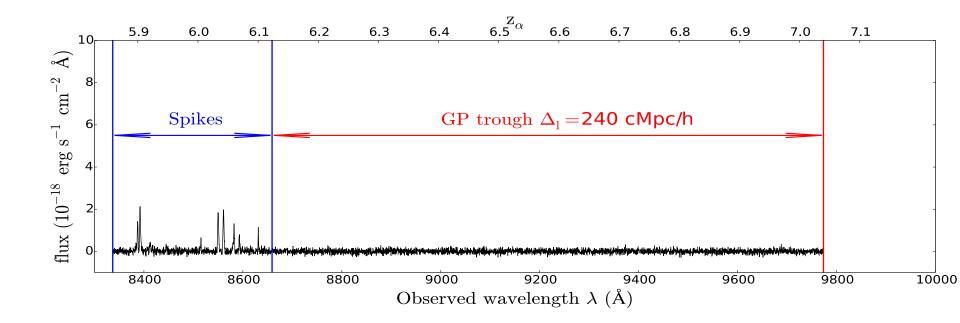
A tale of seven narrow spikes and a long trough: constraining the timing of the percolation of HII bubbles at the tail-end of reionization with ULAS J1120+0641.

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The high resolution spectrum of ULAS J1120+0641

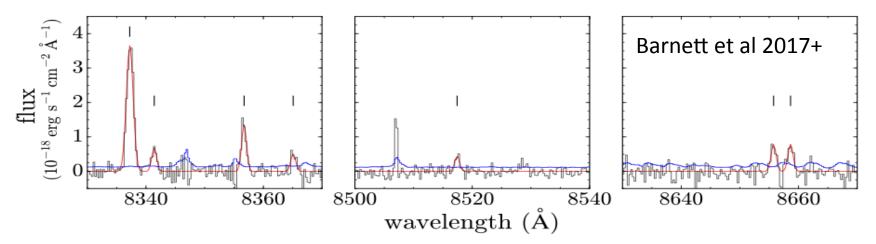


Fig. 5: Significant transmission of flux in the Ly α forest towards ULAS J1120+0641. The unbinned X-shooter spectrum of ULAS J1120+0641 is shown in grey with a pixel size of $10 \,\mathrm{km}\,\mathrm{s}^{-1}$. Shown in blue is the 1σ error spectrum, and the derived transmission spectrum is overlaid in red. The seven detections presented in Table 2 are indicated.

- 30 h integration spectrum of ULAS J1120+0641, the high redshift quasar at z=7.084
- Presence of 7 spikes in the redshift range 5.858<=z<=6.122
- Largest Gunn-Peterson trough ever reported (240 cMpc/h) from z=6.122 up to z=7.04 : the near zone of the quasar.

Using radiative transfer simulation to model the observed spectrum

Model name	Box size	Particle number
512-20-good_h (Fiducial)	20	512^{3}
512-20-good-l (Late)	20	512 ³
512-20-early (Early)	20	512^{3}
512-40 (Low-res)	40	512 ³
512-10 (High-res)	10	512 ³

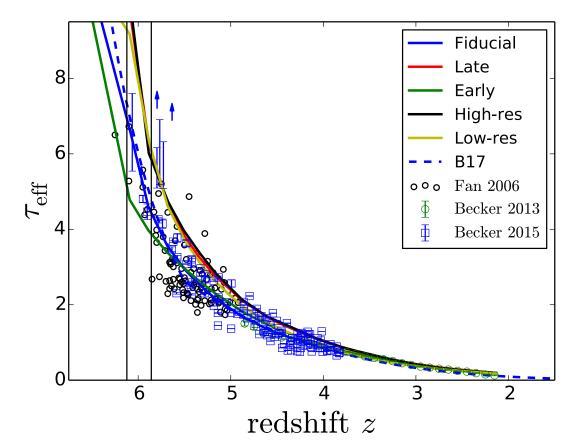
Table 1. Summary of the different simulations studied in this work.

- Five different models from Chardin+2015
- Three 20 cMpc/h boxes with slightly different reionization hisories
- One 40 and one 10 cMpc/h box to test resolution with same reionization history as the Late 20 cMpc/h model
- Post-processed RT simulation with ATON on top of hydro RAMSES simulation
- Calibrated to reproduce well Lyman alpha forest data in the post reionization IGM

The teff evolution in the simulations

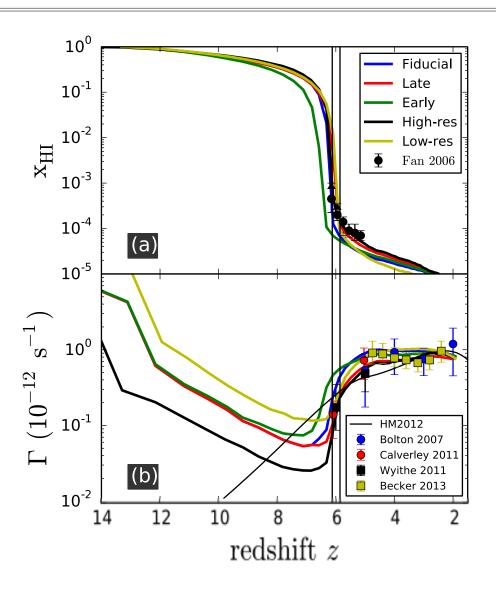
 Evolution of the effective optical depth to Lyman alpha inferred from Barnett et al. 2017 (blue dashed curve)

$$\tau_{\text{GP}}^{\text{eff}}(z) = \begin{cases} 0.85 \left(\frac{1+z}{5}\right)^{4.3} & z \le 5.5 \\ 2.63 \left(\frac{1+z}{6.5}\right)^{11.2} & z > 5.5 \end{cases}$$



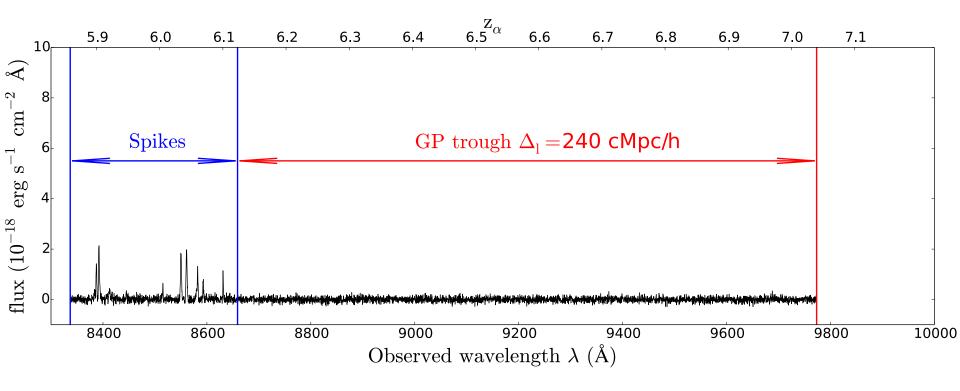
- Fiducial evolution match exactly the Barnett+ power-law
- Other models bracket the inferred evolution from the observations
- Test if we can reproduce the statistic of the observed transmission spikes with the simulations

The reionization histories



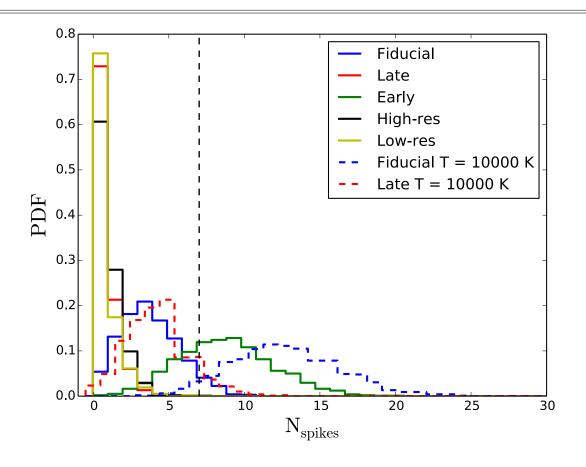
- All models fit well a lot of observational constraints
- Some differences in the redshift range corresponding to the spikes region in the observed spectrum
- Testing how the statistic of spikes vary between the different model could help us to constrain the timing of reionization precisely

Creating mock Lyman alpha forest spectra



- Stitch line-of-sight from different outputs to cover the observed spectrum wavelength range
- Smooth to instrumental resolution + re-bin to 10 km/s + add random Gaussian noise as in the observed spectrum
- Mock spectra resemble the observed one with transmitted spikes and long trough

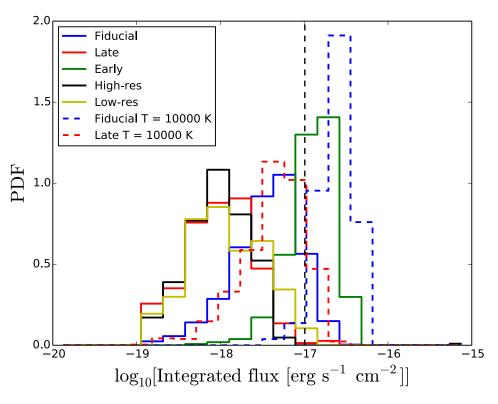
Occurrence of transmission spikes



- Early model best match the data
- Fiducial model still consistent
- Late and low/high res model inconsistent

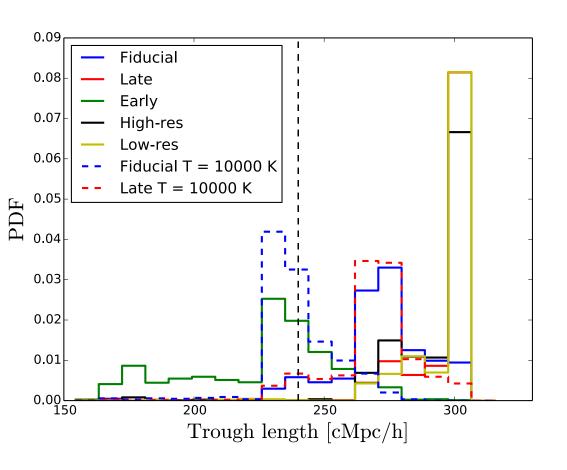
- Changing the IGM temperature can change consequently the results
- Late model could still be consistent with the data

Flux in spikes



- Early model best match the data
- Fiducial model still consistent
- Late and low/high res model inconsistent
- Temperature change the results

GP trough lengths



- Early model best match the data
- Some of the spectra show spikes at higher redshift reducing the length of the trough
- Fiducial model still consistent
- Late low/high res models are not consistent with data
- Increasing temperature in the voids reduces the discrepancies since spikes can be detected at higher redshift

Future works

Simulations

- No QSOs radiation incorporated
- Fluctuation in the photoionization rate will complicate the picture
- Temperature fluctuations would change the results for a given reionization history
- Box size might be too small

Observations

- Now only ~20 QSOs at z>=6.5
- Only the spectrum shown here has such a high resolution enabling to find rare transmission spikes
- Re-observations of already known QSOs spectra would improve the constrain on the reionization history

Conclusion

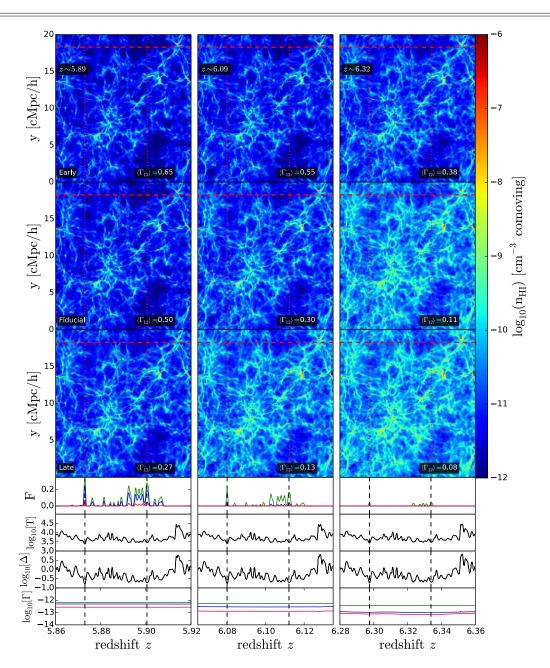
 Our models are successful at reproducing the occurrence of spikes, the integrated transmitted flux and the GP trough length as in the observed QSO spectrum

Our models suggest a somewhat early reionization history

Higher temperatures in under-dense regions favor later reionization history

 Future observation of high resolution QSO spectra could constrain the ionization state of the IGM up to z=6.1 very precisely

Line-of-sights comparison



 Neutral hydrogen density in the three 20 cMpc/h models

Transmitted flux in under-dense regions

 Small Γ variations lead to large differences in the transmitted flux

 Spike statistic should be rather different between the models