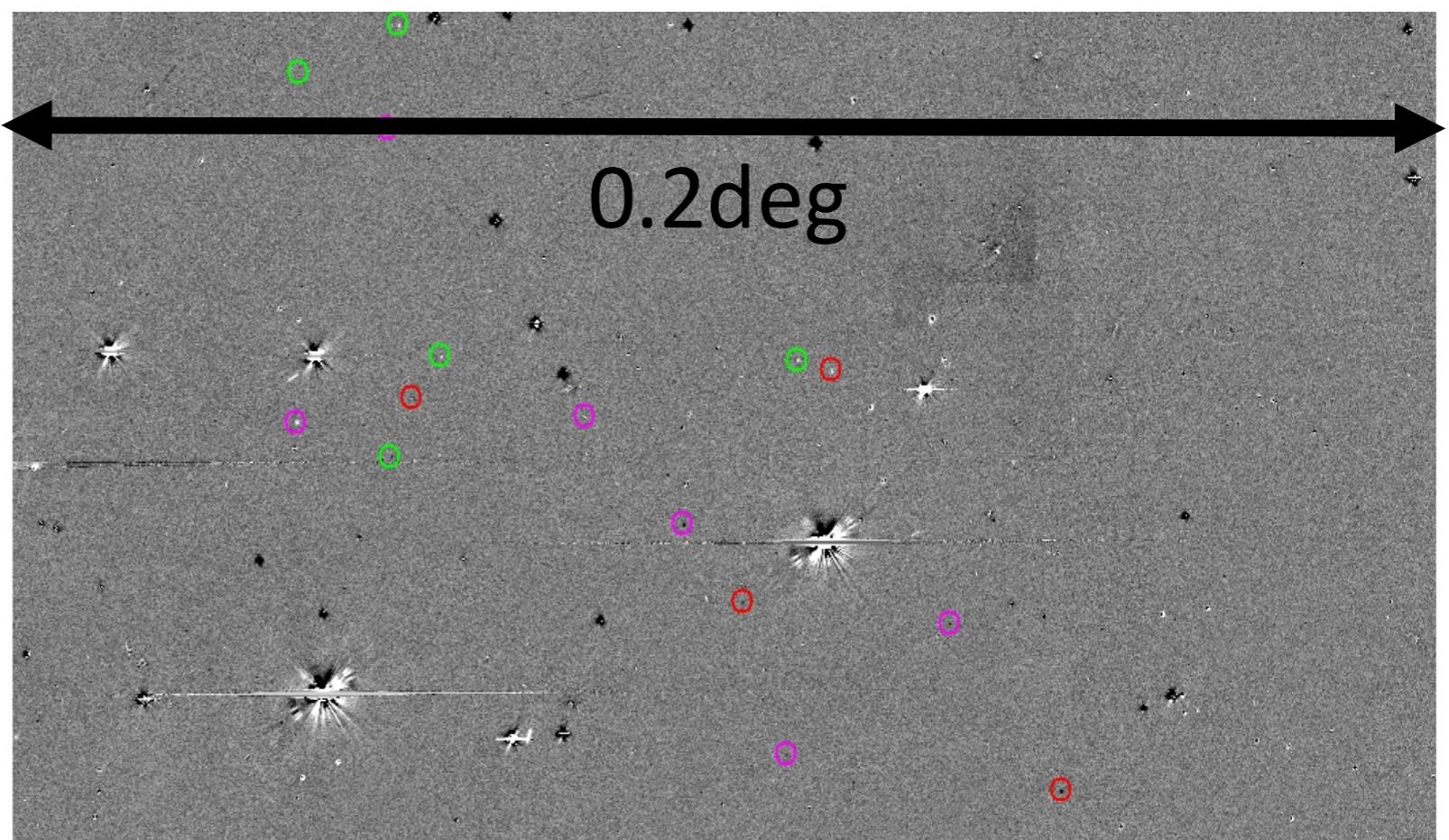
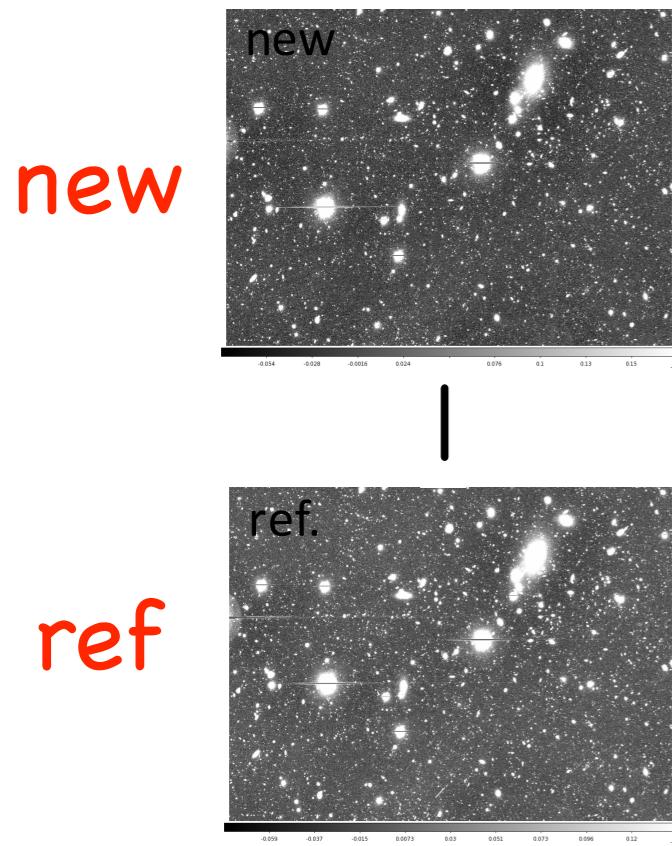
before

after



Example: image subtraction

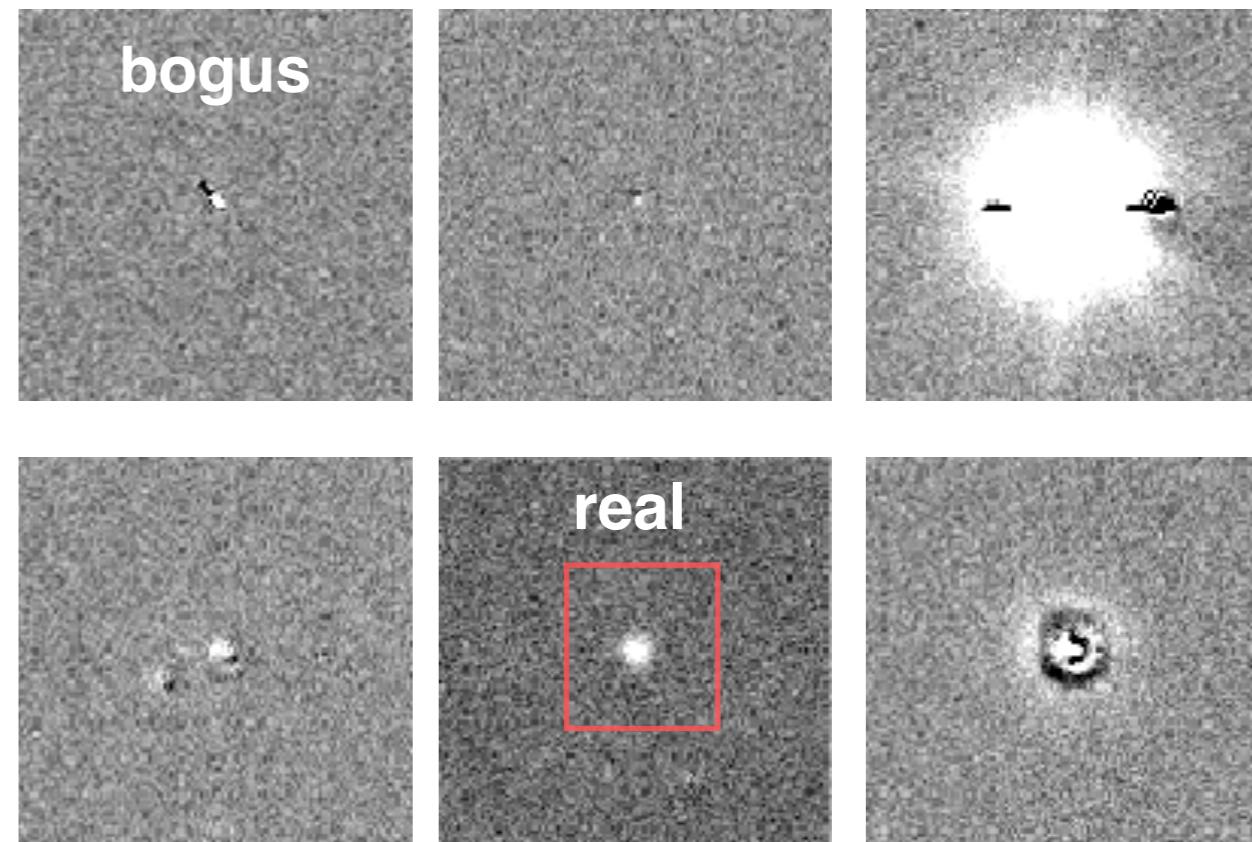




subtracted image

real vs bogus
 $\sim 1 : 1000$

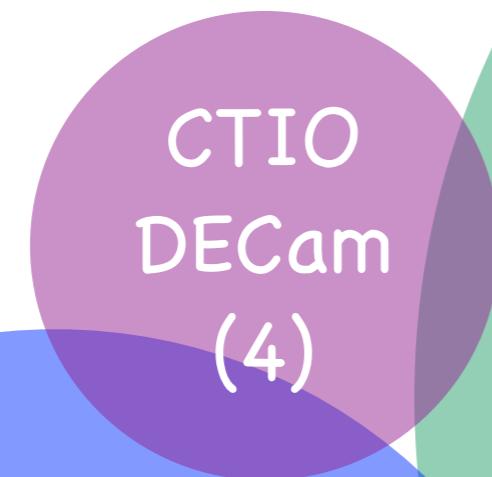
Image subtraction is not perfect,
but this ratio is improved by
machine learning technique (e.g.,
Morii+2016).



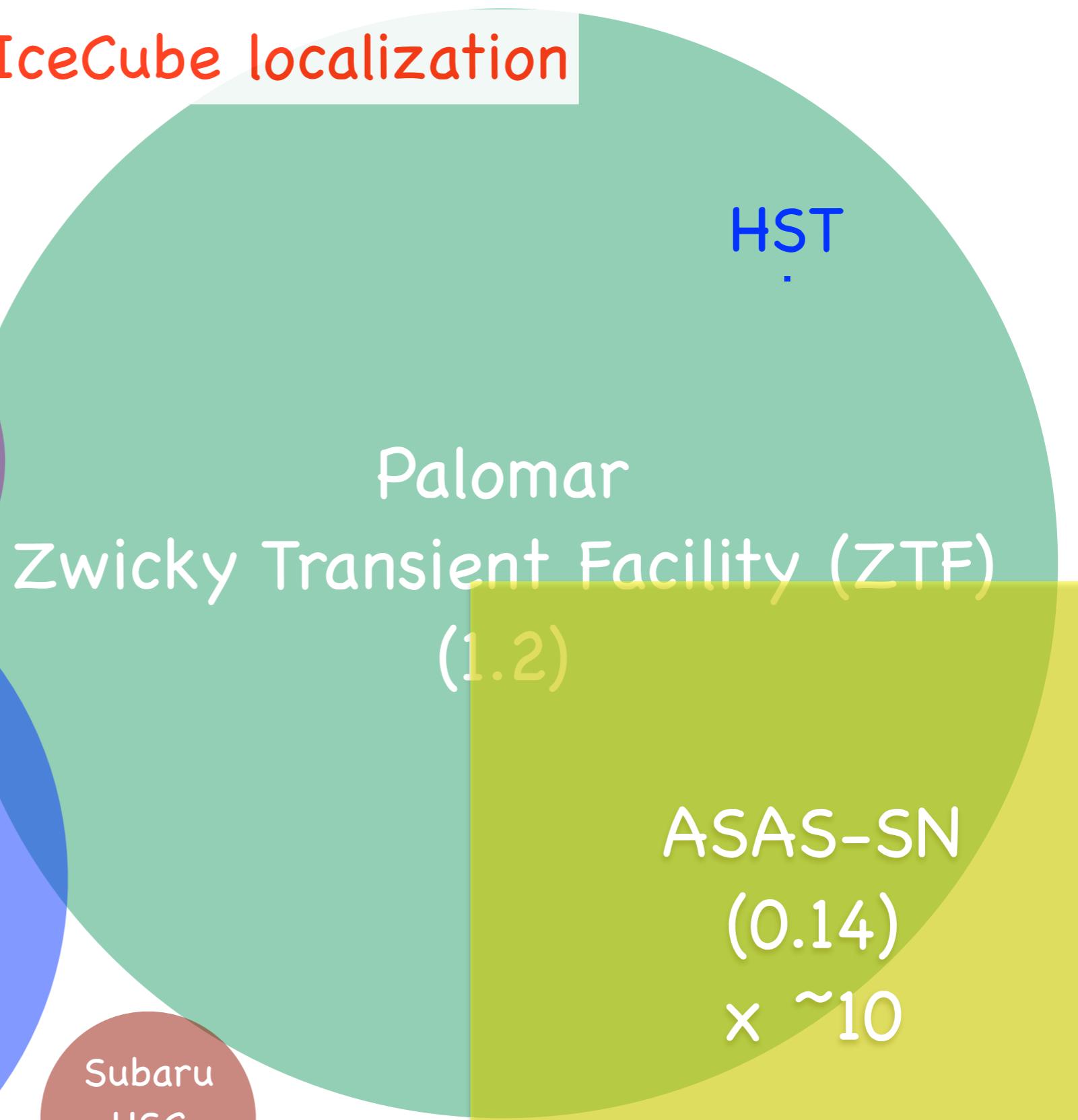
Field-of-View (FoV) of optical/NIR telescopes/instruments

typically ~ 0.1 deg \ll IceCube localization

>1 deg (wide-field)



Kiso Schmidt
Tomo-e Gozen
(1.05)

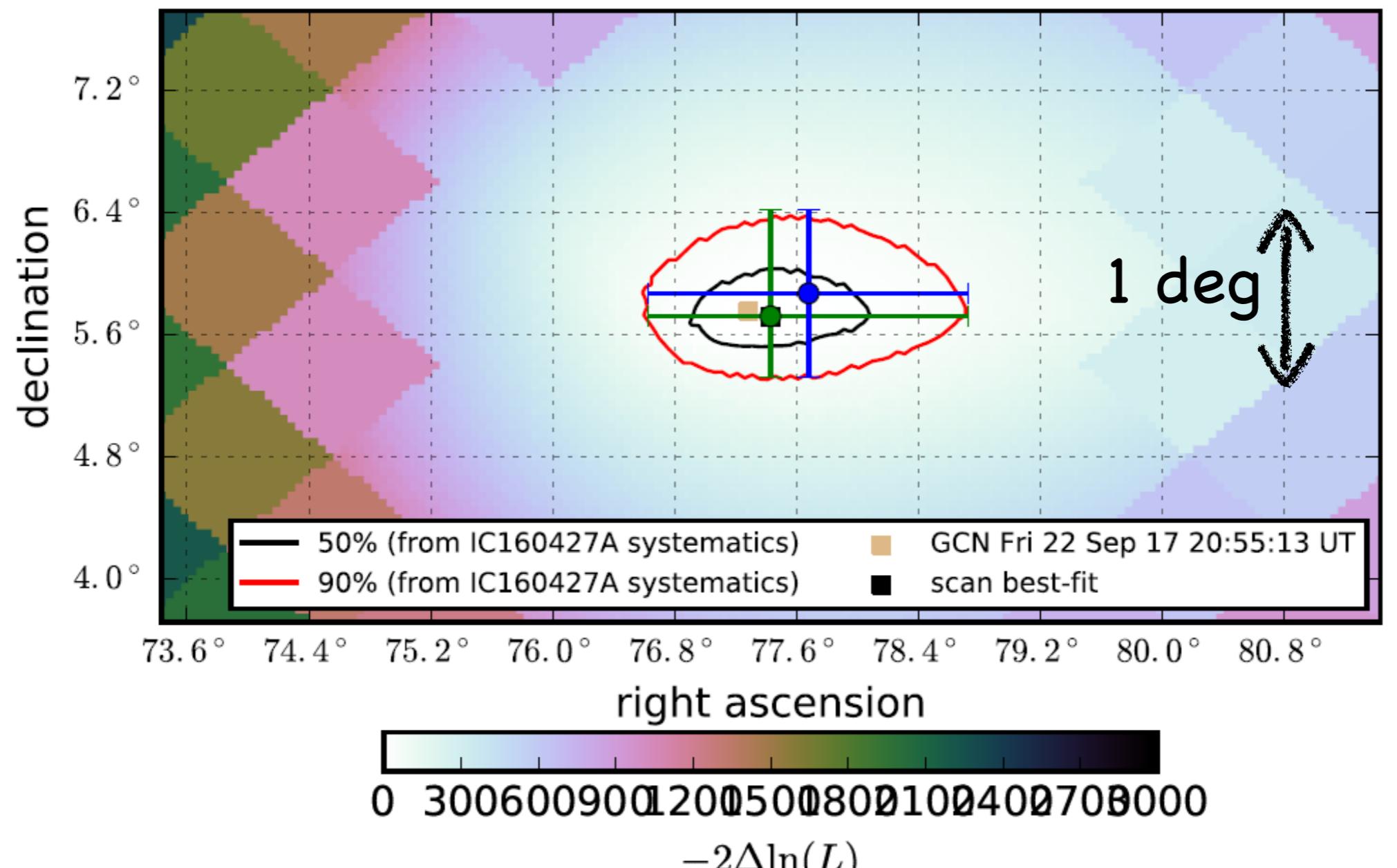


Subaru
HSC
(8.2)

EM Counterpart Search of IceCube events

- ~0.5-0.8 deg localization
- O(10^3) transients@optical (TM+2008, M. Yoshida+2017)

Run: 130033 Event 50579430: Type: EHE MJD: 58018.8711856



Origin of high-energy (TeV) neutrinos (high-energy cosmic ray)

transient (variable)

©U. Tokyo



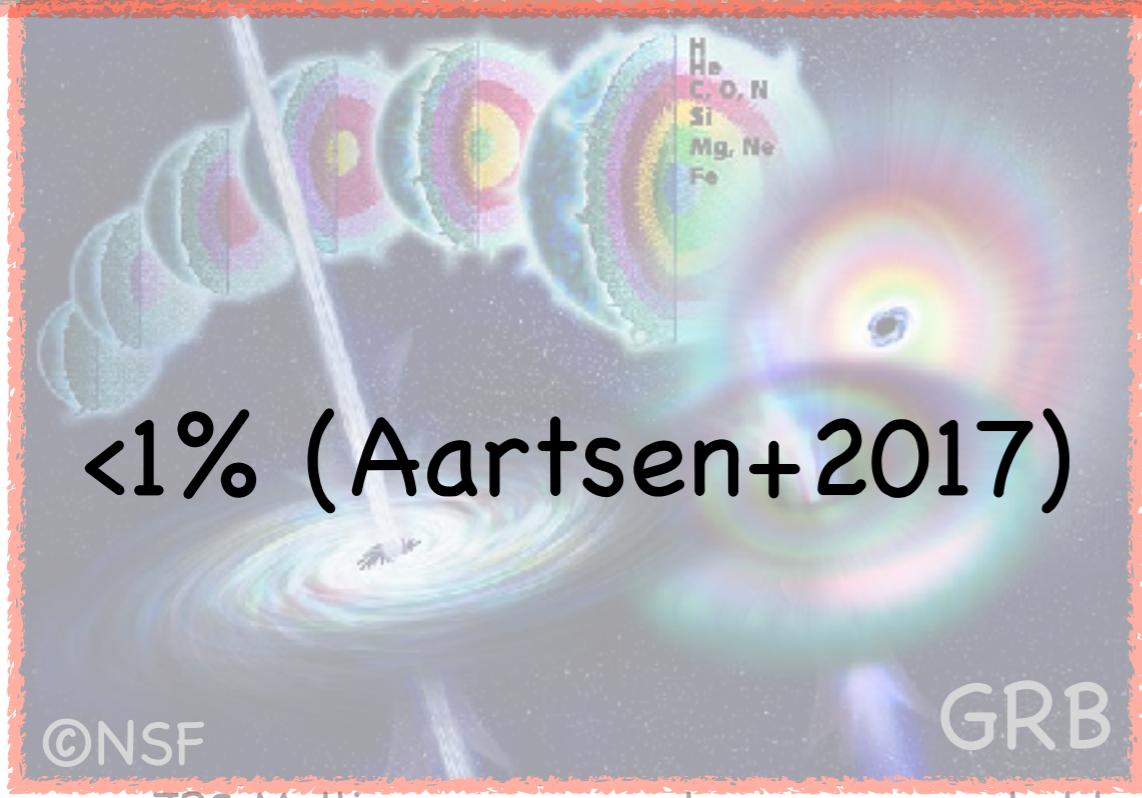
(peculiar) supernova

~ $<27\%$ (Fermi)
(Aartsen+2017)

©NASA/Fermi



blazar: AGN relativistic jet



<1% (Aartsen+2017)

©NSF

GRB

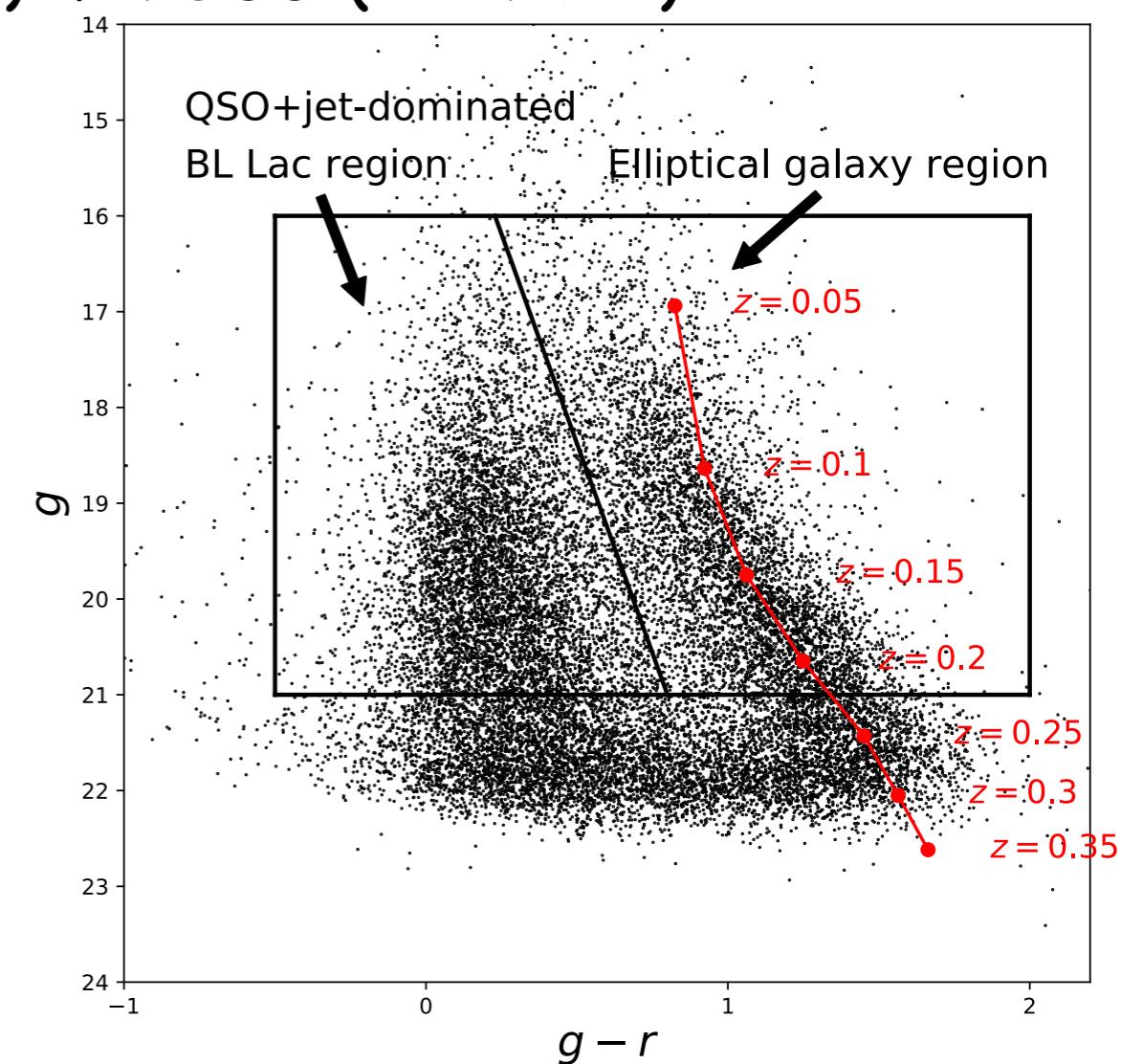
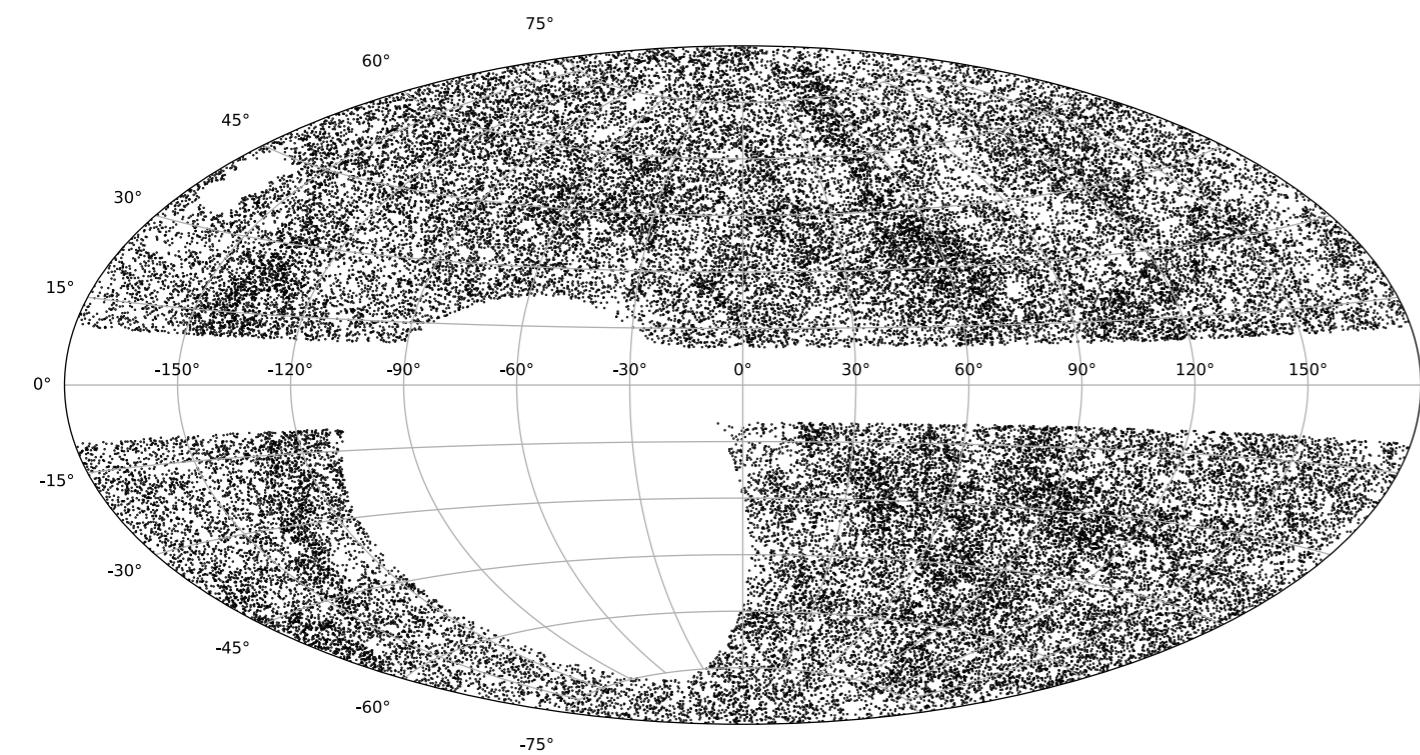
©NASA

~ $<30\%$ @100 TeV
(Bechtol+2017)

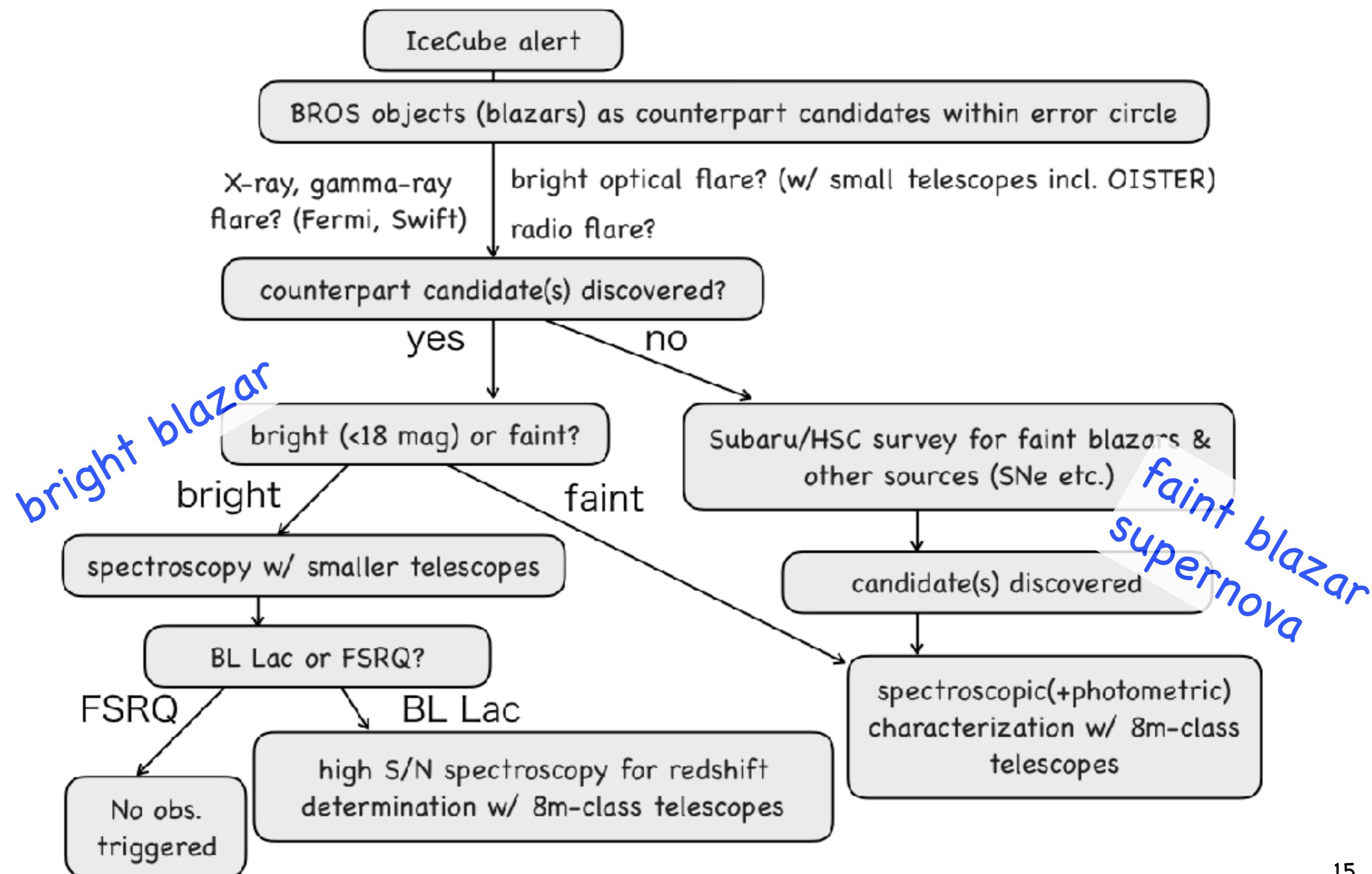
starburst galaxy

New Blazar Catalog (BROS)

- Blazar Radio and Optical Survey (BROS; Itoh et al. in prep.)
- 56,315 sources at Dec.>-40 deg
 - ROMA-BZCAT: 3,516 sources
 - CRATES: ~11,000 sources
 - 3FGL (Fermi): ~1,500 sources
- flat-spectrum@radio: NVSS (1.4 GHz) + TGSS (151 MHz)
- Pan-STARRSS(PS1)@optical
 - ~40% not detected in PS1 ($r > 23$)



Counterpart Search Strategy@optical



OISTER Collaboration: Japanese Optical Telescopes

[name] aperture



Pirka 1.3
(Hokkaido)



Kiso 1.05 (Tokyo)



Kyoto 0.4 ==> 3.8



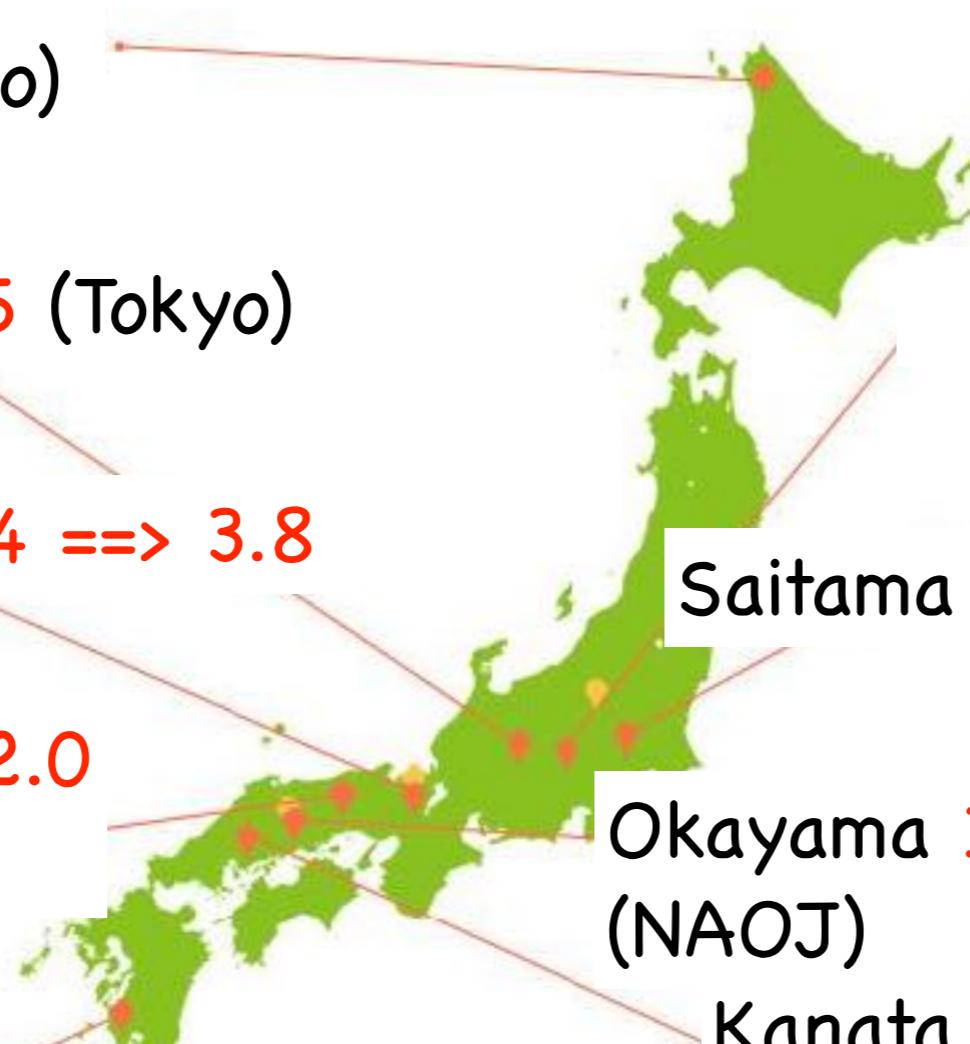
Nayuta 2.0
(Hyogo)



Kagoshima 1.0



MITSuME 1.0
(Ishigaki)



協力機関
群馬県 県立ぐんま天文台
京都産業大学 神山天文台
日本スペースガード協会 美星スペースガードセンター

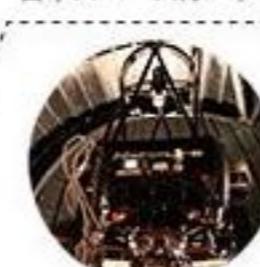
MITSuME 0.5
(Titech)



Okayama 1.88
(NAOJ)



Kanata 1.5
(Hiroshima)



TAO 0.5
==> 6.5
(Tokyo)
@Chile



IRSF 1.4
(Nagoya)
@South Africa



Subaru 8.2
@Hawaii, USA

+@Tibet, China

+@NZ

...

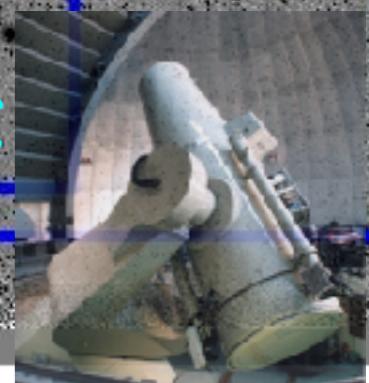
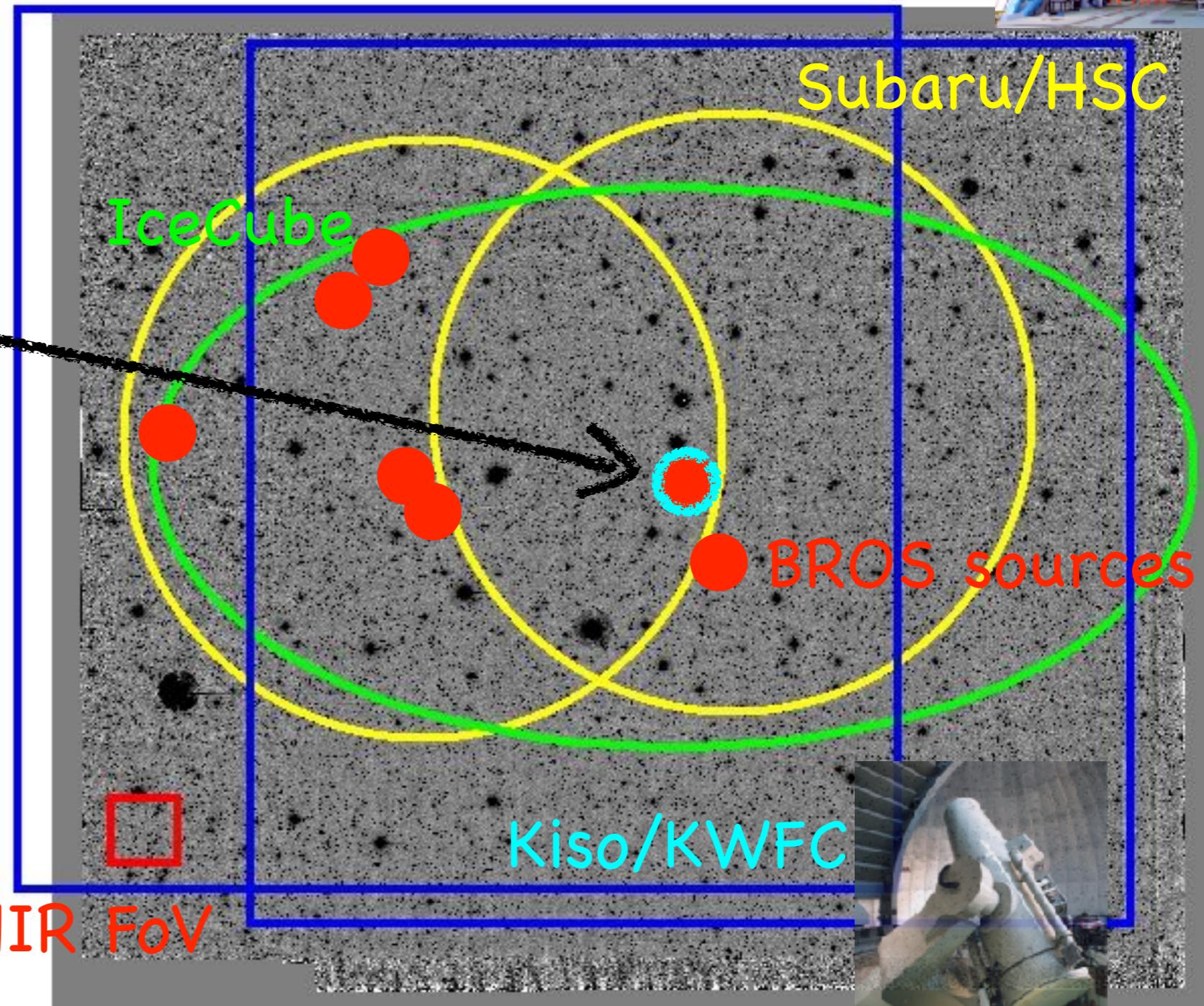
IceCube-170922A case

- 7 BROS sources (blazar candidates) incl. 1 Fermi blazar.
- We observed all these blazars w/ Subaru/Hyper Suprime-Cam (HSC) & Kiso/KWFC.
- covered almost the entire error region.



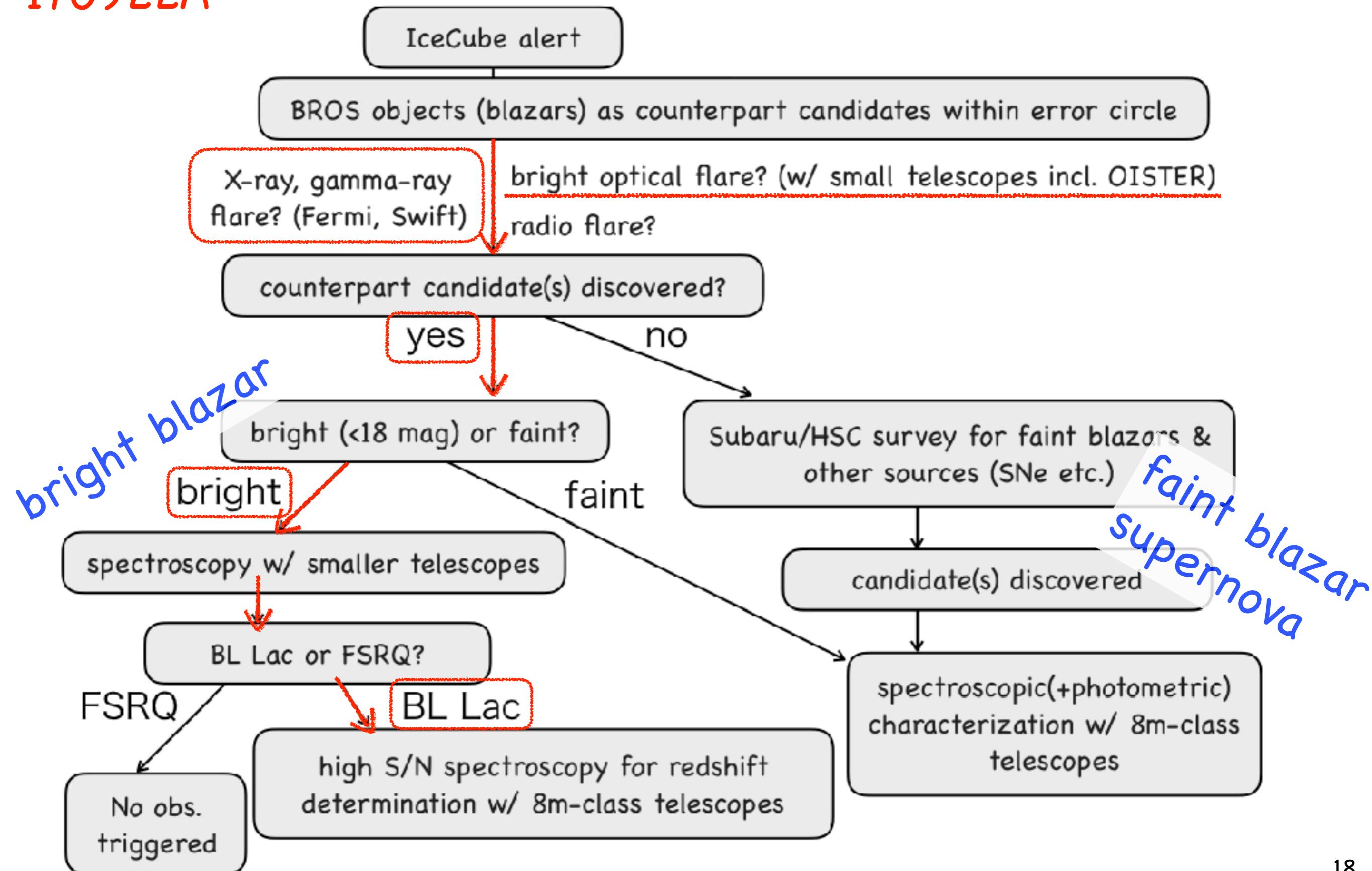
TXS 0506+056

Kanata/HONIR FoV



Counterpart Search Strategy@optical

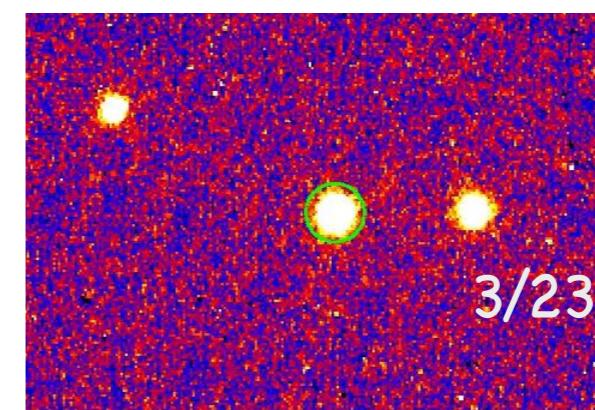
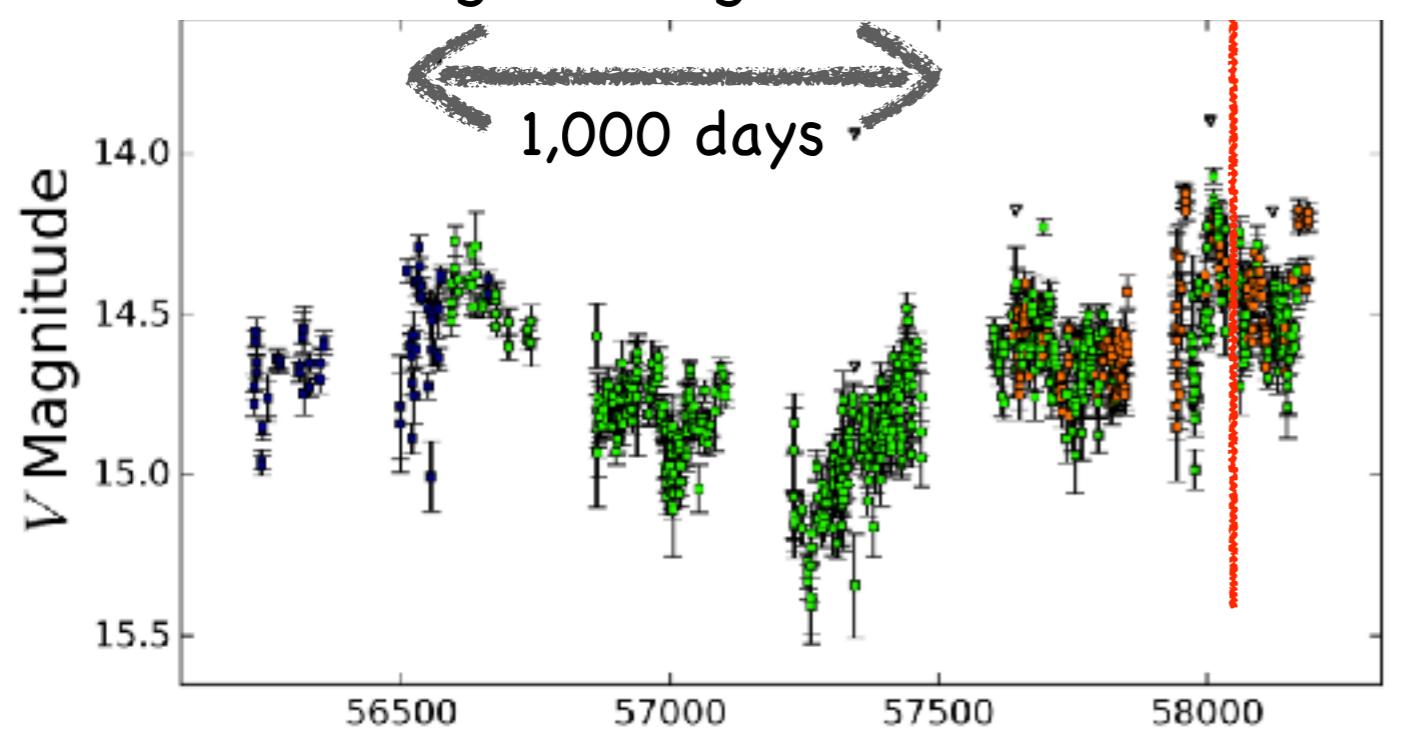
170922A



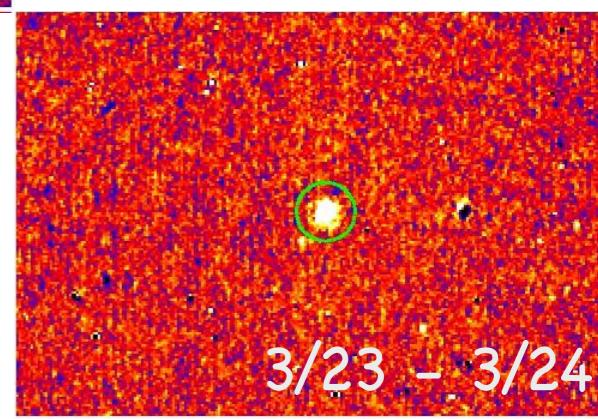
Road to EM Counterpart Discovery of IceCube-170922A

- IceCube alert (GCN 21916): 2017/09/22, 20:54:30 (UT)
- 7 BROS sources within IceCube-170922A error region
 - TXS 0506+056 variability detected with Kanata/HONIR on 2017/09/24
 - Y. Tanaka & TM talked during Subaru ToO observations at remote observing room and noticed LAT variability... (no special care about Subaru...)
- found Fermi/LAT (gamma, ATel #10791, Tanaka+), ASAS-SN (opt) variabilities
==> multi-wavelength follow-up
 - optical/NIR **imaging, spectroscopy, polarization**: incl. **MITSuME, Kiso, Nayuta, Kanata, IRSF (OISTER)** + **Subaru** (TM+ in prep.)
- blue-when-brighter trend: not special
- no detailed studies yet about variability behavior

ASAS-SN long-term light curve (V-band, 0.5um)



Kanata/HONIR
J-band

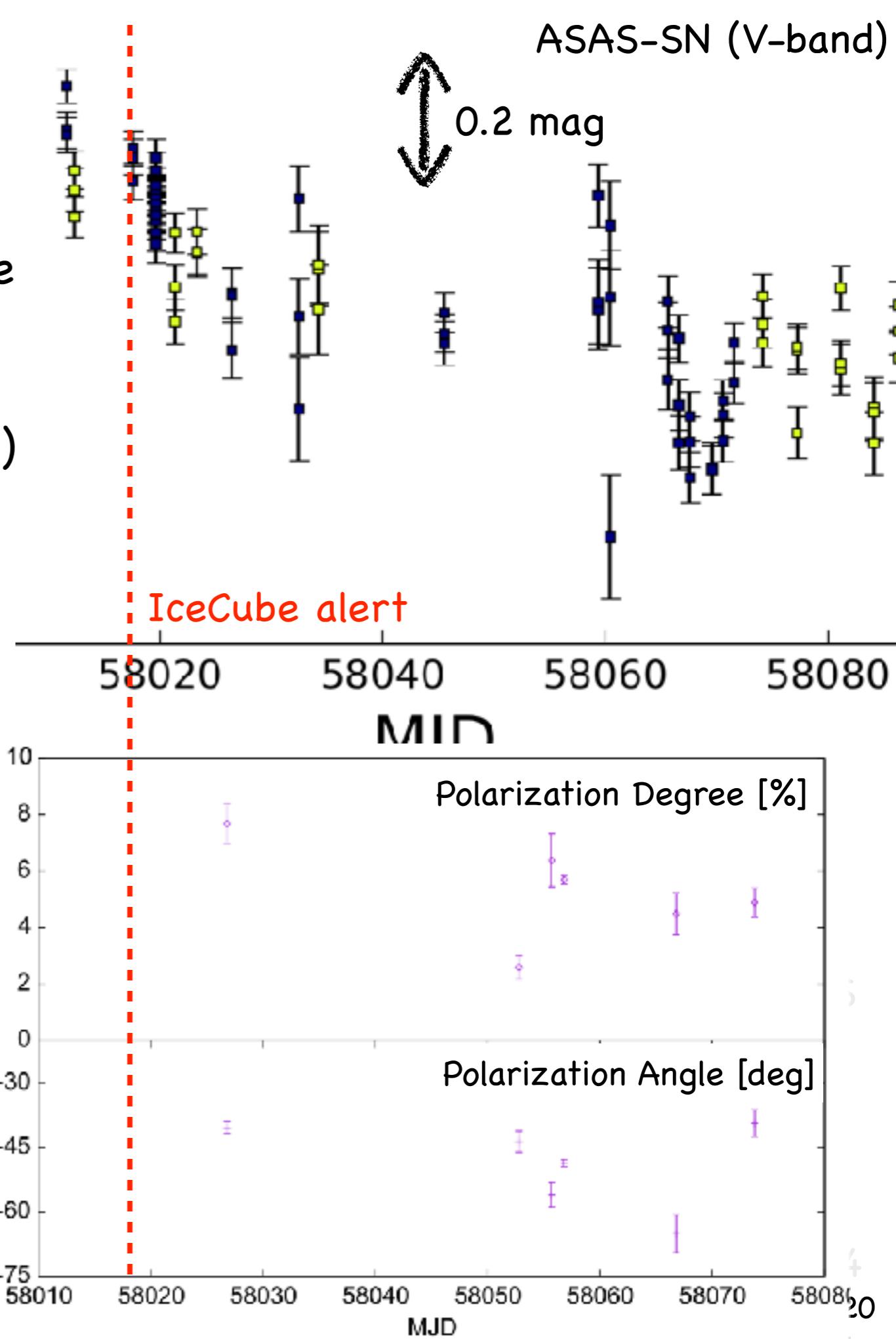


linear polarization of TXS 0506+056

- evidence for highly-ordered large-scale magnetic field (Falomo 2014)
- Kanata/HONIR in R, J-bands
 - 6-8% (Yamanaka+2017, ATel #10844)
 - interstellar pol. < 1%
 - not so variable in time
 - consistent w/ BL Lac
- ~14% polarization measured with Liverpool/RINGO3 recently reported (Steele+2018, ATel #11430)

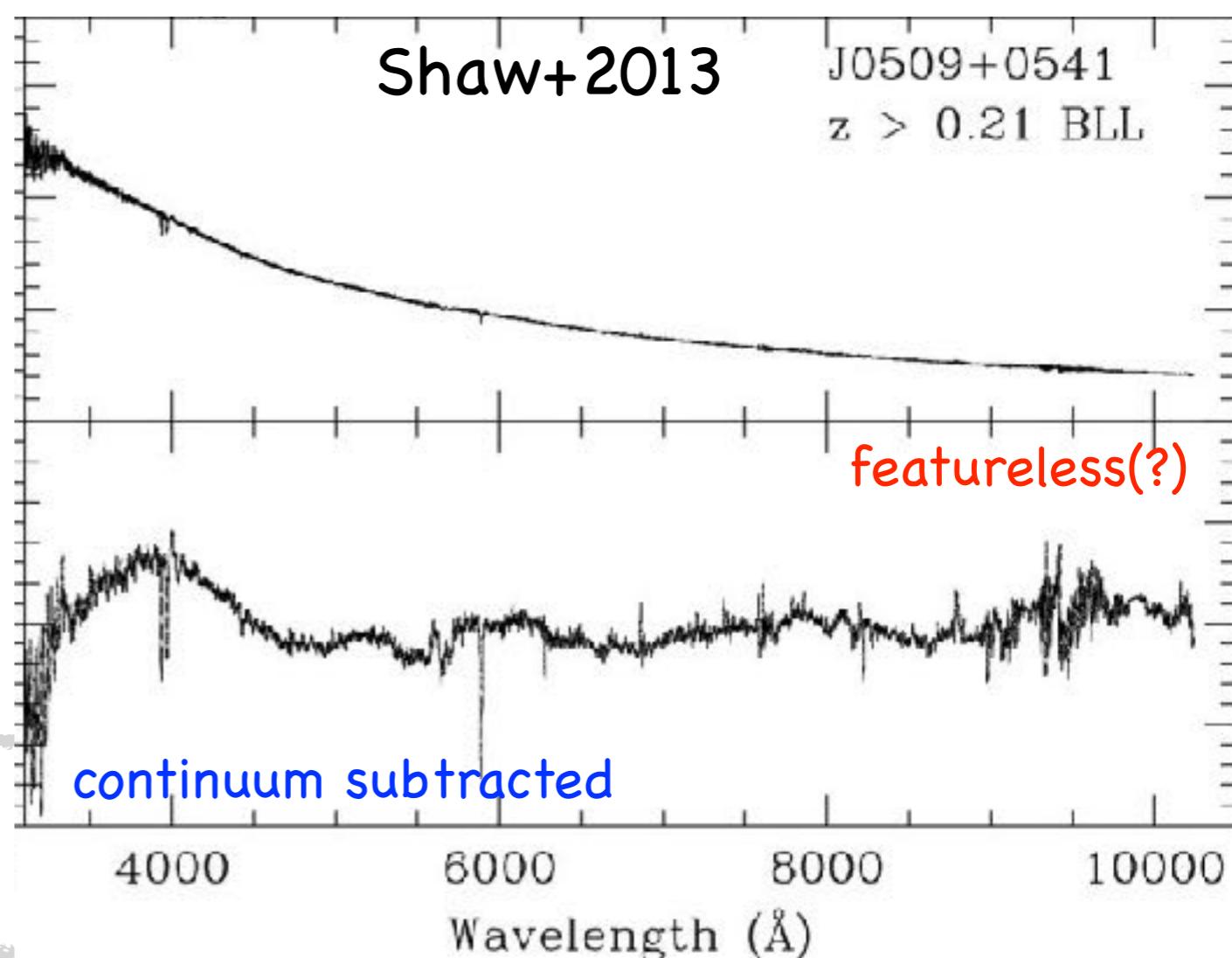
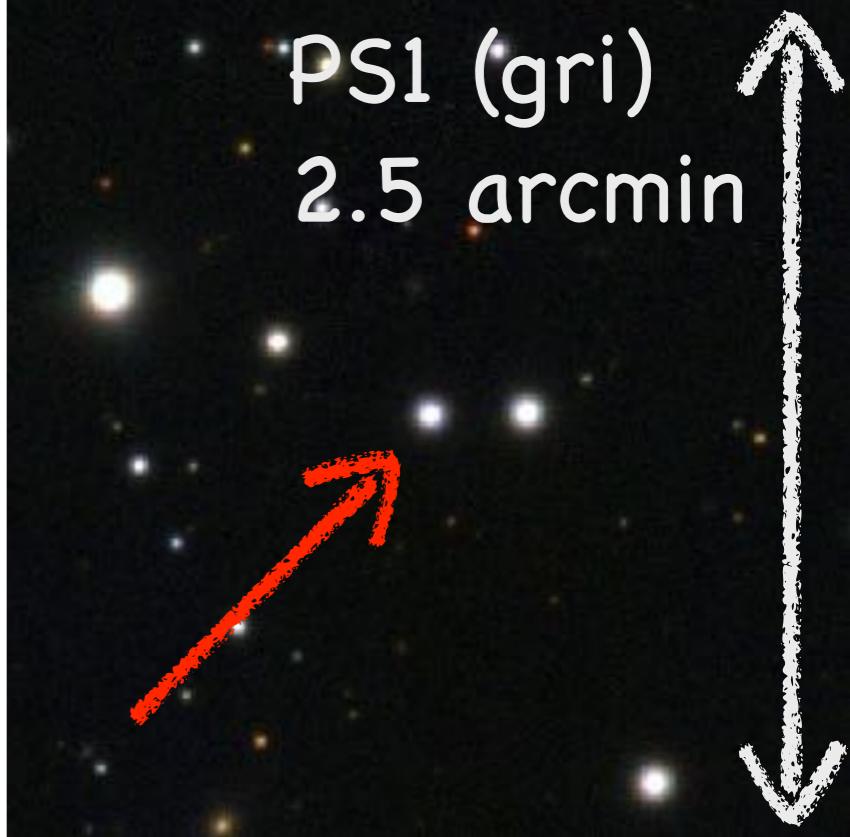


JPS Multi-messenger astronomy prompted by the



redshift (distance) of TXS 0506+056

- BL Lac
 - $z=0.336$ (Ajello+2014)
 - $z>0.21$ (Shaw+2013)
 - no reliable determination?
 - 8m-class telescopes are necessary "in general". (e.g., Landoni+2015)
 - very strong continuum (by relativistic jets)
- new observations
 - Kanata (1.5m)
 - Nayuta (2m)
 - Liverpool (2m)
 - Subaru/FOCAS (8.2m)
 - Gemini-N/GMOS (8.2m)
 - VLT/X-SHOOTER (8.2m)
 - SALT/HRS (10m)
 - GTC/OSIRIS (10m)

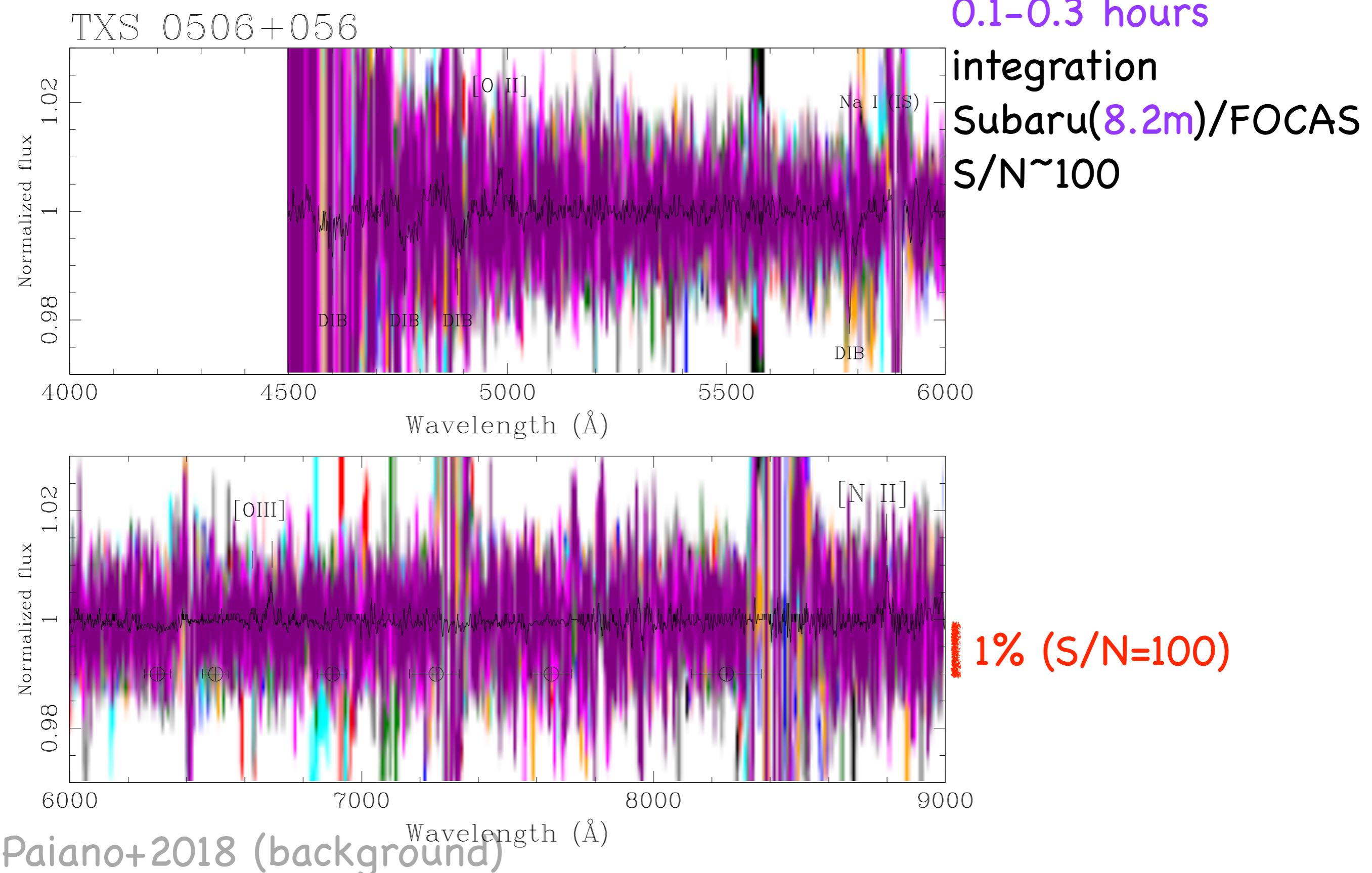


Ajello+2014

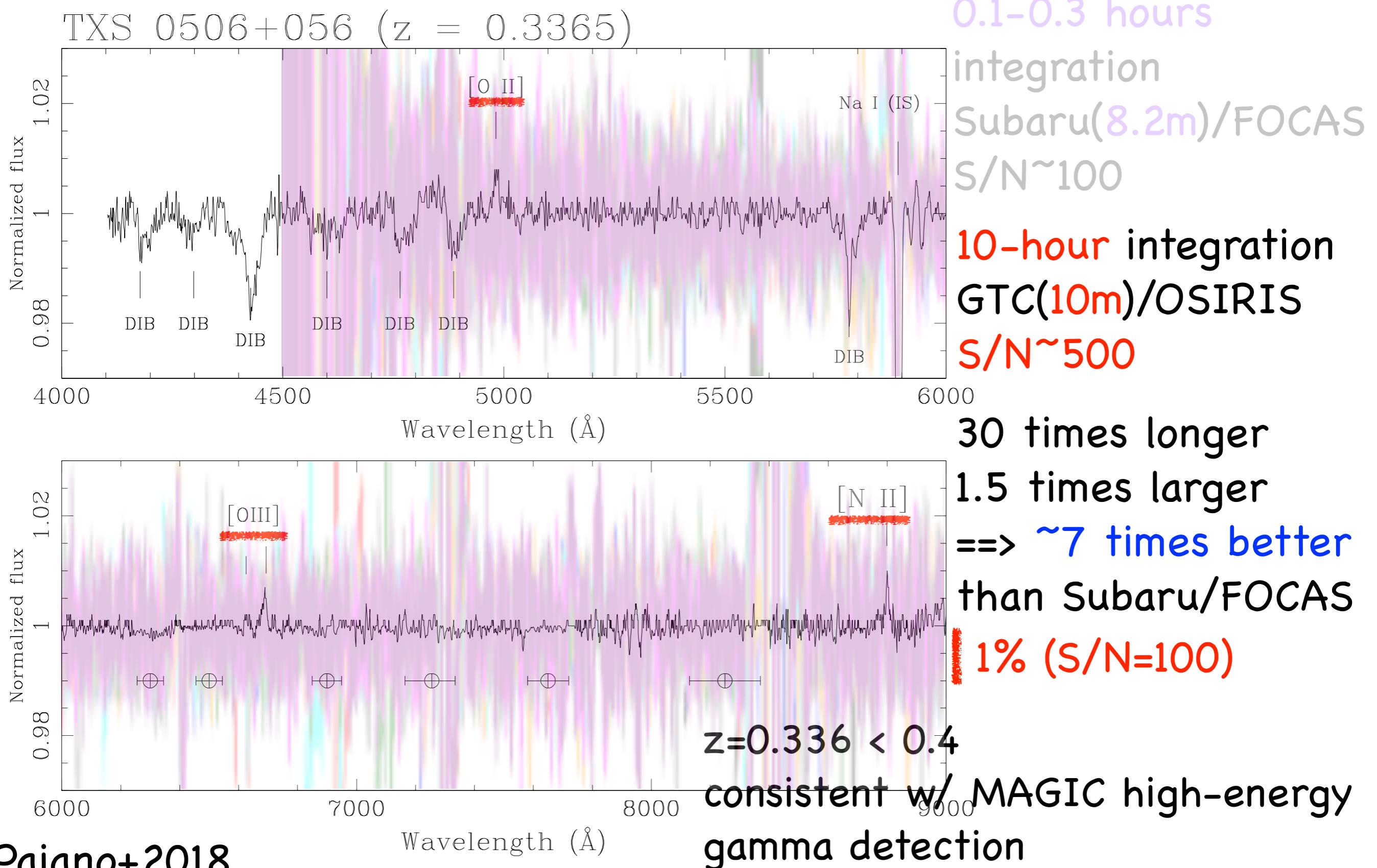
Multi-messenger astronomy prompted by the neutrino alert December 17, 2013

2018, 05, 25

Optical Spectroscopy: redshift determination

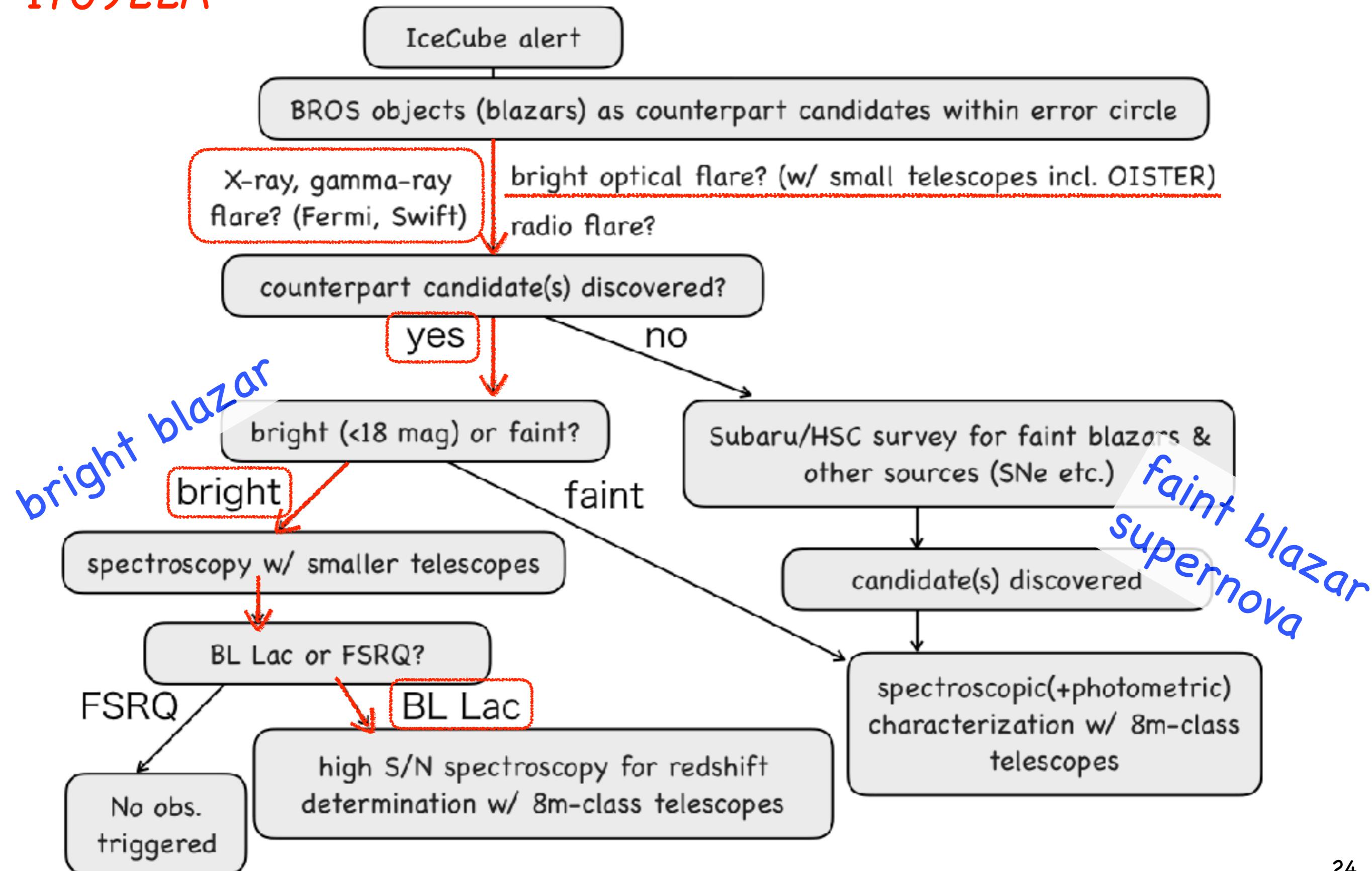


Optical Spectroscopy: redshift determination



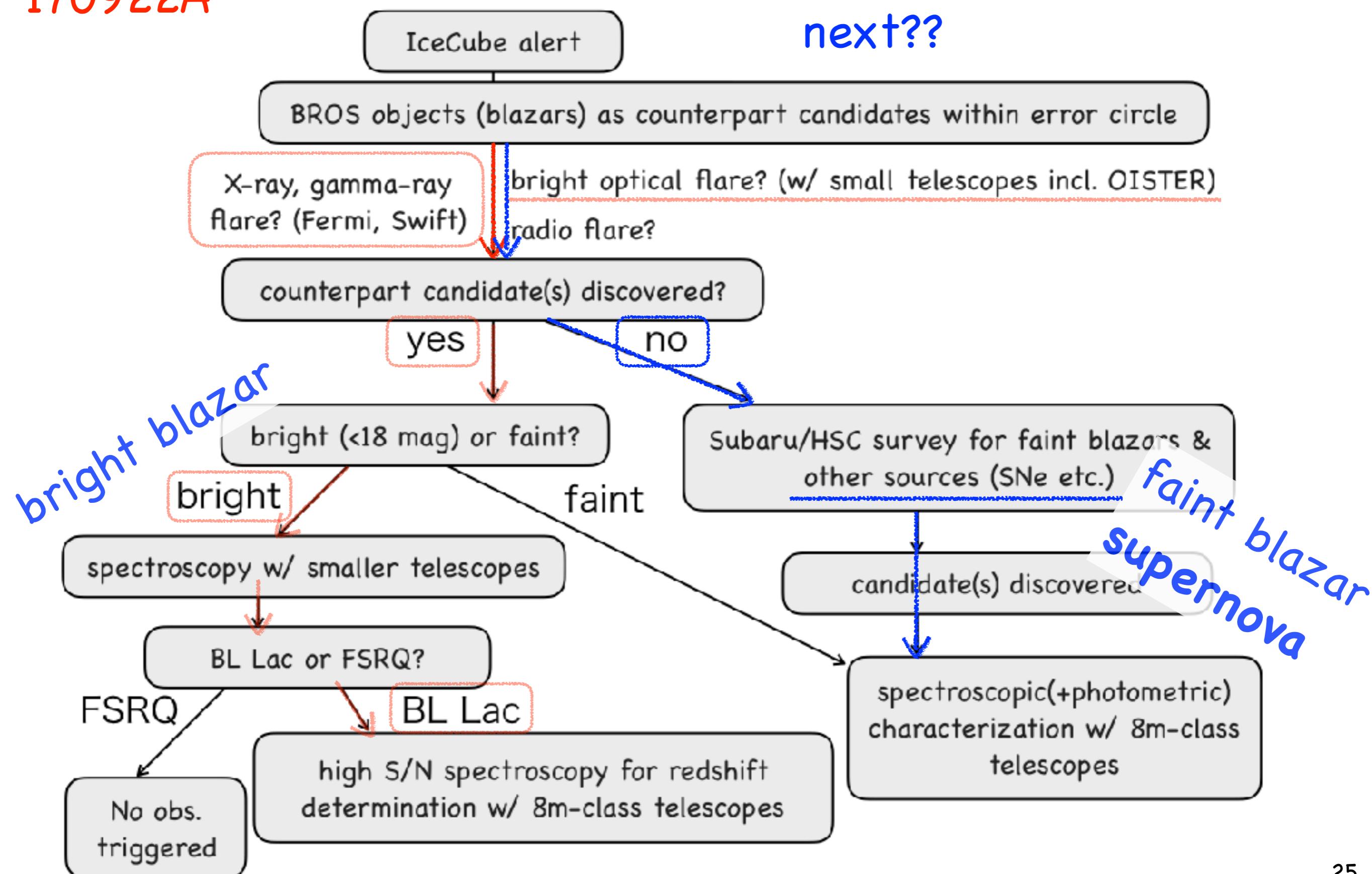
Counterpart Search Strategy@optical

170922A



Counterpart Search Strategy@optical

170922A



Summary & Future Prospects

- Several mechanisms proposed for high-energy neutrinos.
- We developed a new blazar catalog (BROS) incl. \sim 56,000 sources (increased by a factor of \sim 5).
- Our Kanata quick follow-up for BROS (flat-spectrum radio) sources found variability of TXS 0506+056.
==> Fermi/LAT gamma-ray variability ==> further multi-wavelength follow-ups
- Spatial/Temporal coincidence indicates the origin of IceCube-170922A is a blazar (BL Lac), TXS 0506+056.
 - showed optical flare around the Icecube detection.
 - not a special blazar at a glance? (need more examination)
- 10-hour integration w/ 10m telescope GTC determines the redshift $z=0.336$ ($S/N \sim$ several hundreds) of the BL Lac object.
- Future IceCube alerts: blazar again? supernova? need multiple/coordinated observing strategies w/ small (OISTER) & large (Subaru etc.) telescopes