

Tokyo Spring
Cosmic Lyman-Alpha Workshop (Sakura CLAW)

Blueberry galaxies: local analogs of the faint-end LAEs

Huan Yang (杨欢)

Las Campanas Observatory, Carnegie Institution for Science

*Collaborators: Sangeeta Malhotra (NASA), James E. Rhoads (NASA),
TianXing Jiang (ASU), JunXian Wang (USTC), Leopoldo Infante
(Carnegie/LCO)*

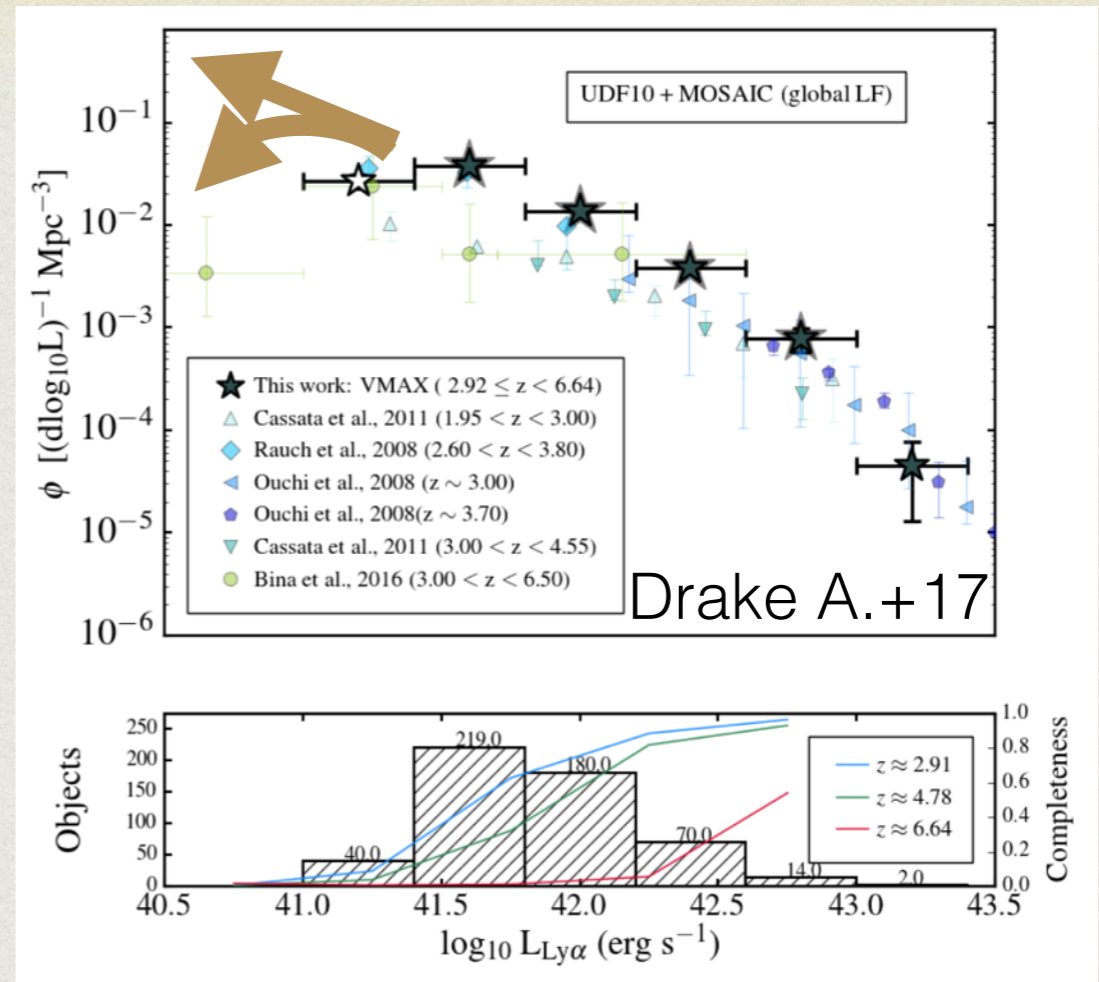
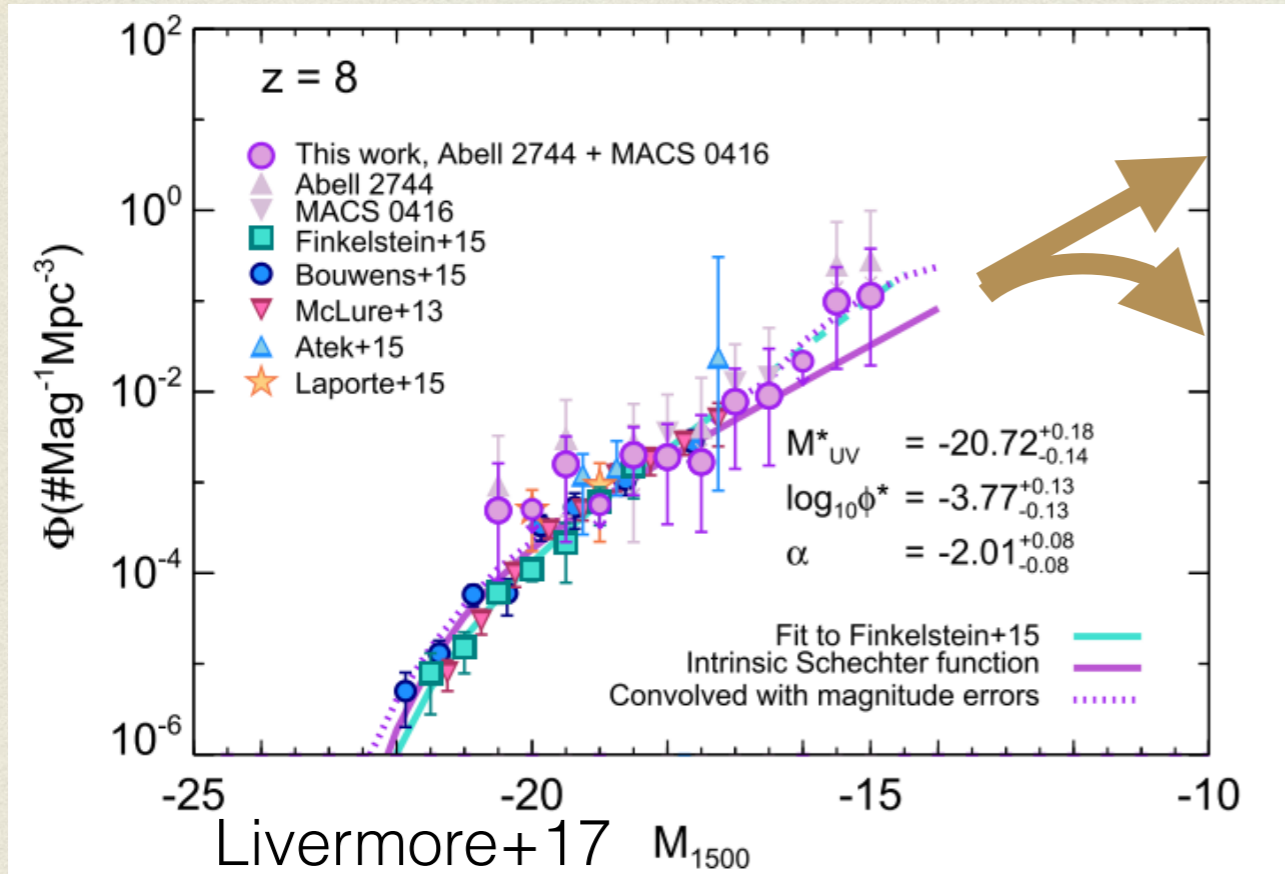


ASU
ARIZONA STATE
UNIVERSITY



中国科学技术大学
University of Science and Technology of China

What are the roles of faint-end starbursts during reionization?



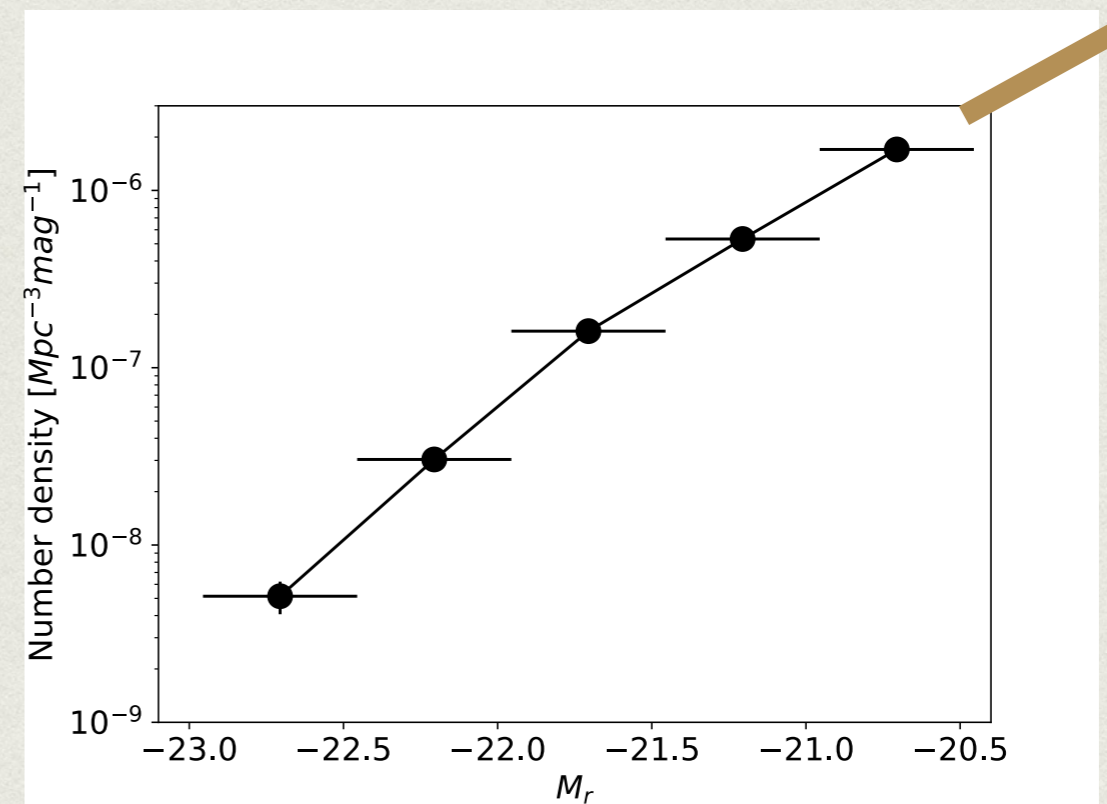
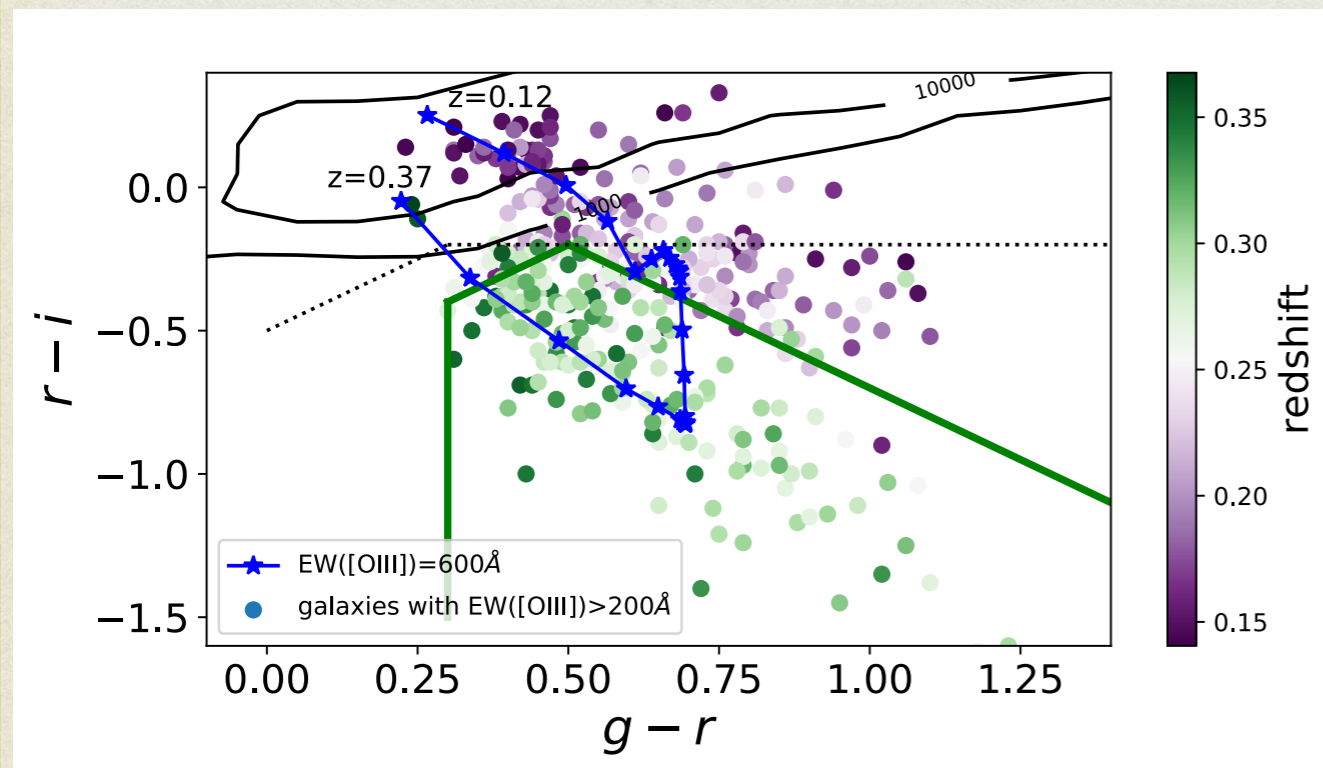
- What are the number densities, faint-end slopes and turn-over luminosities?
- What fraction of faint UV continuum selected galaxies are strong LyA emitters?
- What are the leaking Lyman-Continuum luminosities of these faint-end starbursts?

How can we study these faint-end LAEs?

- HST / JWST / MUSE / KCWI / Narrow-bands (HSC/DECam) ultra-deep surveys ...
- Are there local analogs of faint-end LAEs?

green peas: local analogs of normal/bright LAEs

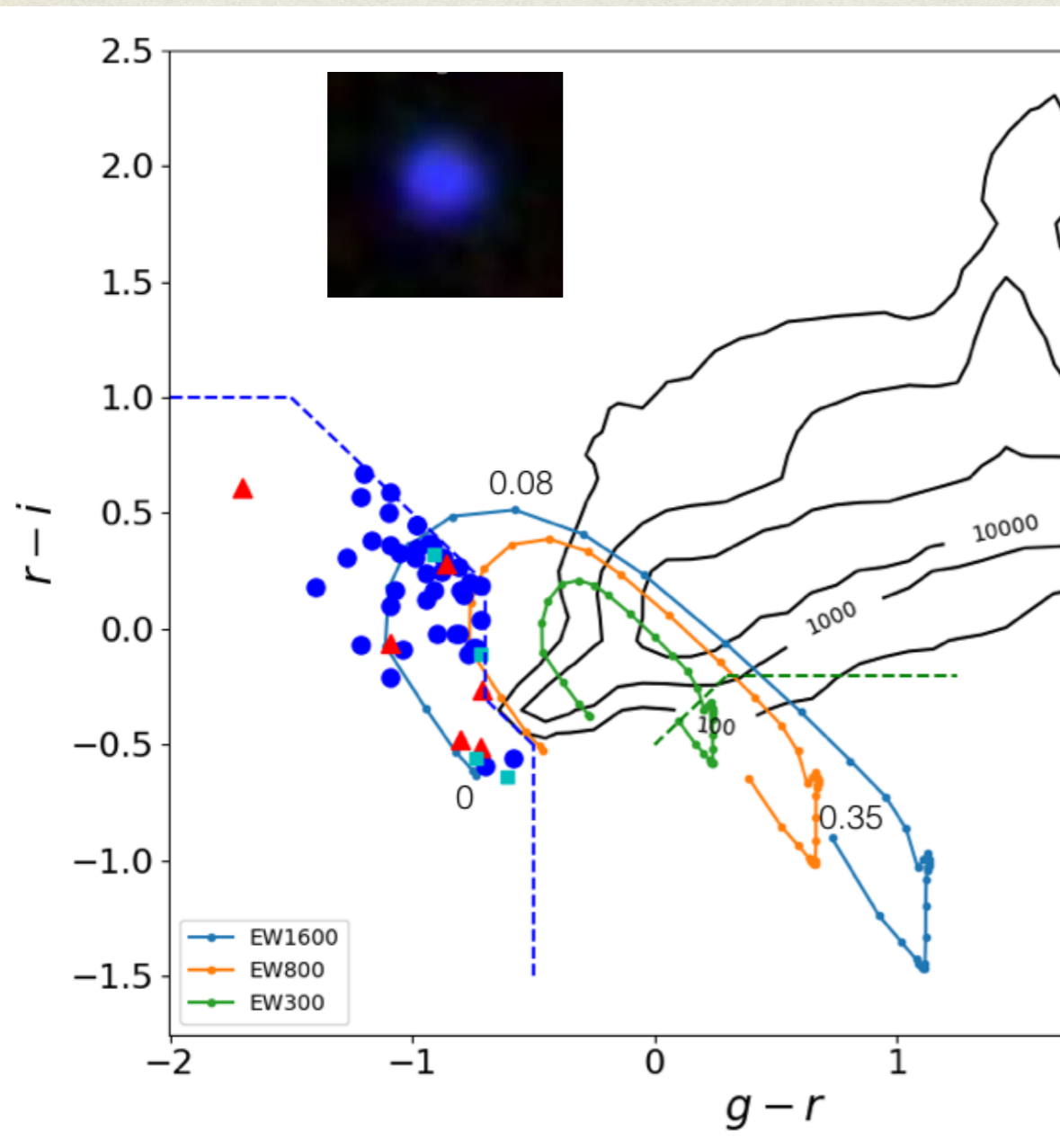
Green pea luminosity function at $z \sim 0.25-0.35$



Yang, H. + in-prep.

very steep slope $\alpha \sim -3.3$

fainter counterparts of green peas — blueberry galaxies



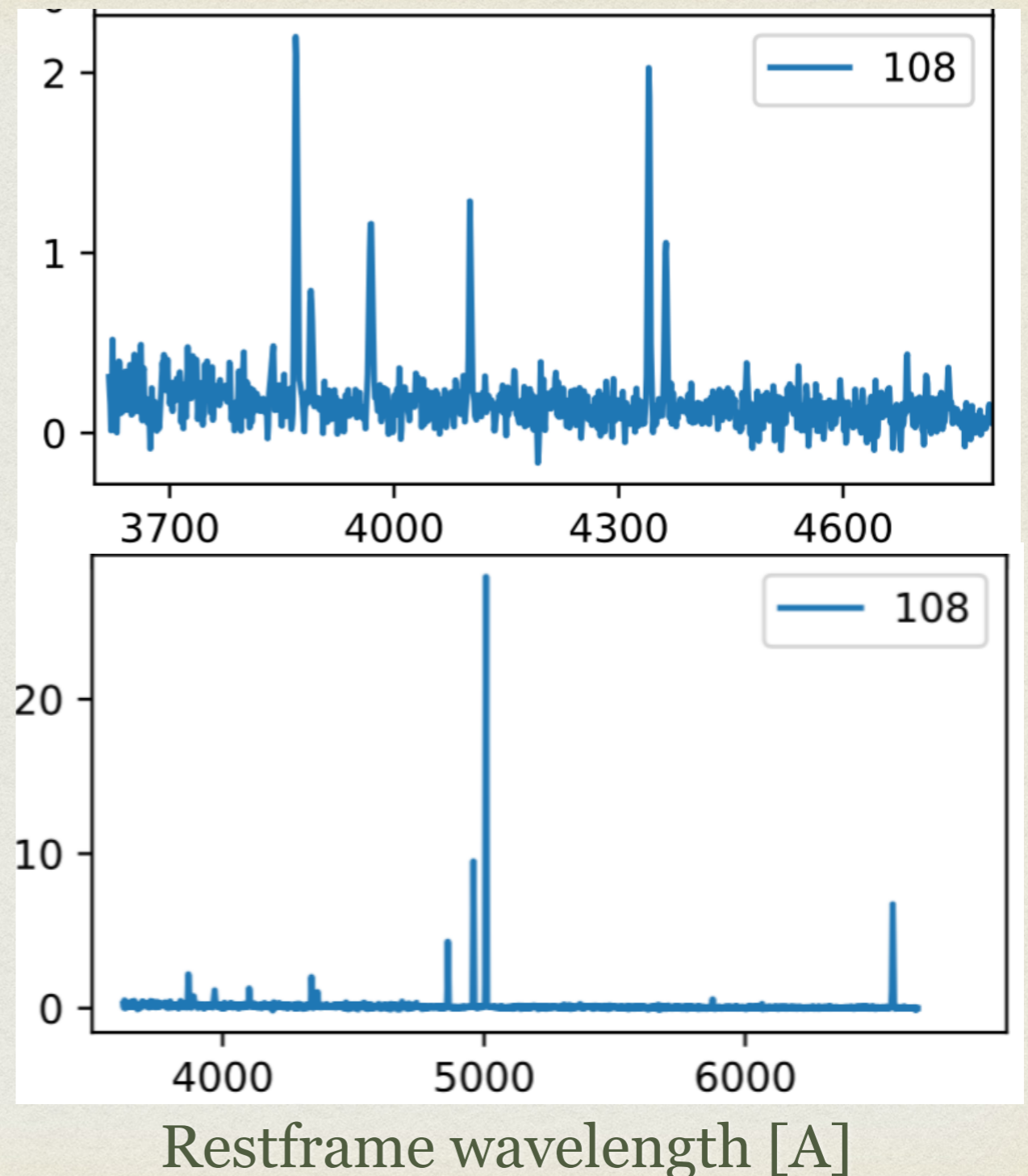
color criteria to select strong [OIII] emission line galaxies at redshift smaller 0.05.

Blueberries have 5-10 times smaller distances, therefore at similar apparent magnitudes, blueberry galaxies reached 100 times lower luminosity than green peas.

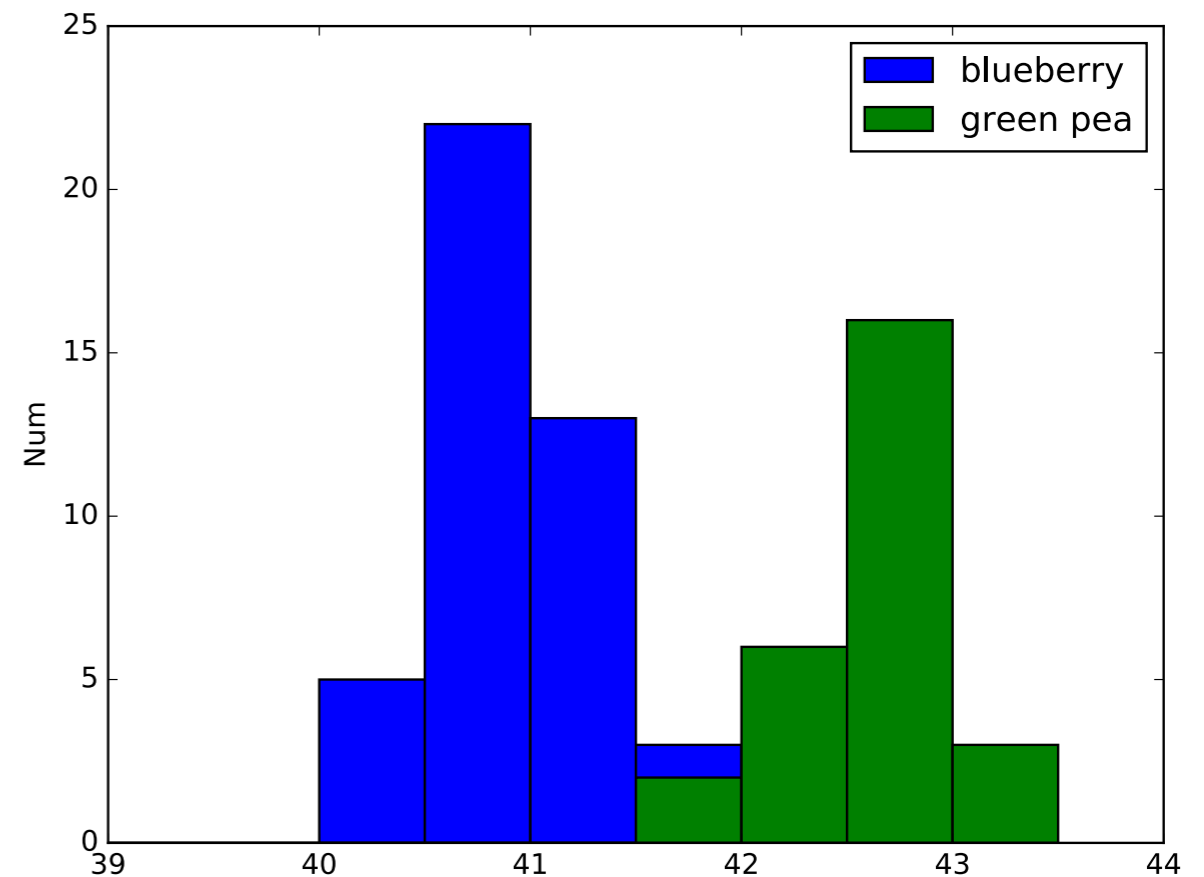
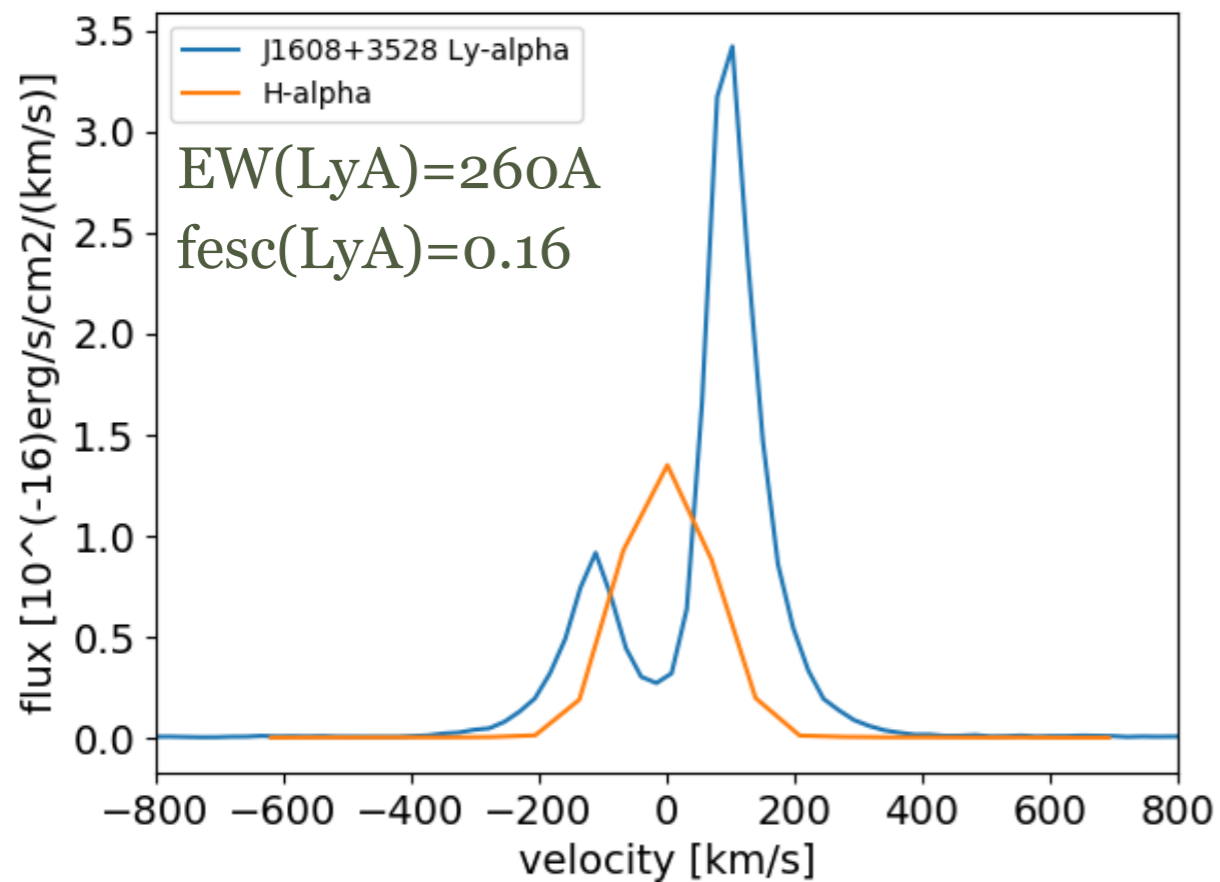
fainter counterparts of green peas — blueberry galaxies

a typical optical spectra.

The spectra is similar to green peas's, showing strong [OIII] emission lines. [OIII]/Hbeta and [OIII]/[OII] ratios are also very large.



Are blueberries LyA emitters?



One blueberry galaxy has HST UV observation and it shows a strong LyA emission line (also see Jaskot+2017).

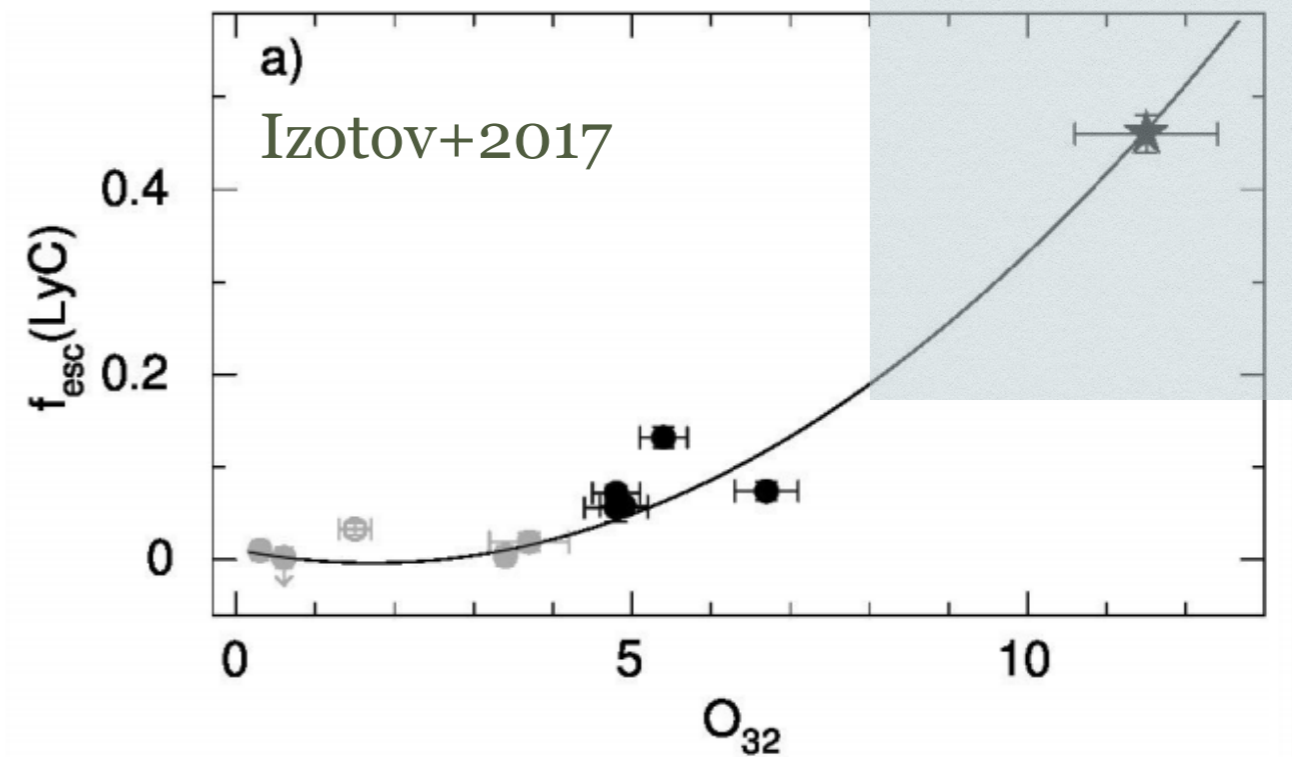
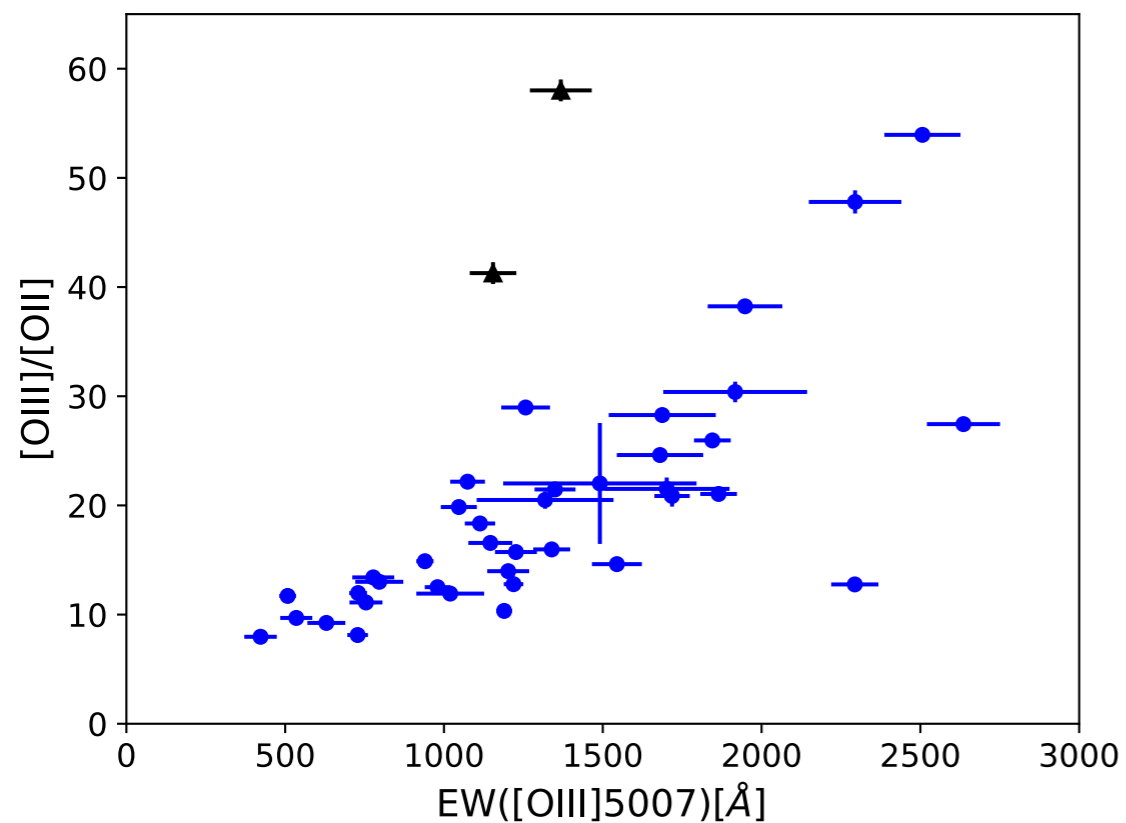
expected Log L(LyA) [erg/s]

Yang, H. + in-prep.

Are blueberries LyC leakers?

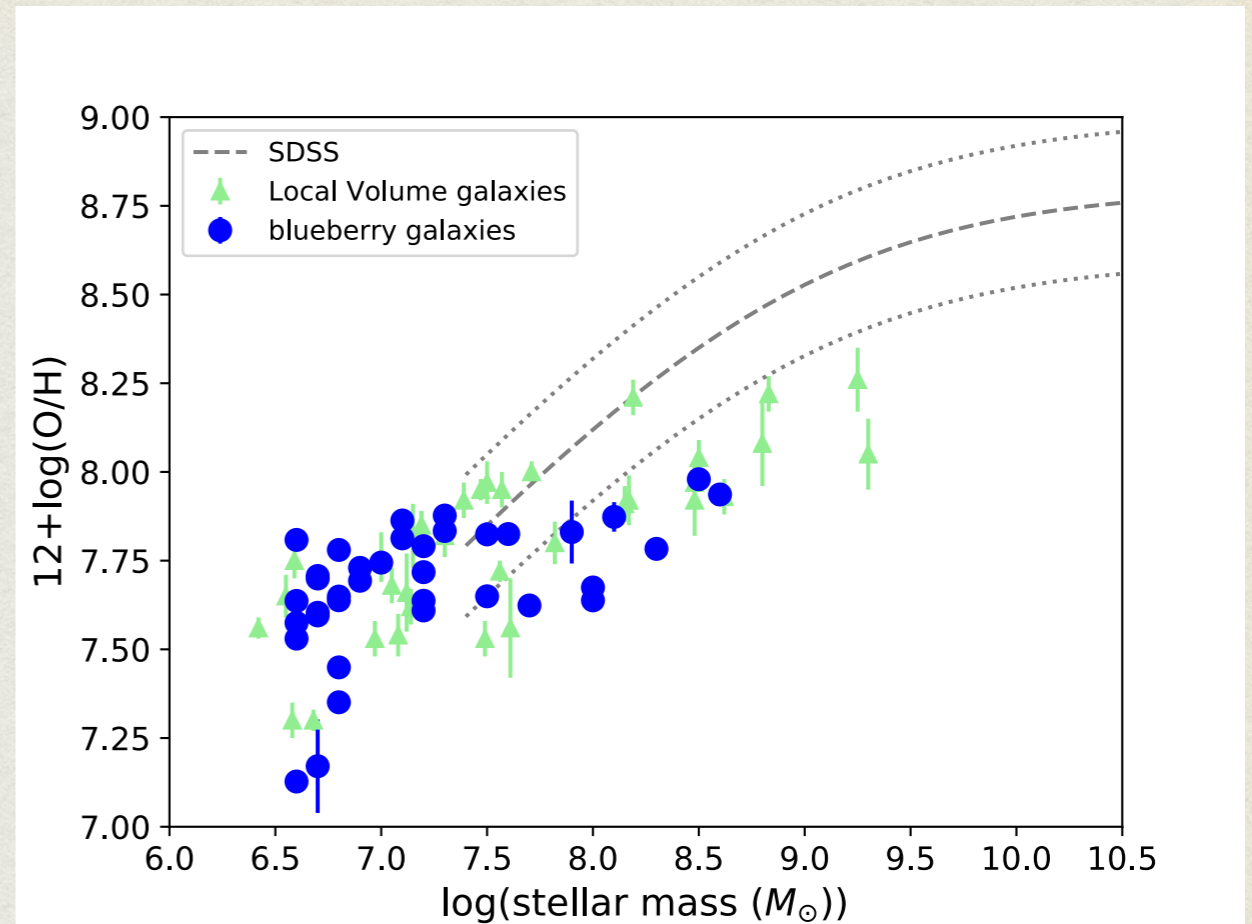
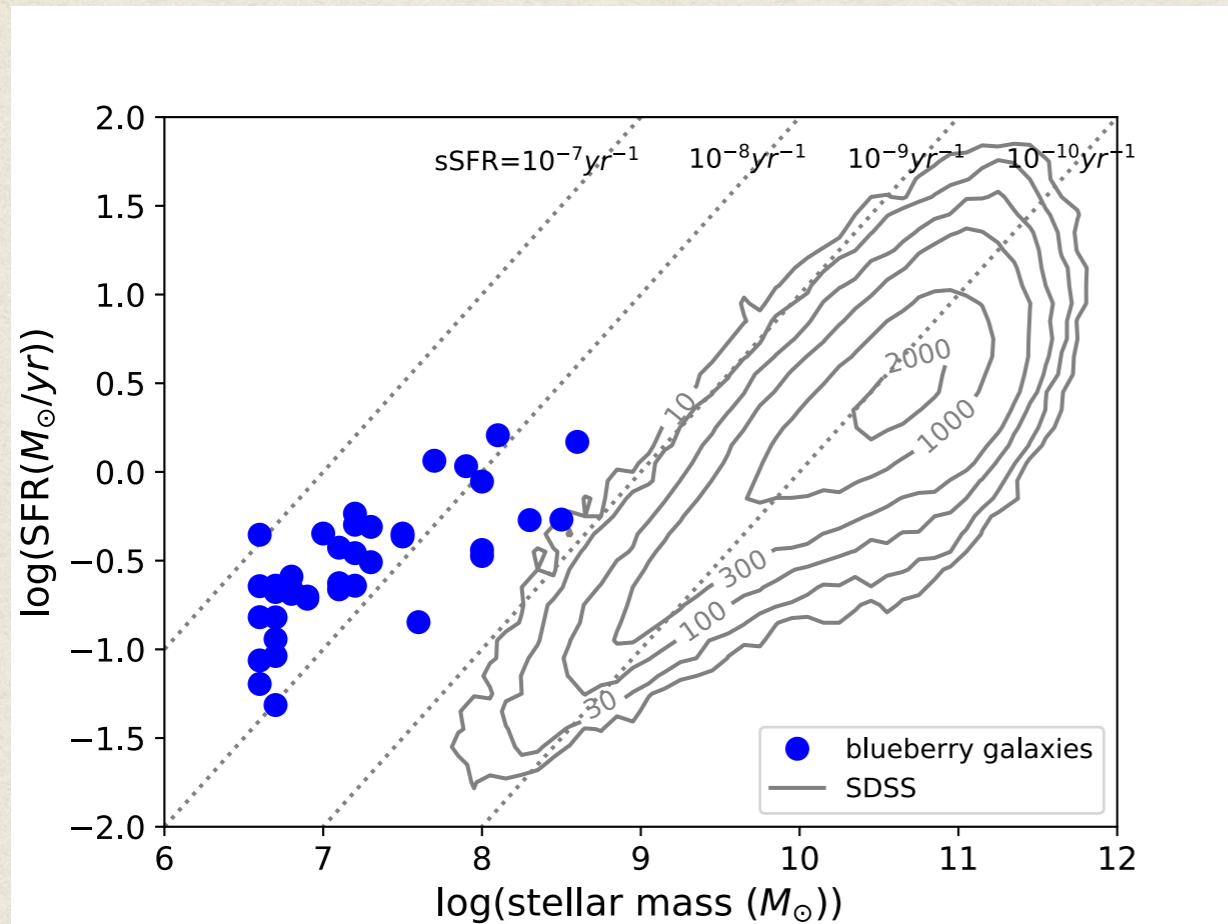
$[\text{OIII}]/[\text{OII}]$ of blueberries are $\sim 8 - 60$.

expected $f_{\text{esc}}(\text{LyC}) > \sim 15\%$



Yang, H. + 2017 arXiv:1706.02819

(1): very young low-metallicity starbursts



stellar mass around 10^7 solar mass

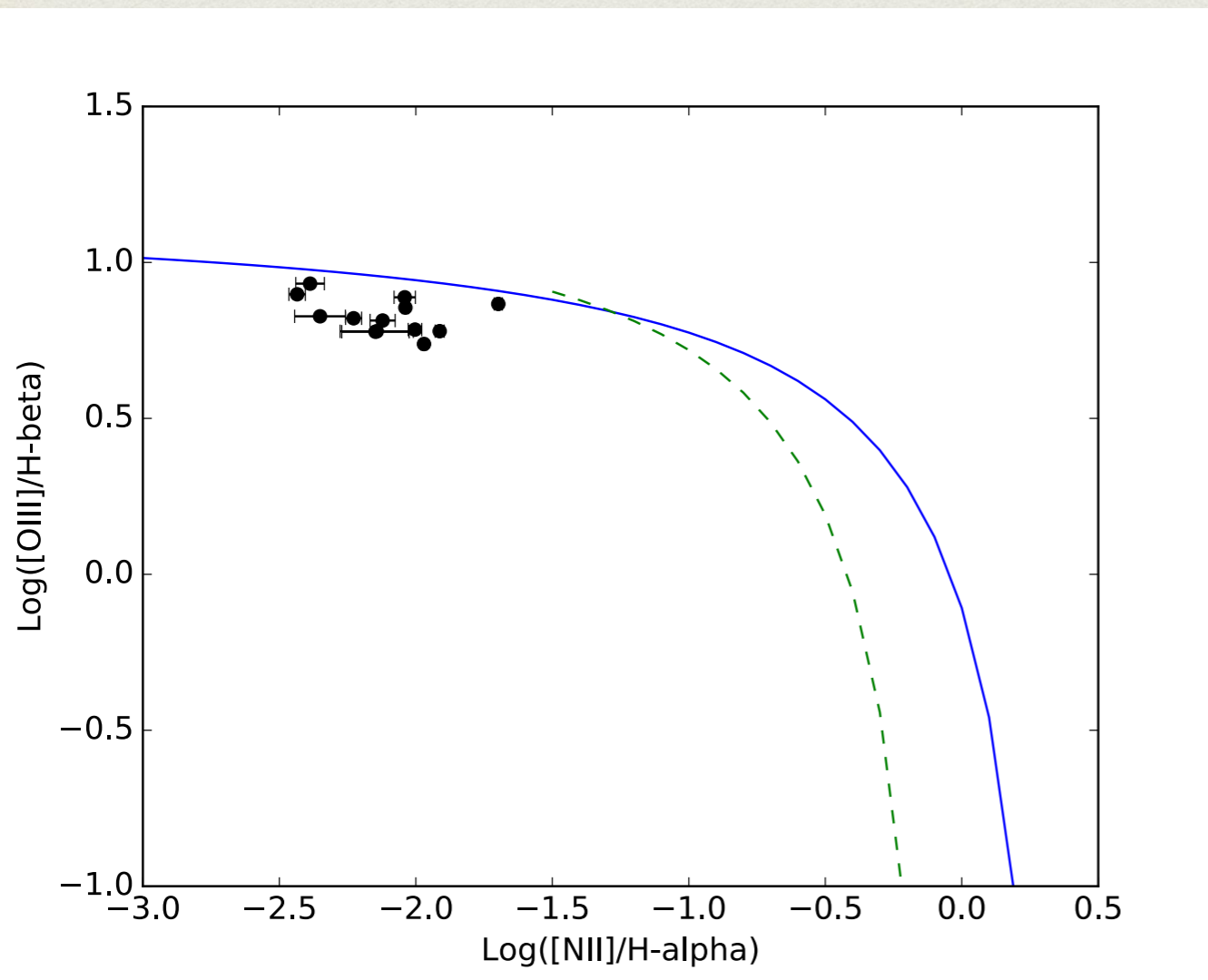
specific star-formation rates ~ 10 - $100 M \text{ yr}^{-1}$

very low gas metallicities ~ 3 to 10 percent solar metallicity

Yang, H. +2017 arXiv:1706.02819

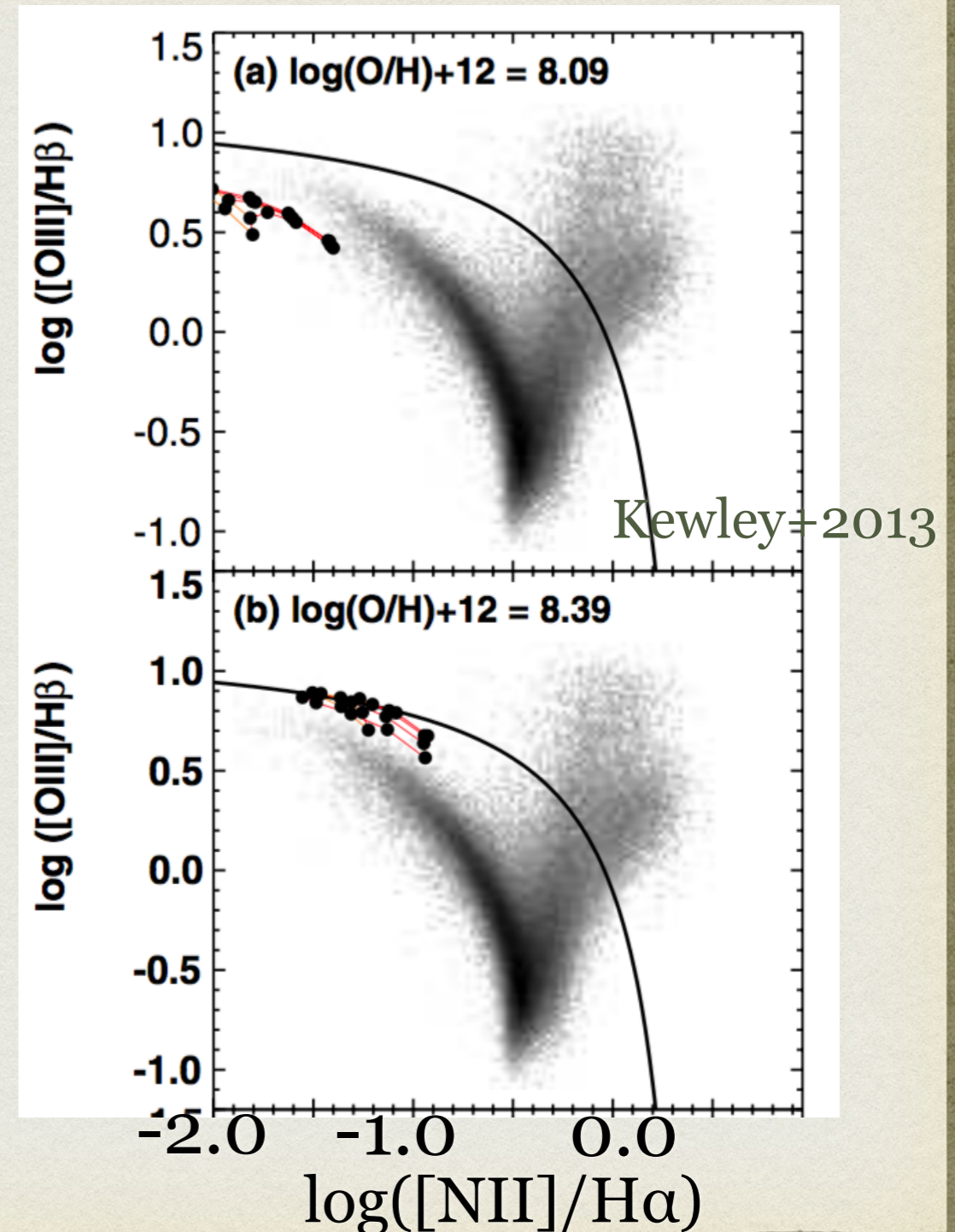
(2) Are there AGNs and black holes in blueberries?

We can not rule out AGN.



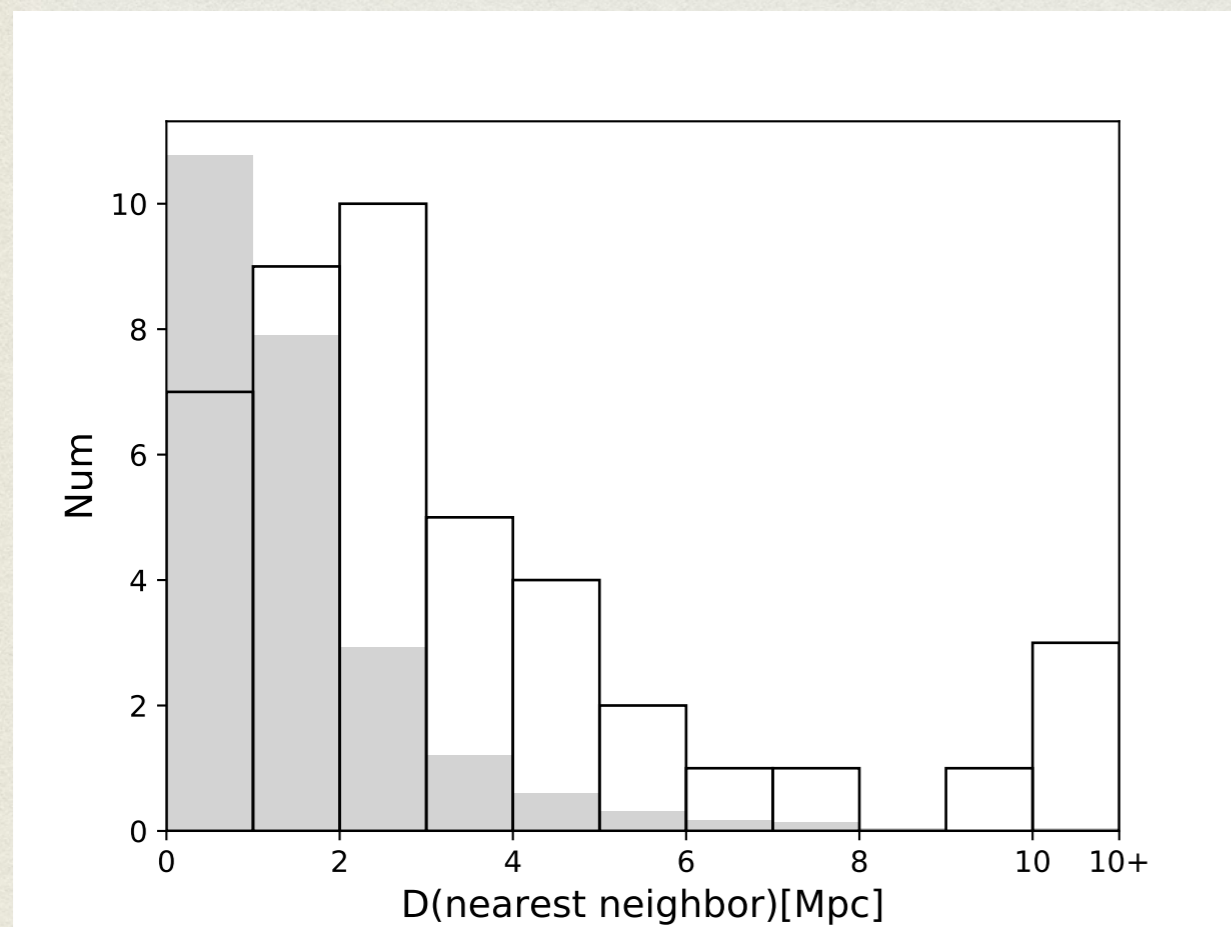
BPT diagram of Blueberry galaxies

Yang, H. + *in-prep.*

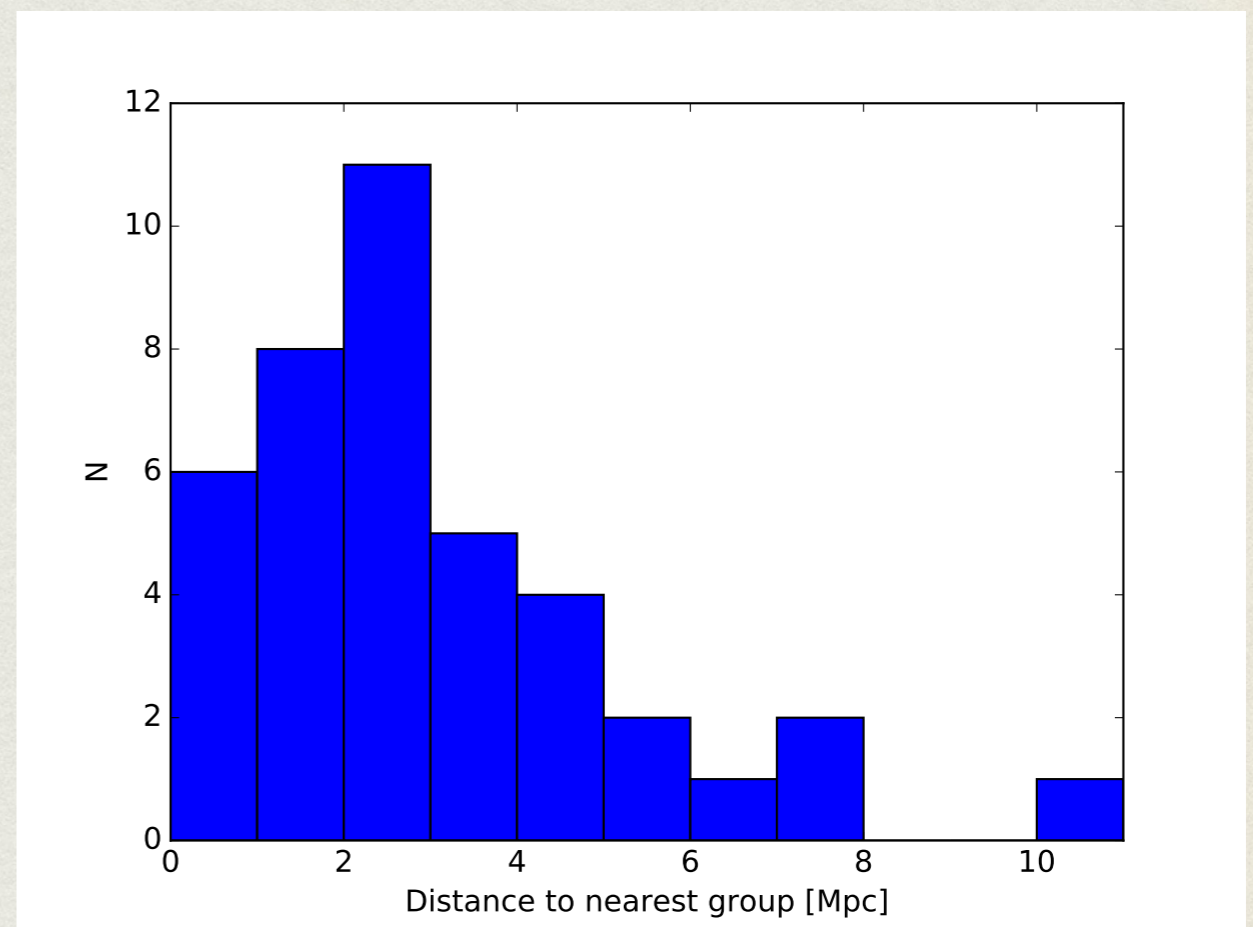


(3): Environment — outskirts of galaxy groups

Blueberry galaxies are at the very outskirts of galaxy groups.



Distance to the nearest
galaxy $\sim 1 - 4$ Mpc



Distance to the nearest
group center $\sim 1 - 4$ Mpc

Yang, H. + 2017 arXiv:1706.02819

(4): HST images

Yang, H. + in-prep.

UV HST-F275W

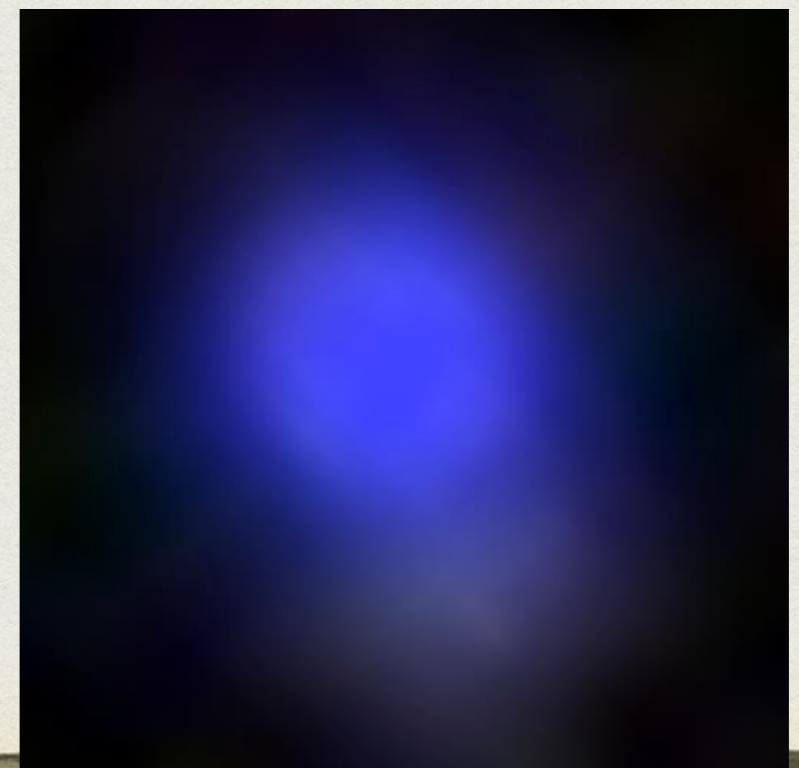
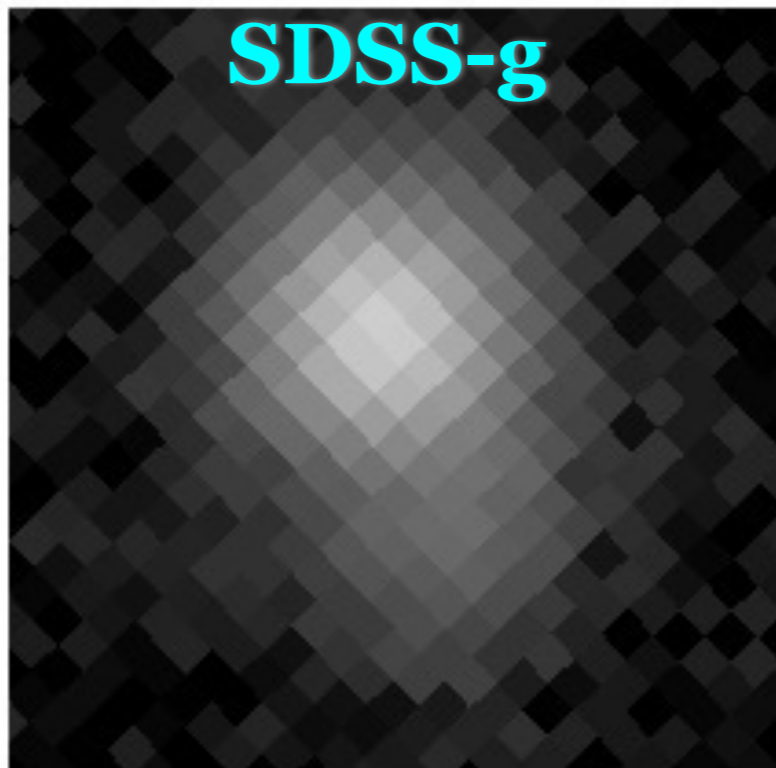
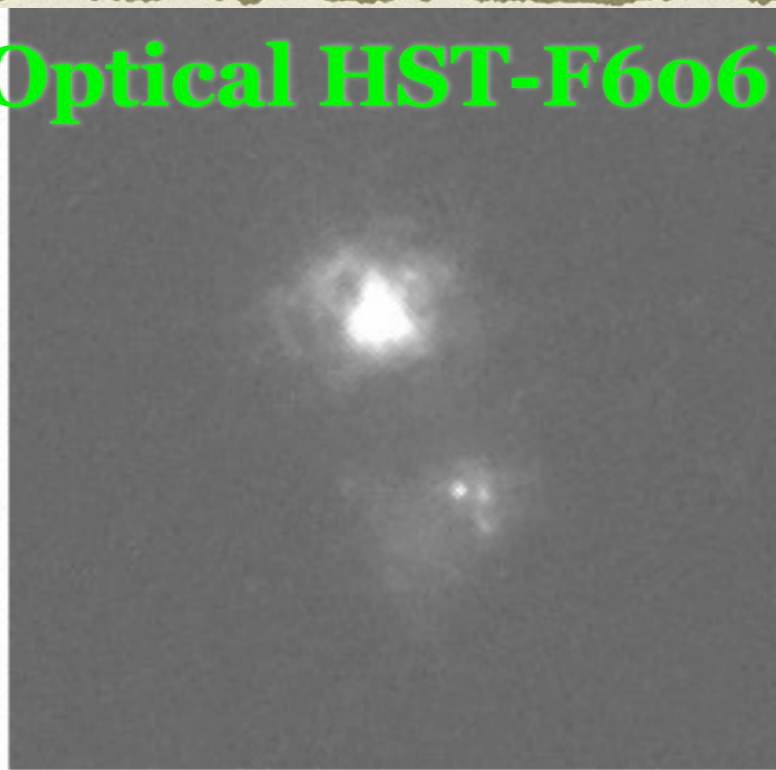
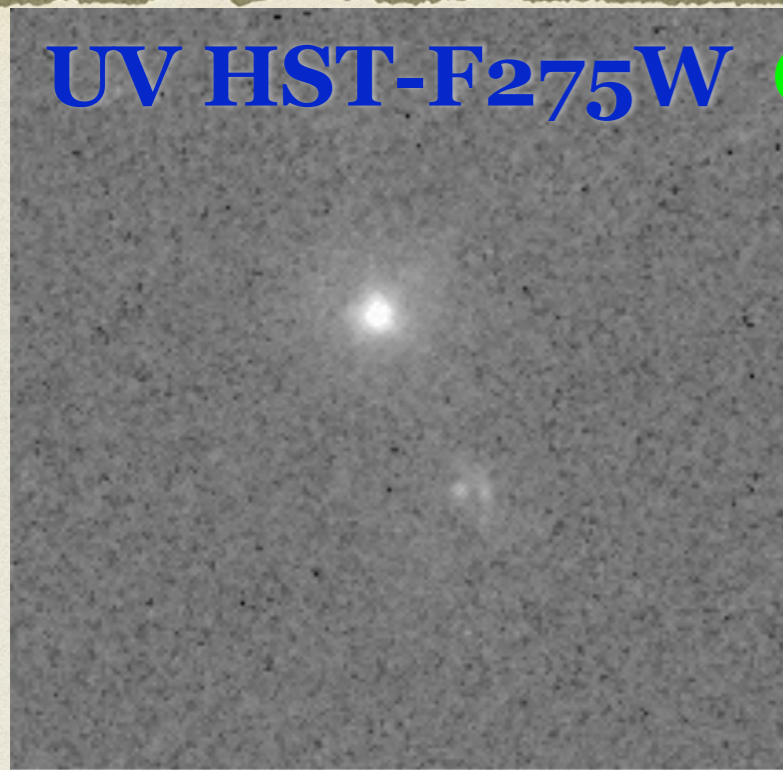
Optical HST-F606W

8x8 arcsec cutout

SDSS g-r-i

NIR HST-F110W

SDSS-g

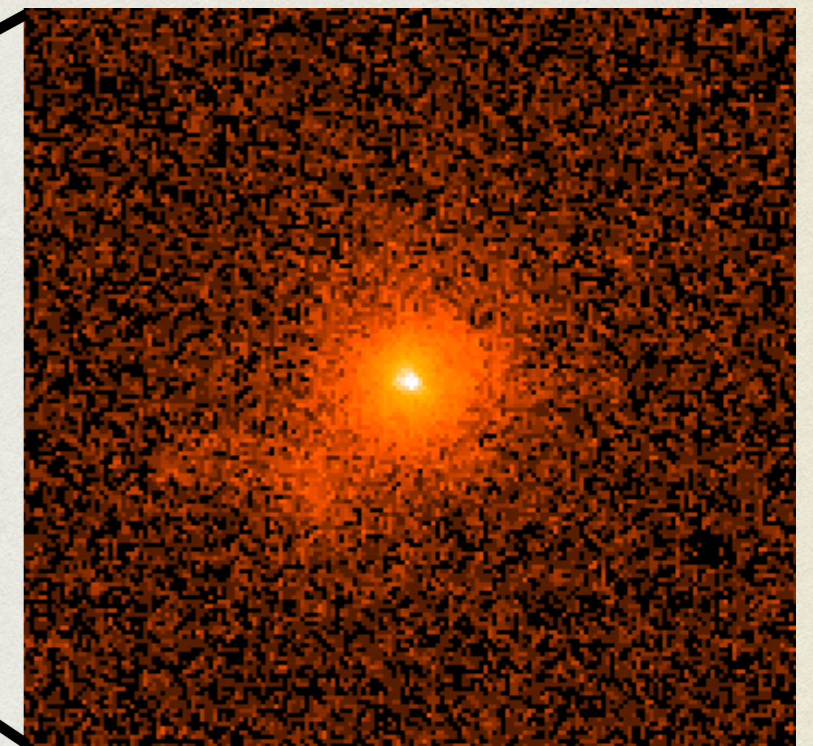
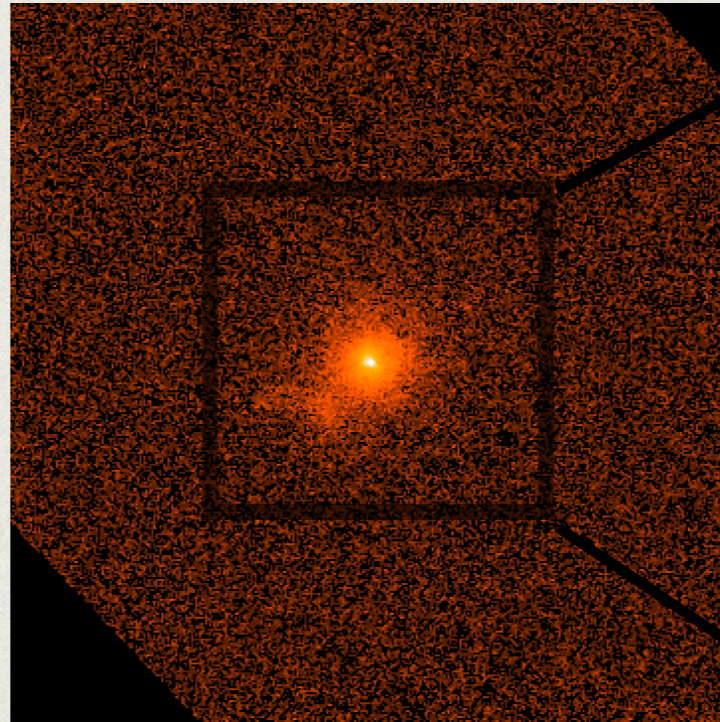


(4): HST images

SDSS g-r-i



UV HST

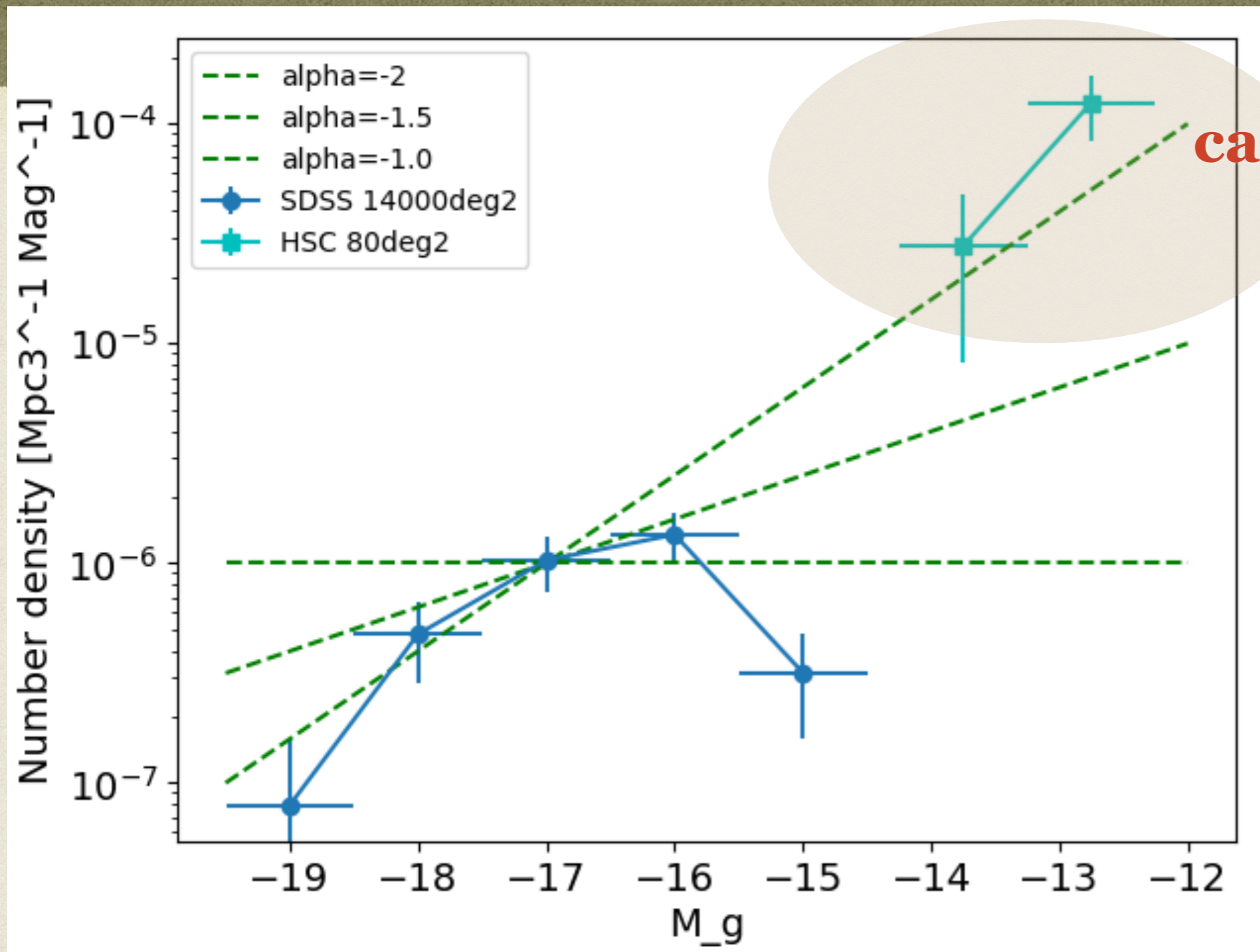


8x8 arcsec cutouts

Some blueberry galaxies may be mergers of two dwarf systems (galaxies or dark gas clumps).

Yang, H. + in-prep.

Number densities of blueberries in SDSS and HSC



candidates!

Assuming a model spectra -> M_{UV} of HSC blueberries ~ -13 to -12.

Yang, H. + in-prep.

Are these HSC blueberries faint dwarf galaxies? Or intergalactic star formation regions without underlying dark matter halos?

Summary

- We calculated the number densities of green peas.
- We selected blueberry galaxies, which are the local analogs of faint-end LAEs.
- These blueberries are metal-poor young starbursts in the outskirts of galaxies groups.
- We selected blueberries in HSC images and are working on the faint-end luminosity function down to $\text{Mag} \sim -12$.