

強重力場からの重力波による 様々な重力理論の検証

Probing Various Gravitational Theories with
Strong-field Gravitational-wave Observations

Kent Yagi

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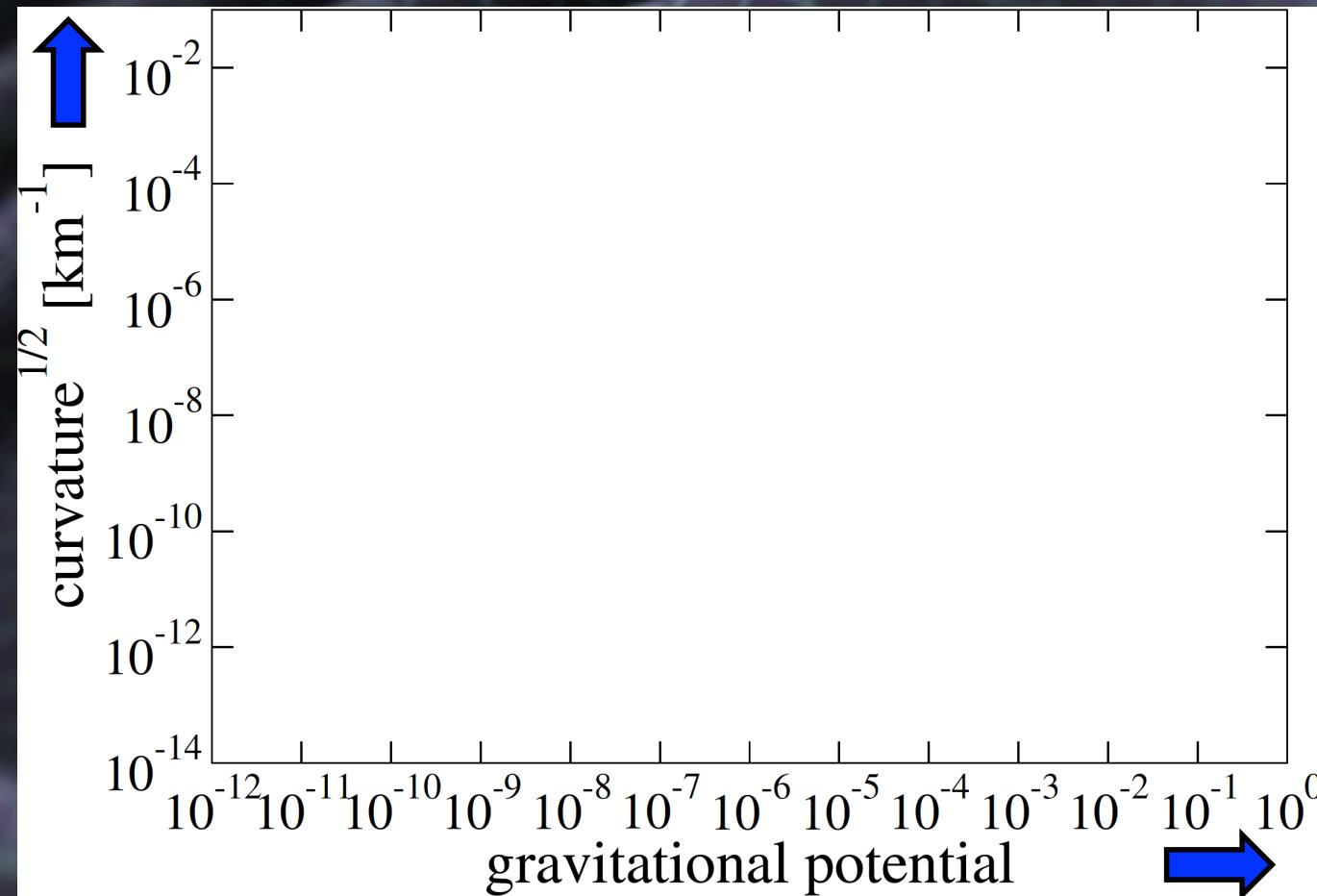
JPS Meeting

Miyazaki U. Sept. 23rd 2016

N. Yunes (Montana State U.), F. Pretorius (Princeton U.)
arXiv: 1603.08955 (to appear in PRD)

Strong/Dynamical Nature of GW Sources

strong

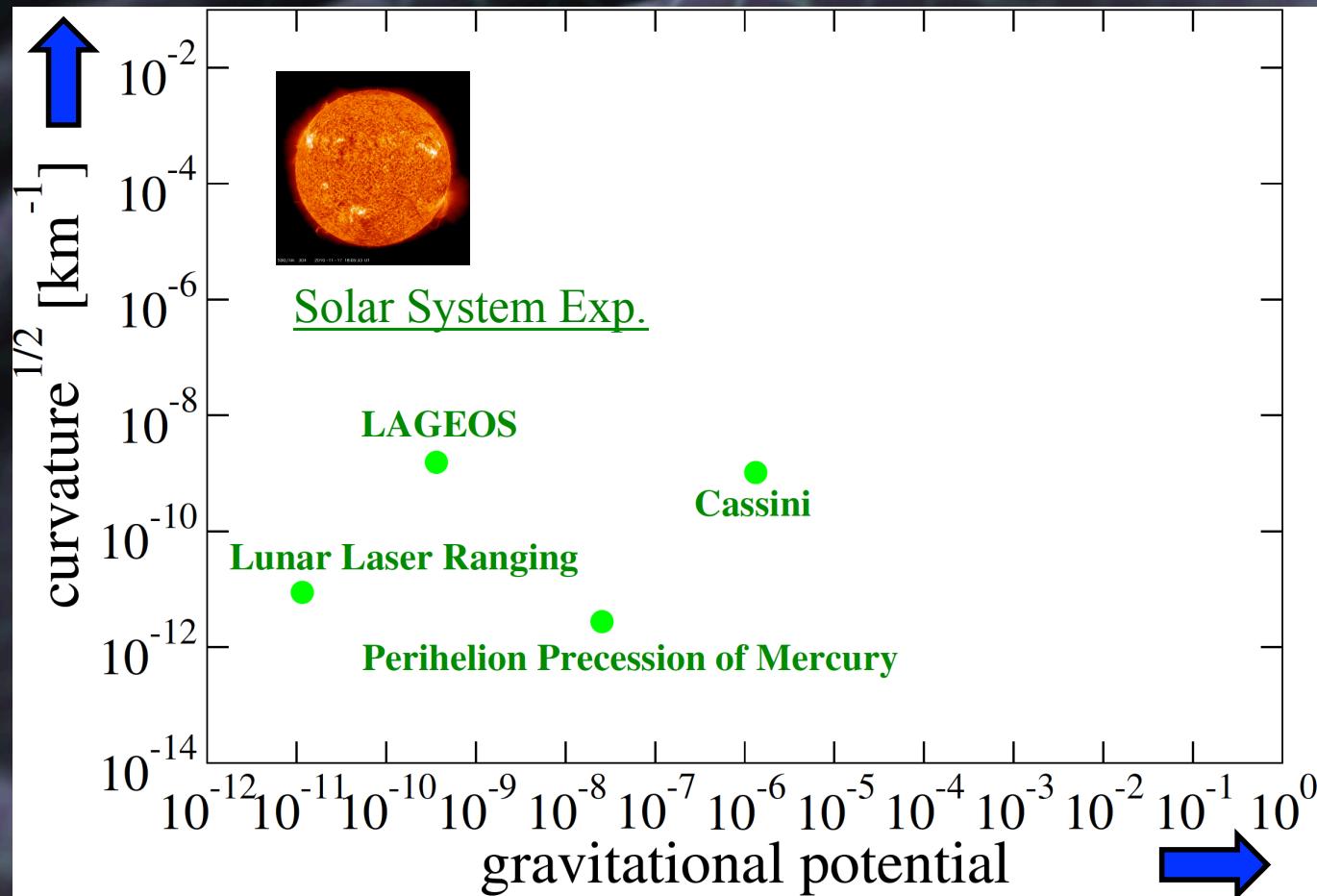


[Yunes, KY & Pretorius arXiv:1603.08955]

large

Strong/Dynamical Nature of GW Sources

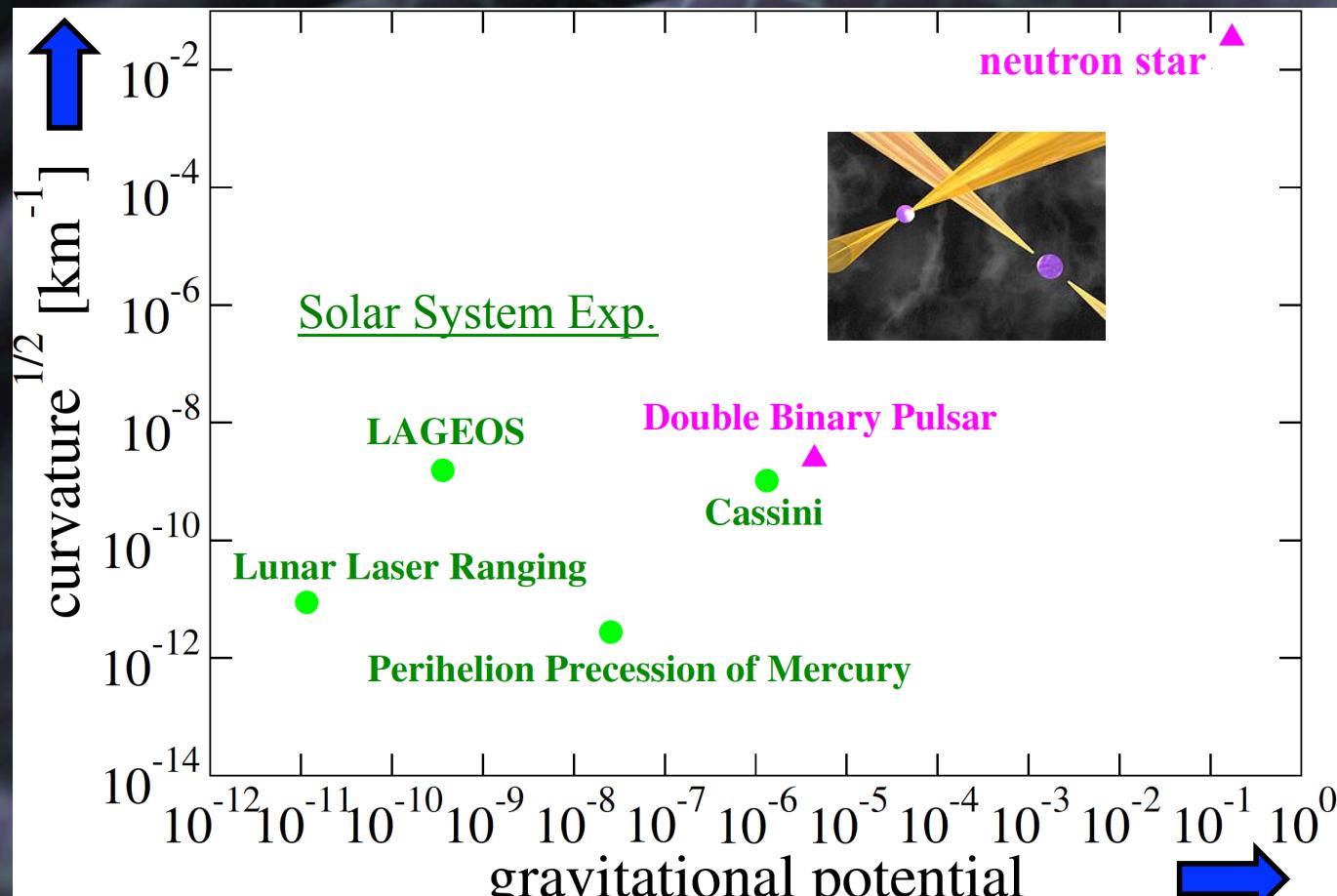
strong



[Yunes, KY & Pretorius arXiv:1603.08955]

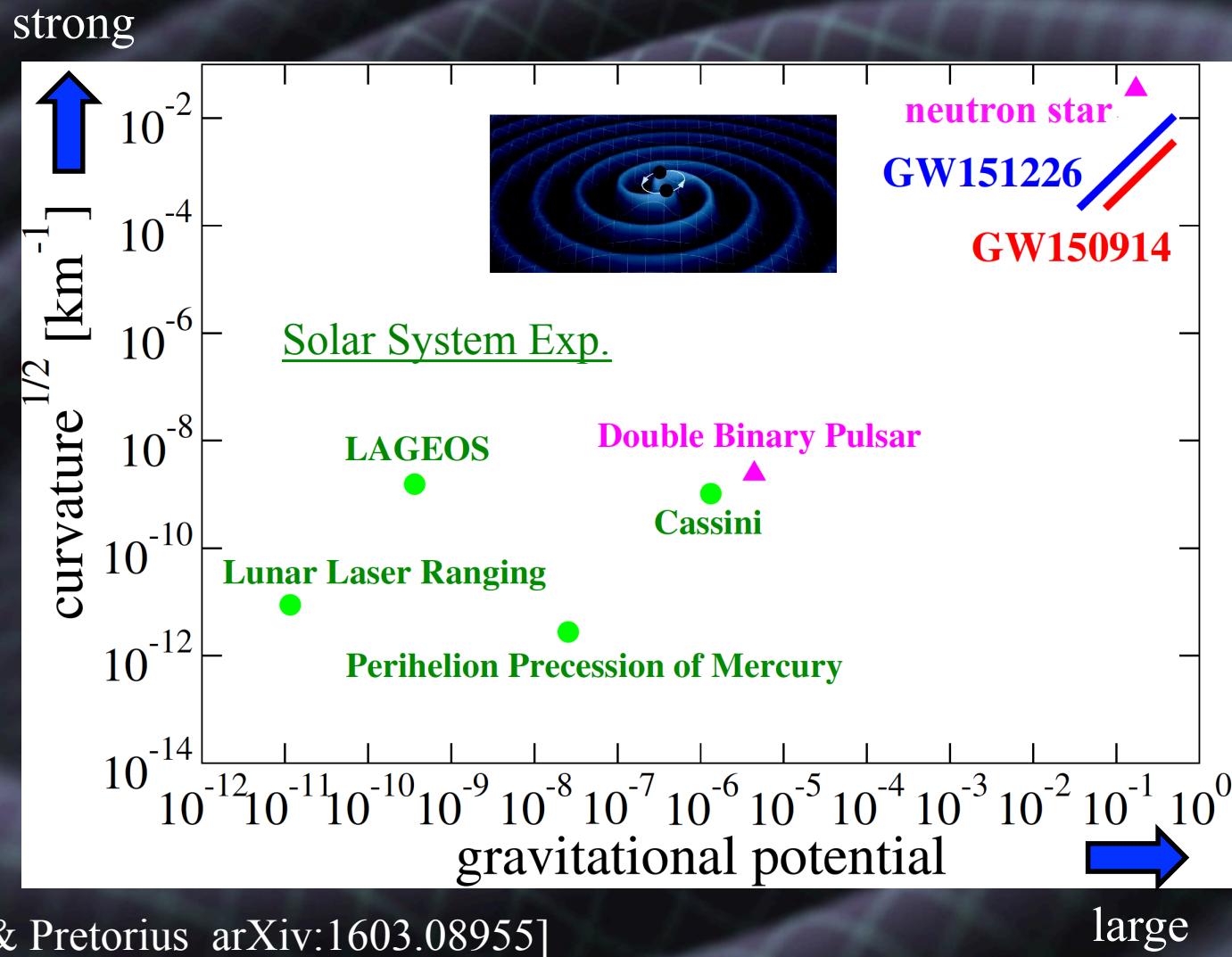
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strong



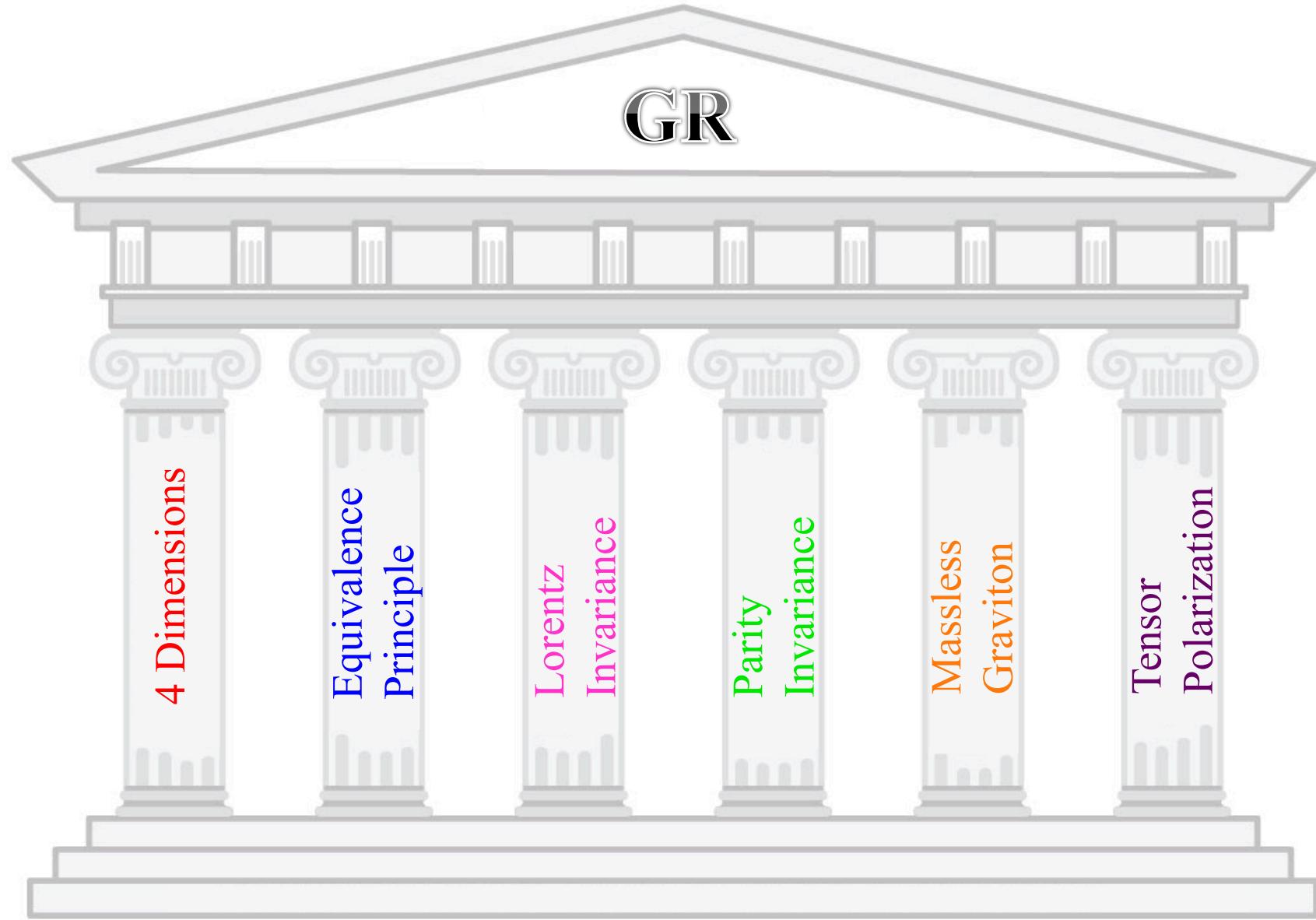
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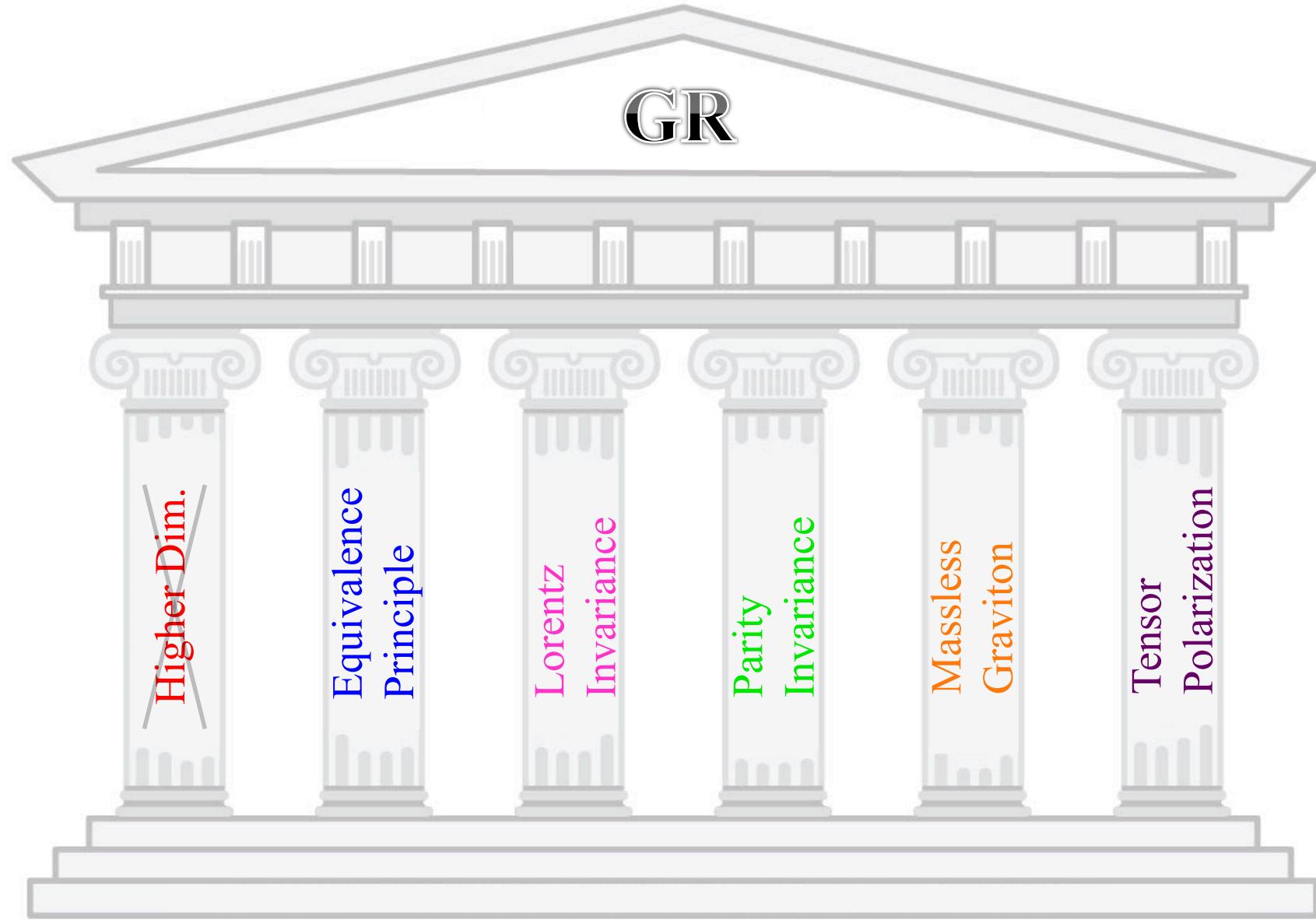


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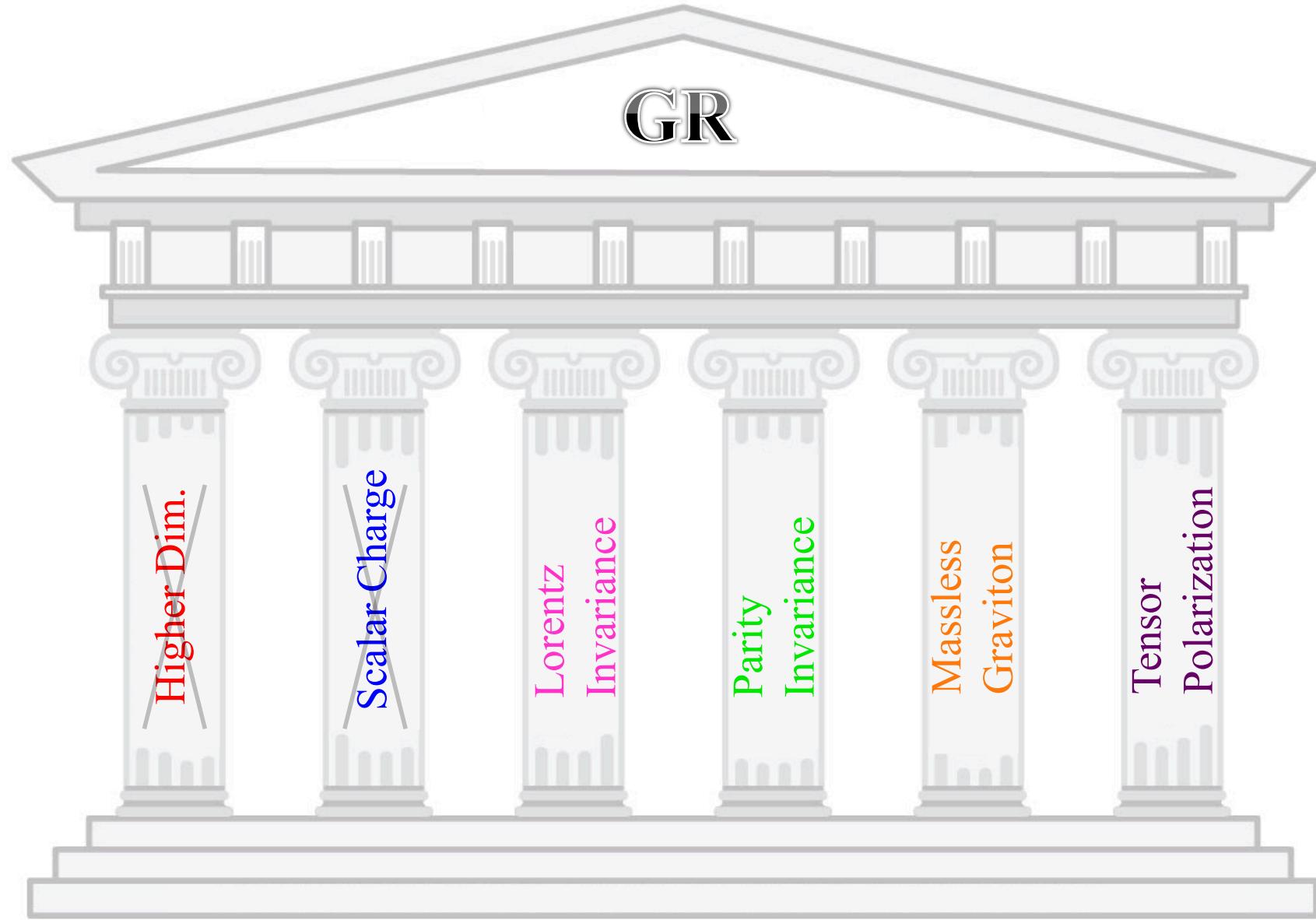
Fundamental Pillars in General Relativity (GR)



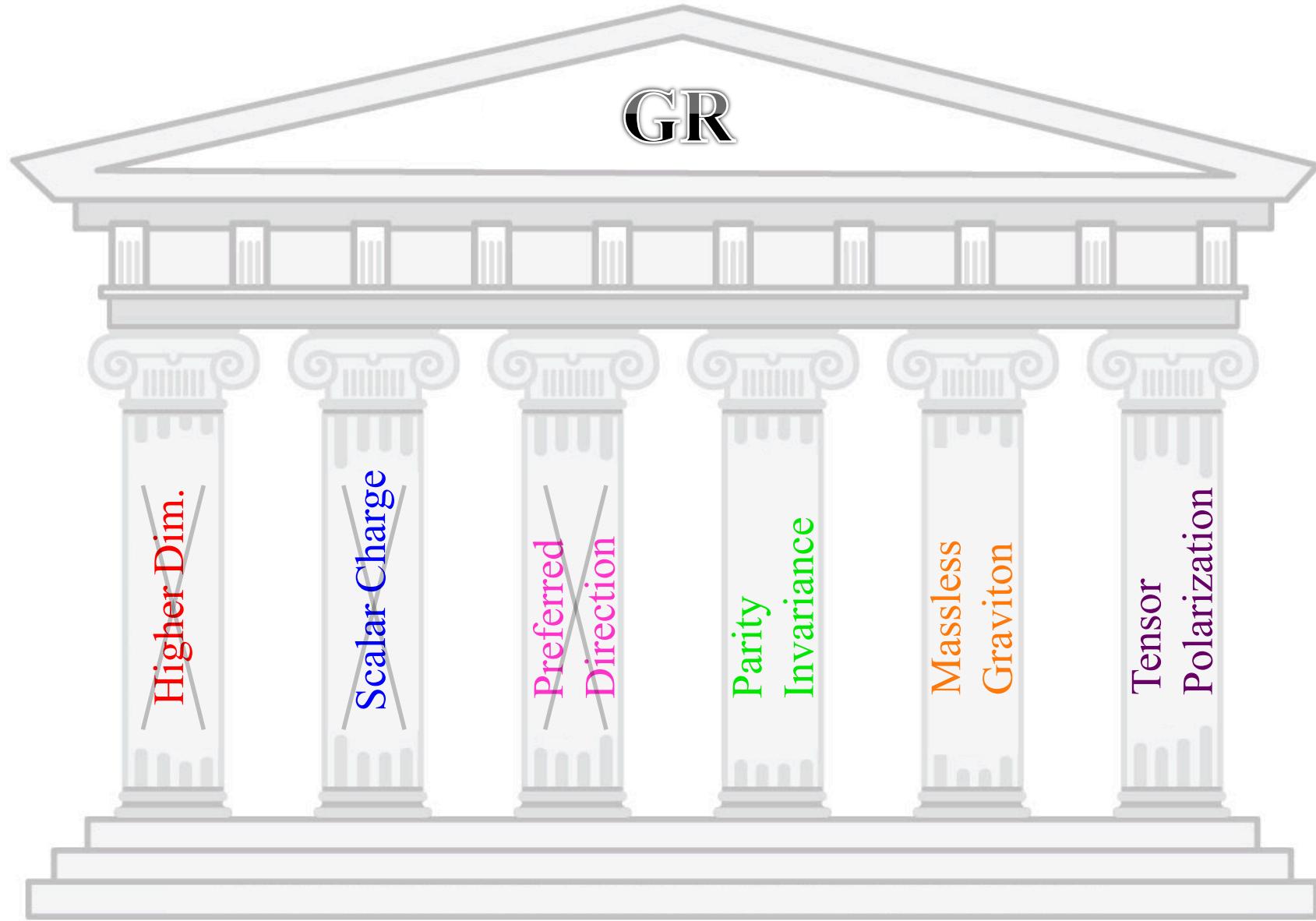
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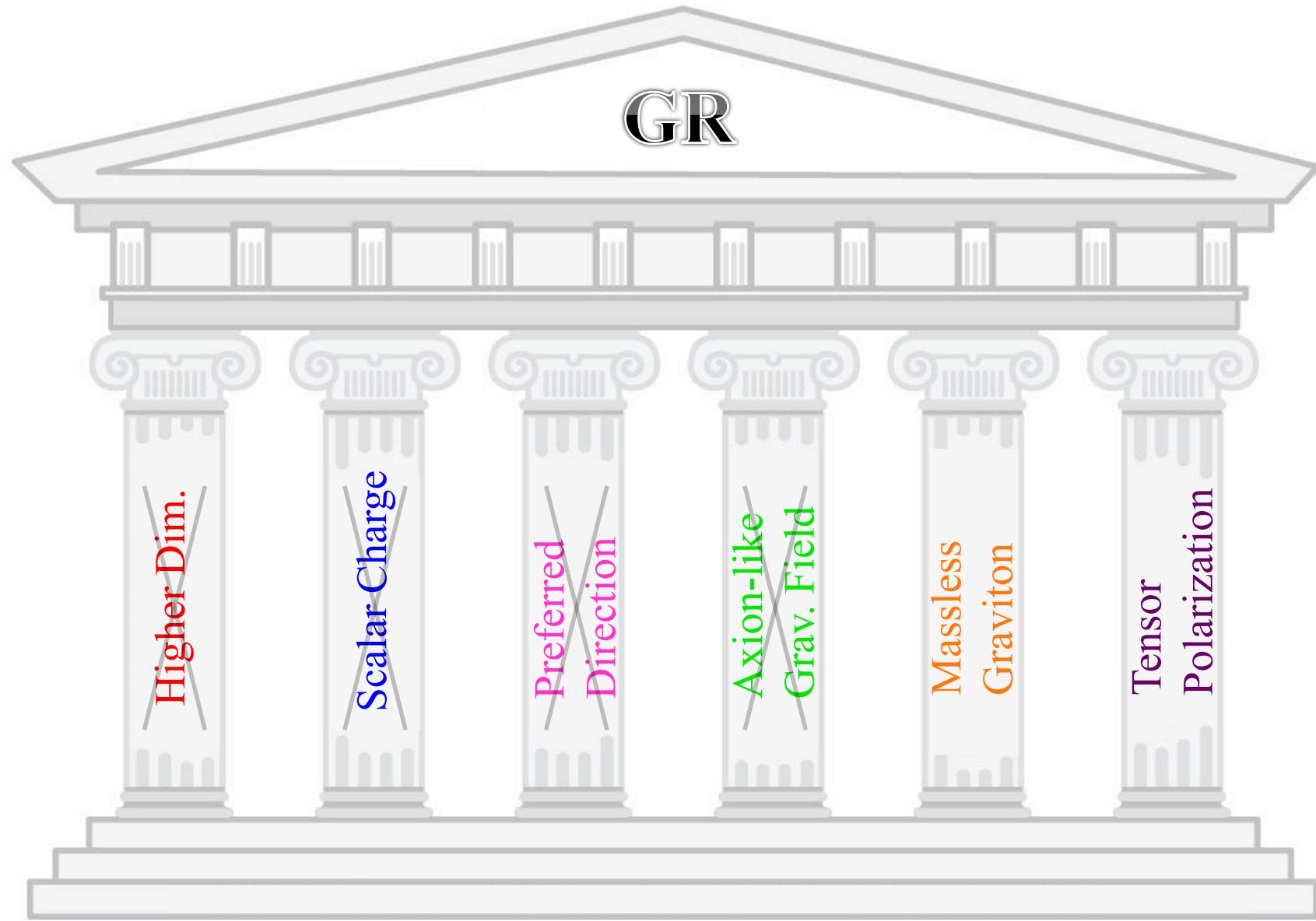
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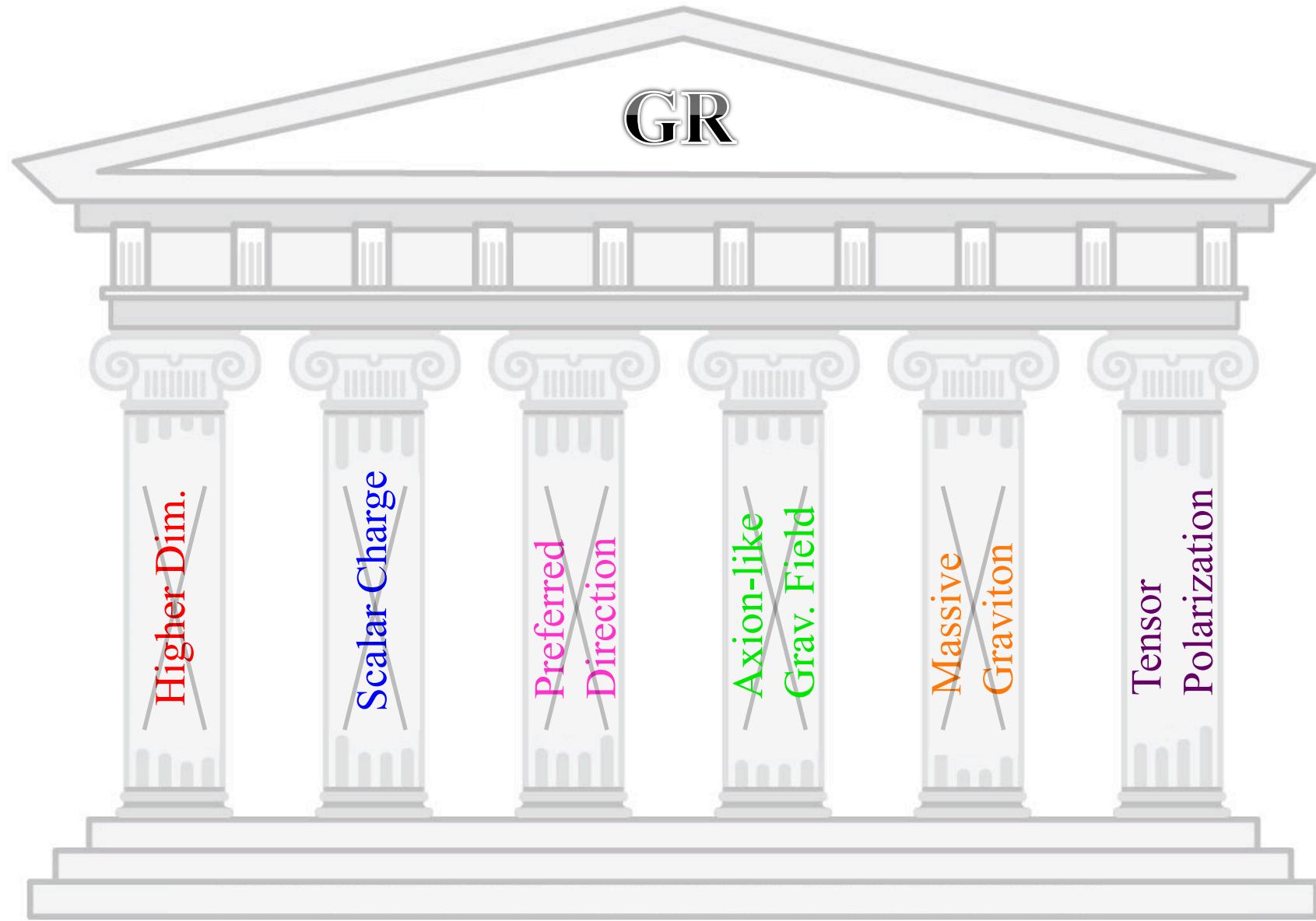
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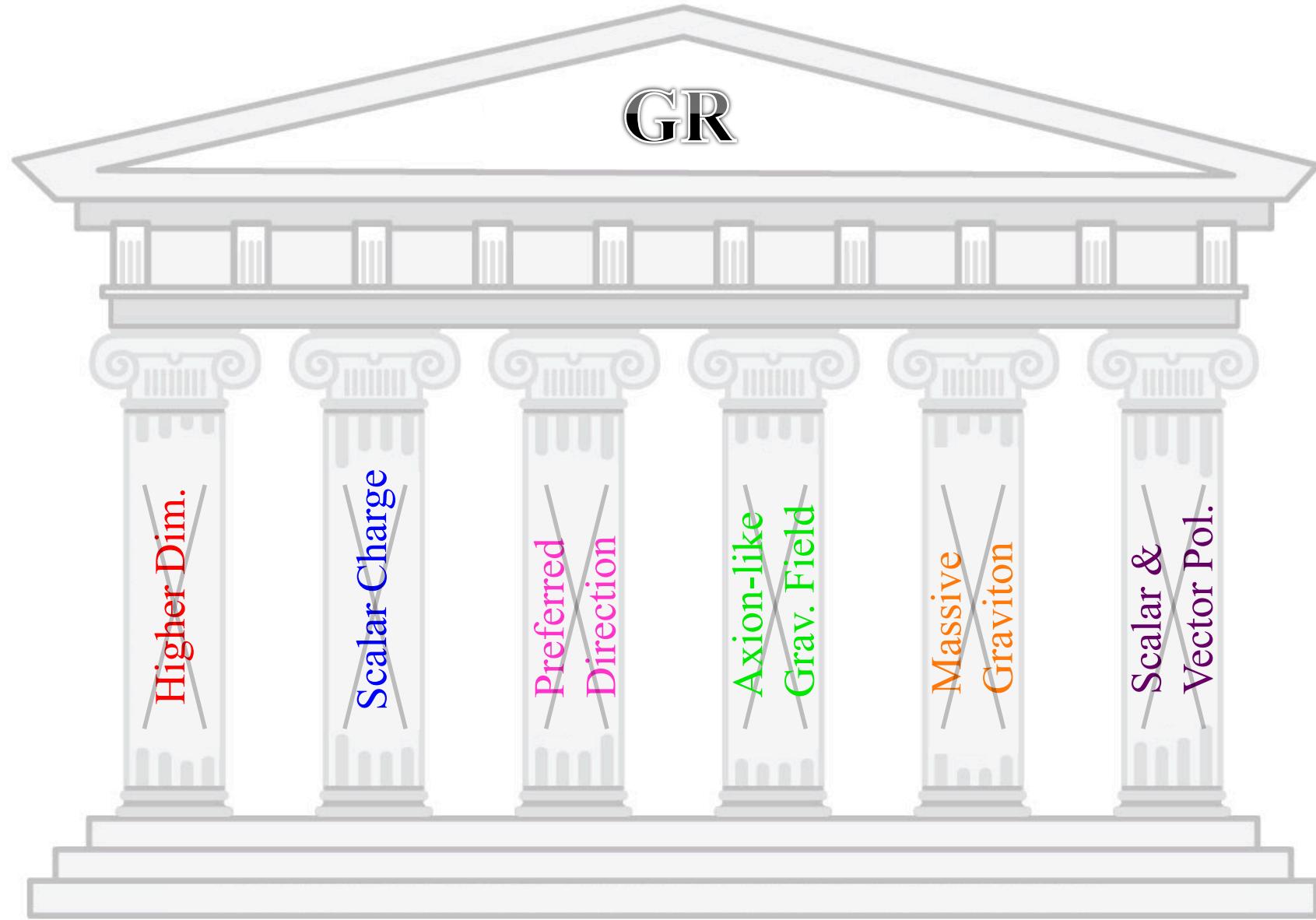
Fundamental Pillars in General Relativity (GR)



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Fundamental Pillars in General Relativity (GR)



parameterized post-Einsteinian (ppE) Formalism

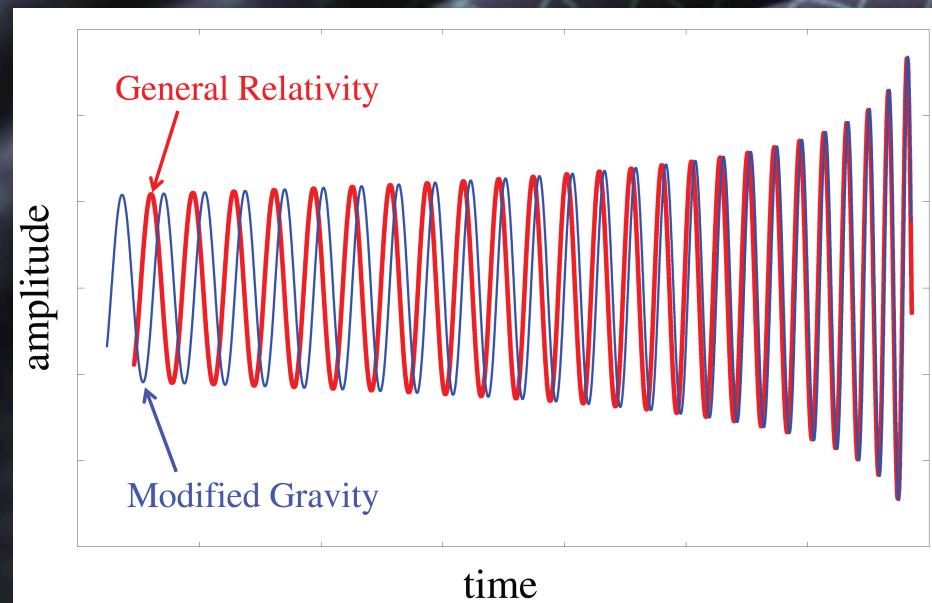
$$c = G = 1$$

gravitational waveform phase:

ppE parameter

[Yunes & Pretorius (2009)]

$$\Psi(f) = \Psi_{\text{GR}}(f) + \beta v^2 n^{-5}$$



n -th post-Newtonian (PN)
correction relative to GR

Time-varying G theories

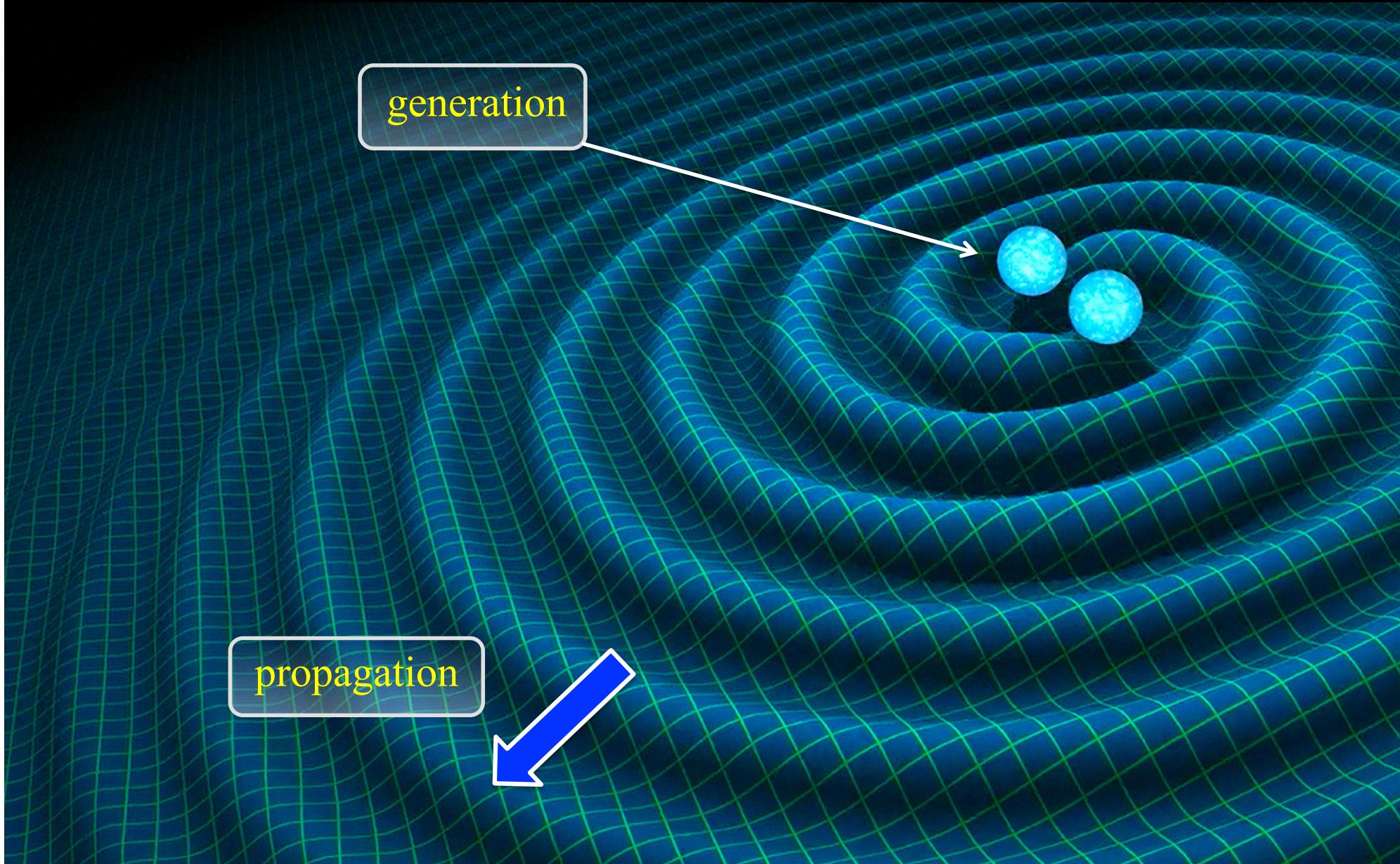
$$\beta_G \sim \dot{G} M_{\text{tot}}$$

Fisher Analysis

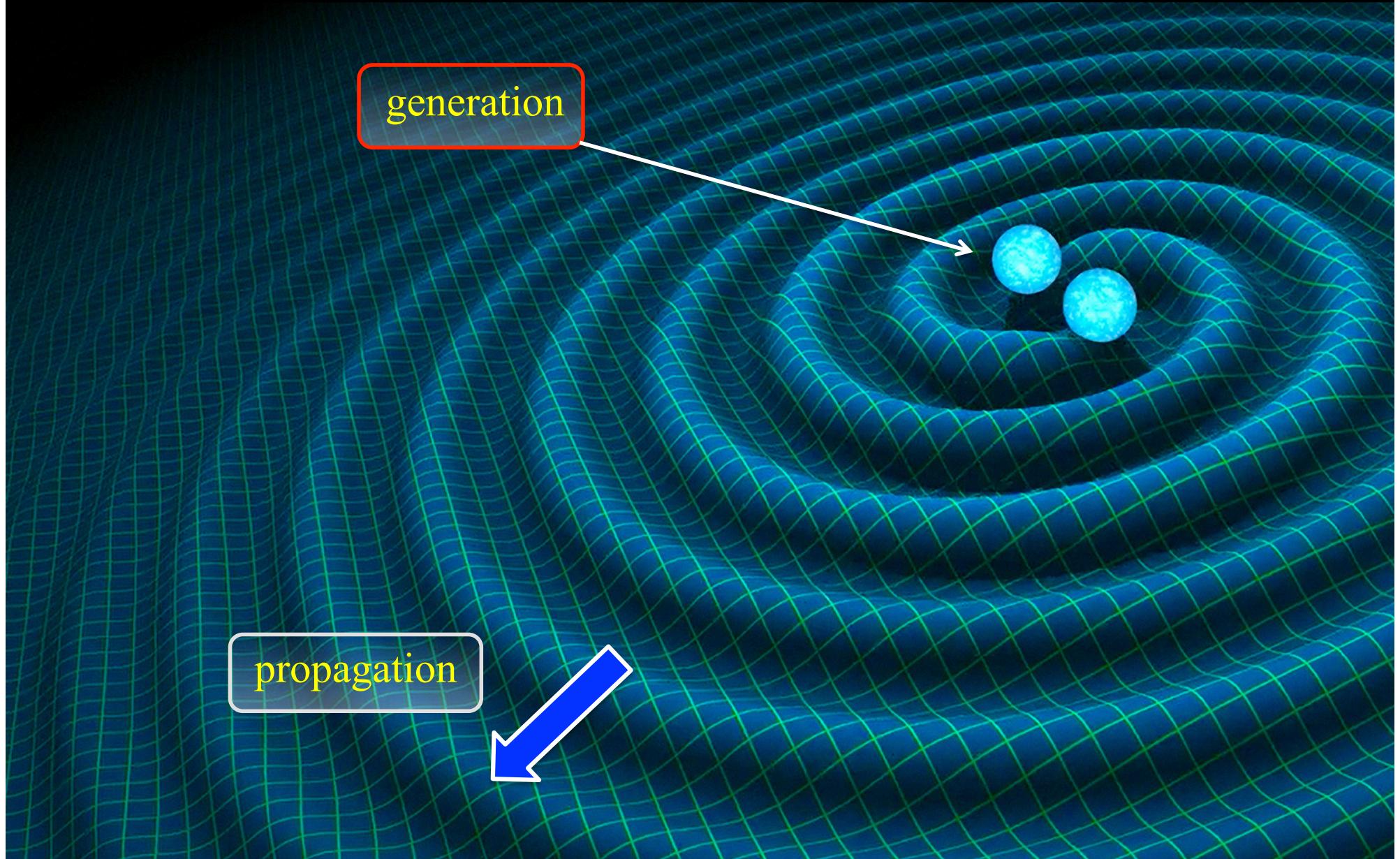
signal: GR waveform consistent with GW150914 & GW151226

template: ppE modified waveform

Different Types of Modifications



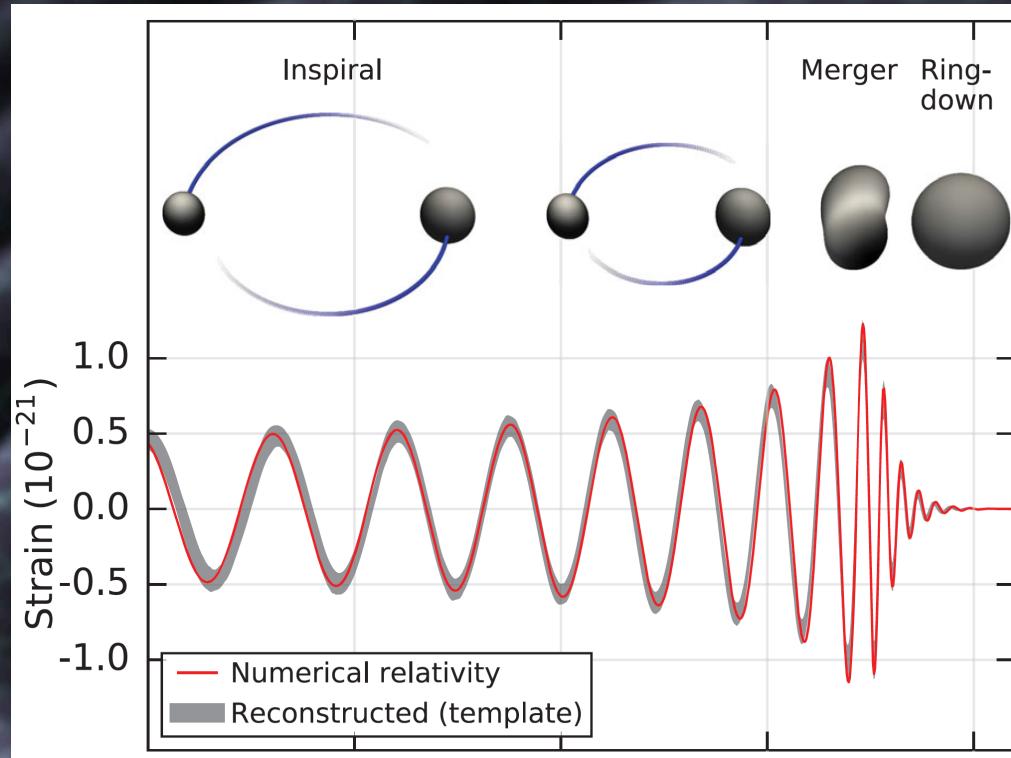
Outline



Outline

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ppE Phase for Modified GW Generation



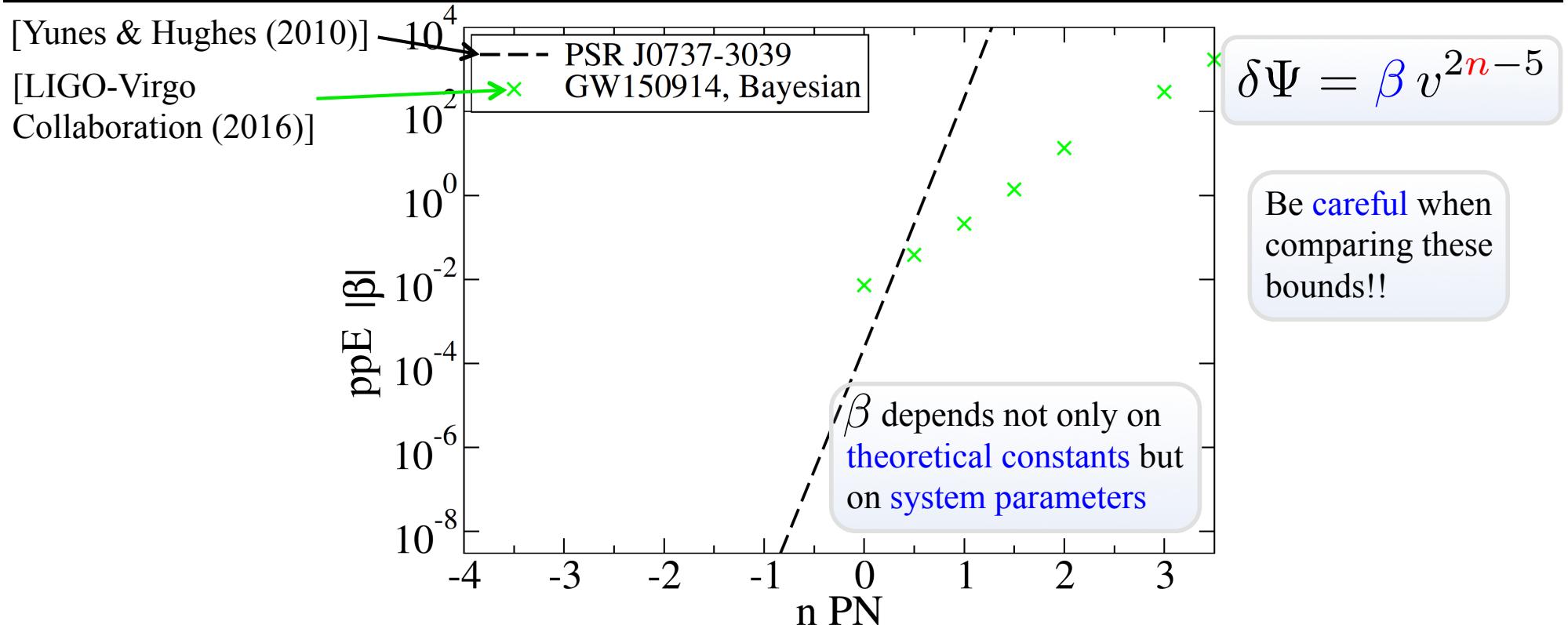
[Abbott et al. PRL 116 061102 (2016)]

non-GR correction in
inspiral only

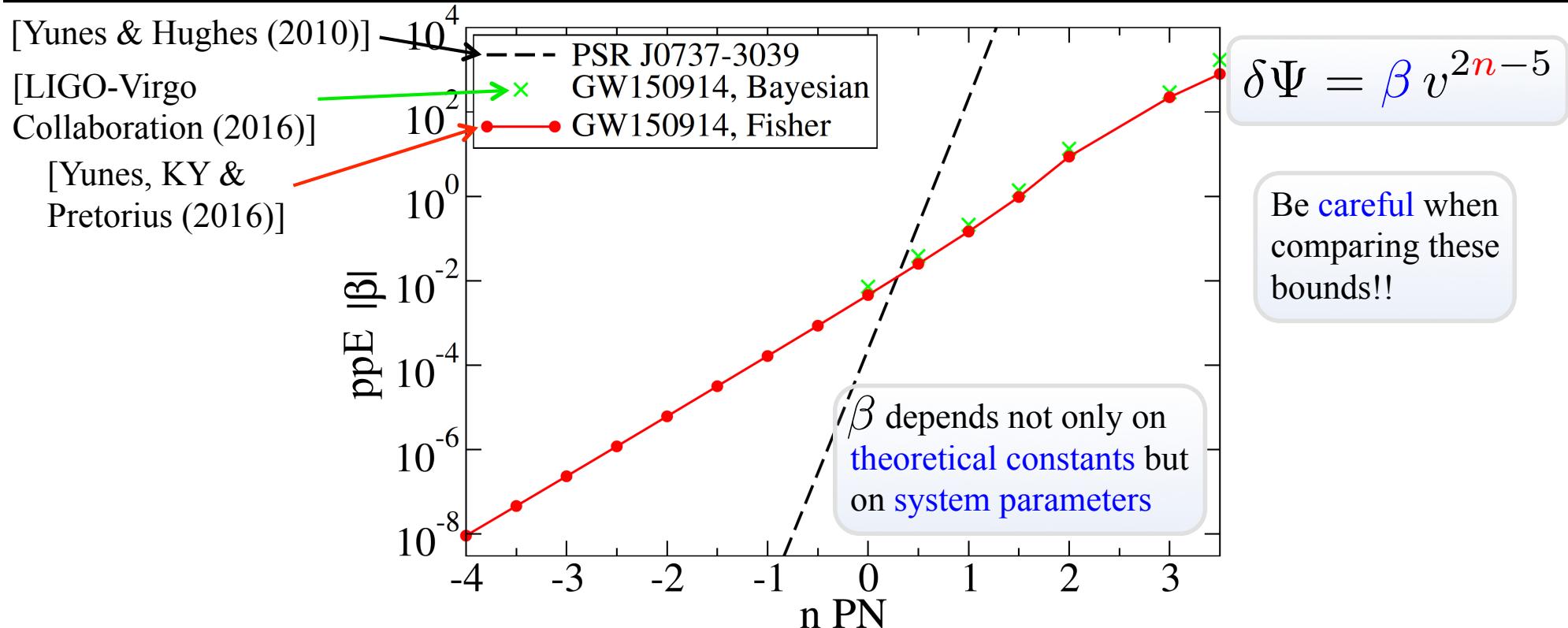
$$\Psi^{(\text{insp})} = \Psi_{\text{GR}}^{(\text{insp})} + \beta v^{2n-5}$$
$$\Psi^{(\text{merg,ring})} = \Psi_{\text{GR}}^{(\text{merg,ring})}$$

(due to the lack of BH binary merger simulations in non-GR theories)

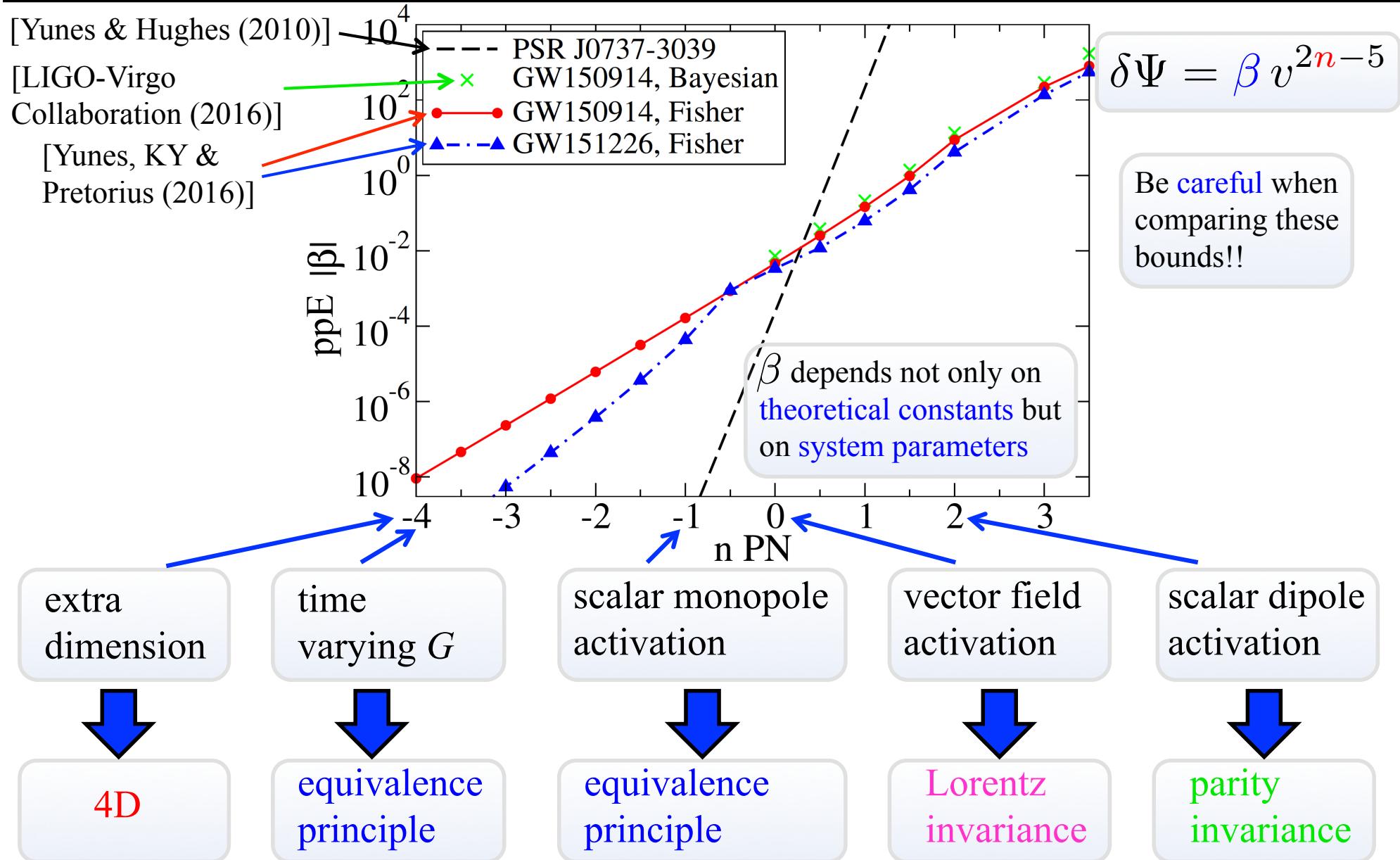
Constraining GR Fundamental Pillars



Constraining GR Fundamental Pillars



Constraining GR Fundamental Pillars



generation

propagation

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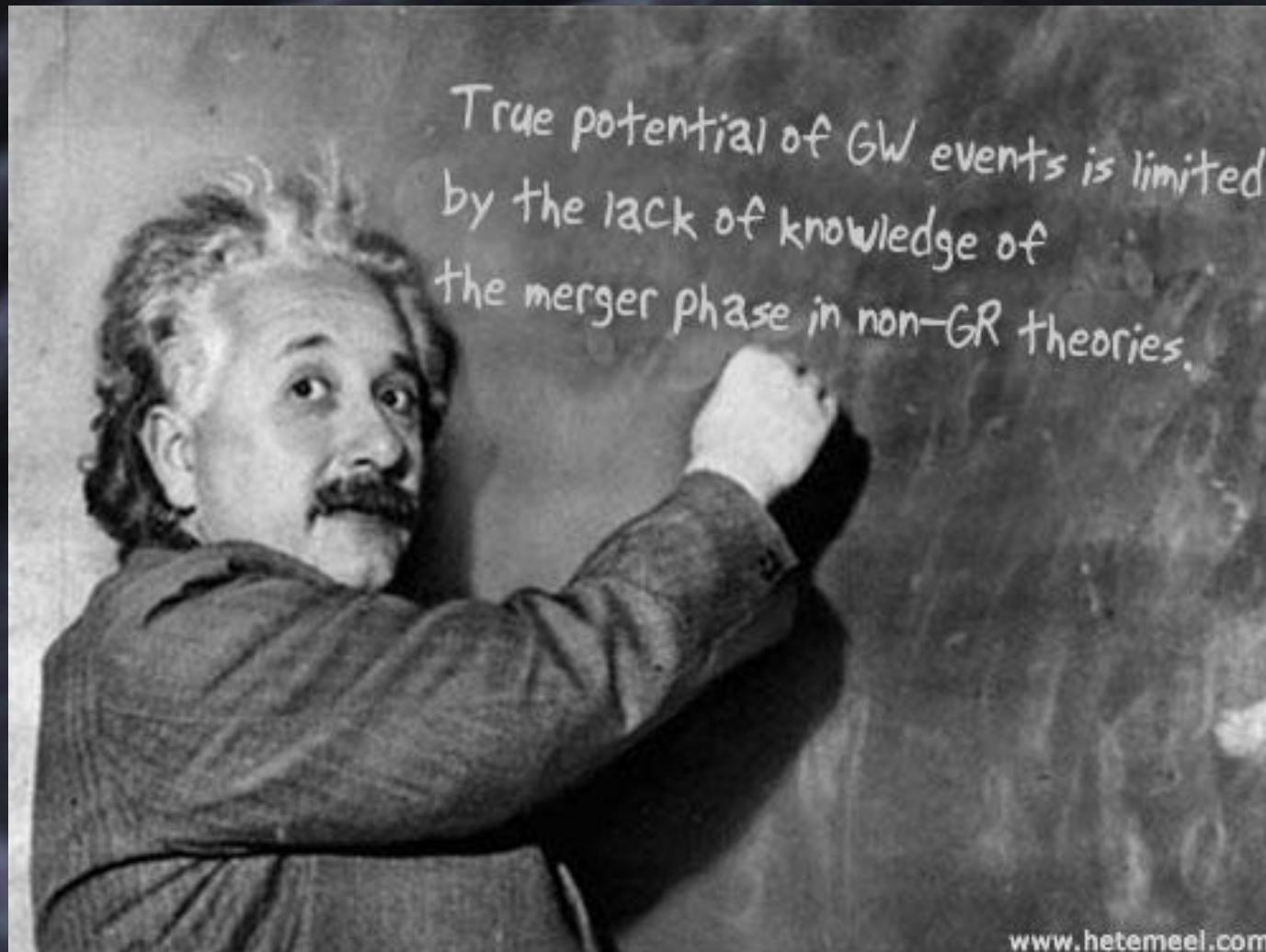
Mapping to Theoretical Constraints

no meaningful constraints
(beyond small-coupling approximation)

Example Theories (Theoretical Parameters)	GR Pillar	PN	Example Theory Constraints		
			GW150914	GW151226	Current
Einstein-dilaton Gauss-Bonnet ($\sqrt{ \alpha_{\text{EdGB}} }$ [km])	Equiv. Princ.	-1	—	—	$10^7, 2$
	scalar-tensor ($ \dot{\phi} $ [1/sec])	Equiv. Princ.	-1	—	10^{-6}
	dynamical Chern-Simons ($\sqrt{ \alpha_{\text{dCS}} }$ [km])	Parity Inv.	+2	—	10^8
Einstein-Æther (c_+, c_-)	Lorentz Inv.	0	(0.9, 2.1)	(0.8, 1.1)	$(0.03, 0.003)$
		4D	5.4×10^{10}	2.0×10^9	$10-10^3$
		Equiv. Princ.	5.4×10^{18}	1.7×10^{17}	0.1–1

weaker than current bounds
first constraint in the strong/dynamical gravity regime

Important Message

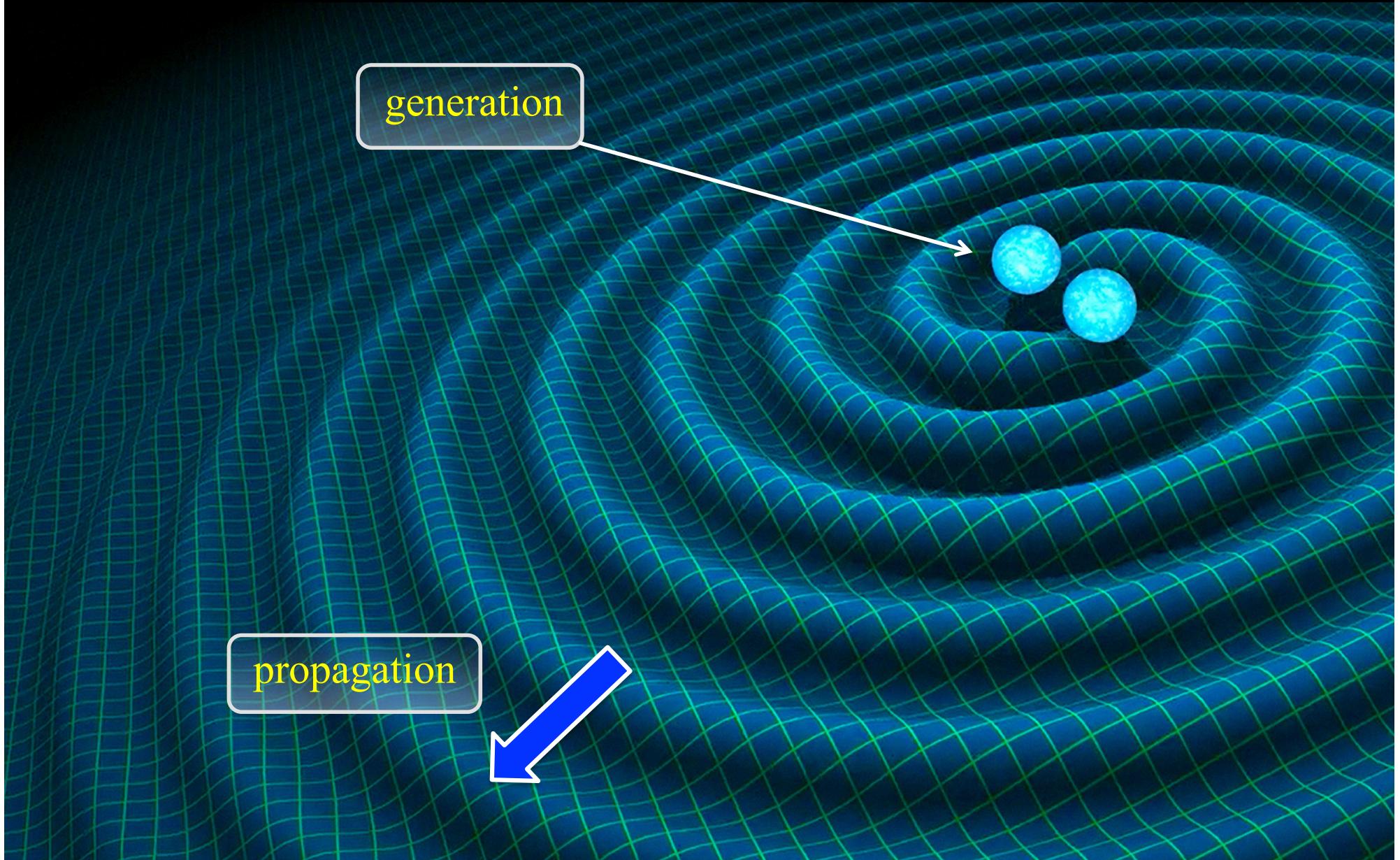


generation

propagation

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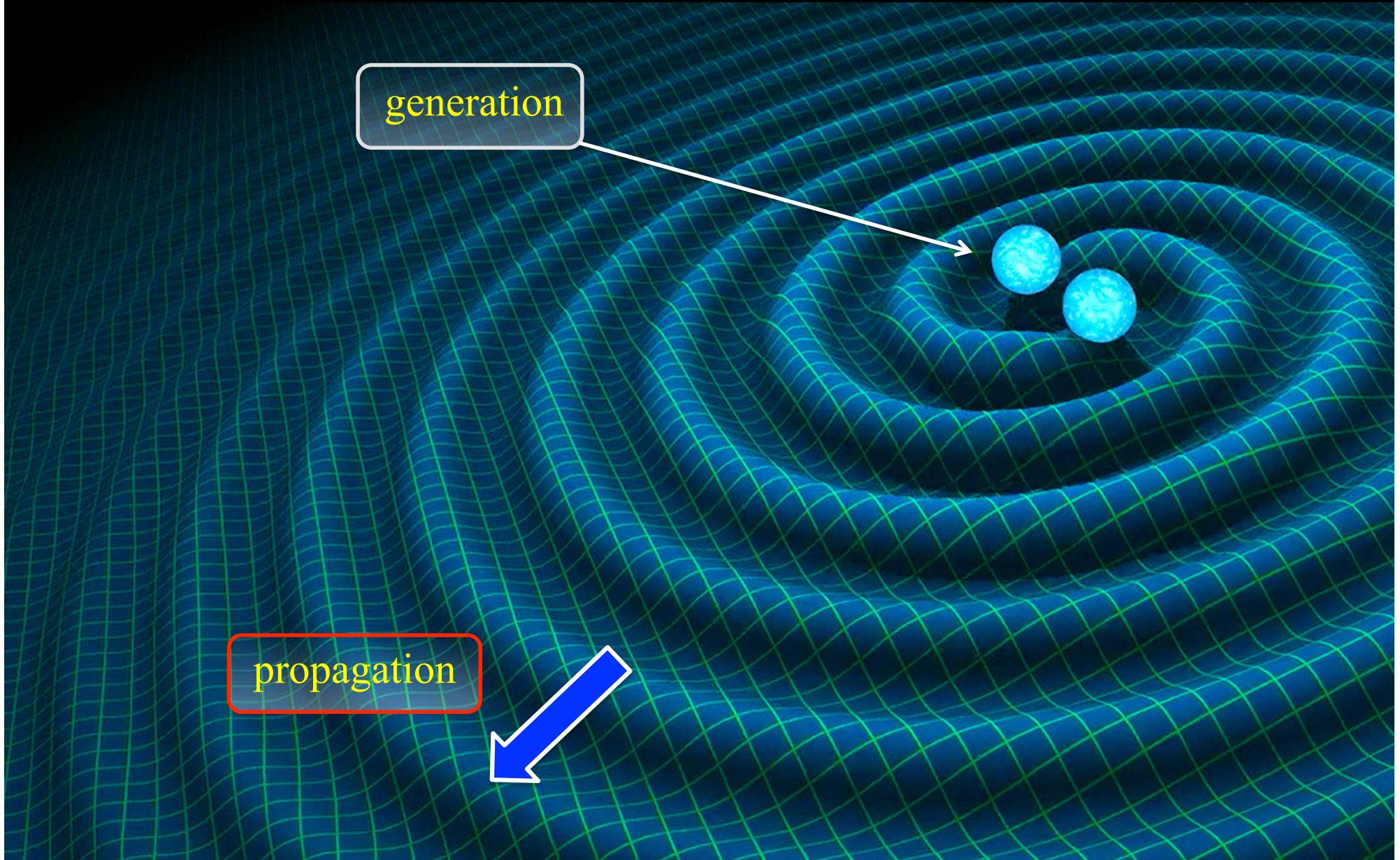
Outline



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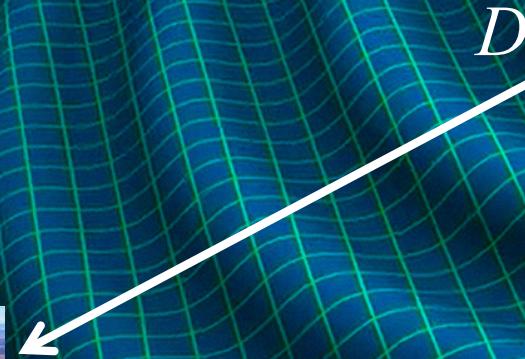
Modified GW Propagation

graviton dispersion relation

$$E^2 = (p c)^2 + A (p c)^\alpha$$



$$\Psi - \Psi_{\text{GR}} \sim A D f^{\alpha-1}$$



This correction is added to full
inspiral-merger-ringdown GR phase.

generation

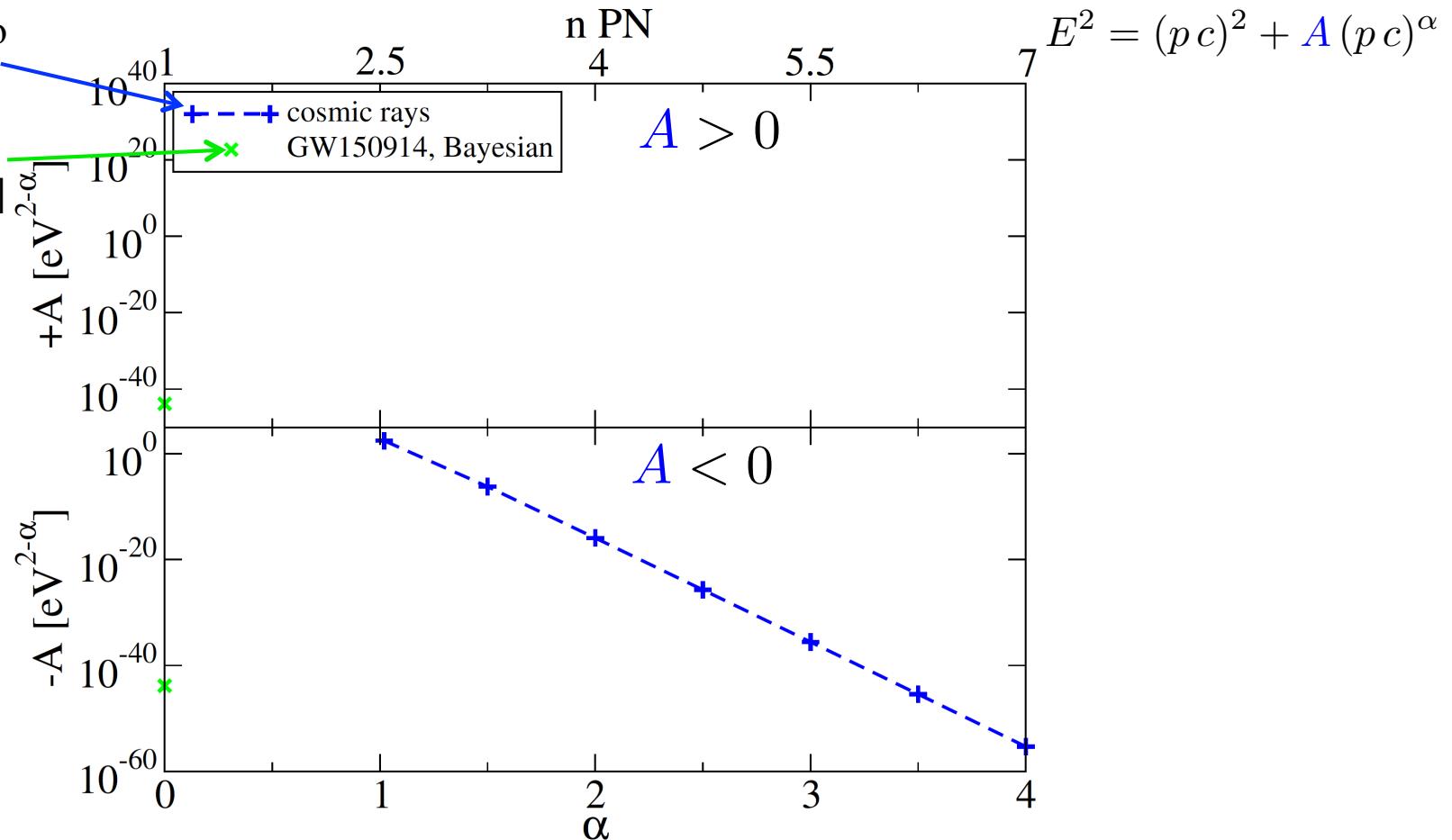
propagation

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Bounds on the Propagation Mechanisms

[Kiyota & Yamamoto
(2015)]

[LIGO-Virgo
Collaboration (2016)]

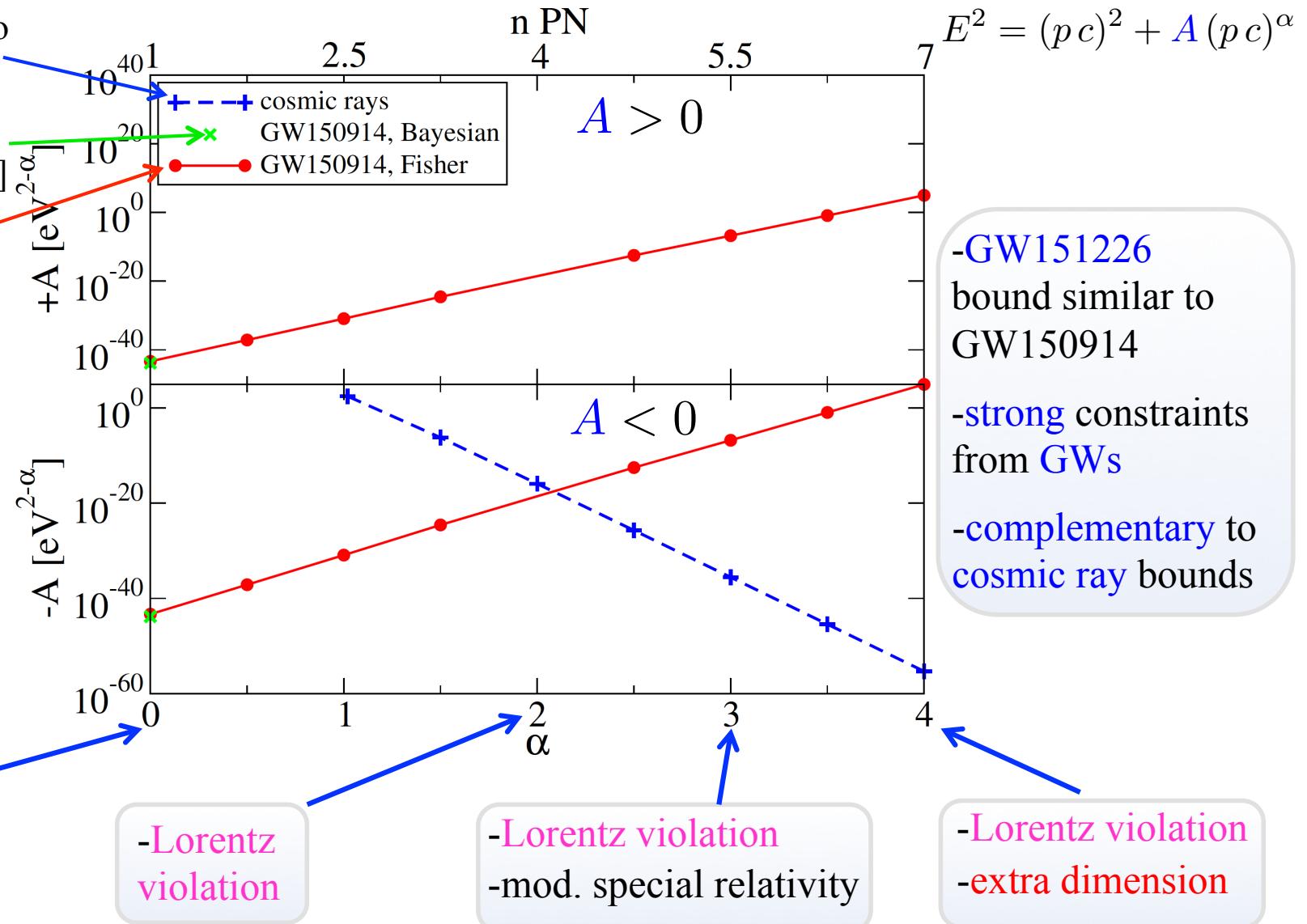


Bounds on the Propagation Mechanisms

[Kiyota & Yamamoto
(2015)]

[LIGO-Virgo
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[Yunes, KY &
Pretorius (2016)]



Mapping to Theoretical Constraints

Example Theories	Repr. Parameters	GR Pillar	PN	Example Theory Constraints		
				GW150914	GW151226	Current
Massive graviton	m_g [eV]	$m_g = 0$	+1	10^{-22}	10^{-22}	$10^{-29} - 10^{-18}$
Multifractional	E_*^{-1} [eV $^{-1}$] (time)	Lor. Inv.	+4.75	5.8×10^{-27}	3.3×10^{-26}	—
	E_*^{-1} [eV $^{-1}$] (space)			1.0×10^{-26}	5.7×10^{-26}	3.9×10^{-53}
Modified Special Rel.	$\eta_{\text{dsrt}}/L_{\text{Pl}} > 0$	Lor. Inv.	+5.5	1.3×10^{22}	3.8×10^{22}	—
	$\eta_{\text{dsrt}}/L_{\text{Pl}} < 0$					2.1×10^{-7}
Standard Model Ext.	$\mathring{k}_{(I)}^{(4)} > 0$		+4	—	—	6.1×10^{-17}
	$\mathring{k}_{(I)}^{(4)} < 0$			0.64	19	—
	$\mathring{k}_{(V)}^{(5)} > 0$ [cm]	Lor. Inv.	+5.5	1.7×10^{-12}	3.1×10^{-11}	1.7×10^{-40}
	$\mathring{k}_{(V)}^{(5)} < 0$ [cm]					—
	$\mathring{k}_{(I)}^{(6)} > 0$ [cm 2]		+7	7.2×10^{-4}	3.3×10^{-3}	3.5×10^{-64}
	$\mathring{k}_{(I)}^{(6)} < 0$ [cm 2]					—
Hořava-Lifshitz	$\kappa_{\text{hl}}^4 \mu_{\text{hl}}^2$ [1/eV 2]	Lor. Inv.	+7	1.5×10^6	6.9×10^6	—
Einstein-Æther	c_+	Lor. Inv.	+4	0.7	0.998	0.03
khronometric	β					0.005

generation

propagation

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Conclusions

Takeaway

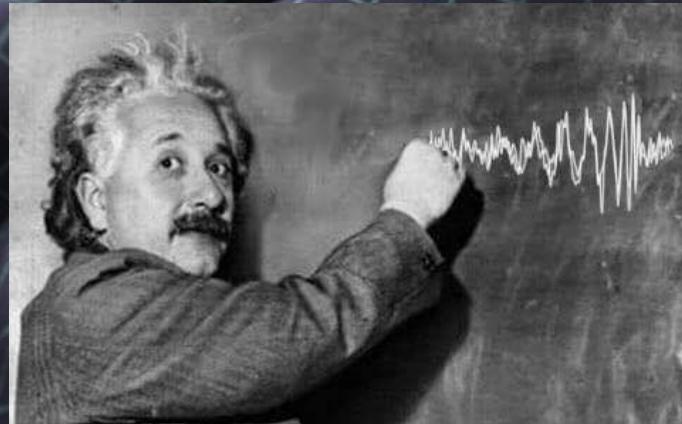
generation mechanisms:

- weak bound
- first strong/dyn. gravity constraint

propagation mechanisms:

- strong bound
- complementary to cosmic rays

Thank You



[https://igoligo.wordpress.com/
2016/02/12/einstein-was-right/](https://igoligo.wordpress.com/2016/02/12/einstein-was-right/)

- Merger simulations in non-GR theories
- Stacking multiple signals
- Electromagnetic counterpart

Back Up

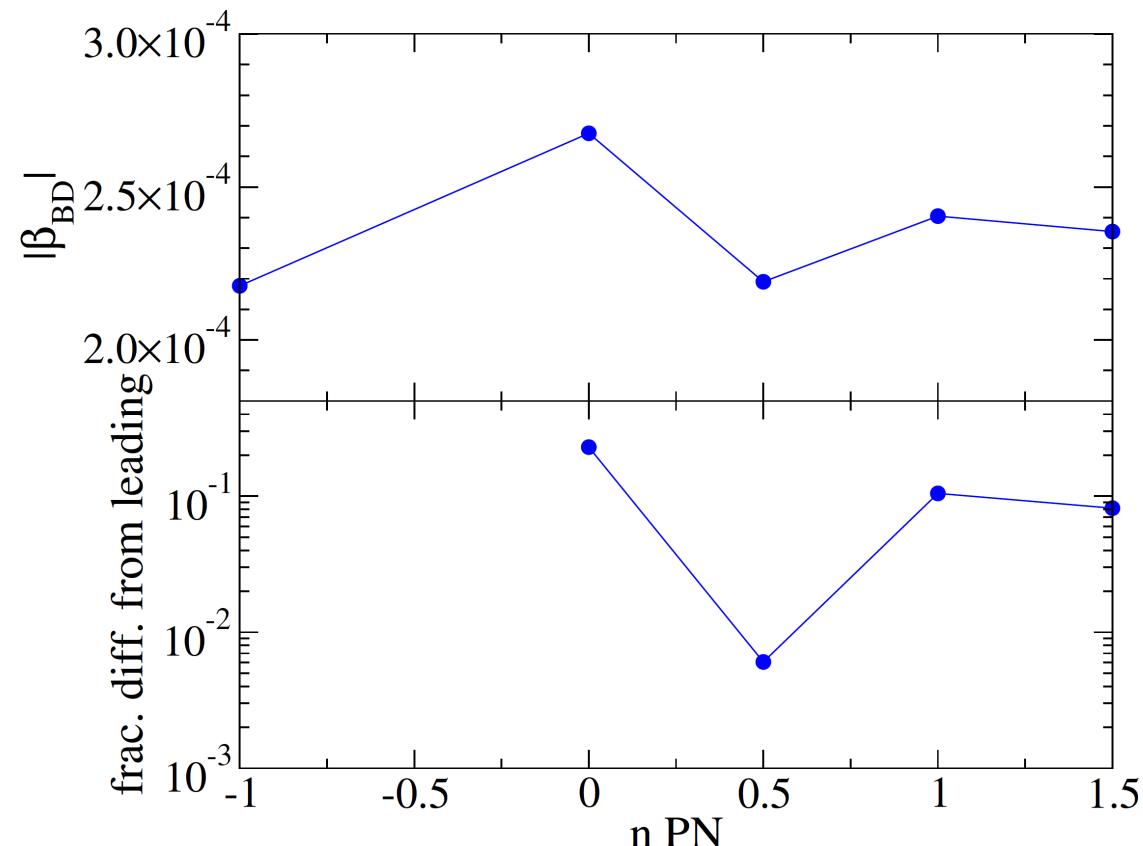
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Effect of Higher PN Corrections

Brans-Dicke (BD) theory as an example

$$\Psi_{\text{BD}}(f) = \Psi_{\text{GR}}(f) + \beta_{\text{BD}} (\pi \mathcal{M} f)^{-7/3} \left[1 + \sum_{k=2}^5 \delta \psi_k^{\text{BD}}(\eta) (\pi \mathcal{M} f)^{k/3} \right]$$

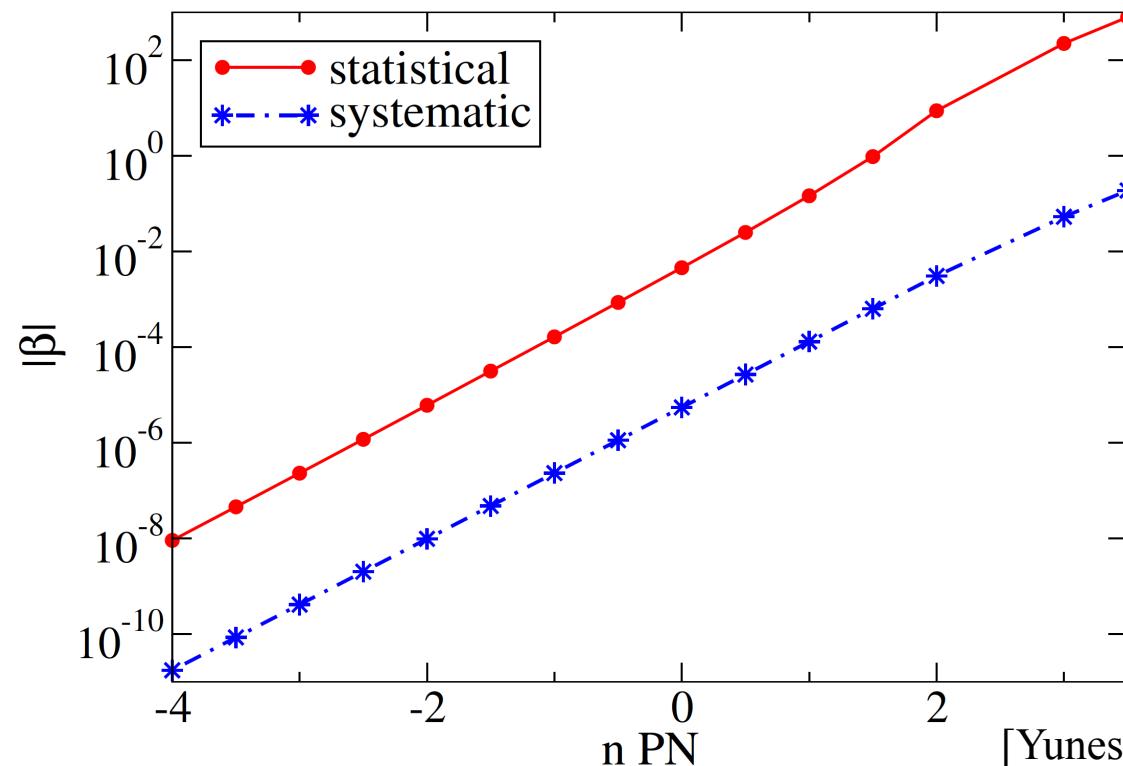
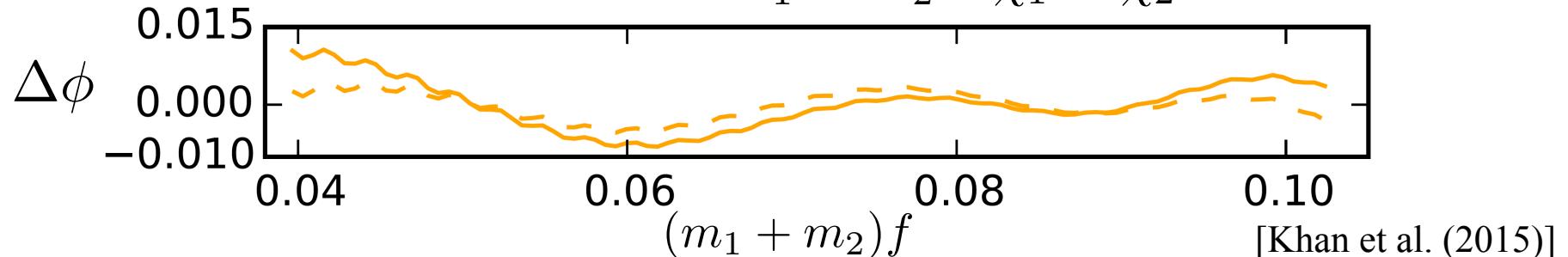
Higher PN corrections known for small mass ratio limit



Mismodeling Template Systematics

Waveform mismodeling error in phase

$$m_1 = m_2 \quad \chi_1 = \chi_2 = 0$$



Mismodeling systematics
much smaller than
statistical errors

Strong Equivalence Principle

[Will, Living Review (2006)]

1. The trajectories of freely-falling test bodies, including self-gravitating ones, are independent of their **internal structure** and **composition**.

2. Results of any local experiment, including gravitational ones, are independent of **when** and **where** they are performed (**local position invariance**), and of the **velocity** of the experimental apparatus (**local Lorentz invariance**).

Constraining GW Propagation with EM Counterparts

$$\left| \frac{v_g - c}{c} \right| < \frac{c \Delta \tau_{\text{int}}}{D}$$

[Nishizawa &
Nakamura (2014)]

- Fermi GBM
- super-luminal GWs
- prompt emission

