

# Census of Ly $\alpha$ , [OIII]5007, Ha, and [CII]158 $\mu$ m Line Emission with 1000 LAEs at $z=4.9-7.0$ Revealed with Subaru/HSC

arXiv:1711.03735

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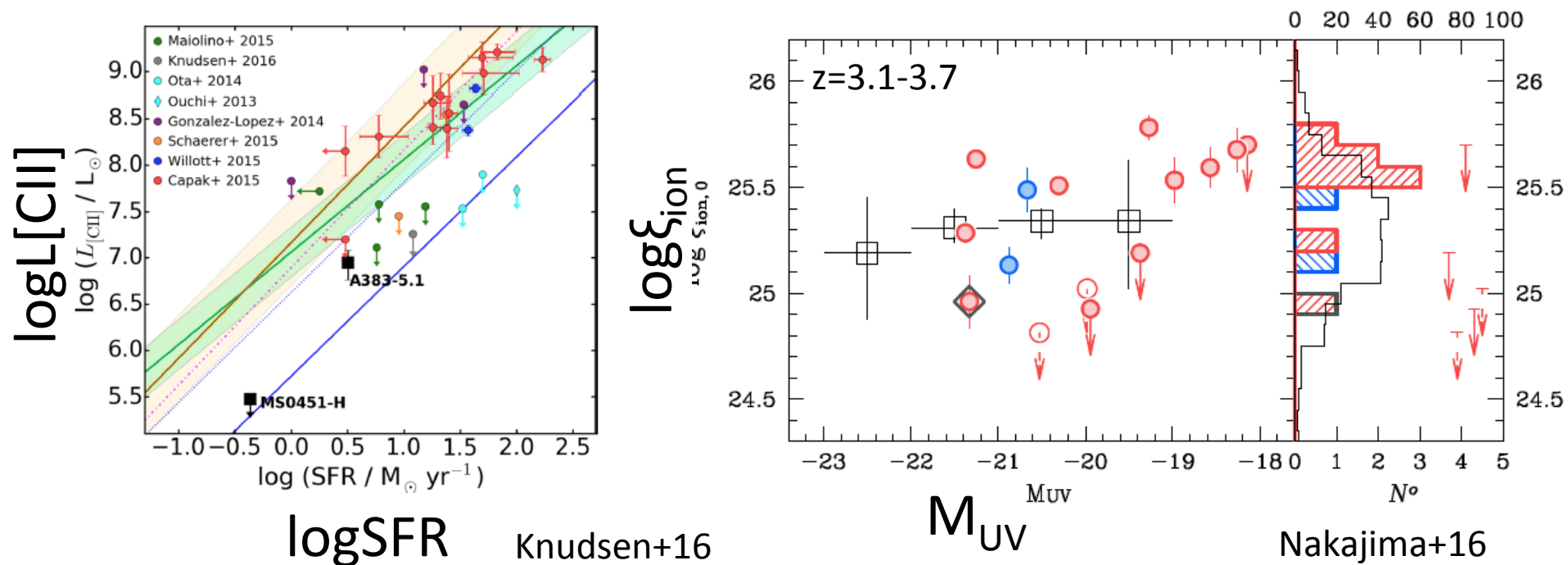
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Ryohei Itoh, Yoshiaki Ono, Ryo Higuchi, Akio K. Inoue, Jacopo Chevallard,  
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# Outline

- Introduction
- Sample
- IRAC excess technique
- Results and Discussion

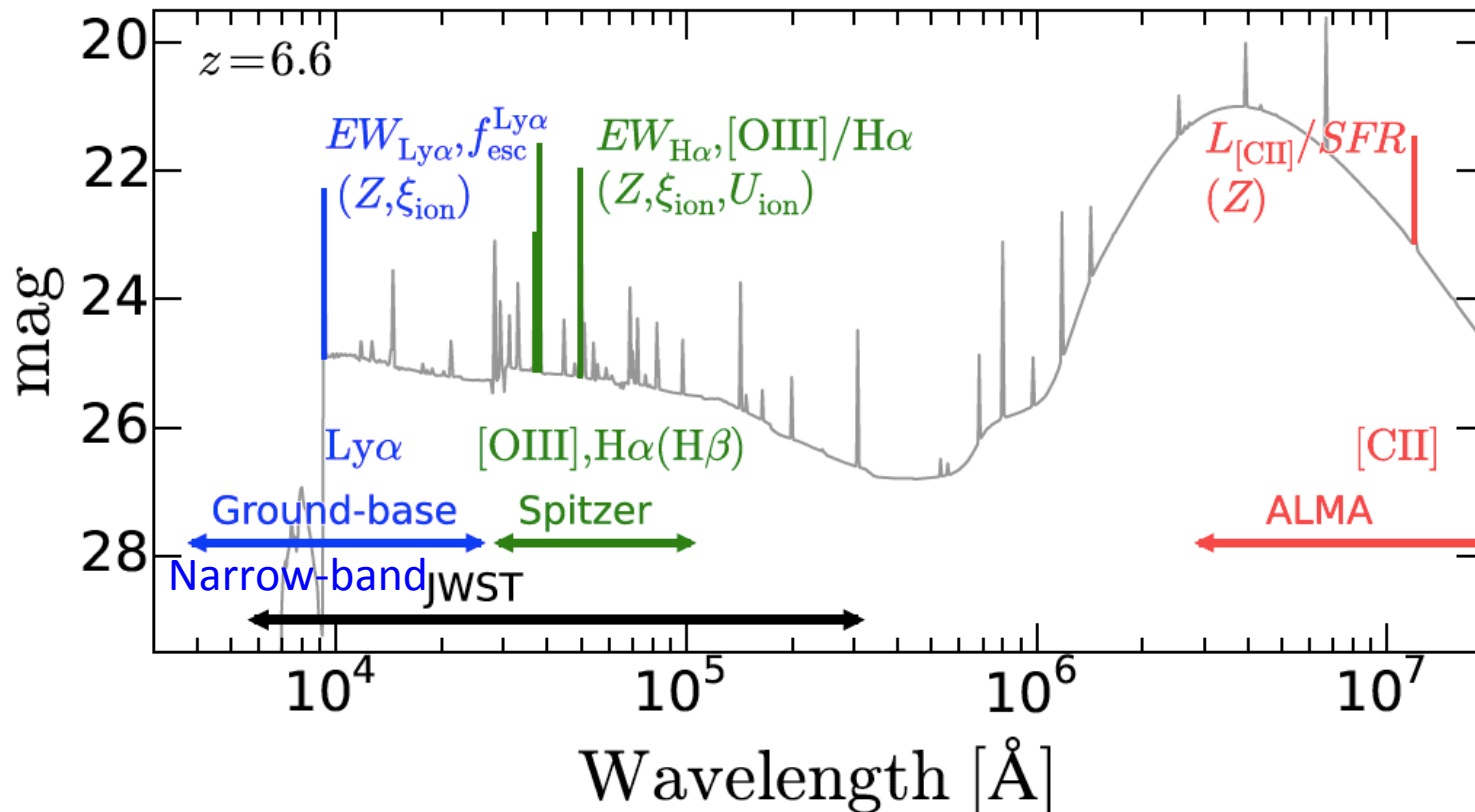
# ISM of $z > 4$ LAEs

- Weaker [CII] emission in high- $z$  galaxies
  - What is the origin?
- Reionization by star-forming galaxies (LAEs)?
  - Few measurements of  $\xi_{\text{ion}}$  for LAEs at  $z > 4$



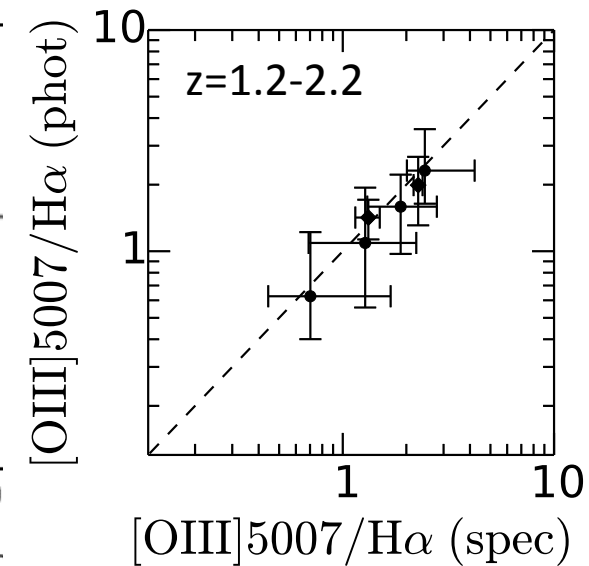
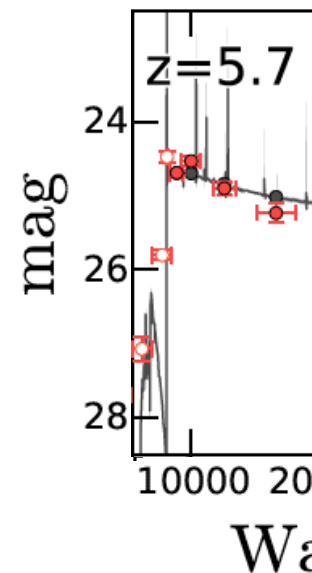
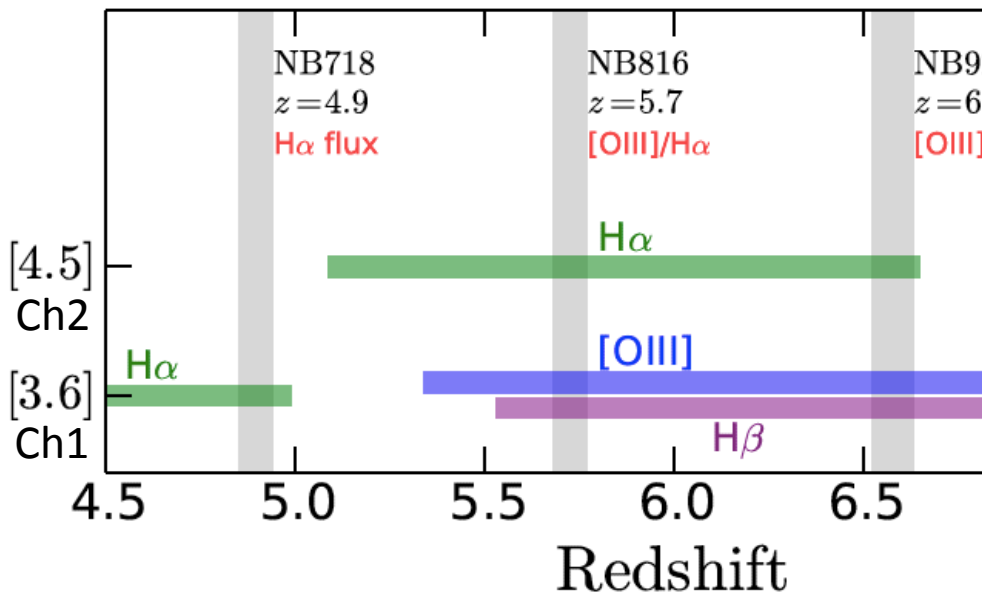
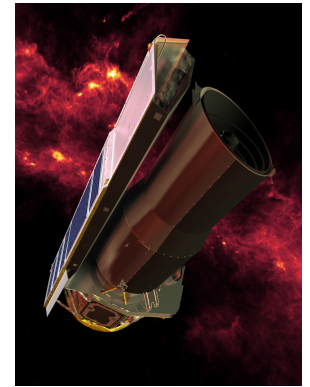
# Our Sample: Subaru/HSC

- 1092 LAEs with Spitzer (Shibuya+18, Itoh+, Zheng+ in prep.)
  - NB816, NB921 (HSC-SSP), NB718, NB973 (CHORUS)
- 34 galaxies with ALMA/PdBI [CII] data (literature)



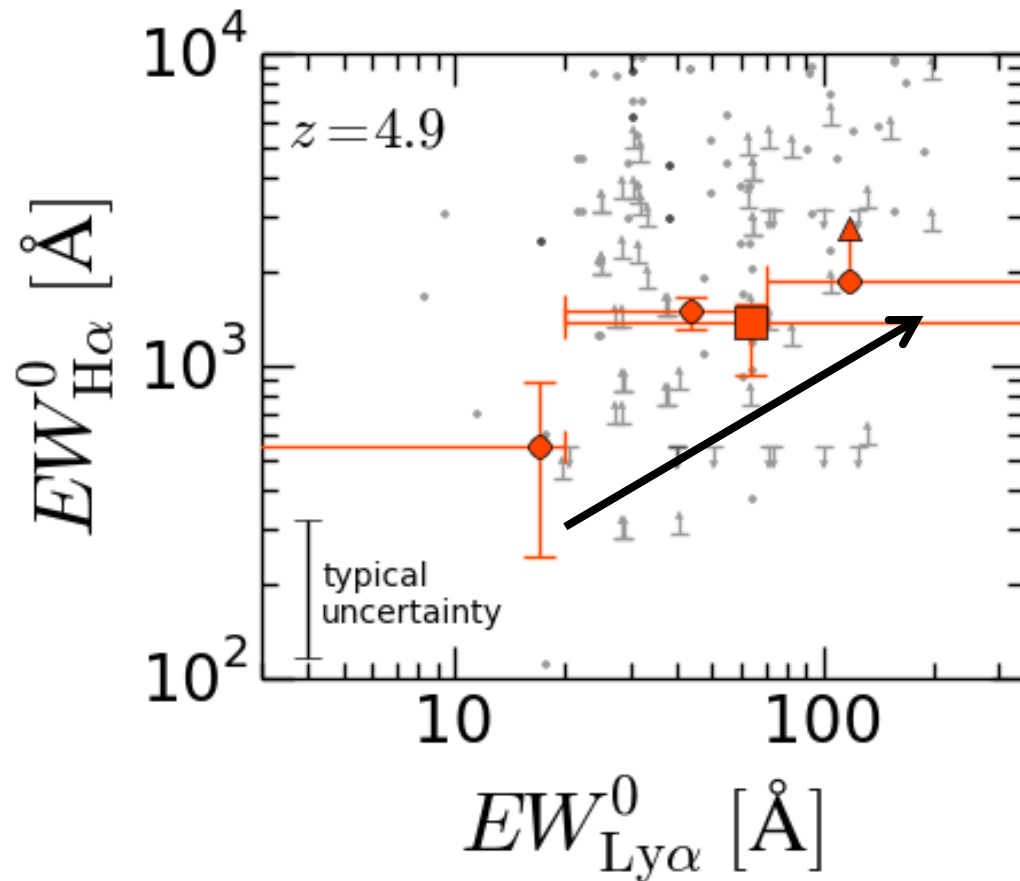
# Spitzer/IRAC Excess Technique

- IRAC excess by strong rest-optical emission
  - e.g., Roberts-Borsani+16, Stark+17
- Constraining H $\alpha$ , [OIII], H $\beta$  at  $z=4.9-7.0$ 
  - SPLASH data (PI: P. Capak)



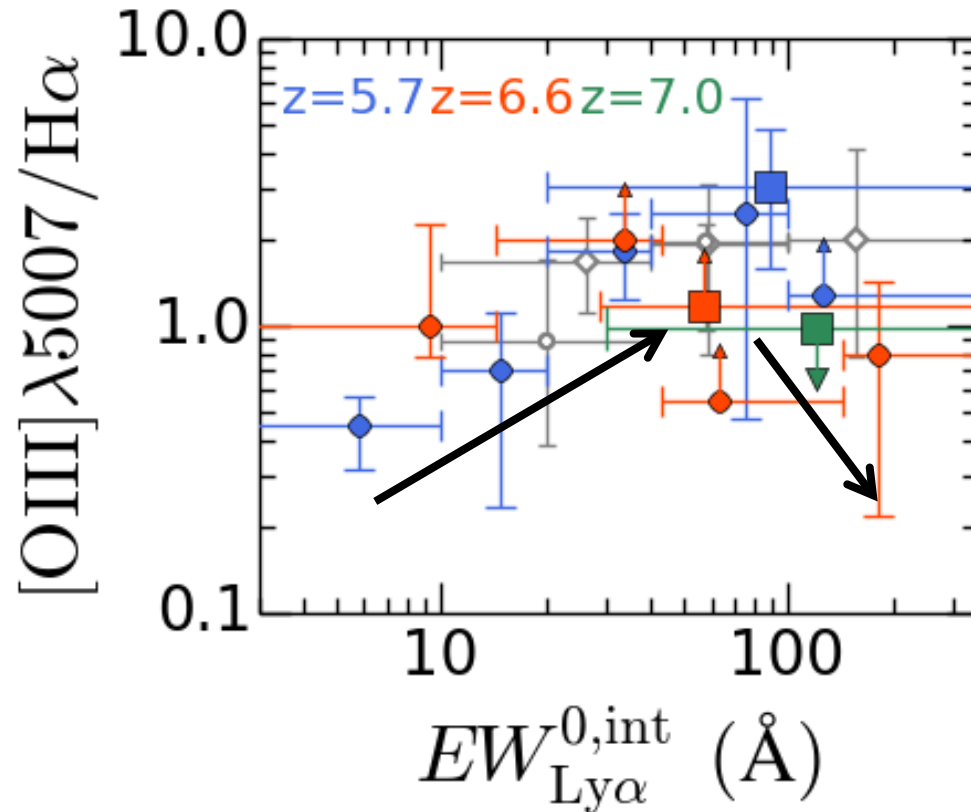
# Result 1: Ha EW

- Ha EW positively correlates with Ly $\alpha$  EW
  - $EW_{\text{Ha}} > 2000 \text{ \AA}$  with  $EW_{\text{Ly}\alpha} \sim 100 \text{ \AA}$



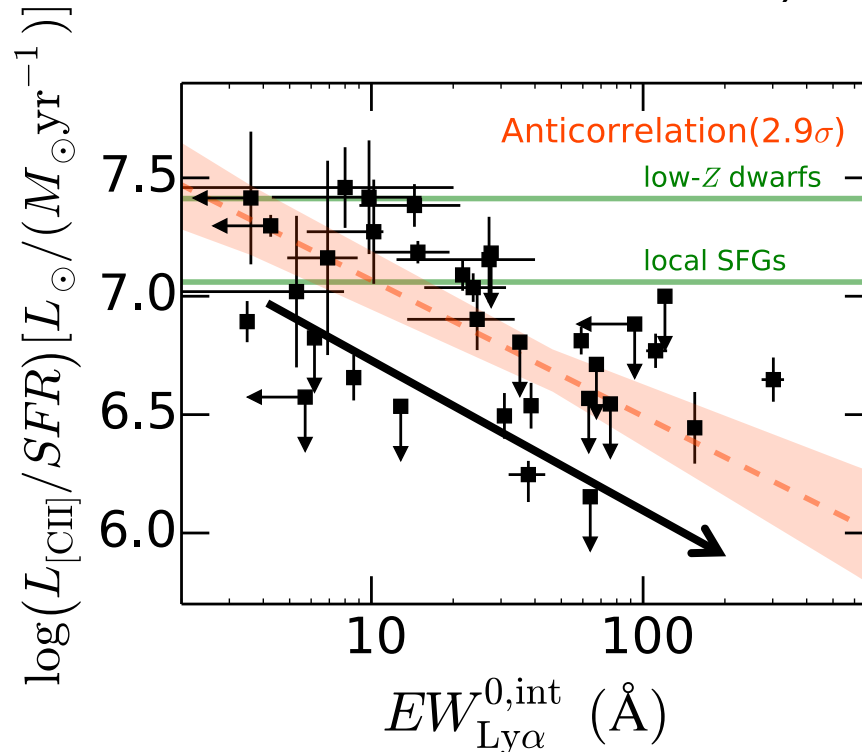
# Result 2: [OIII]/Ha

- Turn-over trend in [OIII]/Ha- $EW_{Ly\alpha}$  plane
  - Indicating low-metallicity in high- $EW_{Ly\alpha}$  LAEs



# Result 3: [CII]-Ly $\alpha$

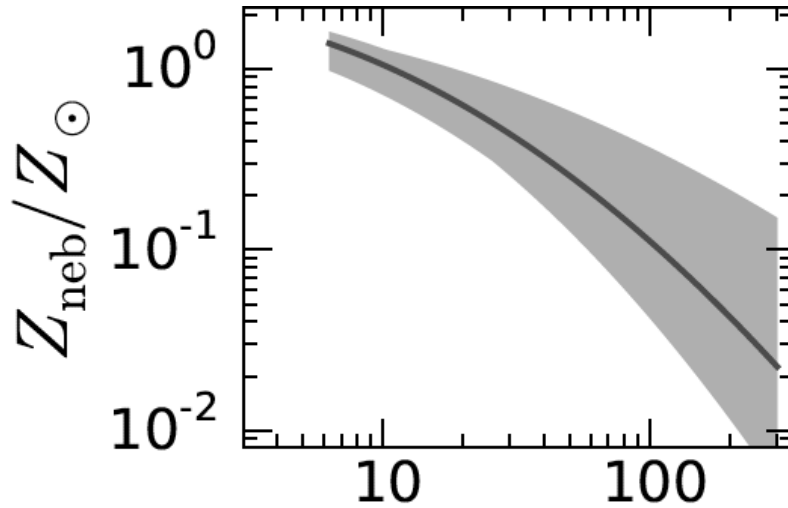
- [CII]/SFR negatively correlates with Ly $\alpha$  EW
  - Including results of Carniani+17, 18
  - Confirming [CII] deficit in LAEs for the first time  
(see also discussion in Pentericci+16)





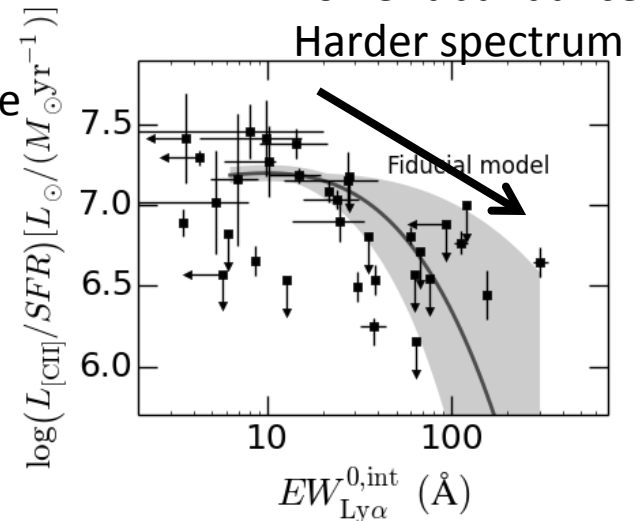
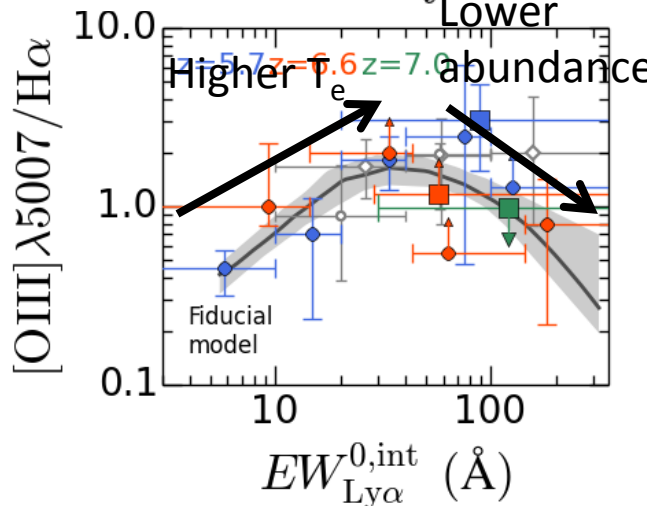
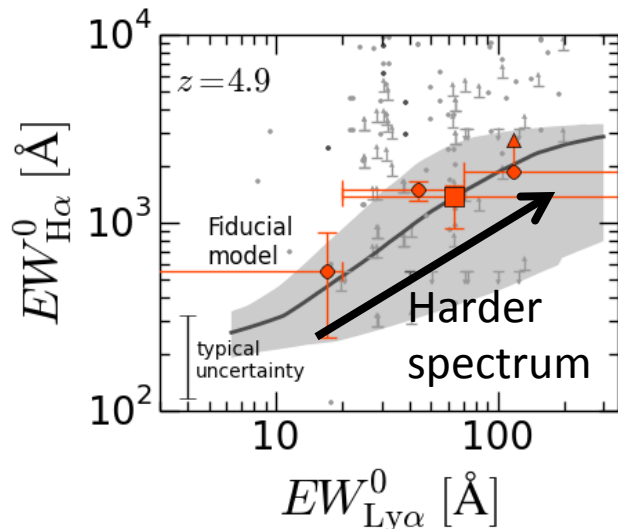
# Metallicity-EW Ly $\alpha$ Anti-Correlation

- BEAGLE (Cenarro et al. 2016)
  - Stellar synthesis models
  - $-2 < \log(Z/Z_{\odot}) < 1$
- Metallicity-EW Ly $\alpha$  anti-correlation
- three relations



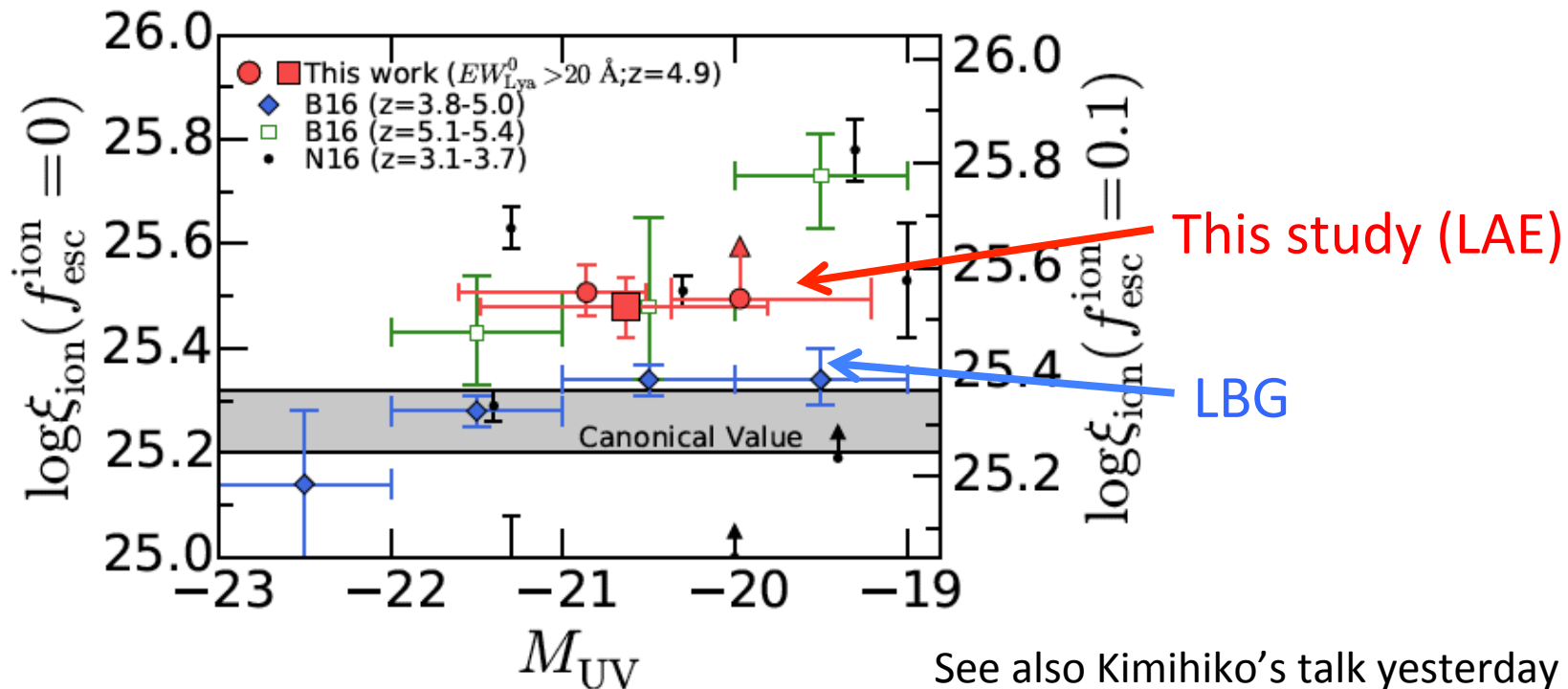
models  
 age/yr < 9.1  
 can explain the

$EW_{Ly\alpha}^0$  ( $\text{\AA}$ )



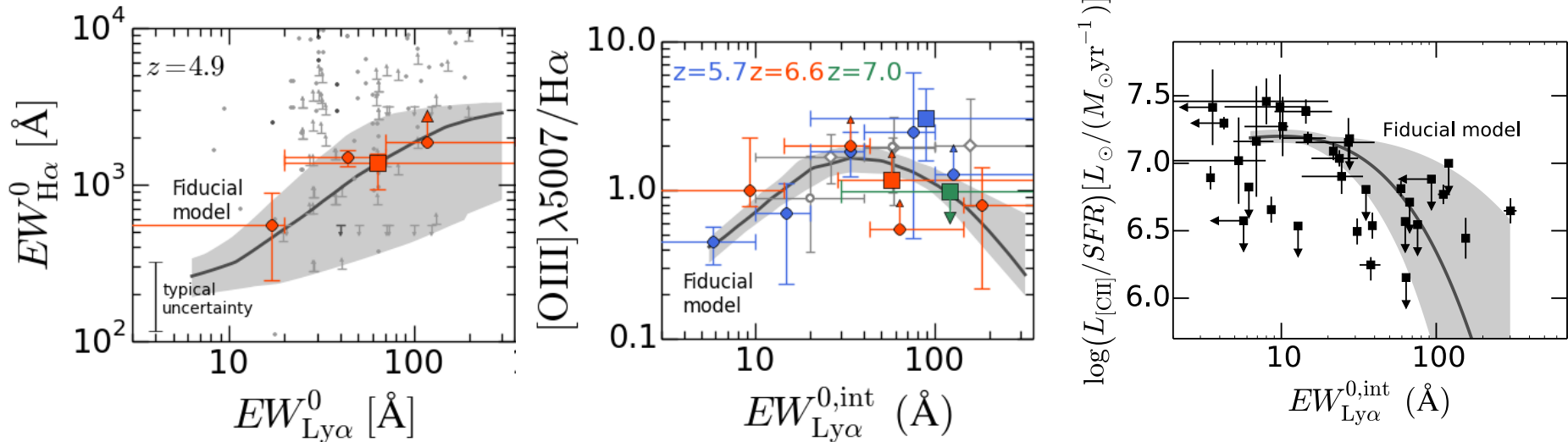
# Implication for Cosmic Reionization

- Ionizing photon production efficiency  $\xi_{\text{ion}}$ 
  - $\xi_{\text{ion}} = N(\text{H}^0)/L_{\text{UV}}$ ,  $L_{\text{Ha}} = 1.36 \times 10^{-12} N(\text{H}^0) (1 - f_{\text{esc}})$
- $\log \xi_{\text{ion}} \sim 25.53$ , higher than LBGs by 60-100%
  - Enough ionizing photons ( $f_{\text{esc}} = 10\%$ , UVLFs)



# Summary

- ISM properties from  $\sim 1000$  LAEs at  $z=4.9-7.0$



- Metallicity- $EW_{Ly\alpha}$  anti-correlation can explain all results.
  - LAE: metal poor and high  $\xi_{ion}$

