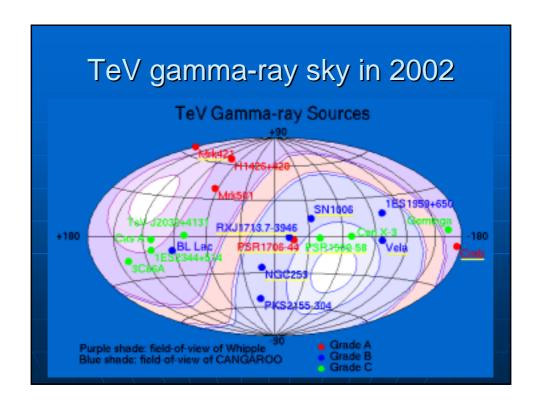
# CANGAROO

## Masaki Mori Institute for Cosmic Ray Research

シンポジウム 法人化後の宇宙線研究所研究プロジェクトについて」
Symposium "ICRR Projects after university reform in 2004"
February 01, 2003 @ICRR



#### **CANGAROO** team University of Adelaide Institute of Physical and Chemical Research Australian National University Shinshu University Ibaraki University Institute for Space and Aeronautical Science Ibaraki Prefectural Tokai University University Kanagawa University Tokyo Institute of Technology Konan University Yamagata University Kyoto University Yamanashi Gakuin Nagoya University University National Astronomical Observatory of Japan Osaka city University



(Yellow: CANGAROO)  TeV sources		
4 Pulsar nebulae	Crab Vela PSR 1706-44 PSR1509-58	
8 Blazars	Mrk 421 Mrk501 1ES2344+514 PKS2155-	
	304 3C66A BL Lac 1H1426+428 1ES1959+65	
3 Supernova remnants	SN1006 Cas A RX J1713.7-3946	
1 X-ray Binary 1 Starburst	Cen X-3 NGC253	

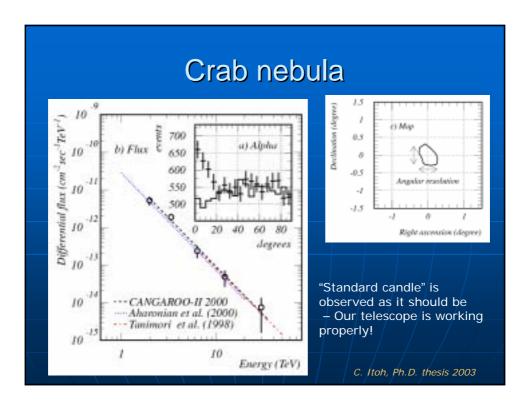
### CANGAROO 10m telescope

Upgraded in 2000

galaxy

- 114 x 80cm CFRP mirror segments (first plastic-base mirror in the world!)
- Focal length 8m
- Alt-azimuth mount
- 552ch imaging camera
- Charge and timing electronics

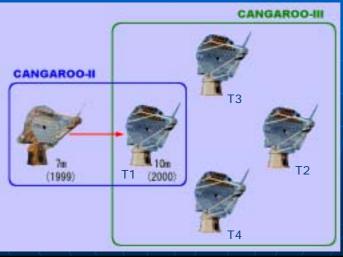




CANGAROO-II observations			
<ul> <li>SNR/Pulsar Crab</li> <li>SNR RX J1713.7-3946</li> <li>AGN Mrk421</li> <li>Starburst galaxy NGC253</li> <li>Pulsar PSR 1706-44</li> </ul>	Signal Publish O O (Nature) O (ApJL) O (AApL) O Δ		
<ul> <li>SNR SN1006</li> <li>PSR 1259-63/SS2833</li> <li>AGN PKS2155-304, PKS2005-489</li> <li>SNR RX J0852-4622</li> </ul>	Ο Δ Ψ Δ Ψ Δ		
<ul> <li>SNR RCW86</li> <li>Galactic Center/Sgr A*</li> <li>Galactic jet object SS433</li> <li>EGRET unID 3EG J1234-1318</li> <li>Galaxy Small Magellanic Cloud</li> </ul>			
<ul> <li>Vela pulsar</li> <li>Signal: O detected, ✓ upper limit, △ under a</li> <li>Publish: O published, △ in preparation</li> </ul>	<u> </u>		

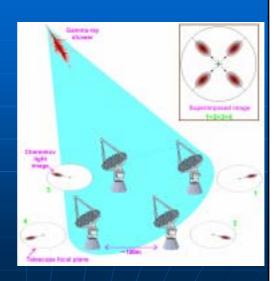
#### **CANGAROO-III** project

4 x 10m telescopes to be completed in 2003

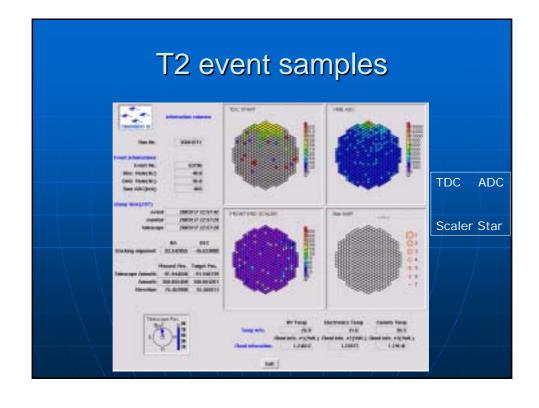


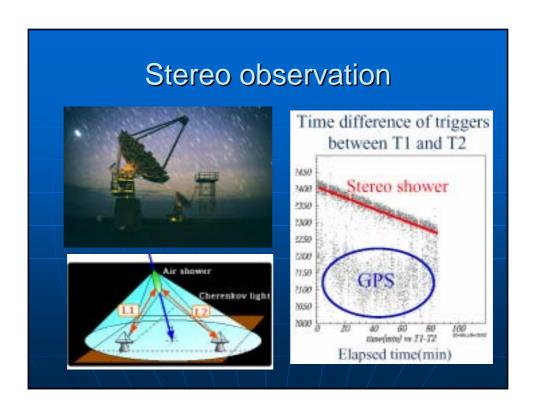
#### Merit of stereo observation

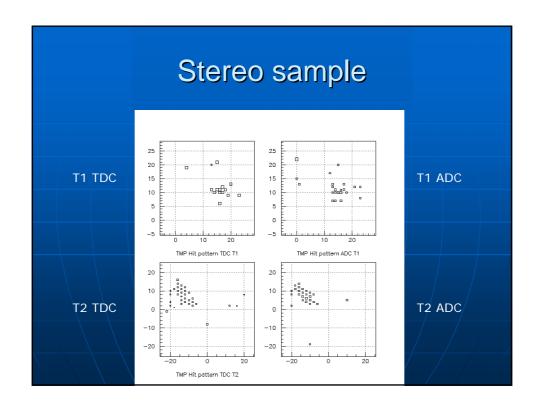
- Cherenkov shower pool: ~300mø
- Stereo ⇒ Info. on distance to showers
- Better angular resolution ∆θ=0.2°→0.05°
- Better energy resolution ΔE/E=30%→15%

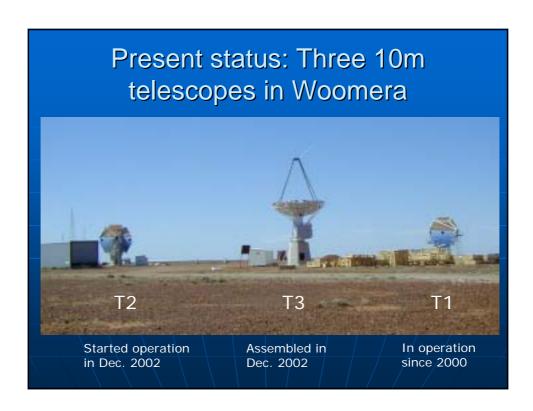


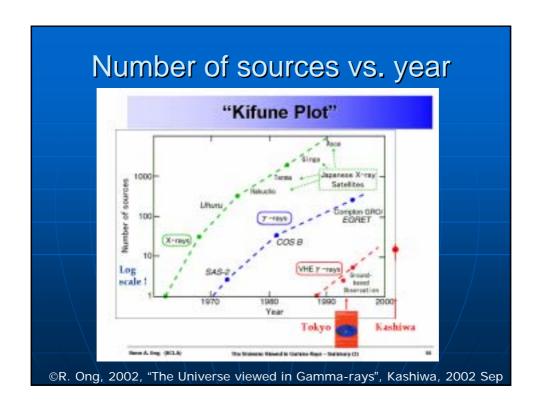






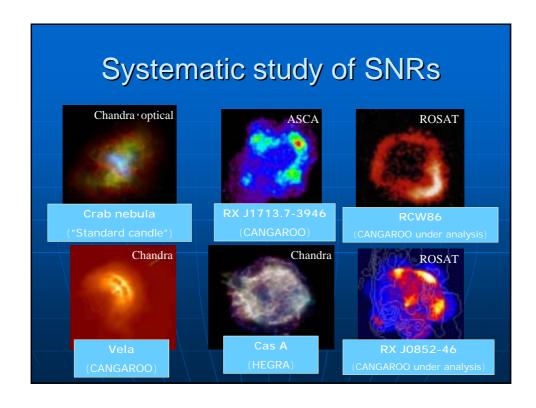


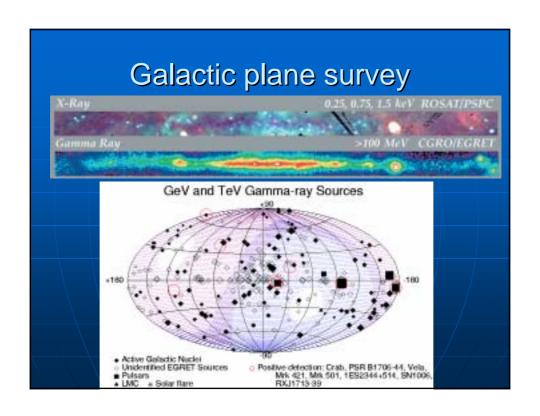




#### Next several years

- Systematic study of SNRs
- Survey of the galactic plane
- International/multiwavelength coordination
- Present fund ends in 2004 March
- We need running fund!
- T1 mirror & electronics: replacement is necessary!







#### **Next CANGAROO**

- R&D for 15m class telescope
  - Light, durable plastic mirror
     → the reflector can be larger
  - Mountain altitude? ← Lower threshold
  - Lower energy gamma-rays, overlapping the satellite (GLAST) region

GLAST	IACT
50MeV-100GeV	100(→20?)GeV-100TeV
All sky survey	Deep survey