

Probing Cosmic Reionization at $z \sim 7$ with Ly α Galaxies from the LAGER Survey

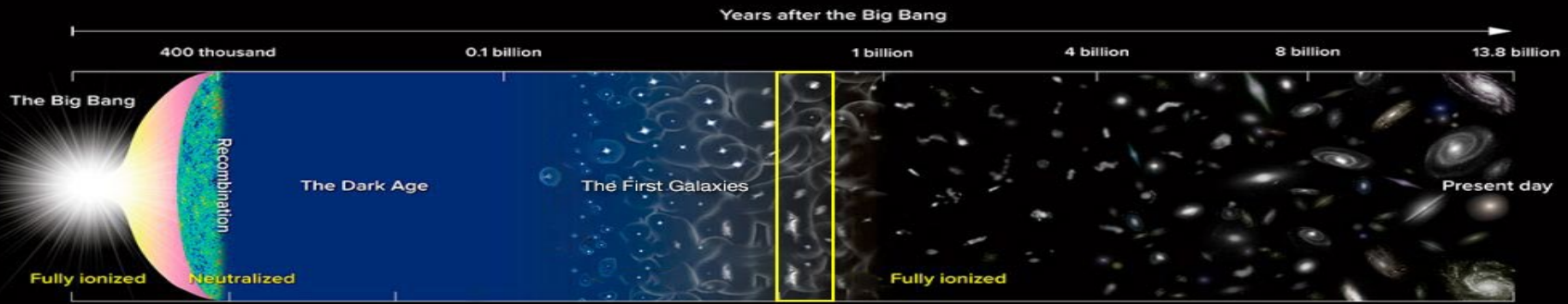
LAGER: Lyman Alpha Galaxies in the Epoch of Reionization

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Shanghai Astronomical Observatory

On behalf of the LAGER Team

**Tokyo Spring Cosmic Lyman-Alpha Workshop
(Sakura CLAW)
March 26-30, 2018**



Lyman Alpha Galaxies in the Epoch of Reionization (LAGER)

CHINA

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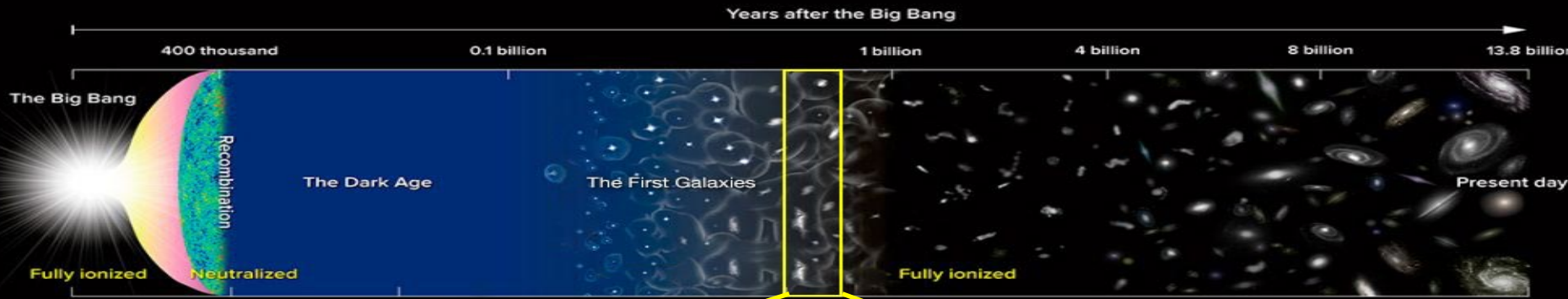
USA

Sangeeta Malhotra (ASU, GSFC)*,
James Rhoads (ASU, GSFC)*,
 Alistair Walker (CTIO),
 Alicia Gonzalez (ASU),
 Vithal Tilvi (ASU),
Steven Finkelstein (U. Texas), ...

CHILE

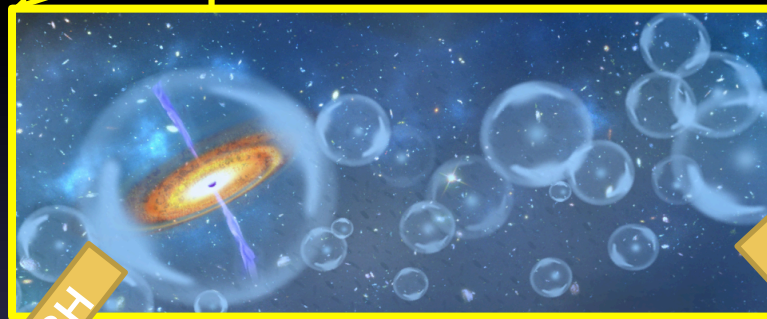
Leopoldo Infante (LCO, PUC)*,
 Felipe Barrientos (PUC),
Huan Yang (LCO),
 Pascale Hibon (ESO),
 Gaspar Galaz (PUC),
 Franz Bauer (PUC), ...





1. Reionization History
2. Topology
3. Contributing Sources

Epoch of Reionization



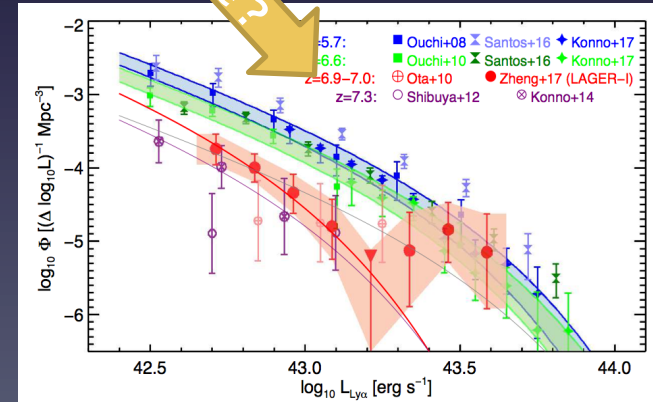
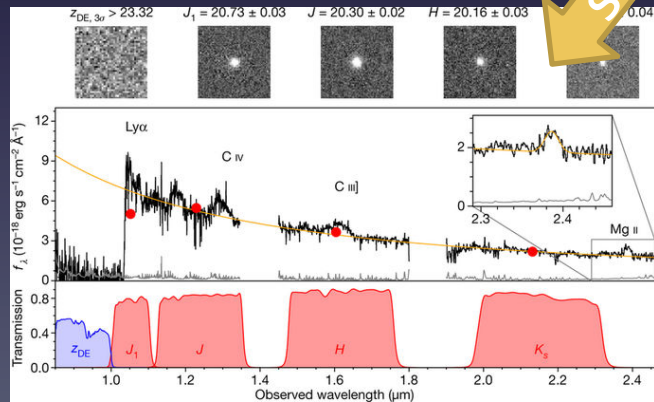
Banados+2017,
J1342+0928, $z=7.54$

Zheng+2017, $z=6.94$
Bright-end Excess
LAGER-COSMOS

SMBH

Bubble?

Quasars,
GRBs,
Galaxies,
21cm signals,
CMB.



Probing Reionization with High-z Galaxies

◆ Lyman- α luminosity function (LF) test

(Rhoads & Malhotra 2001, Malhotra & Rhoads 2004, Ouchi+2010, Kashikawa+2011, Konno+2014, Ota+2017, Zheng+2017, Konno+2017)

□ Lyman- α visibility test

(On LBGs, e.g., Stark et al. 2011, Schenker+14, Bian+2014, Tilvi+2016, etc.; On Ly α EW distribution of LAEs, e.g., Choudhury+2015, Kakiichi+2016)

◆ Clustering test of Lyman- α galaxies

(Theory: Furlanetto+2006, McQuinn+2007, Jensen+2013; Observation: Ouchi+2010, Ouchi+2017)

◆ Volume test (Malhotra & Rhoads 2006, 2018)

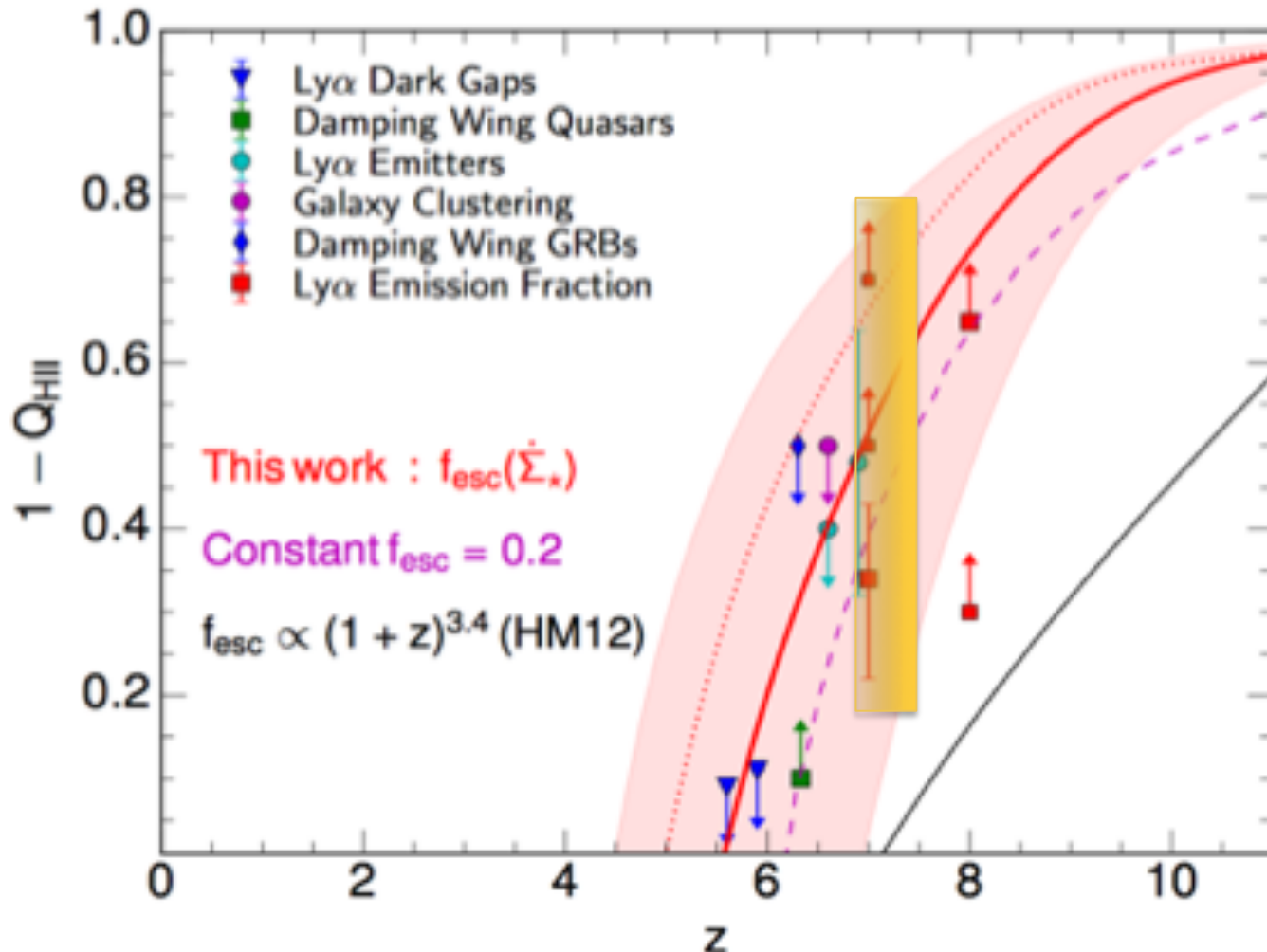
❖ GRB & QSO spectroscopy at $z > 6$

(e.g., Salvaterra+2015, Banados+2017)

Zhenya Zheng @ SHAO

Image Credit: Jón Pálmason

Global History of Reionization



Sharma+16, Robertson+15, Mitra+15, Bouwens+15, Kakiichi+16, Greig & Mesinger17 ...

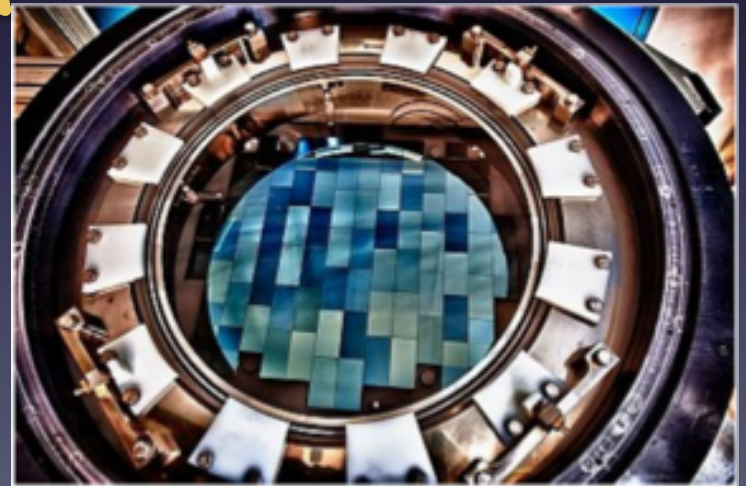
The LAGER Project:

NB964
Narrowband
Filter

$\lambda_c = 9642 \text{ \AA}$ & $\text{FWHM} = 92 \text{ \AA}$
 $\rightarrow z(\text{Ly-}\alpha) = 6.93 \pm 0.04$

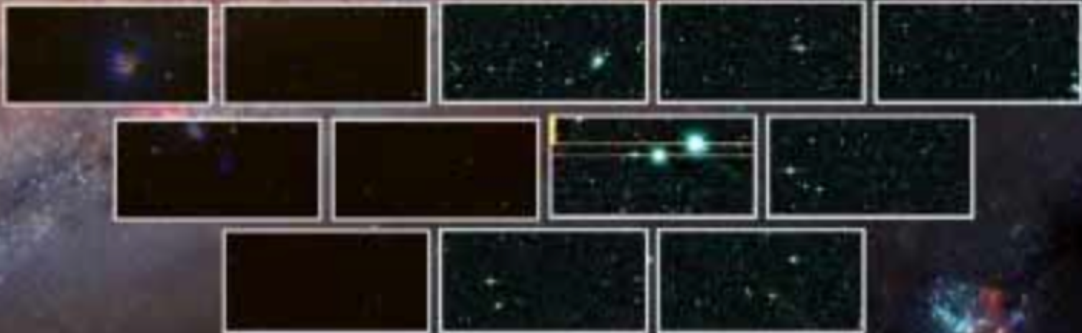
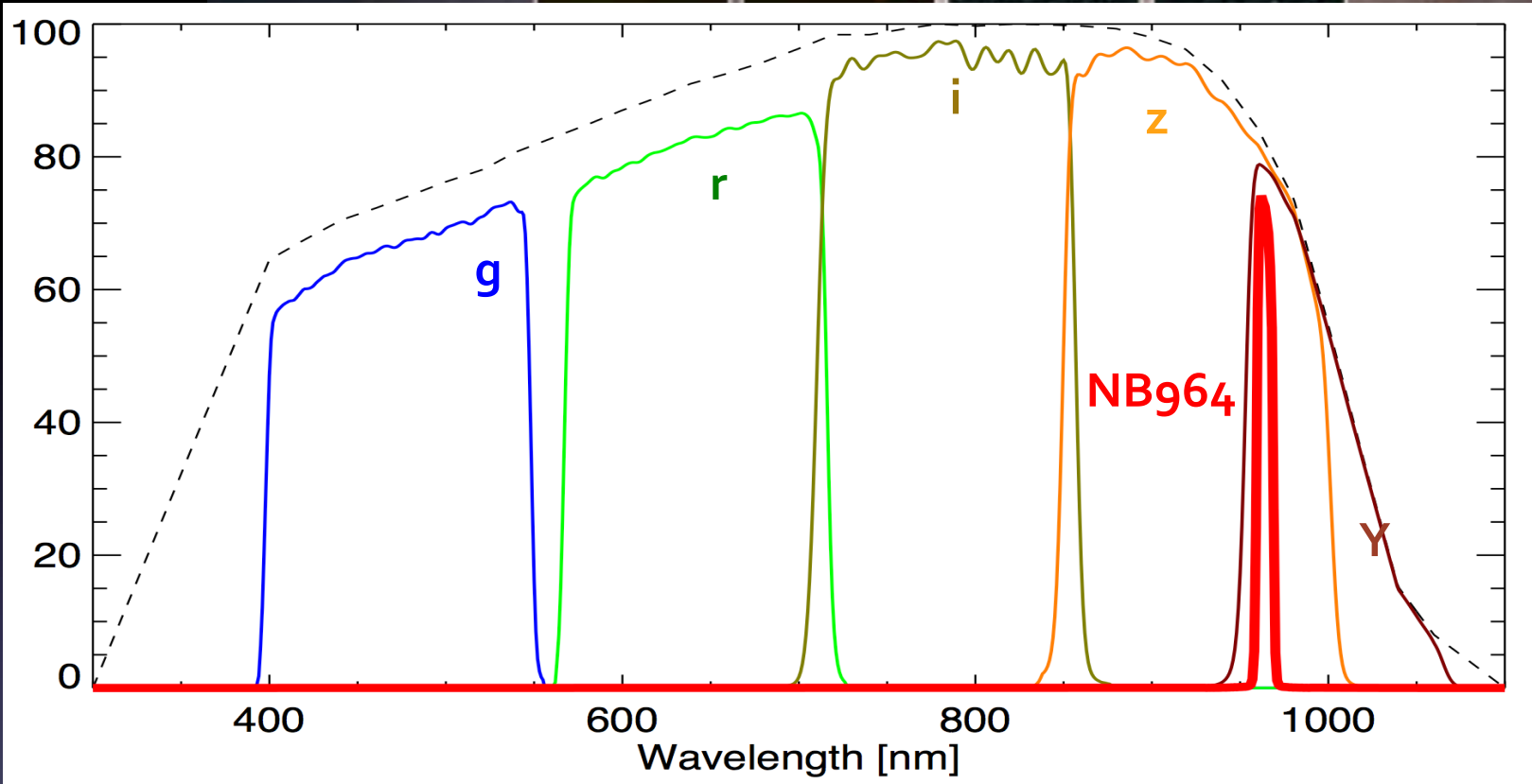


CTIO 4m Blanco Telescope
(Cerro Tololo, Chile)

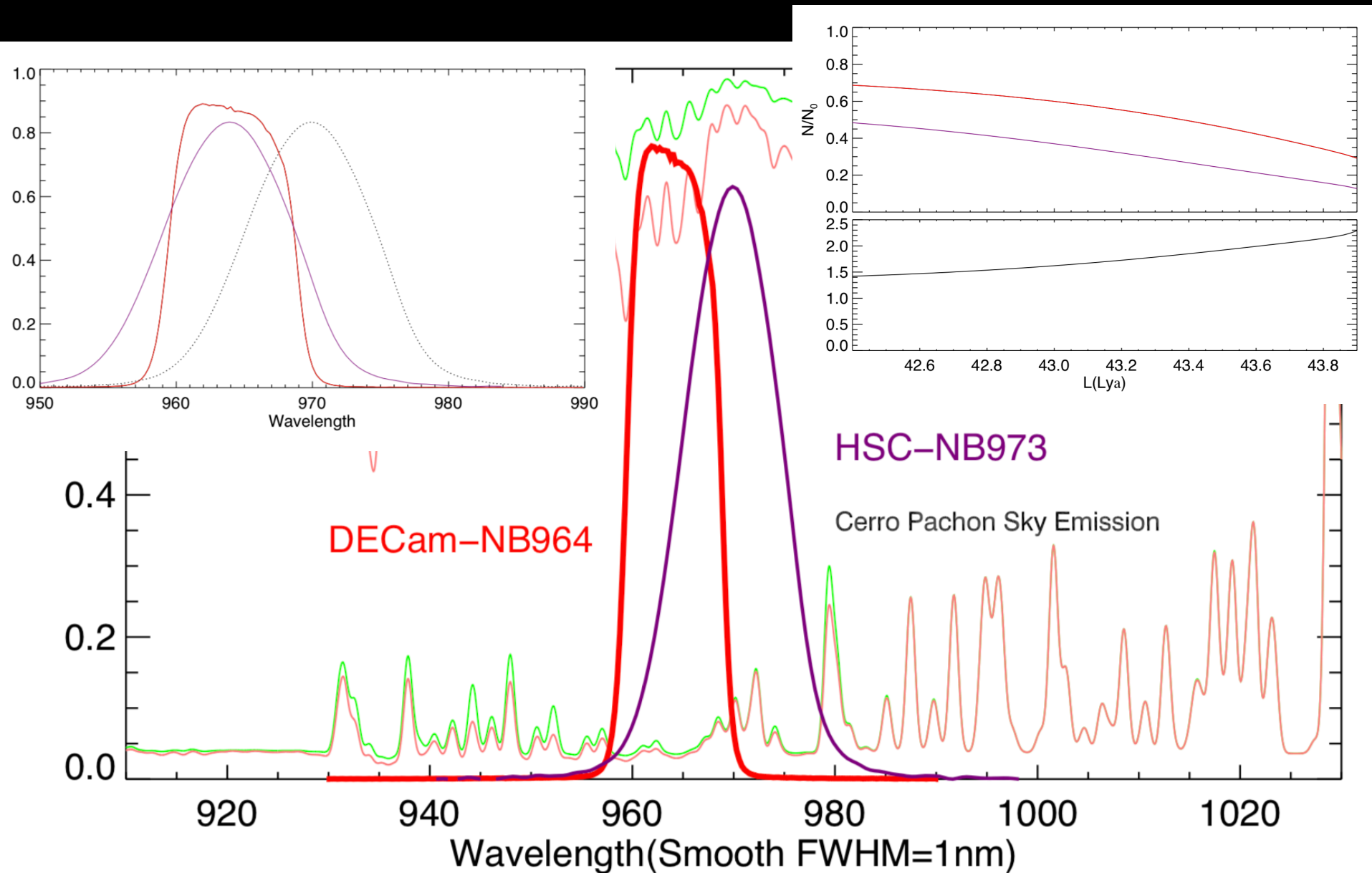


Dark Energy Camera DECam
(FOV = 3 sq-deg)

Why DECam?



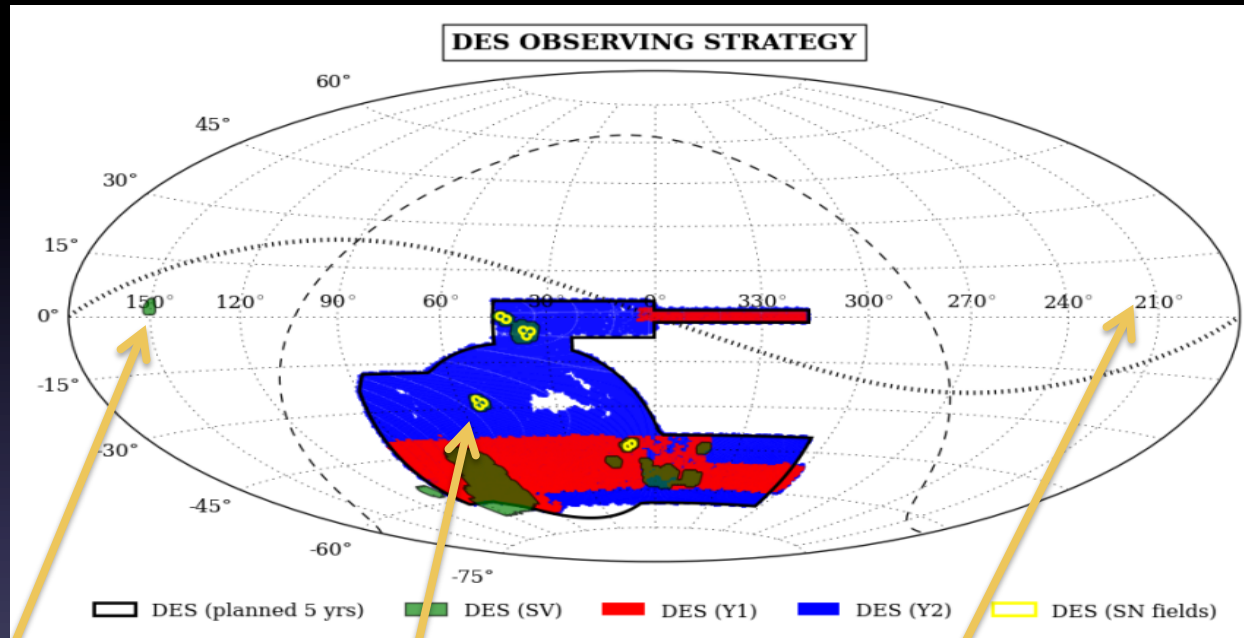
NB964 Filter Profile vs. Sky Lines



LAGER Runs

Summary (2015 Dec. – 2018 Mar.):

- **21 nights** awarded & Observed with DECam & NBg64, but 5 with bad weather.
- **4 fields** with NBg64
T_{exp} ~ **20-40 hrs each.**



Fields	COSMOS	CDF-S	GAMA15a,b
NB Exp. T	44 hrs	35 hrs	20 hrs x 2
Date Obs.	Dec. 7 15; Feb. 4-9 16; Mar. 9-12 16; Dec. 24-27 17	Dec. 7 15; Mar. 9-12 16; Nov. 25-26 16	May 24-29 17 (w. bad weather); Mar. 5-10 18

First results from LAGER: LAGER-COSMOS in 34 hrs NB964

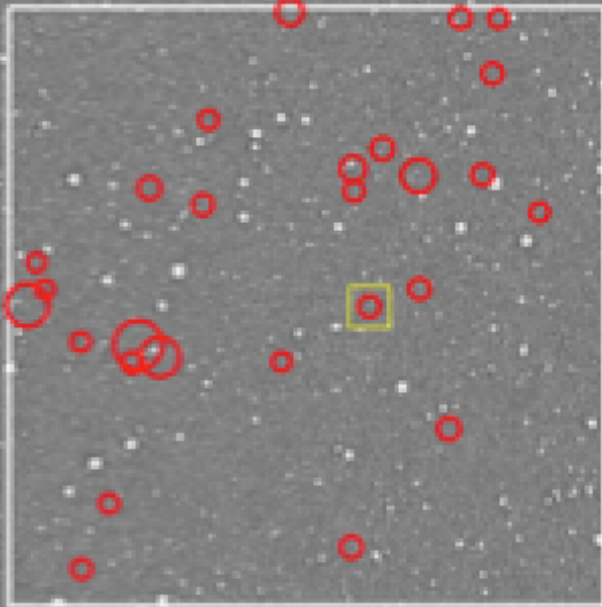
Inhomogeneous Distribution of
23 $z \sim 7$ LAEs in LAGER-COSMOS:

Patchy Reionization?

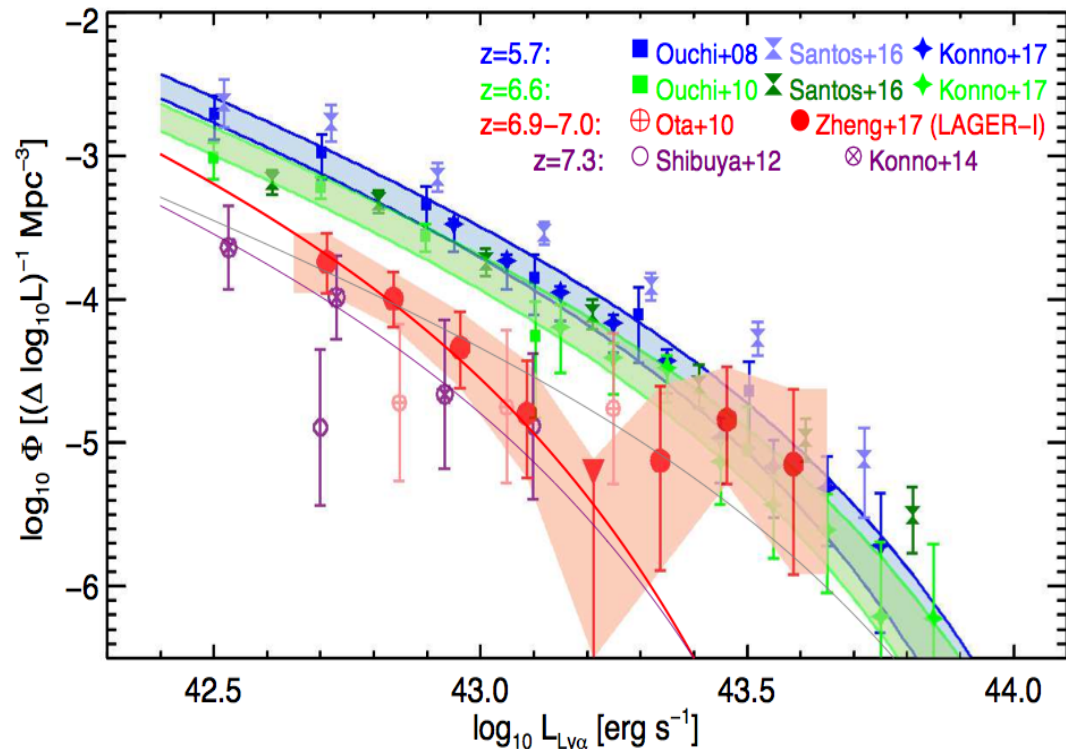
- Little Evolution of Ly α LF at $z \sim 3-6$:
(Ouchi+08, Faisst+2014, Zheng+2016, ..)

At $z \sim 7$ (Zheng+2017) :

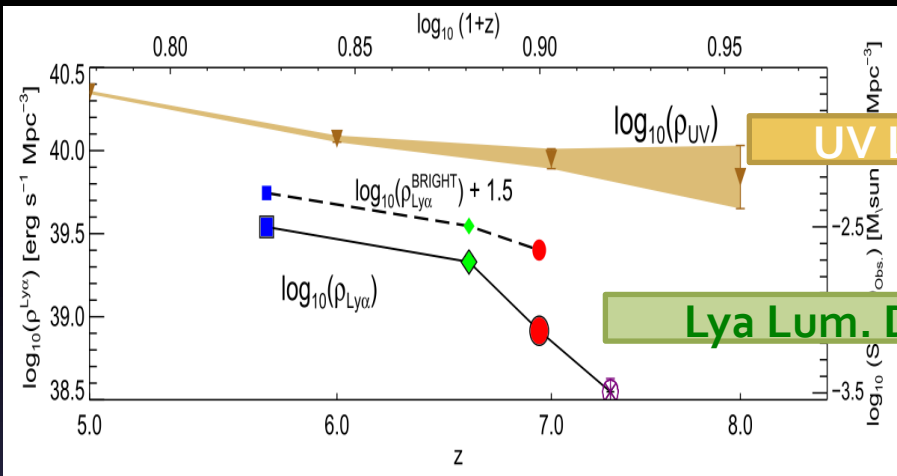
1. Different Evolution at Bright & Faint Ends.
2. Bright-End Excess.



Ly α emitters at $z \sim 7$



1. Probing IGM Neutral Fraction at z~7

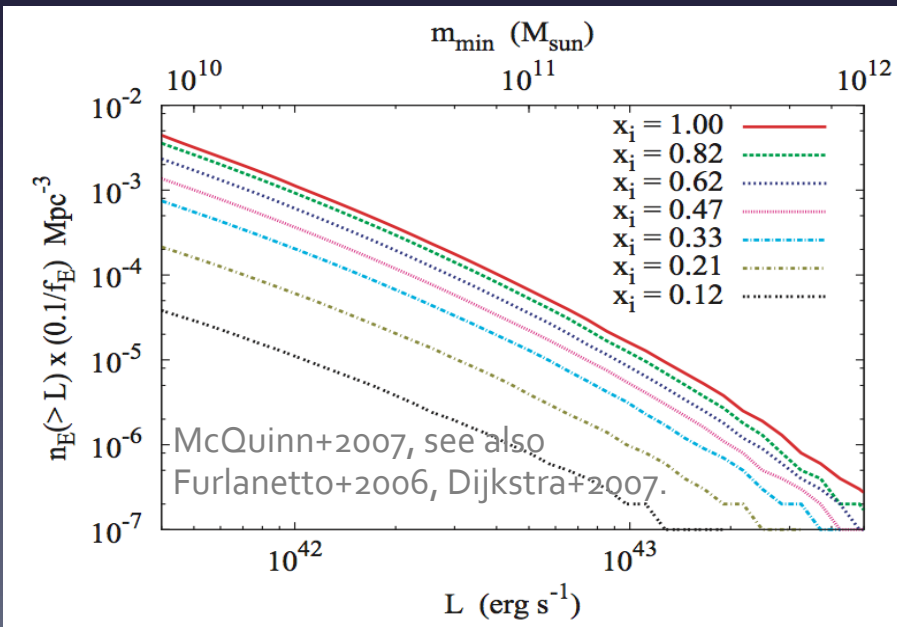


UV Lum. D.

UV LD represents the Galaxy Evolution

Lya Lum. Density

Lya LD represents the joint effect of
1) Galaxy Evolution &
2) resonant scattering by neutral IGM.



IGM Transmission factor T':
(UV from Finkelstein+15)

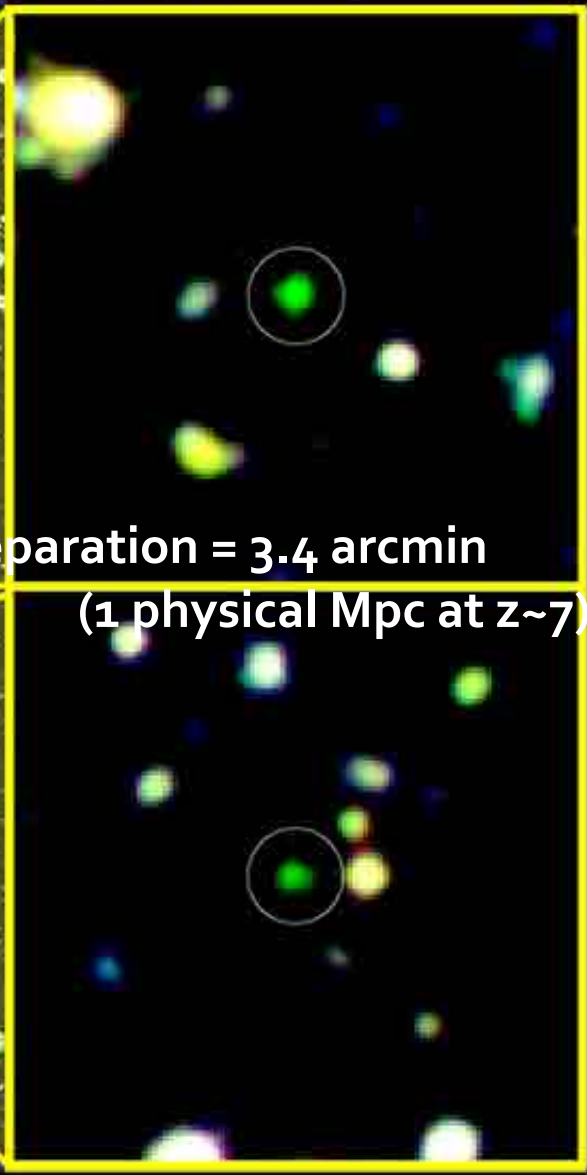
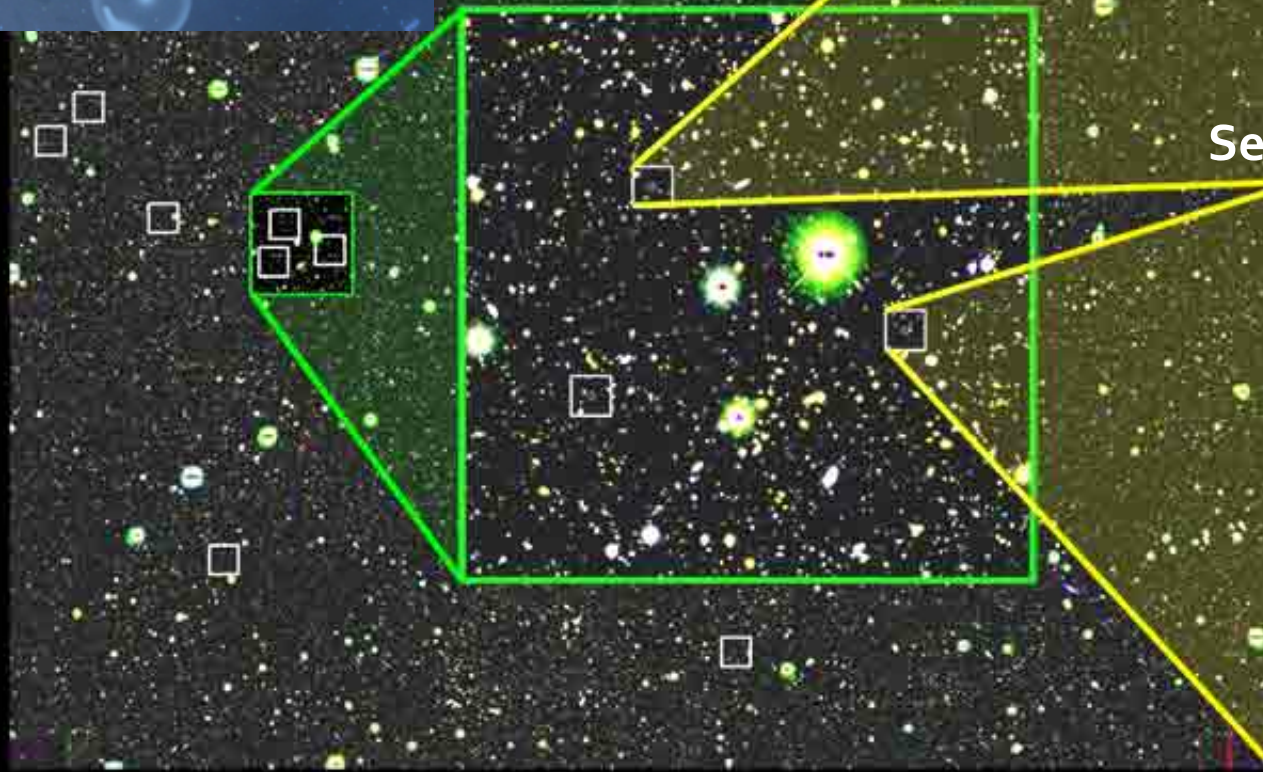
$$T'_{z=6.9} = \frac{T_{Ly\alpha, z=6.9}^{IGM}}{T_{Ly\alpha, z=5.7}^{IGM}} = \frac{\rho_{z=6.9}^{Ly\alpha, tot} / \rho_{z=5.7}^{Ly\alpha, tot}}{\rho_{z=6.9}^{UV} / \rho_{z=5.7}^{UV}}$$

T' = 0.37 →
x_HI ~ 0.4-0.6 at z~7



2. Bubbles at $z \sim 7$?

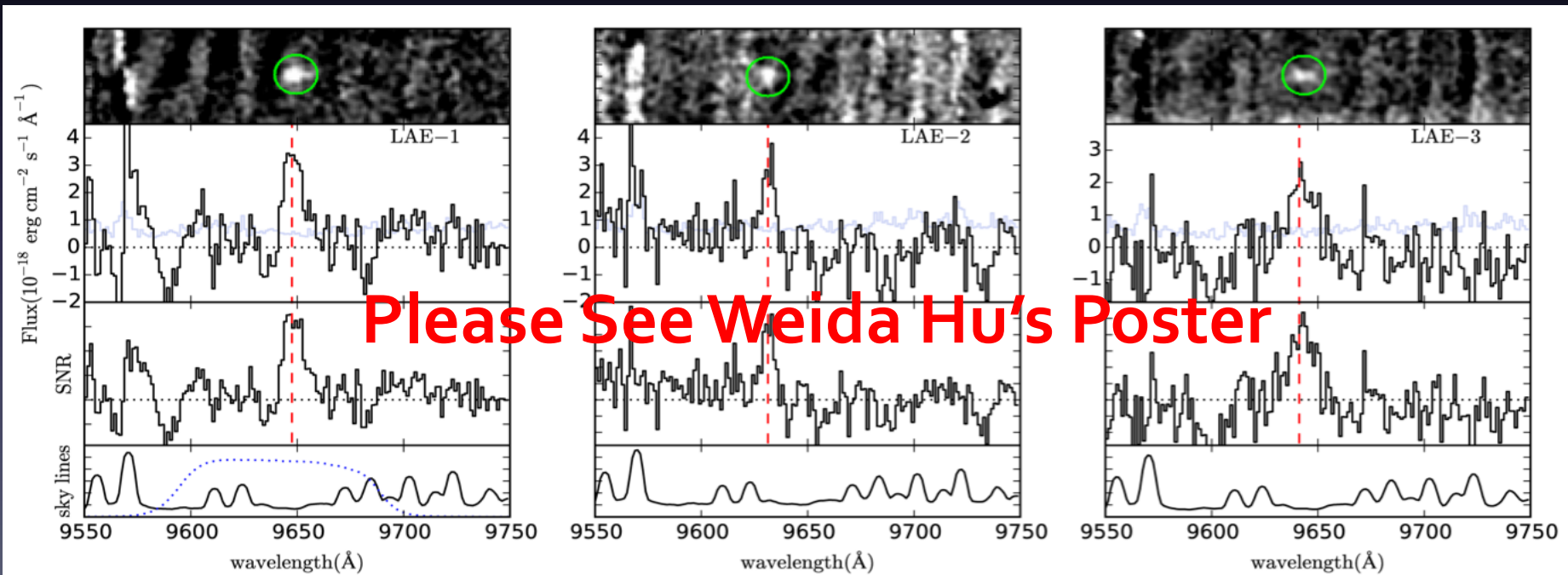
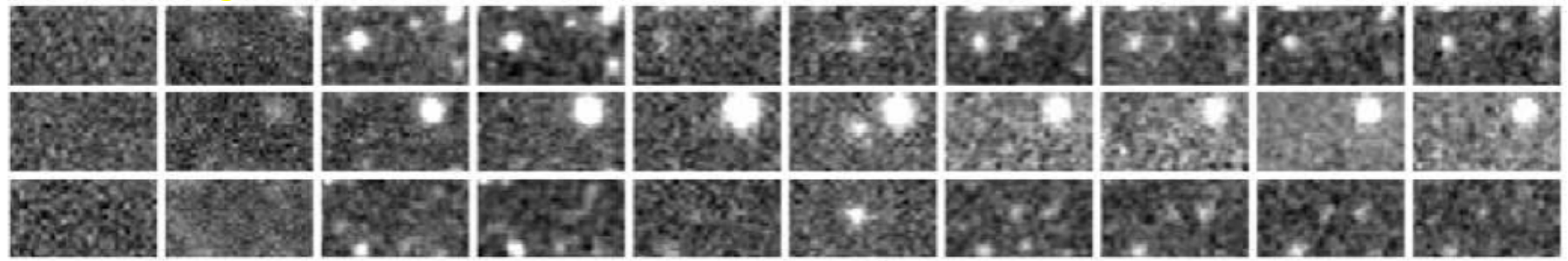
LAGER-COSMOS
2 sq-deg • 23 Ly- α galaxies at $z \sim 7$



Separation = 3.4 arcmin
(1 physical Mpc at $z \sim 7$)

Spec. Confirmation of the 3 Brightest LAEs

u g' r' i' z NB964 Y J H Ks



6 hrs Magellan/IMACS obs. on Feb 6-8, 2017 (FWHM~200-300 km/s)

Hu et al., 2017, ApJL, 845, L16 (arXiv:1706.03586).

A faint-AGN at z~7 ?

Tokyo Spring Cosmic Lyman-Alpha Workshop (Sakura CLAW)
 March 26-30, 2018 The University of Tokyo, Hongo, Japan

Science Topics

- Lyman-alpha emission as a probe of galaxy formation and evolution
- Lyman-alpha emission and absorption as a probe of the ISM/CGM/IGM/LSS
- Lyman-alpha emission as a probe of reionization

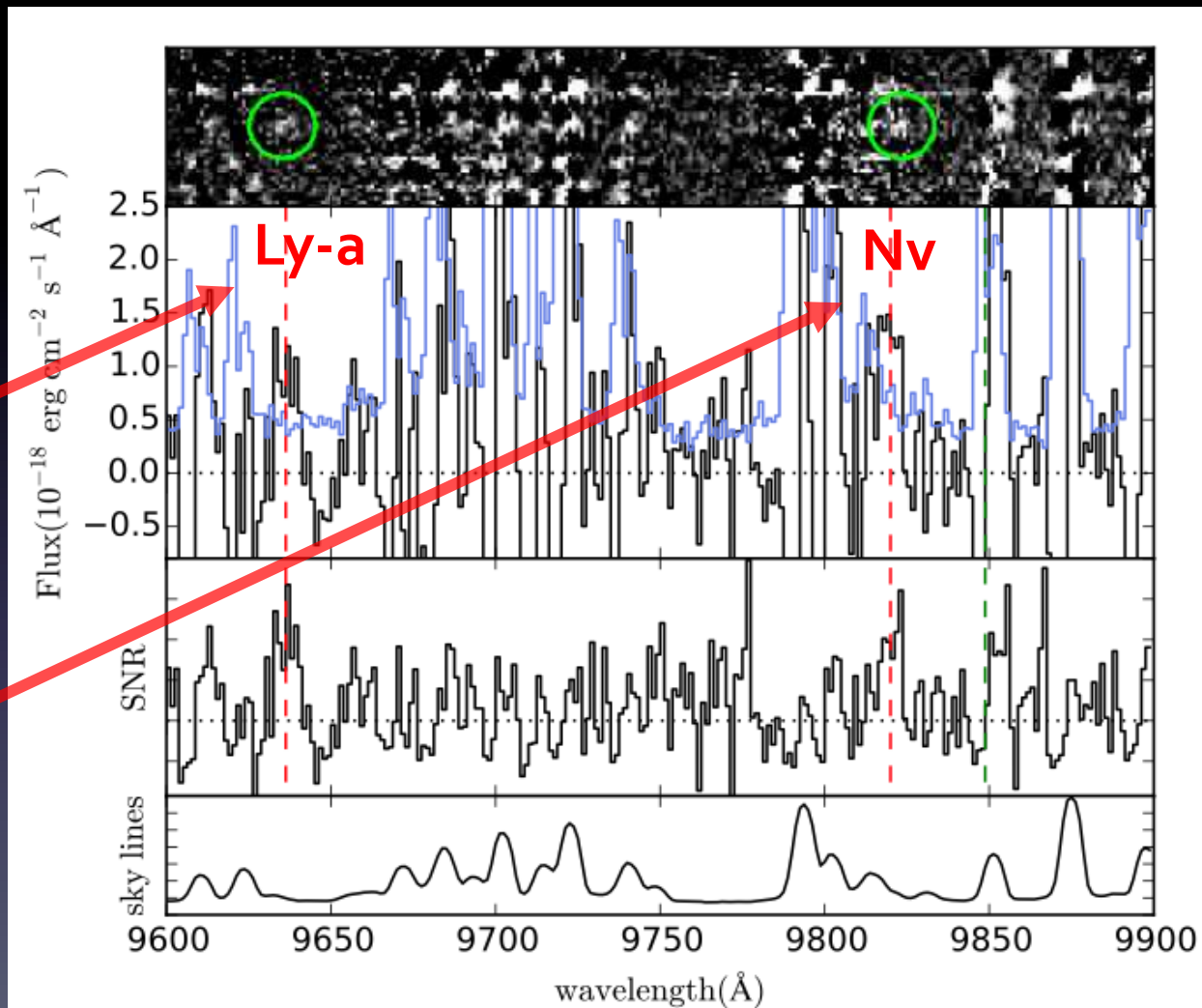
Confirmed Speakers

Roland Bacon	(Iyon)
Jeremy Blaizot	(CRAI)
Sebastiano Cantalupo	(ETH Zurich)
Len Cowie	(Hawaii)
Daren Erb	(Wisconsin)
Andrea Ferrara	(Ipsa)
Steve Finkelstein	(UT Austin)
Matthew Hayes	(Stockholm)
Shao-Gan Lee	(BNL)
Sergio Malhotra	(Arizona State)
Chris Martin	(Caltech)
Masumi Ouchi	(Tokyo)
Dan Stark	(Arizona)
Tommaso Trieloff	(Ipsa)
Anne Verhamme	(Geneva)
Hidenori Wajima	(Tohoku)

Scientific Organizing Committee

Len Cowie	(Hawaii)
Andrea Ferrara	(Ipsa)
Anne Jaekel	(Massachusetts)
Yoshi Matsuda	(NAGI)
Masumi Ouchi	(Tokyo) [Chair]
Lauro Parenti	(Rome)
Alice Shapley	(Ipsa)
Anne Verhamme	(Geneva) [Co-Chair]
Lutz Wisotzki	(Potsdam)
Zheng Zhang	(Urb) [Co-Chair]
Zhenya Zheng	(Shanghai)

http://cos.icrr.u-tokyo.ac.jp/lya_conference



6 hrs Magellan/IMACS obs. on Feb 6-8, 2017 (FWHM~200-300 km/s)

Hu et al., 2017, ApJL, 845, L16 (arXiv:1706.03586).

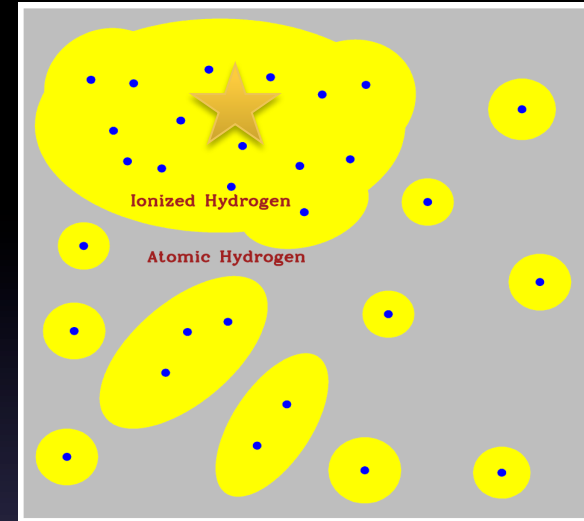
Bubbles at $z \sim 7$?

– A group of galaxies?

- Bright-end excess of UVLF at $z \sim 7$ (Bowler+2014)
- Bright LAEs from HST grism survey (Bagley+2017)
- HSC Ly α LF at $z \sim 6.6$ (Konno+2017)
- LAGER Ly α LF at $z \sim 7$ (Zheng+2017)

– Strong ionizing power?

- Faint AGN? (Matsuoka+2016, Jiang+2016, Hu+2017)
- PopIII G., Sobral+15; or DCBH, Smith+16. (CR7 is ruled out Shibuya+2017).
- Strong instant SF? (Mainali+2017, high ionized CIV & OIII] at $z=6.11$.)
- Large velocity offsets for Ly α (Willott+2015, Mainali+2017)

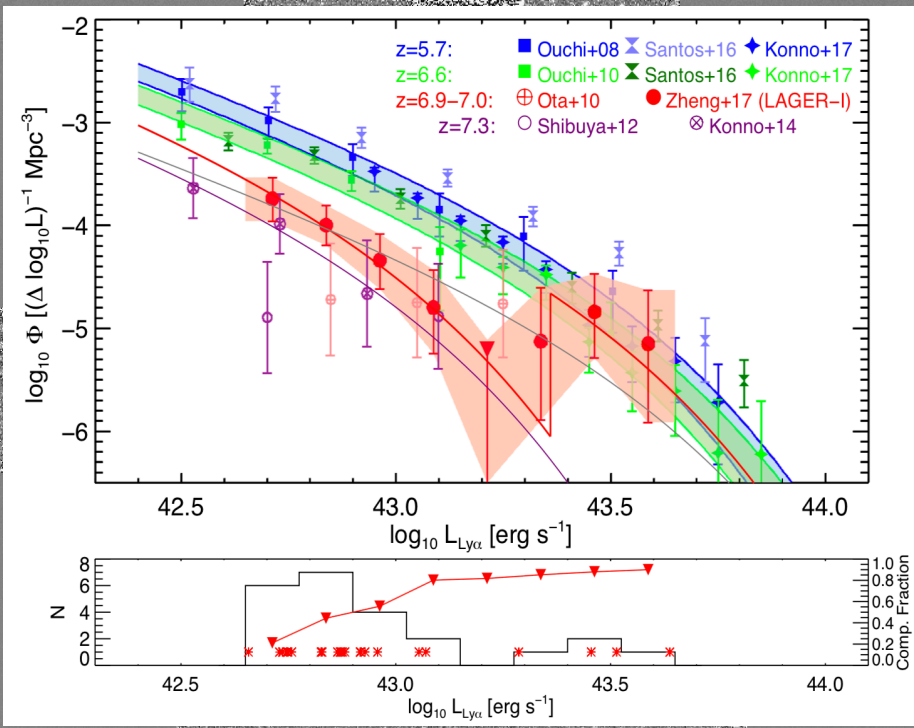


We Need: 1. Deeper NB α 64 Exposure; 2. NIR Spec. Obs. Of $z \sim 7$ LAEs

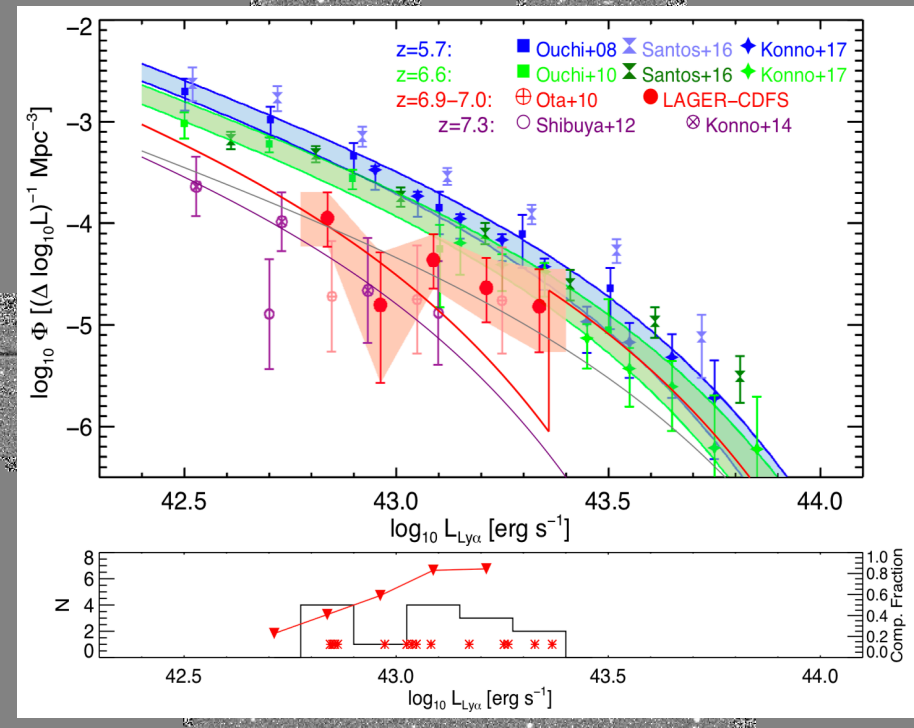
3. Ongoing Process

- More NB_g64 imaging data
- Improved image stacking methods
- Better & deeper broadband images (HSC & DES)
- Opt. & IR spectroscopic followup (Magellan, VLT, Keck, Gemini, HET)

COSMOS vs. CDFS



Zheng et al. 2017



Preliminary Results

Improve the Ly α LF at $z \sim 7$ and give better constraints on x_{HI}

Probing Reionization with High-z Galaxies

➤ Lyman- α luminosity function (LF) test

(Rhoads & Malhotra 2001, Malhotra & Rhoads 2004, Ouchi+2010, Kashikawa+2011, Konno+2014, Ota+2017, Zheng+2017, Konno+2017)

□ Lyman- α visibility test

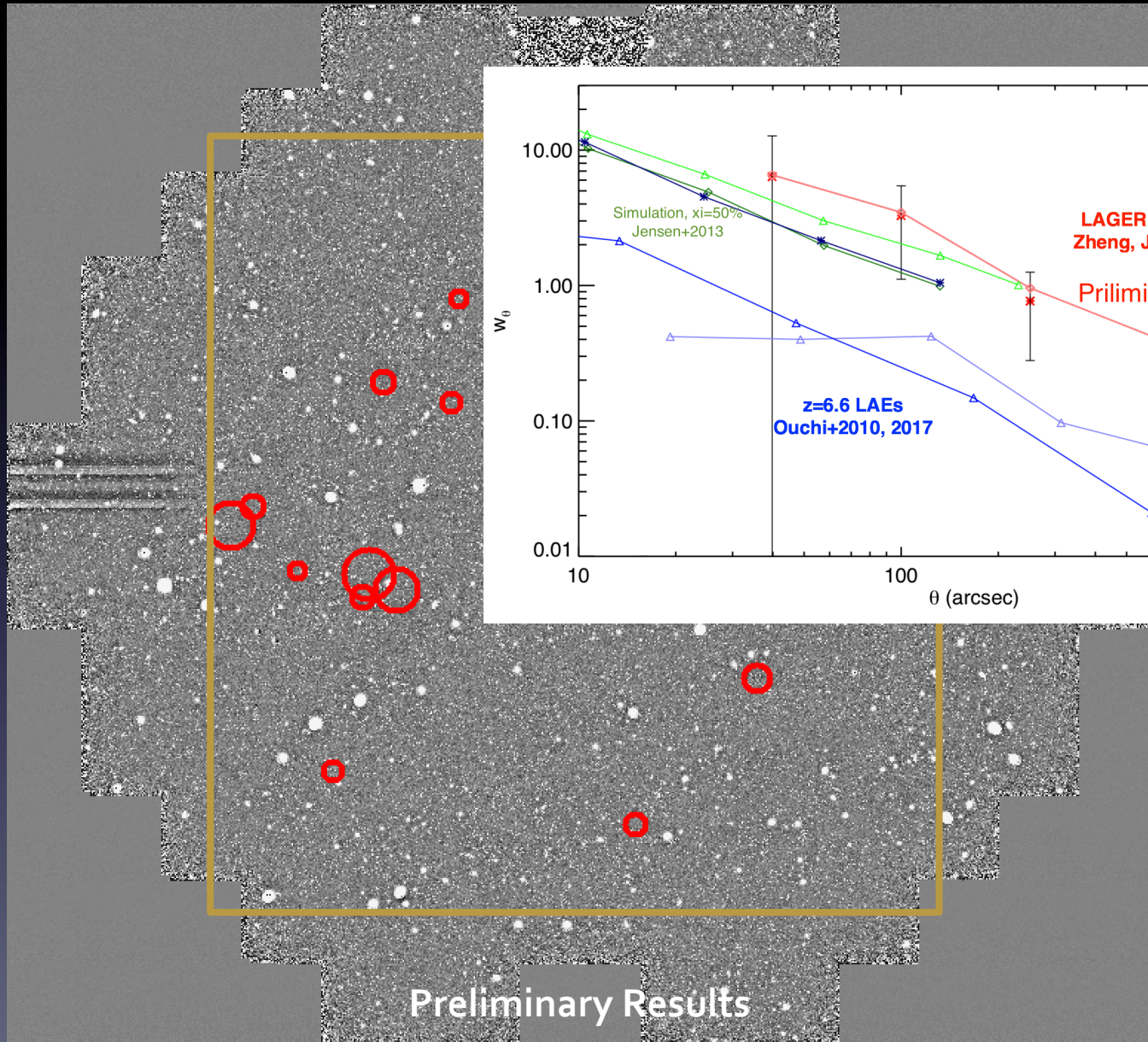
(On LBGs, e.g., Stark et al. 2011, Schenker+14, Bian+2014, Tilvi+2016, etc.;
On Ly α EW distribution of LAEs, e.g., Choudhury+2015, Kakiichi+2016)

◆ Clustering test of Lyman- α galaxies

(Theory: Furlanetto+2006, McQuinn+2007, Jensen+2013; Observation: Ouchi+2010, Ouchi+2017)

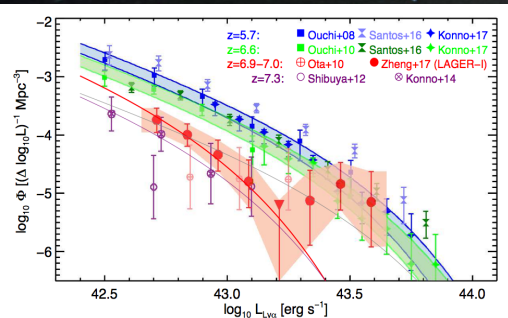
Need a careful treatment of: 1) faint (and fake) objects;
2) pixel-to-pixel depth on both NB & BB images; and
3) Filter profile.

Improved COSMOS Sample

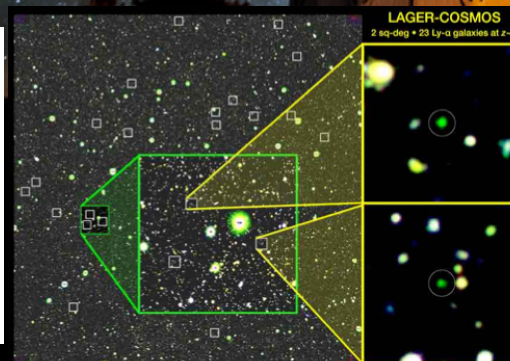


Conclusion

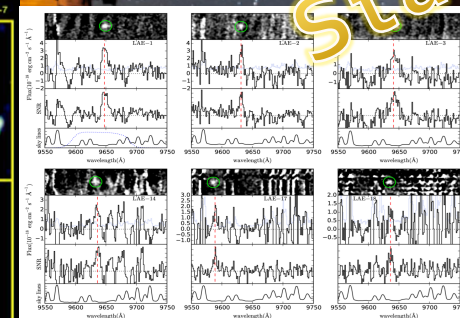
- LAGER is the largest narrowband survey for LAEs at $z \sim 7$ so far.
- In the first LAGER field (COSMOS), 23 (22 new) candidate LAEs, the largest sample to date of candidate LAEs at $z \sim 7$, were discovered.
- LAGER helps us to find 4 most luminous LAEs at $z \sim 7$, of which 3 were spec. confirmed.
- The new Ly α LF from LAGER LAEs shows different evolution at the faint-end and at the bright-end $\rightarrow x_{\text{HI}} \sim 0.4\text{--}0.6$ and ionized bubbles at $z \sim 7$.
- More are coming.....



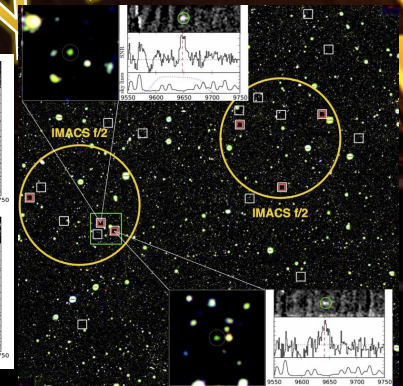
Zheng et al. 2017



NOAO Press Release 1703



Hu et al. 2017



Carnegie Press Release

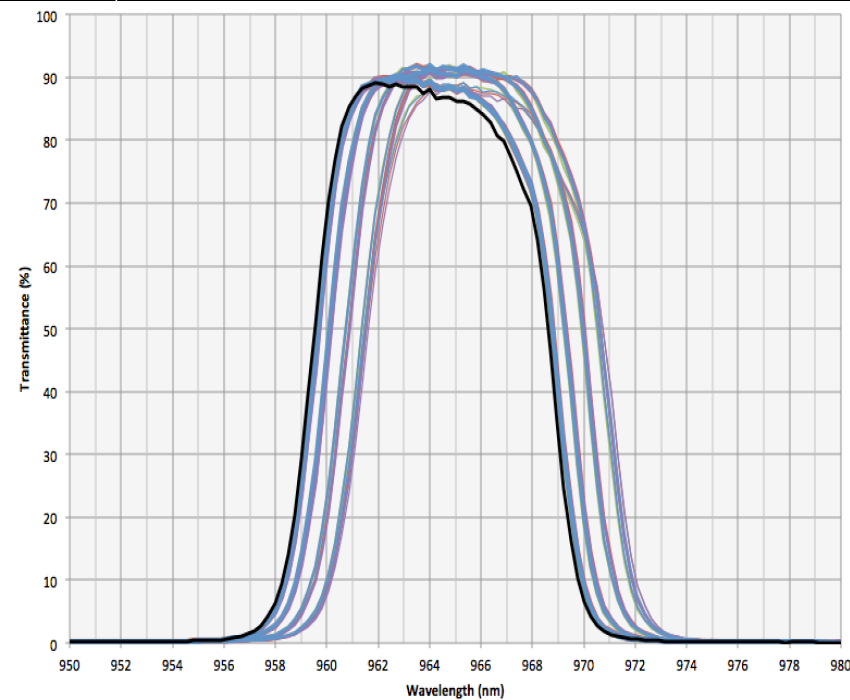
Stay Tuned!

Thank You!

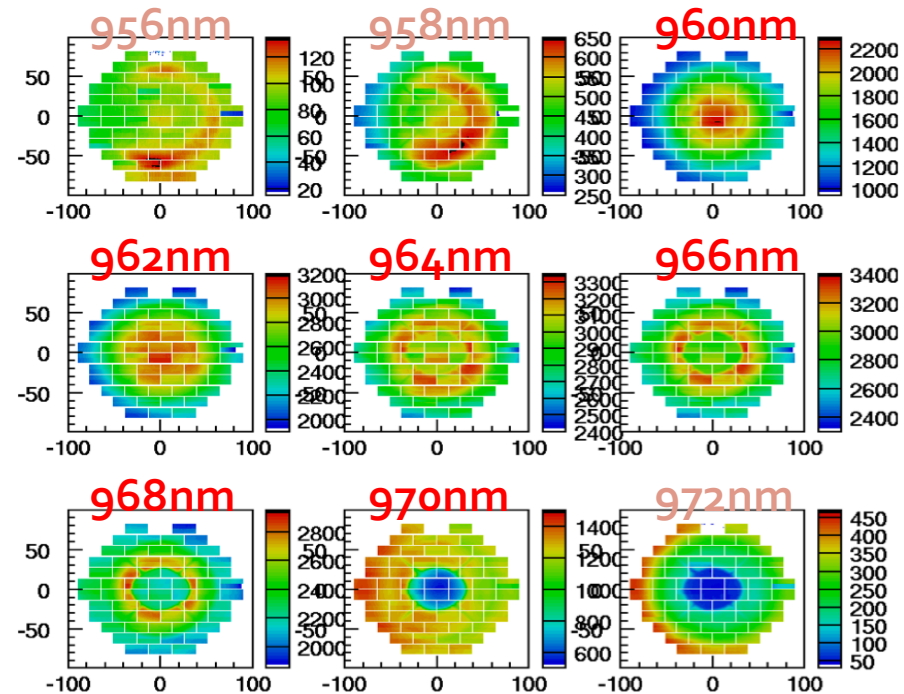
NB₉₆₄ Filter Performance

1. NB filter lab-test (Multi-curves for different Radii)

2. NB filter on-site test (Spatial dist. as a function of λ)



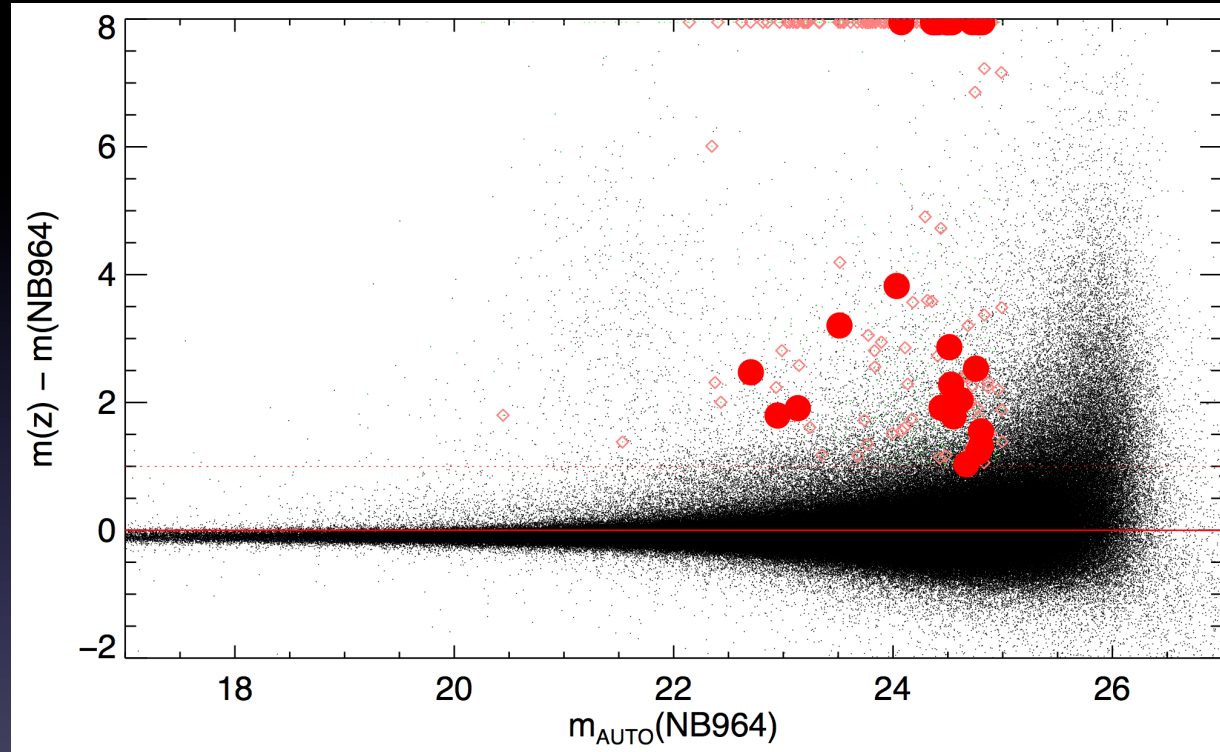
- 55MM A
- 110MM A
- 165MM A
- 220MM A
- 275MM A
- 55MM B
- 110MM B
- 165MM B
- 220MM B
- 275MM B
- 55MM C
- 110MM C
- 165MM C
- 220MM C
- 275MM C
- 55MM D
- 110MM D
- 165MM D
- 220MM D
- 275MM D
- CTR



Candidate LAEs at $z \sim 7$

Selection Criteria:

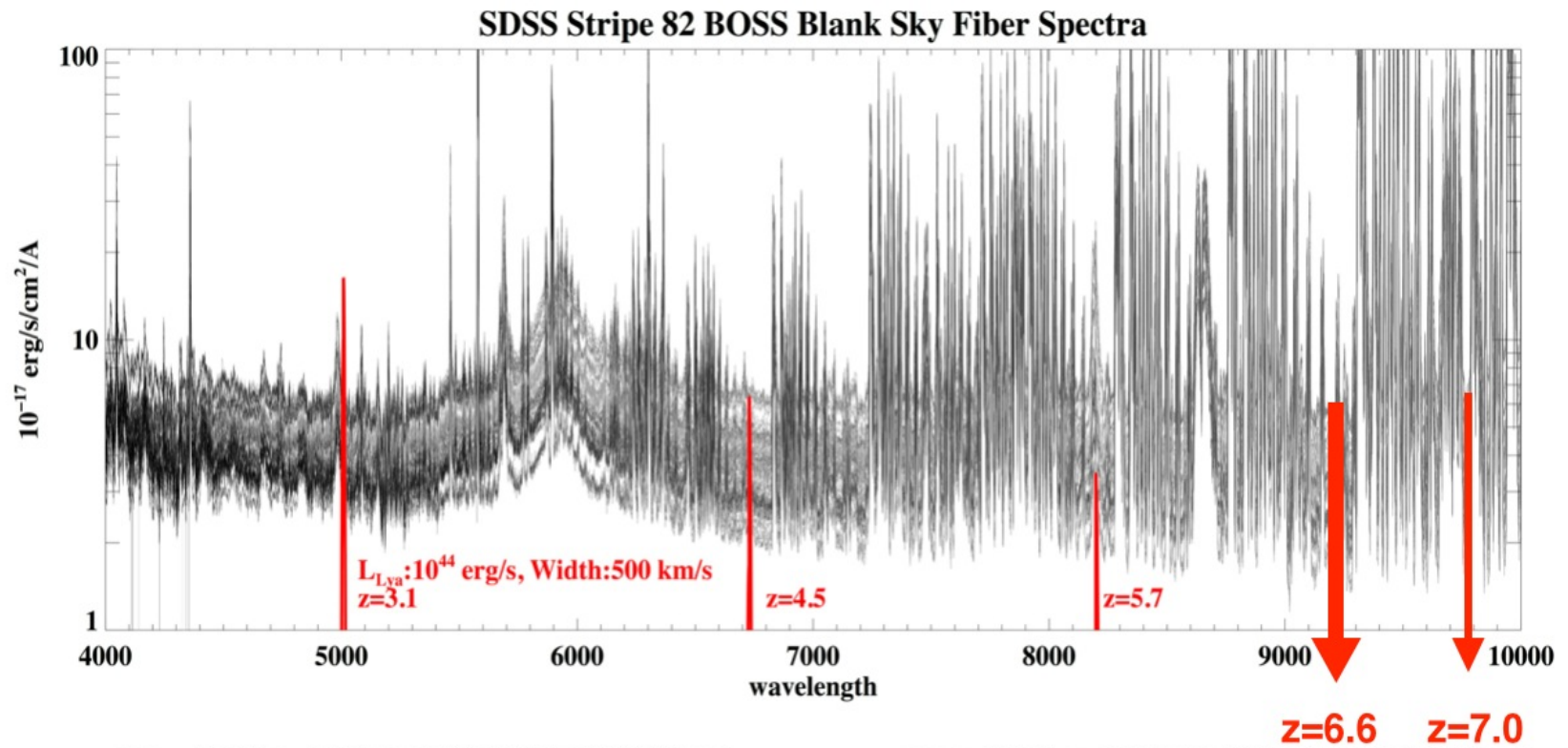
- **Non-Detection in Blue Bands:**
DECam-ugri $< 3 \sigma$ & Subaru-BVgri $< 3 \sigma$ & Subaru-NB711, NB816, NB921 $< 3 \sigma$
- **NB Significant:**
NB964 ($> 5 \sigma$) < 25
- **Line Significant:**
DECam-z - NB964 ≥ 1
& $EW_r(\text{Ly}\alpha) > 10 \text{ \AA}$



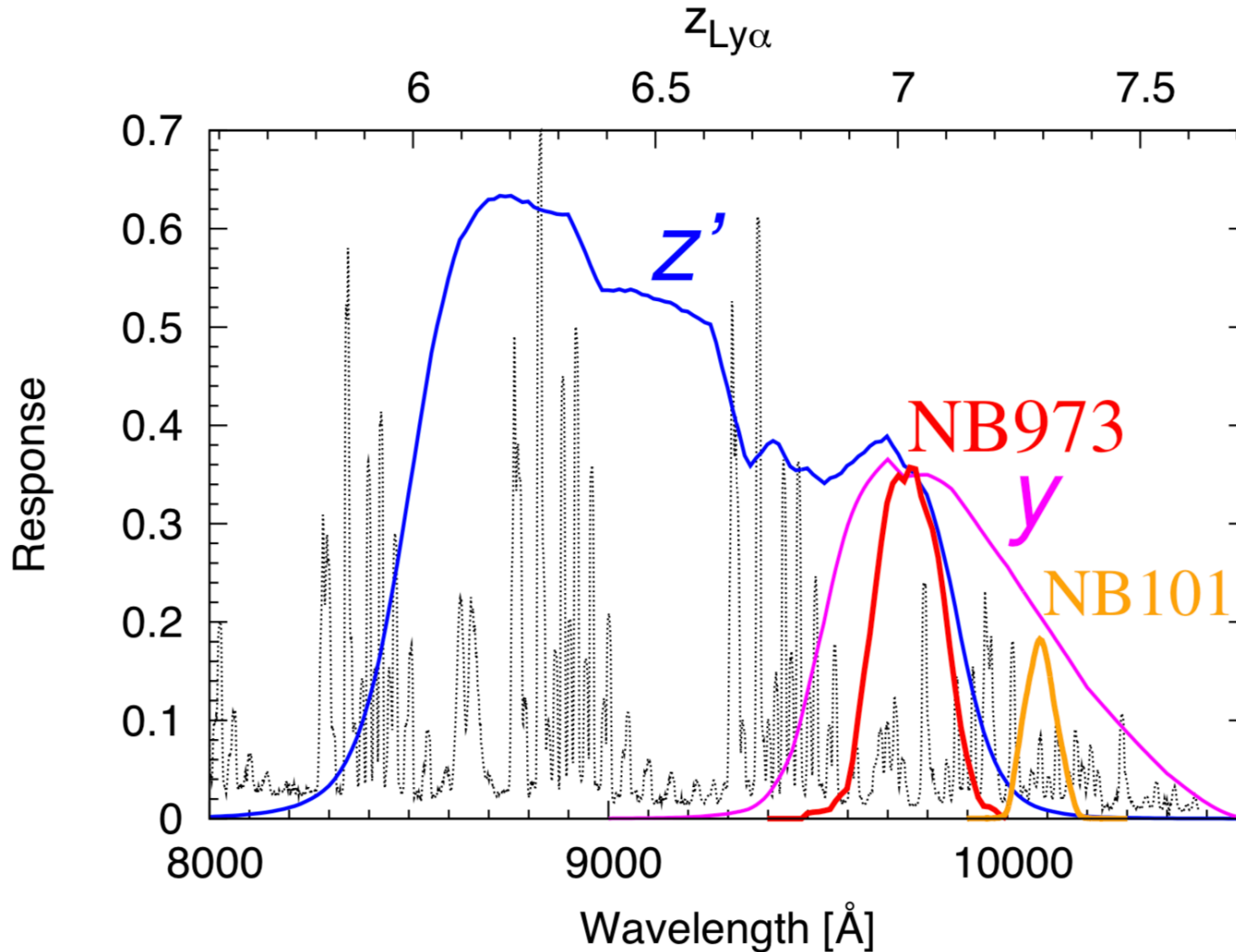
We find 23 (22 new) candidate LAEs at $z \sim 7$ in the COSMOS field.
Survey Volume = $1.26 \times 10^6 \text{ cMpc}^3$ ($> 4 V_{\text{other}_z7}$)

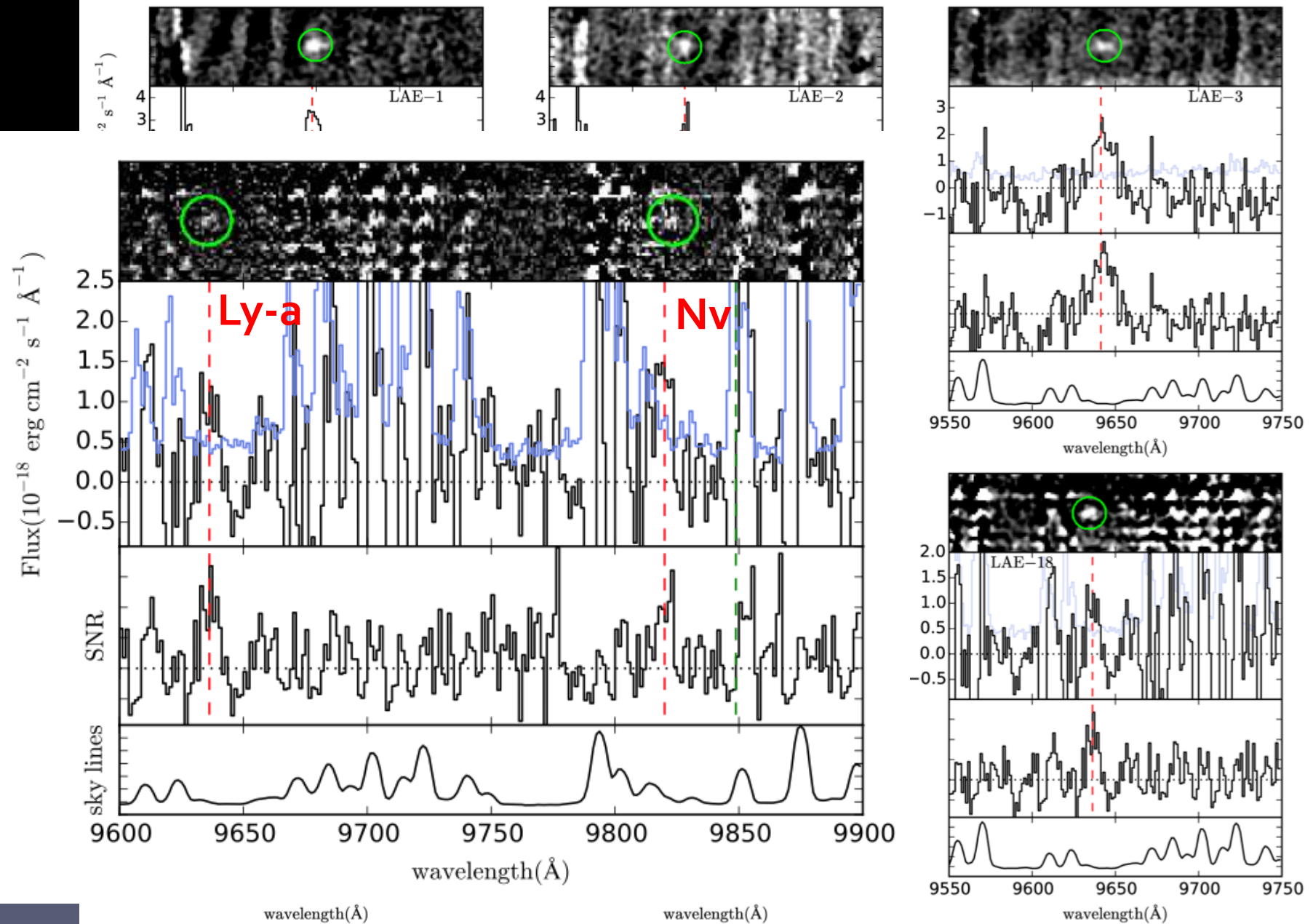
Observational challenges at $z > \sim 7$

Sky Background



Subaru SC NB973 (Ota et al.)





9 targeted, 6 confirmed (successful rate $\sim 67\%$).

Hu et al., 2017, *ApJL*, 845, L16 (arXiv:1706.03586).