

宇宙線が気候変動および 気象現象に及ぼす影響について

宮原 ひろ子 (一次線グループ)
Hiroko MIYAHARA,

Institute for Cosmic Ray Research,
The University of Tokyo, Japan

hmiya@icrr.u-tokyo.ac.jp

Collaborators

Yusuke Yokoyama (AORI, The Univ. of Tokyo)

Yasuhiko T. Yamaguchi (AORI, The Univ. of Tokyo)

Takeshi Nakatsuka (Nagoya University)

Hong K. Peng (The Univ. of Tokyo)

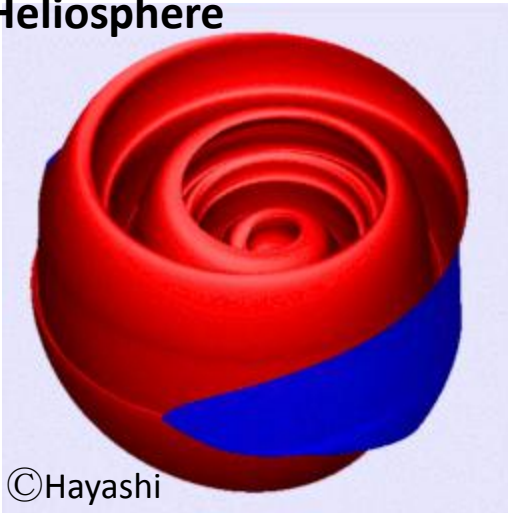
Yukihiro Takahashi (Hokkaido University)

Mitsuteru Sato (Hokkaido University)

Hiroyuki Matsuzaki (MALT, The Univ. of Tokyo)

Fuyuki Tokanai (Yamagata Univ)

Heliosphere



©Hayashi

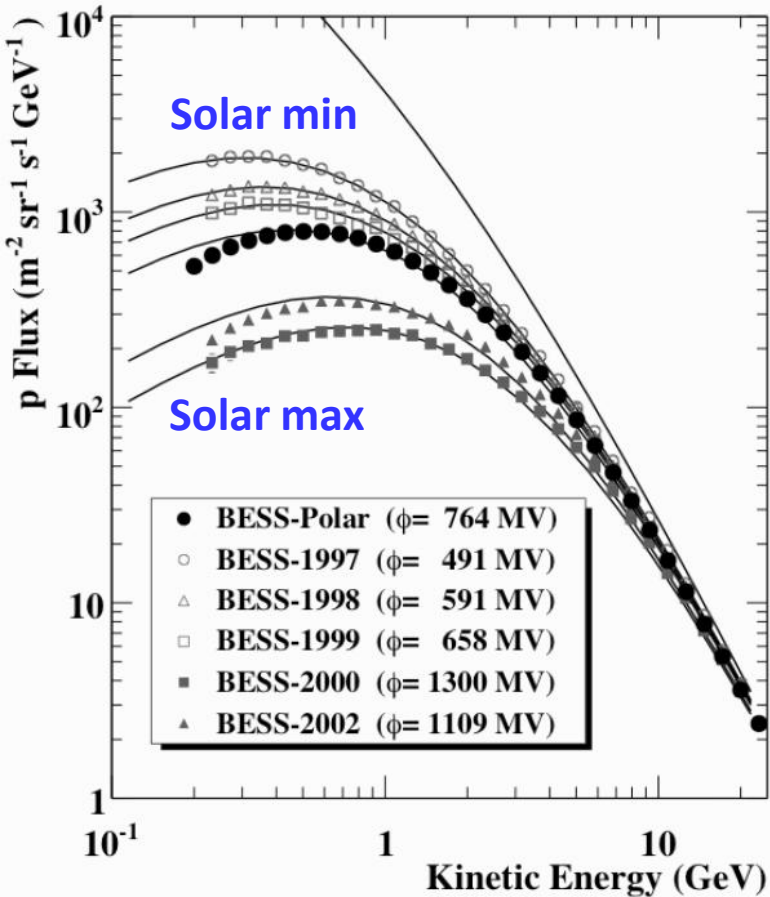
Contents

宇宙線の22年周期変動

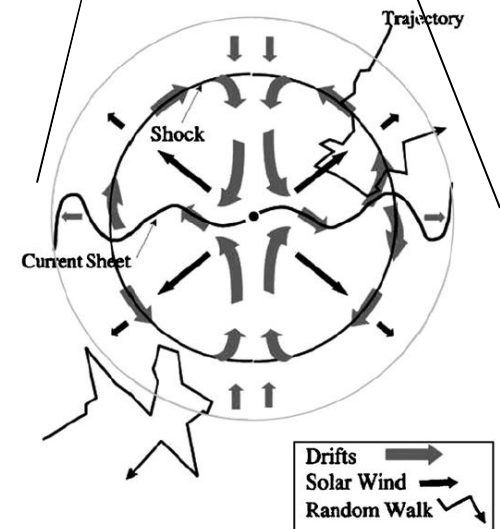
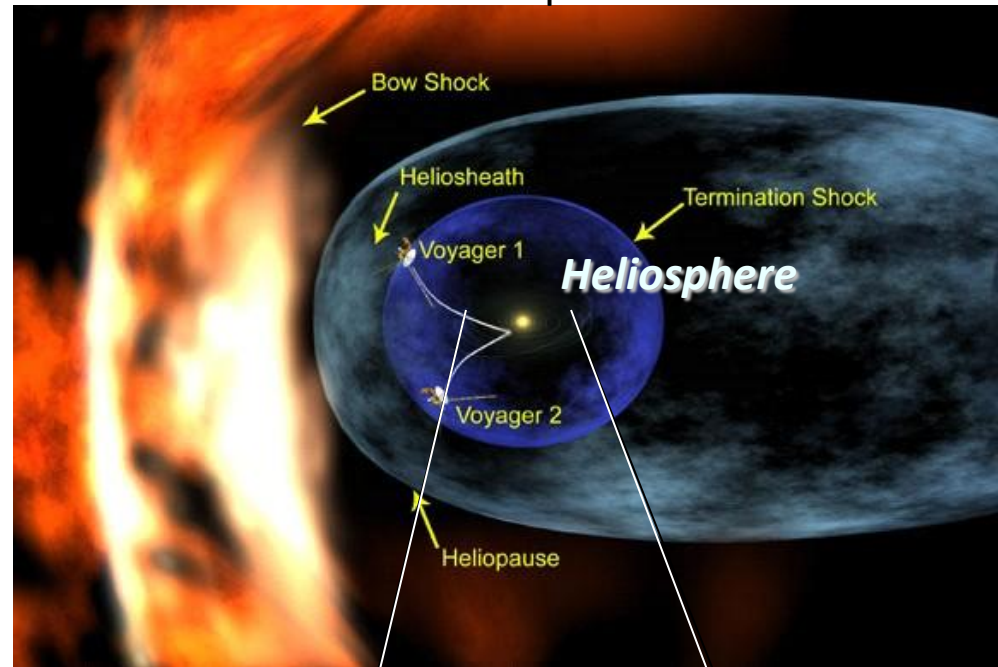
マウンダー極小期における宇宙線変動と気候変動

宇宙線の27日周期変動と赤道熱帯域の雲活動の変動
(宇宙線が気候システムに及ぼす変動のトレースの観点、
気象への影響の観点から)

Solar modulation of Galactic Cosmic Rays (GCRs)



- Charged particles (mainly protons)
- Accelerated at supernova remnant

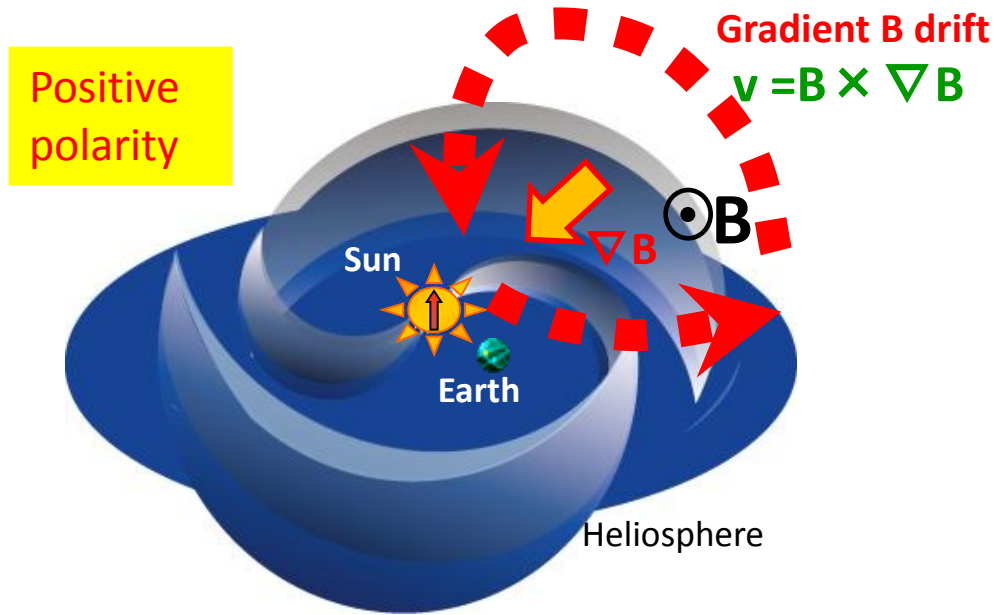
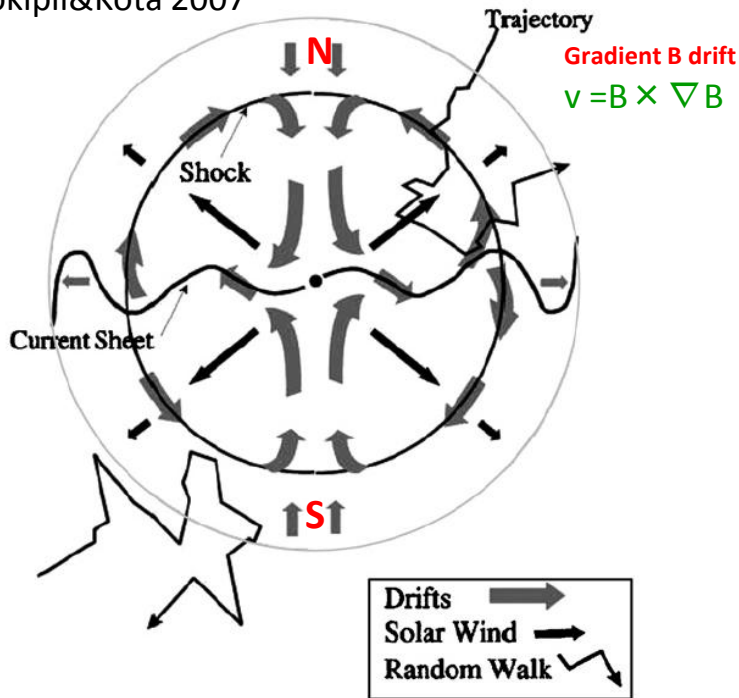


- diffusion
- advection by solar wind
- drift ($B \times \nabla B$ ドリフト)

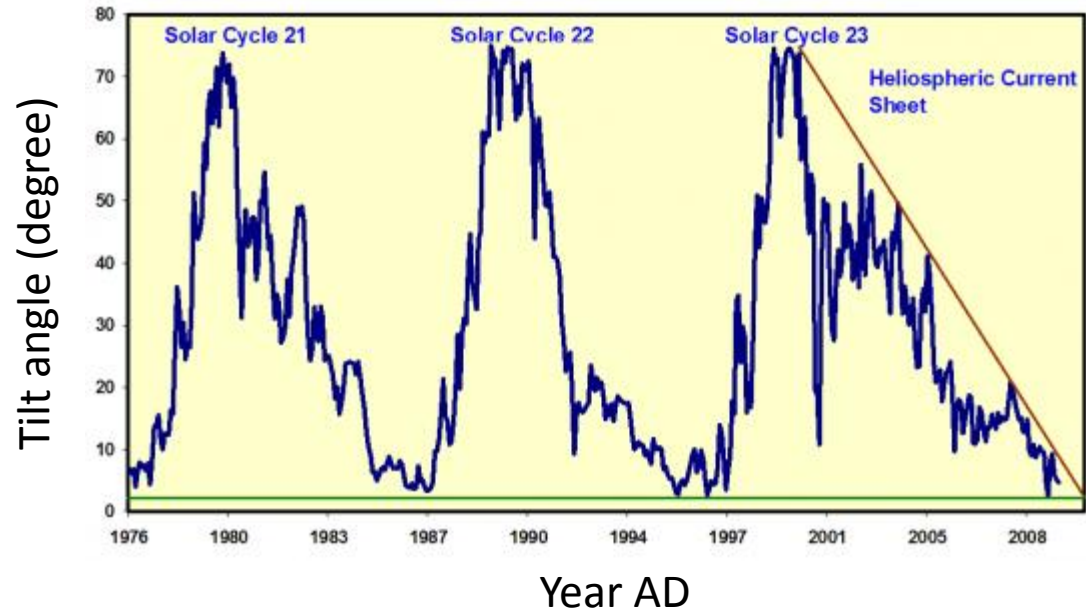
Solar modulation of cosmic rays & Drift effect

Heliosphere

Jokipii&Kota 2007

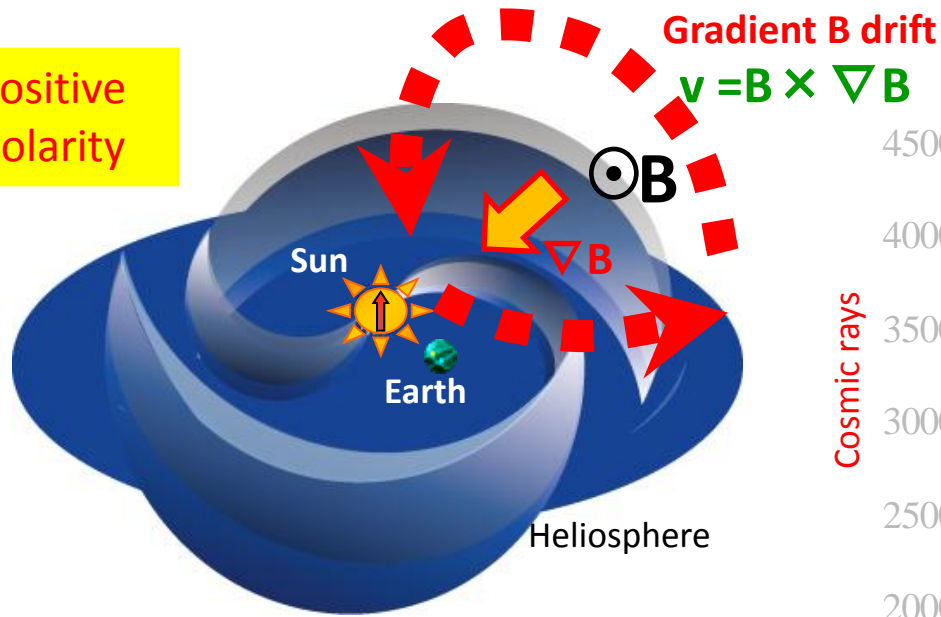


- diffusion
- advection by solar wind
- drift ($B \times \nabla B$ ドリフト)

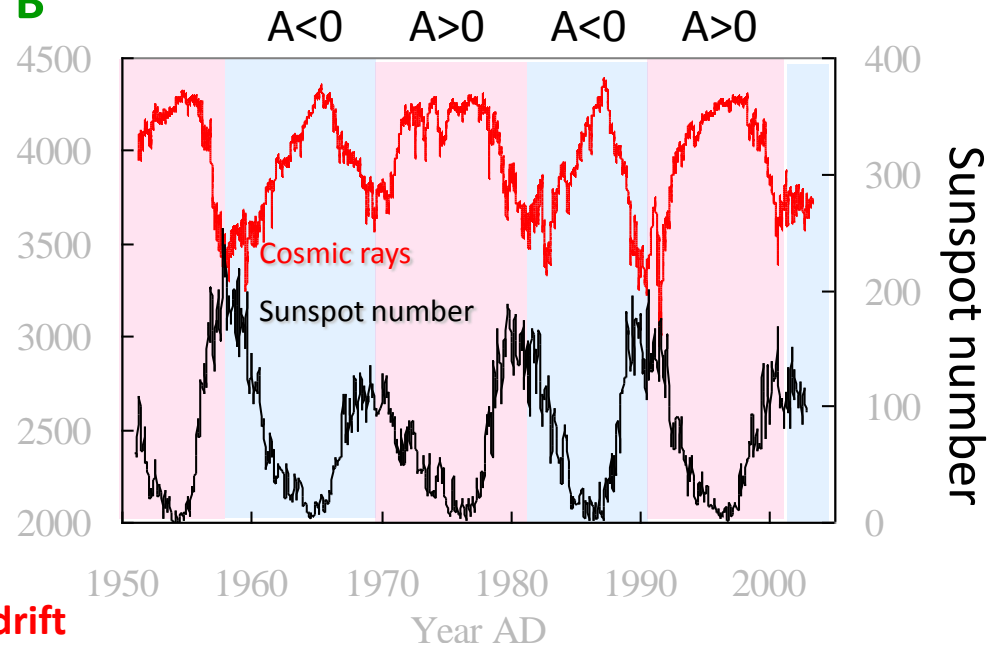


Cosmic ray variation & Solar magnetic polarity

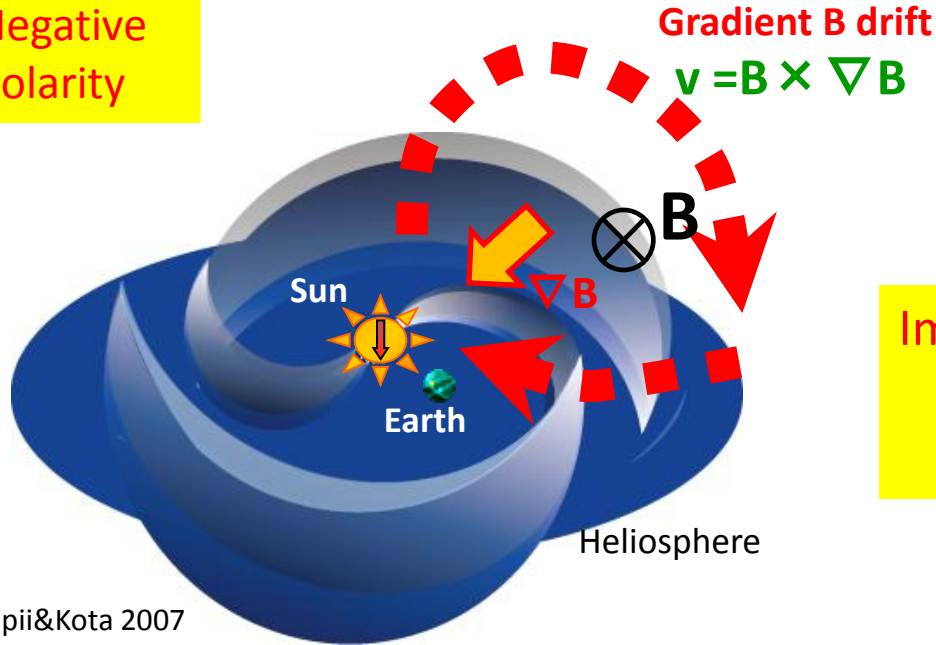
Positive polarity



Cosmic rays



Negative polarity

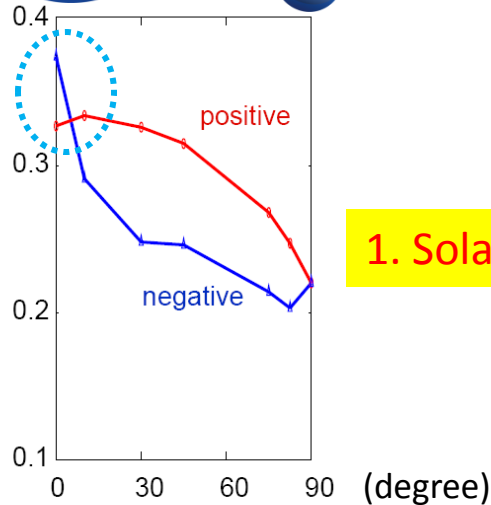


- Important parameters for solar modulation
- solar dipole magnetic polarity
 - tilt angle of heliospheric current sheet

Variable "22-year" variation of cosmic rays

Miyahara et al., 2009

Expected cosmic rays

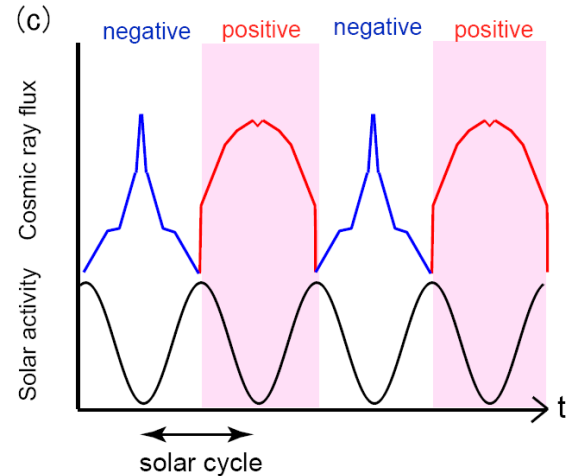
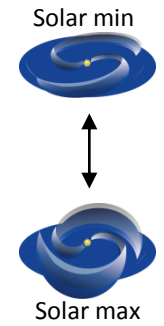


1. Solar polarity

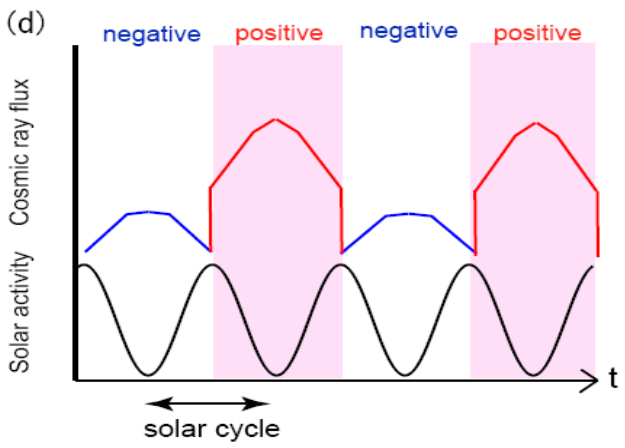
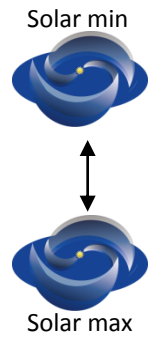
2. Tilt angle

Kota&Jokipii, 2003

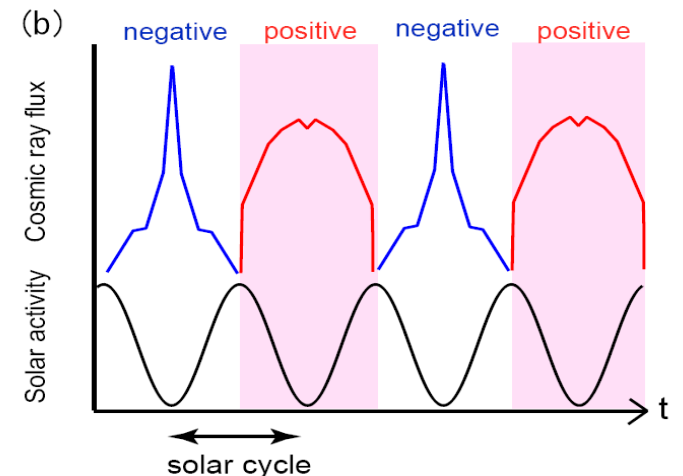
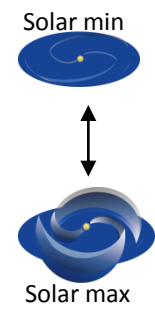
Modern: 5-75 degrees



If 30-75 degrees



If 0-75 degrees



Production of cosmogenic nuclides: ^{14}C and ^{10}Be

Galactic cosmic rays



Attenuation by solar/geo-magnetic field

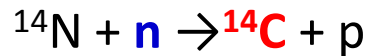


Air shower in the atmosphere



Secondary neutron

Spallation



Atmospheric circulation



Photo-synthesis

Tree ring

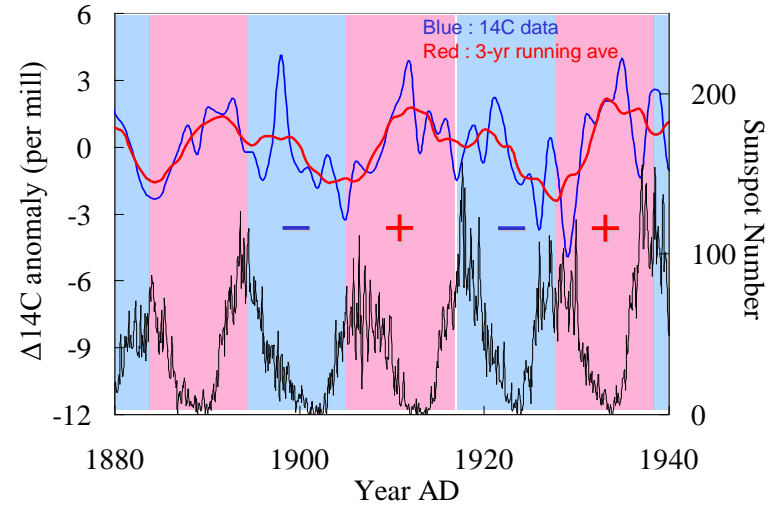
^{10}Be etc.



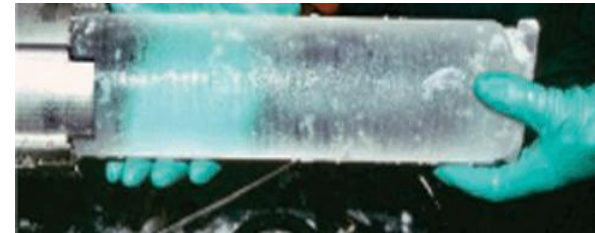
Precipitation

Antarctic/Greenland Ice sheets

^{14}C anomaly in tree rings

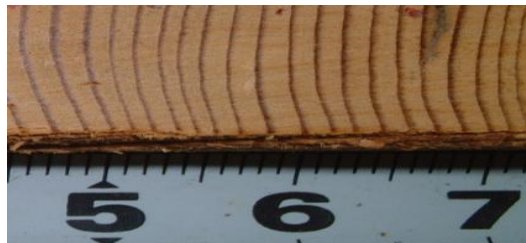


3-yrs of lag in carbon cycle is corrected



Clear signal

A few years of dating error

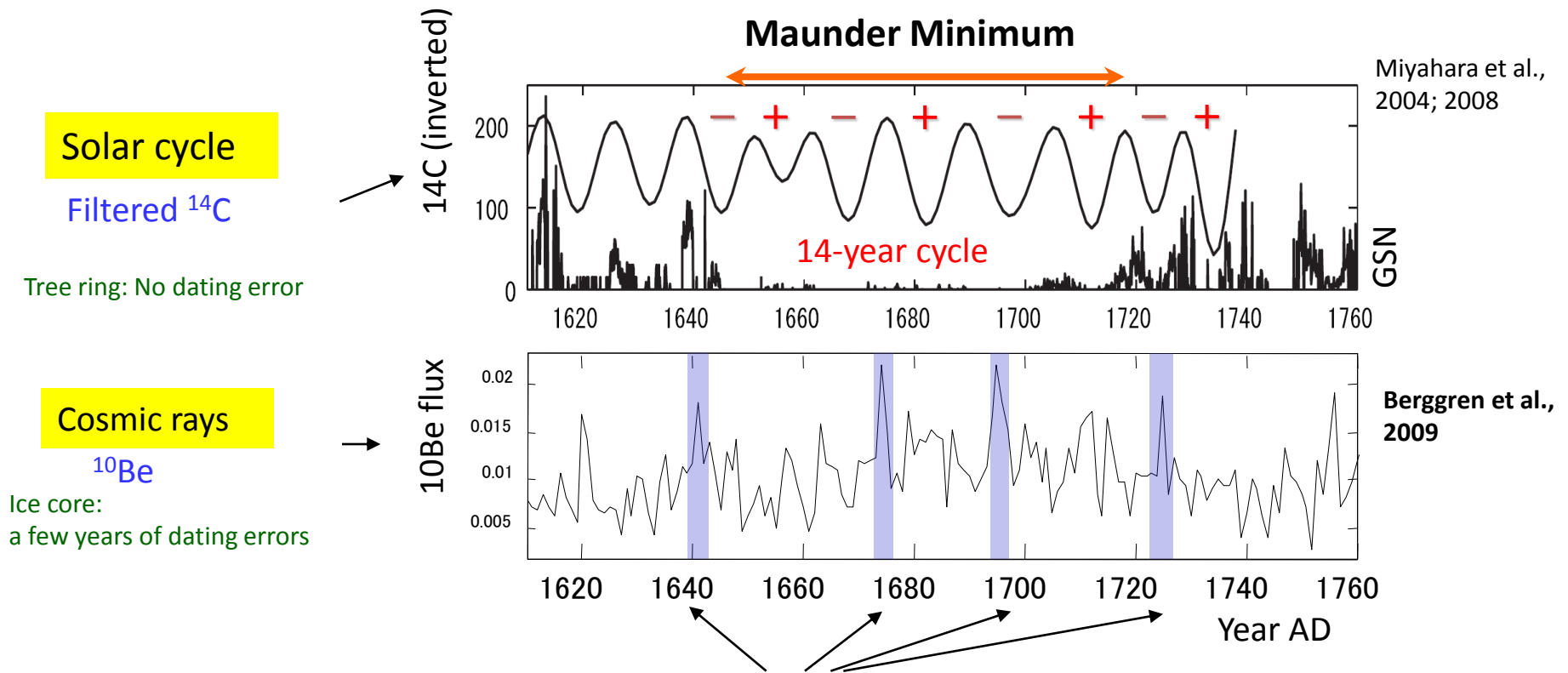


Absolute age

Strongly attenuated signal

Cosmic-ray “22-year (28-year)” variation at the Maunder Minimum

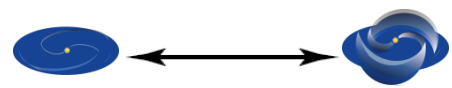
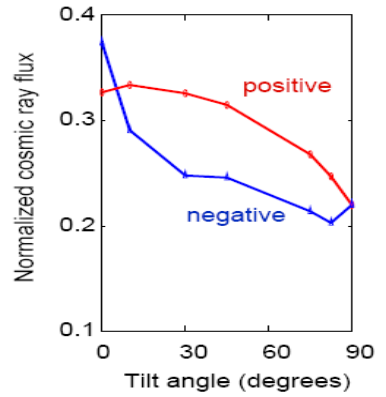
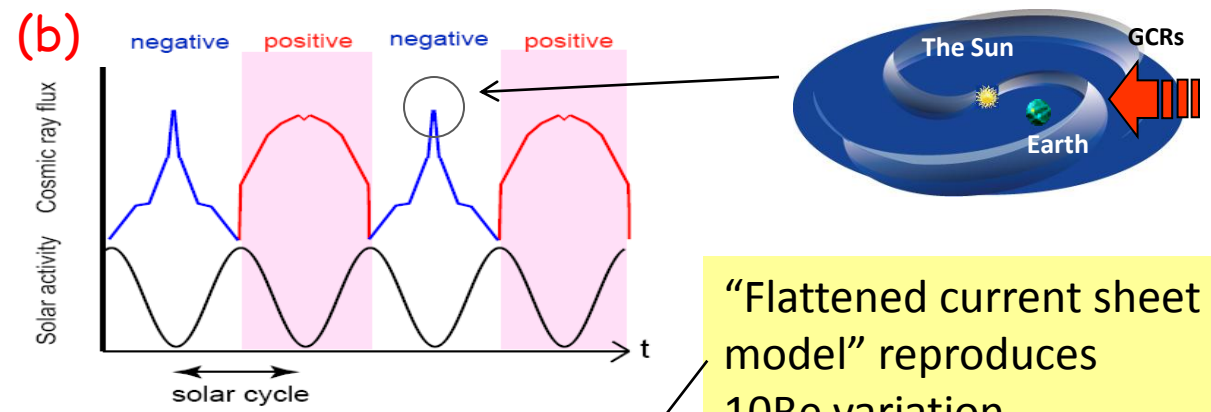
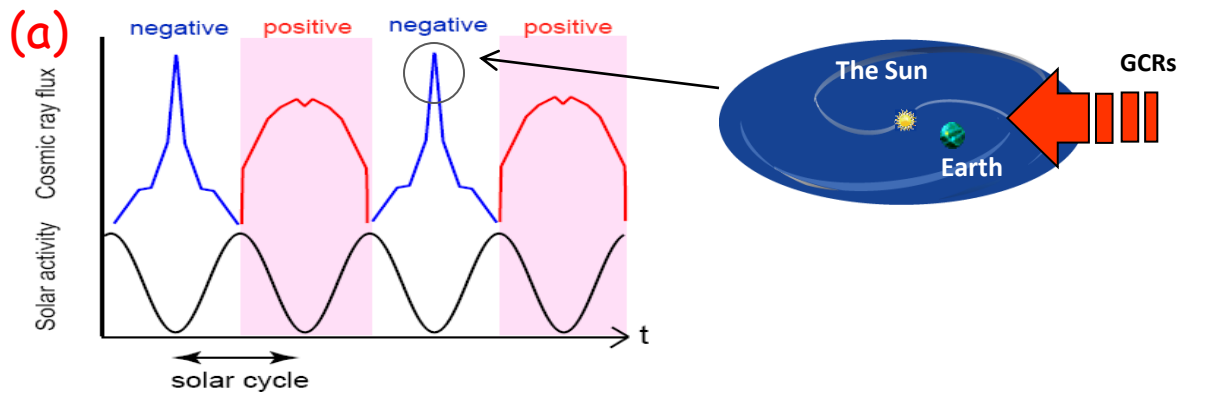
Miyahara et al., IAU proc., 2009, Yamaguchi et al., PNAS, 2010



- Periodic cosmic ray enhancements, only for negative polarity (~28-year period)
- **1-year scale enhancement, 30-50% higher than the peak for positive polarity**
- Significant manifestation of drift effect

Based on Kota&Jokipii, 1983; 2003

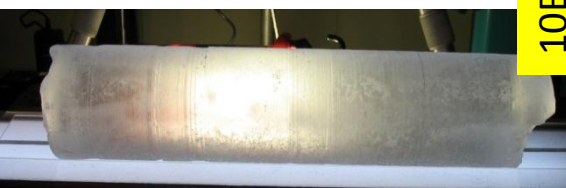
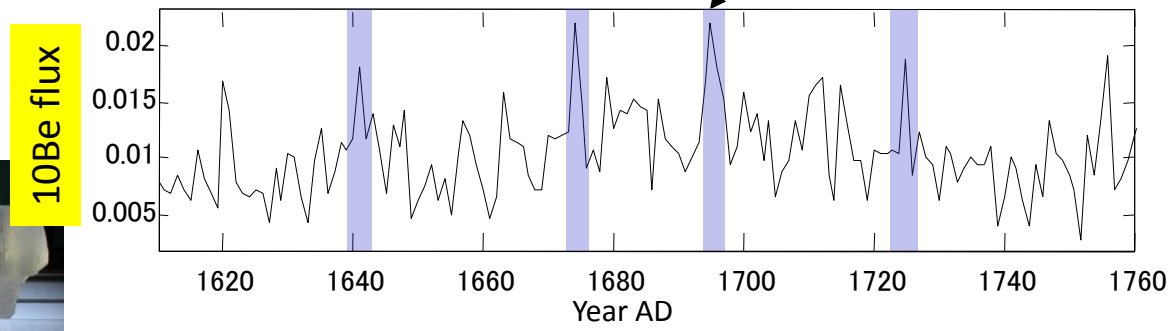
(a) 0 deg. at cycle min
(b) 5 degs. at cycle min



Heliospheric Magnetic field

“Flattened current sheet model” reproduces 10Be variation

Berggren et al., 2009



10Be flux

What ^{14}C and ^{10}Be suggests for the Maunder Minimum

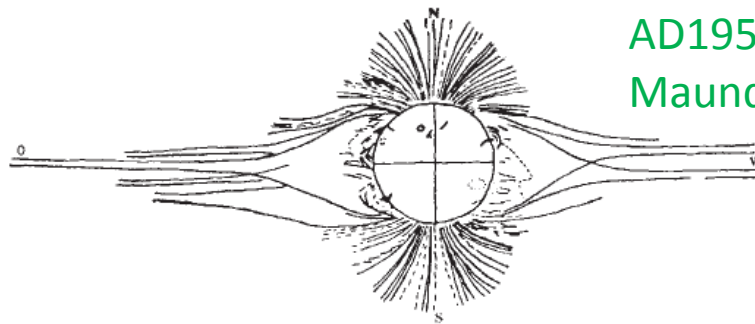
Solar Cycle length : ~ 11 years

Magnetic polarity reversal : YES (~ 22 -year period)

Onset : two preceding 11-year cycles

Cosmic ray variations : Strong 11-year component

Heliospheric current sheet : More flattened



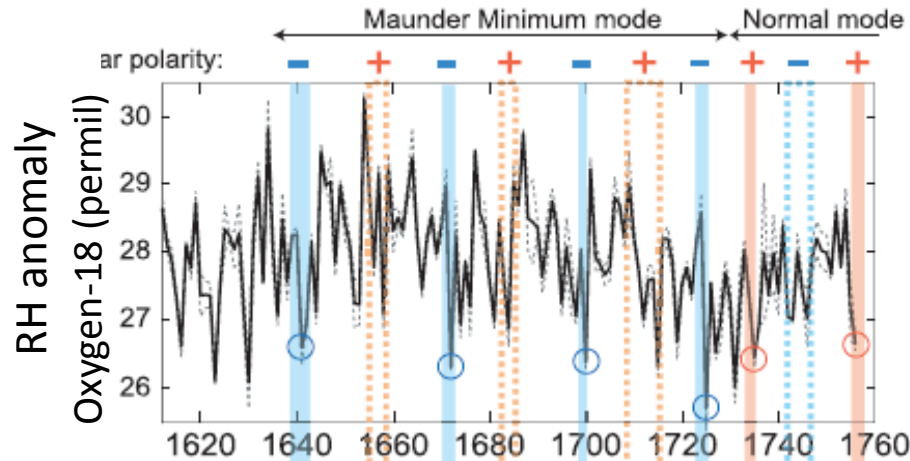
AD1954 case: stronger polar field
Maunder Min: weaker equatorial field

Fig. 2 The structure of a sunspot minimum solar corona drawn from eclipse photographs¹¹ (June 30, 1954) obtained in Kozeletsk.

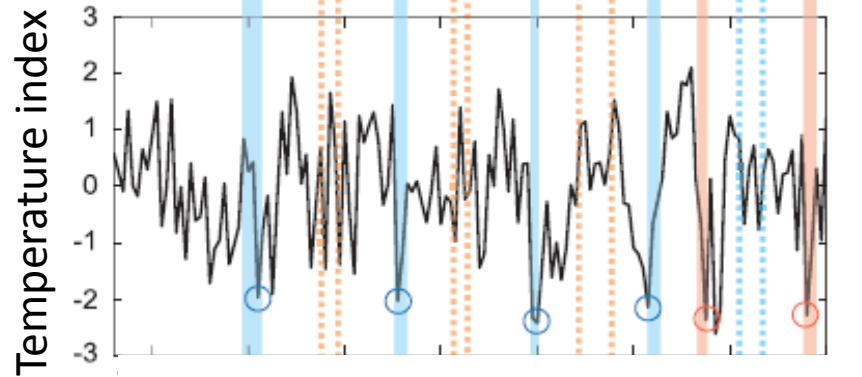
Any impact on climate? : GCR spikes can be the tracer

Climate response to cosmic-ray spikes during the Maunder Minimum

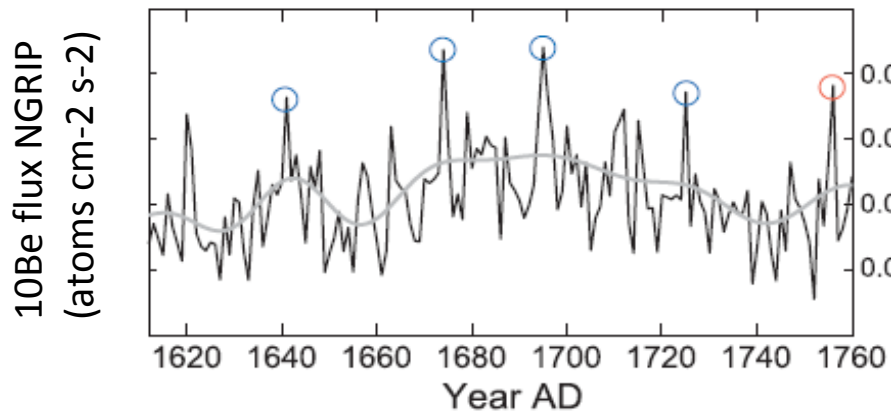
Yamaguchi, Yokoyama,
Miyahara et al., PNAS, 2010



Humidity in Japan around June



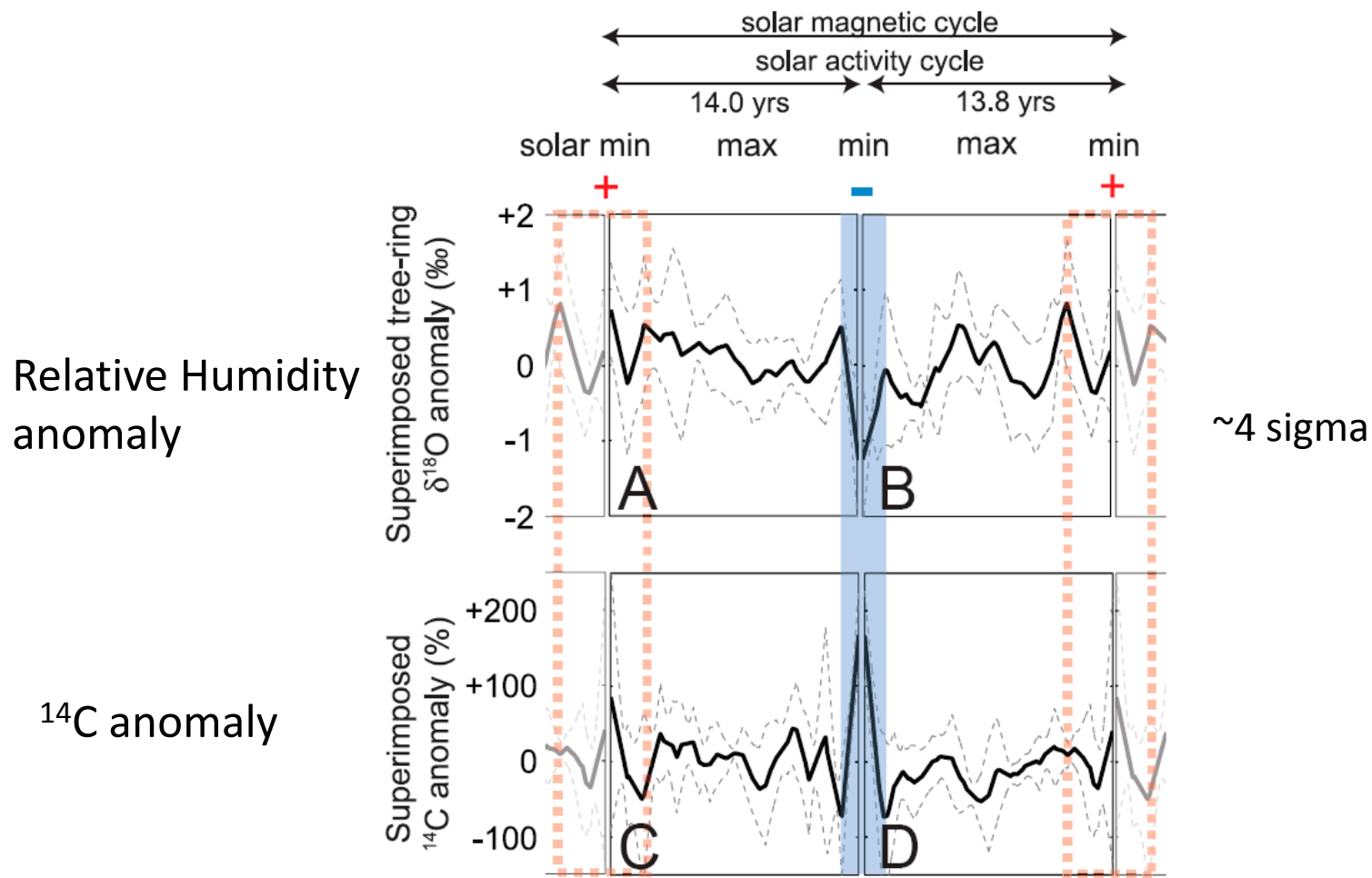
Greenland winter temperature



GCR flux
(Berggren et al., 2009)

Ice core:
a few years of dating errors

Superposition of four 1-year spikes for ^{14}C (GCR) and ^{18}O (climate)



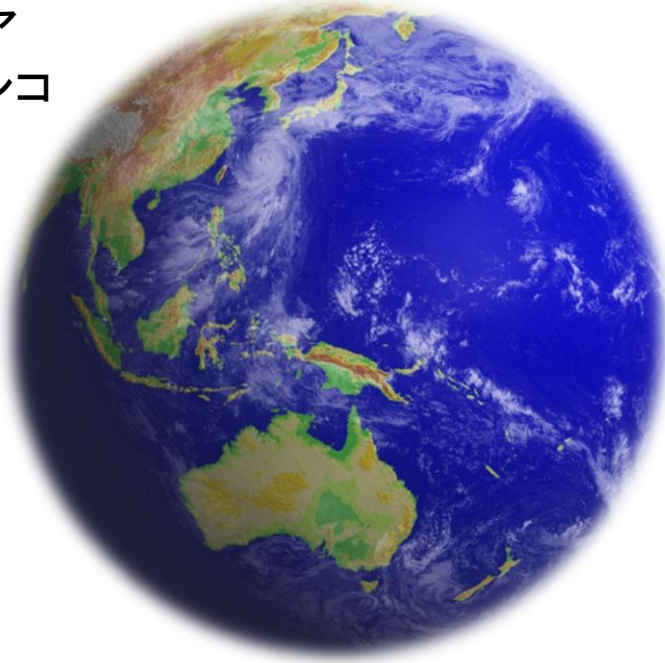
No time lag!

現在進行中の計画 : 宇宙線スパイクをトレーサーとした 全球気候マッピング
(計4イベント)

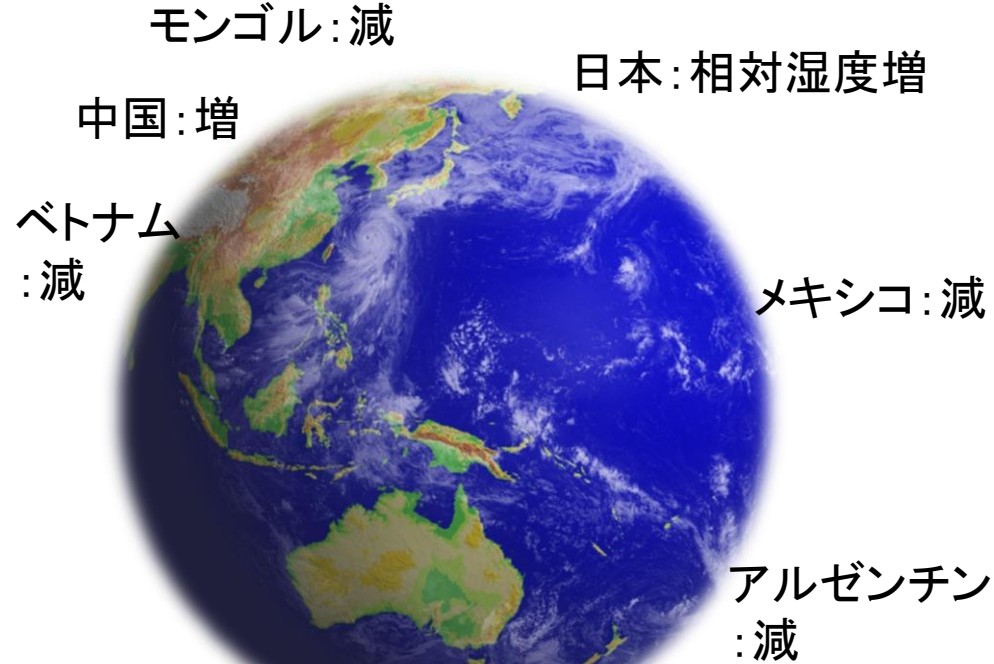
炭素14の超高精度分析も山形大学で実施中 (従来の1/3の統計誤差)

気温: 全球で寒冷傾向

ヨーロッパ全域
ロシア
メキシコ
日本
など



降水: 強い地域性?(モンスーンを介した影響?)

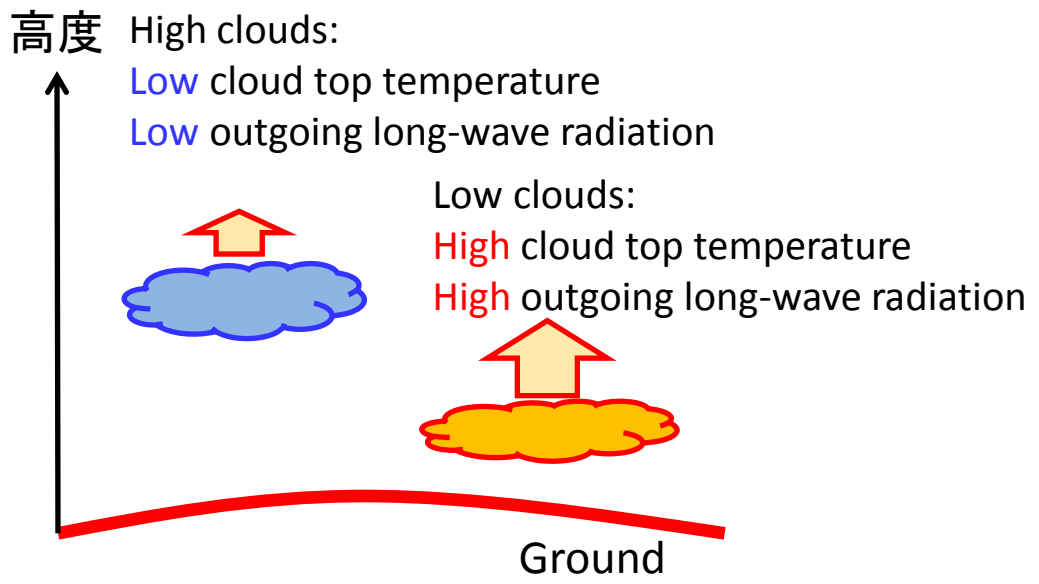


宇宙線に対する地球気候システムの複雑な応答解明

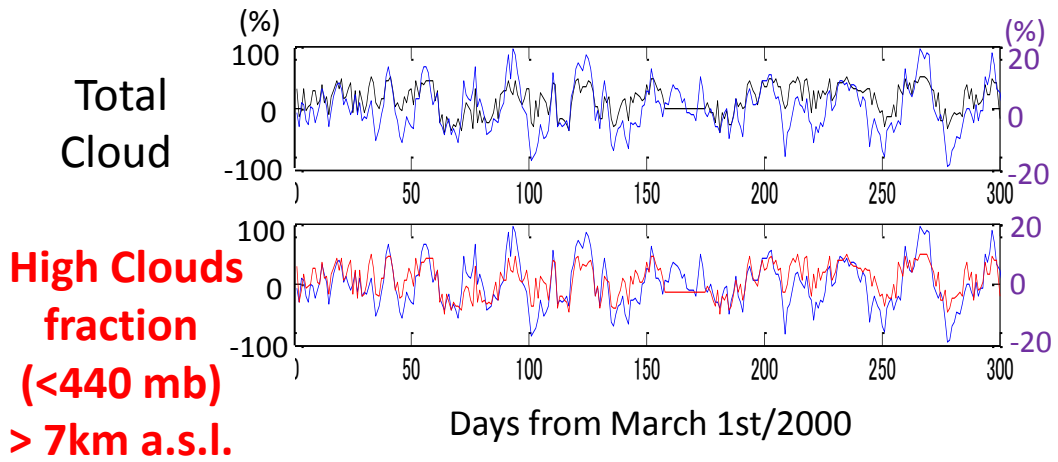
宇宙線はどのように気候を変えるのか？

日日スケールからの検証

- Data :**
- Daily Outgoing Long-wave Radiation**
Duration : 1979/Jan – 2004/Dec
10 x 10 degrees grid
 - MODIS terra/aqua cloud fraction
Duration : 2000/March –



OLR (inverted) vs MODIS cloud fraction

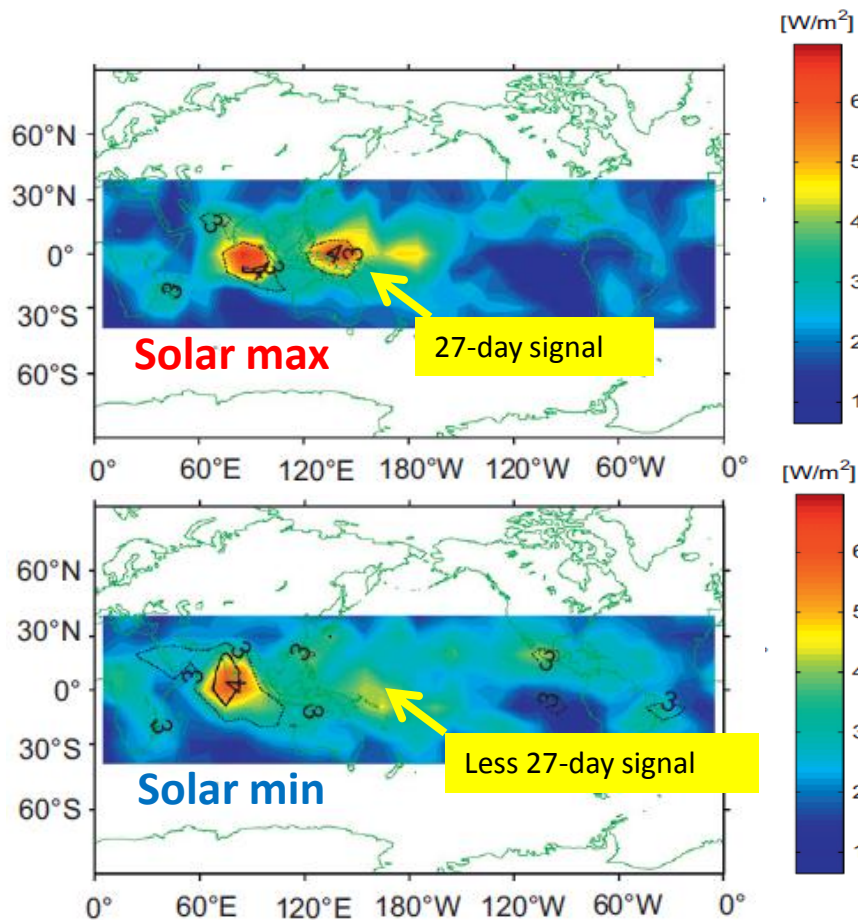


Outgoing Longwave Radiation(OLR)
> monitoring high-altitude clouds

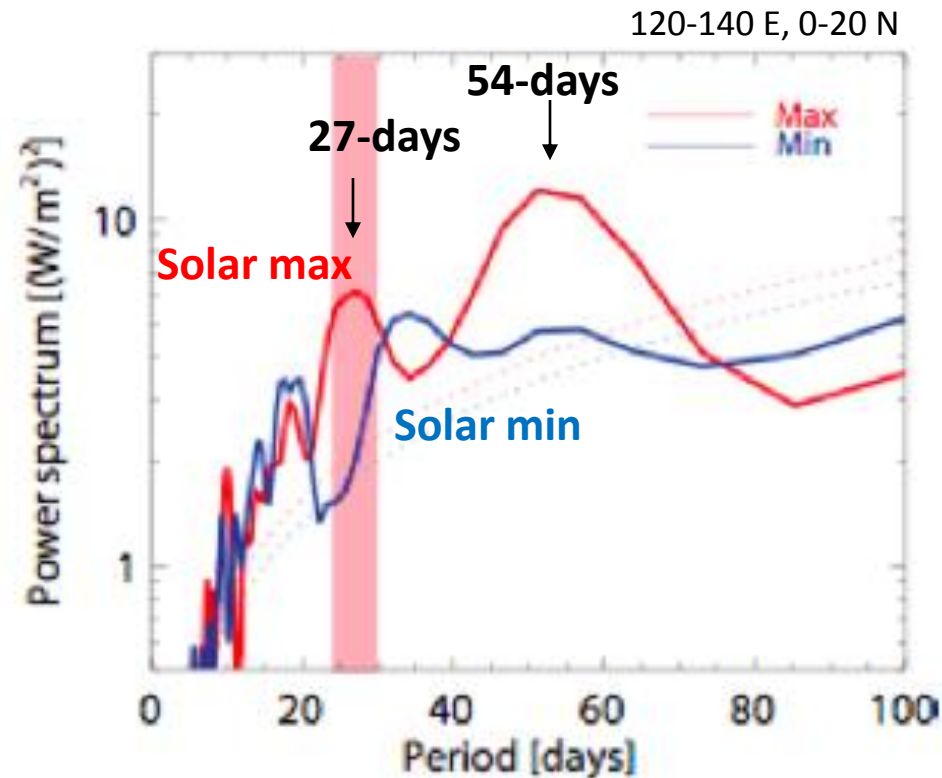
赤道熱帯域の高層雲の27日周期

Takahashi et al., ACP, 2010

Hong, Miyahara et al., JASTP, 2010



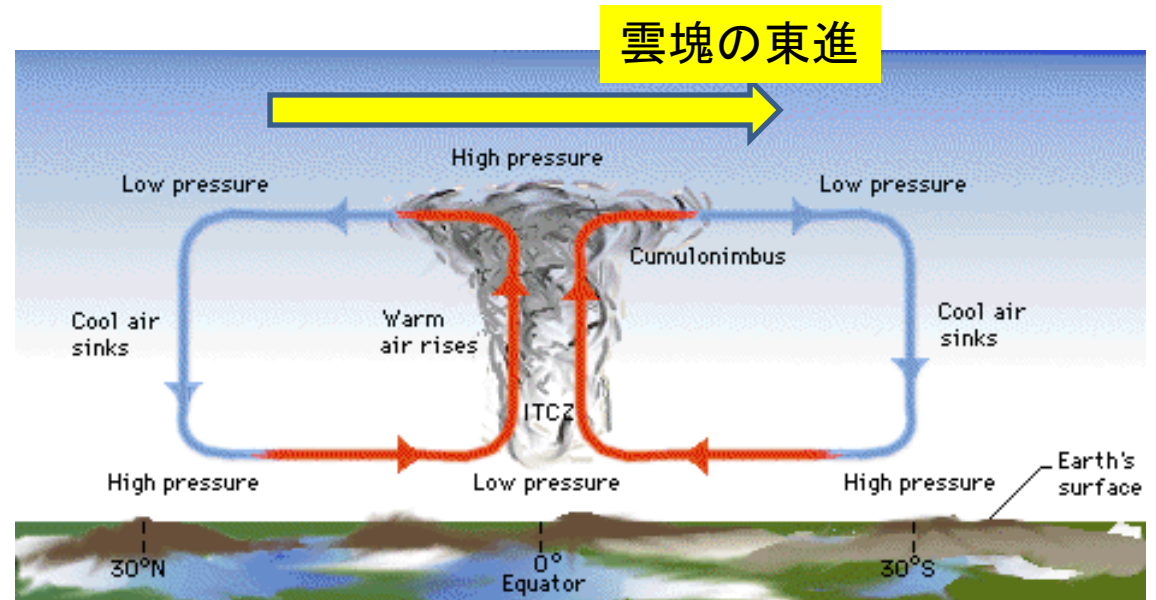
Frequency spectrum of OLR



※27日周期は、11年変動の極大のみで顕著

Dynamic cloud activity at equatorial region (Madden-Julian Oscillation) has 30-60 day periodicity. Intrinsic period (30-50 day period) is modulated to be 27-day and 54-day periods at solar max

マッデン・ジュリアン振動とは

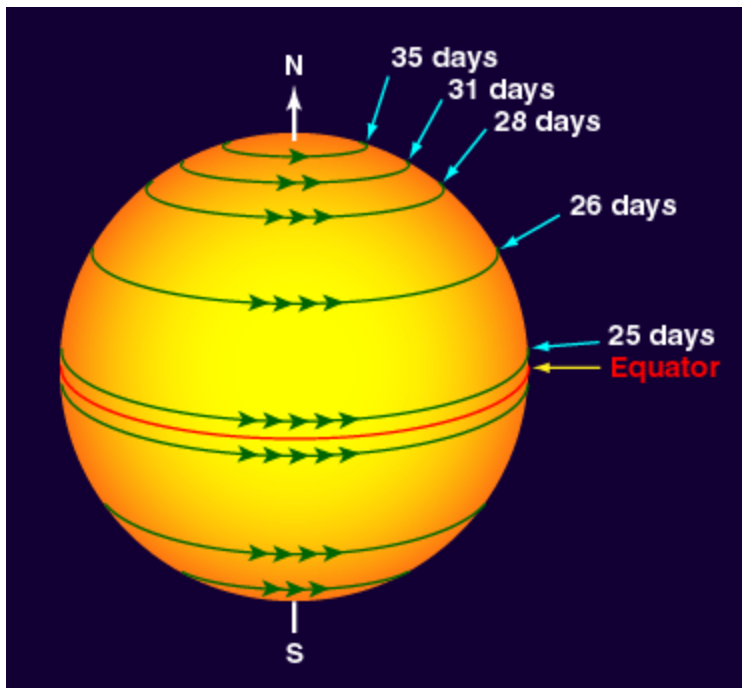


<http://www.ccsr.u-tokyo.ac.jp/~sato/ncam/index.html>

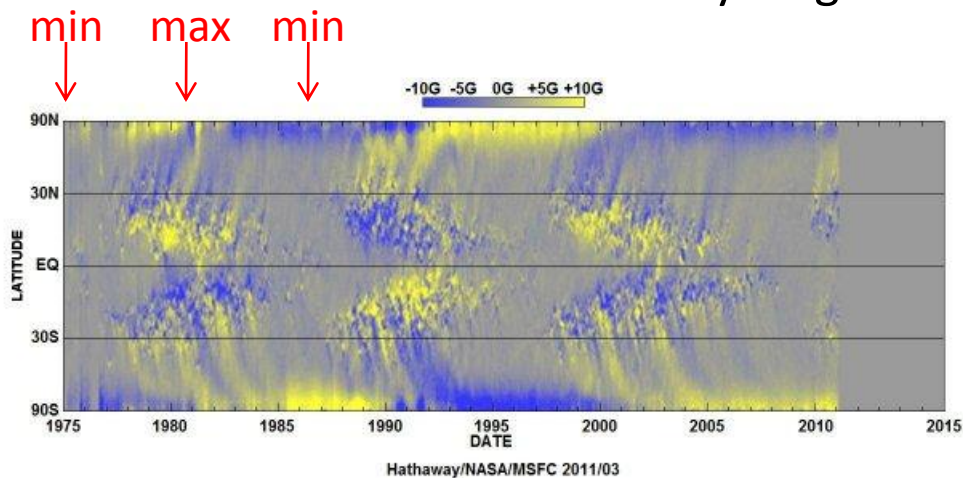
- ・赤道熱帯域の雲活動の30－60日周期(周期性の決定因子は未解明)
- ・エルニーニョの開始、終焉をコントロール
- ・南北半球のモンスーンを介して、中高緯度の気候にも影響

Solar differential rotation

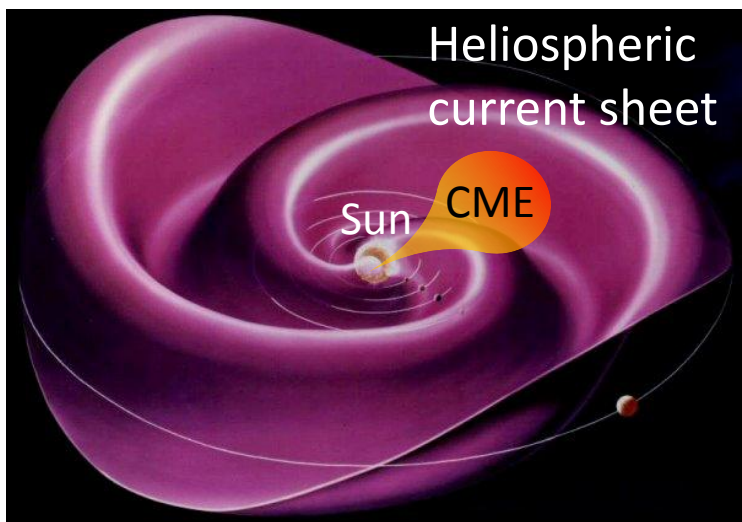
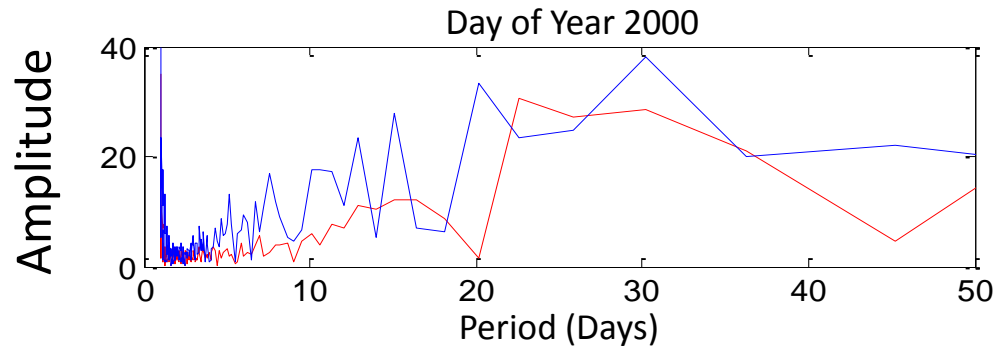
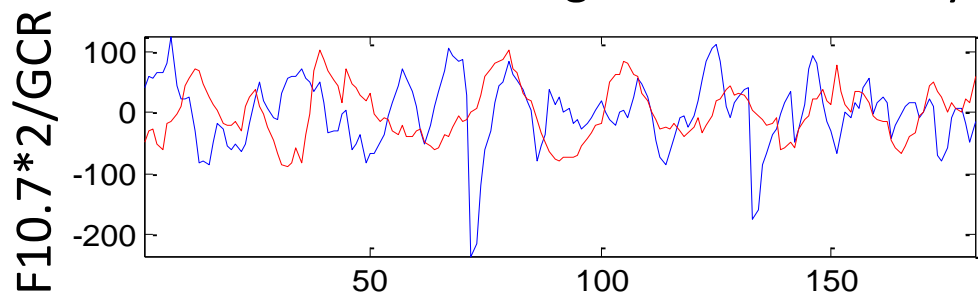
Differential Rotation



Butterfly diagram

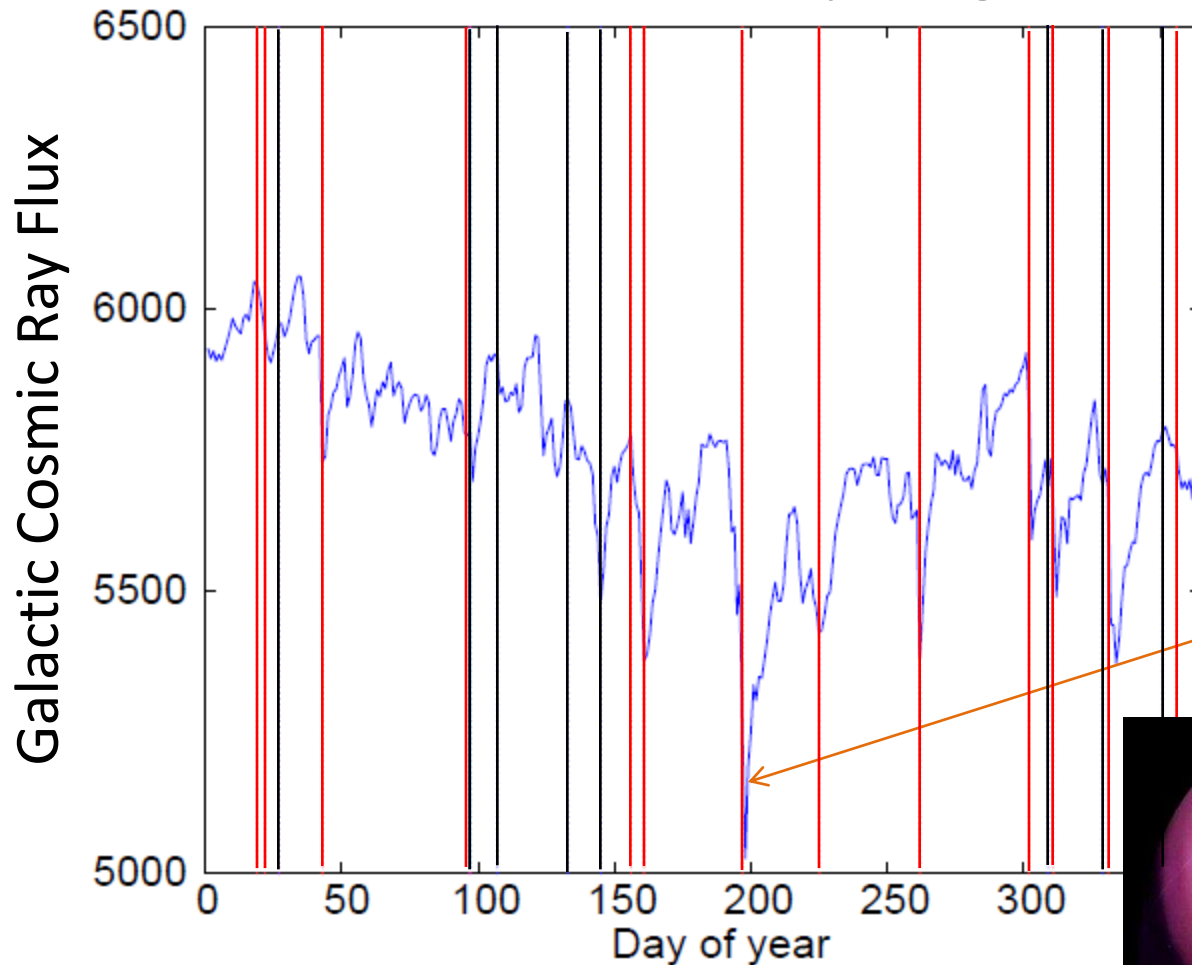


Red : F10.7 solar radio flux
Blue : galactic cosmic rays



Red: Solar flares (Coronal Mass Ejections (CMEs))

Black: Current sheet passage



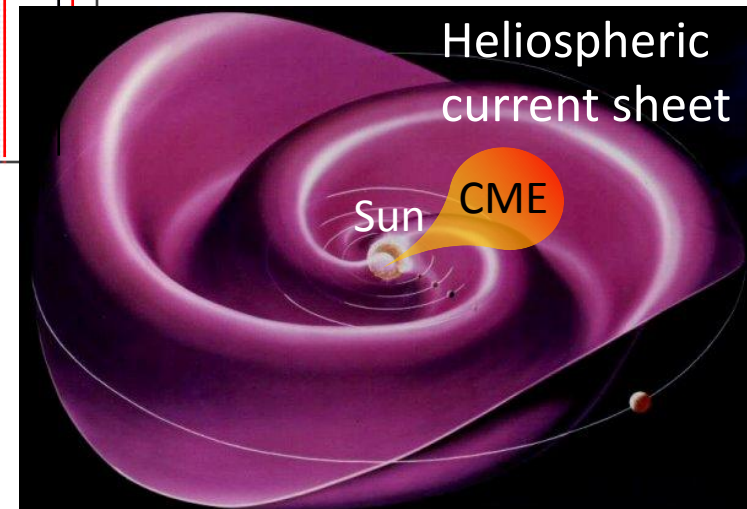
太陽極大期の変動は
ほとんどがCME起源

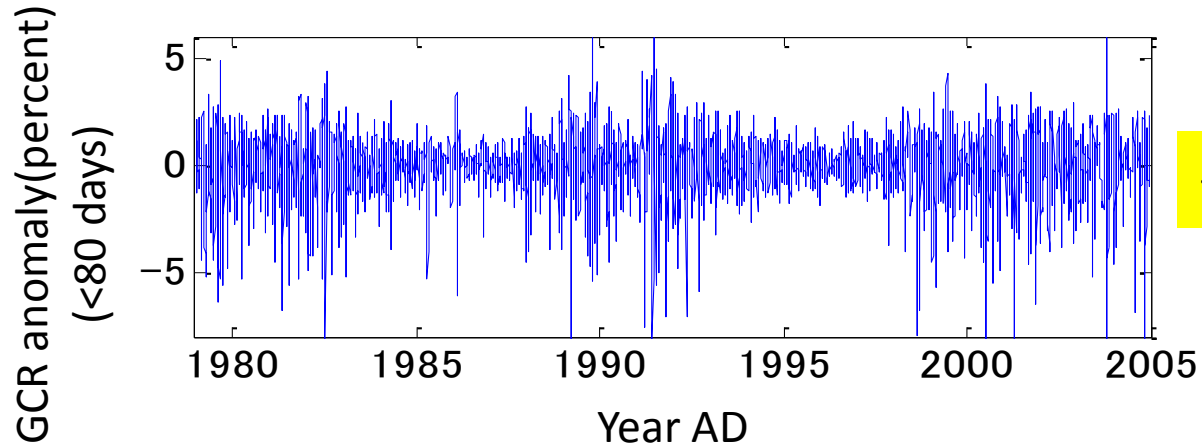
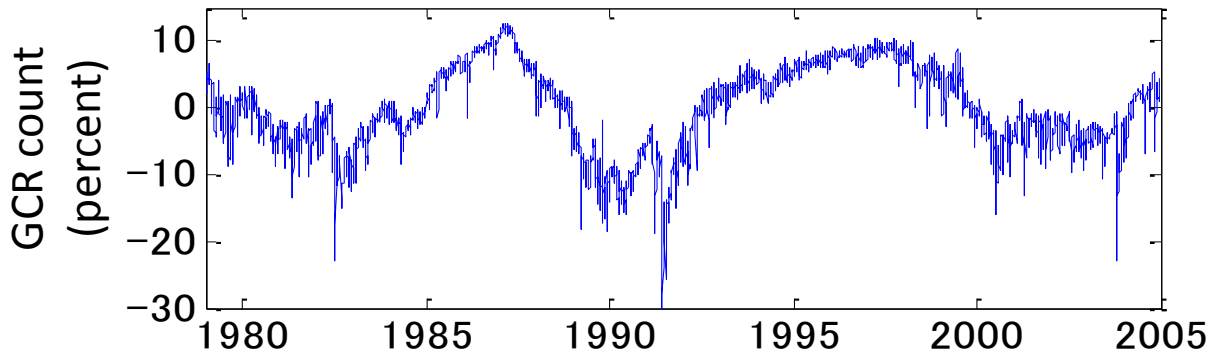
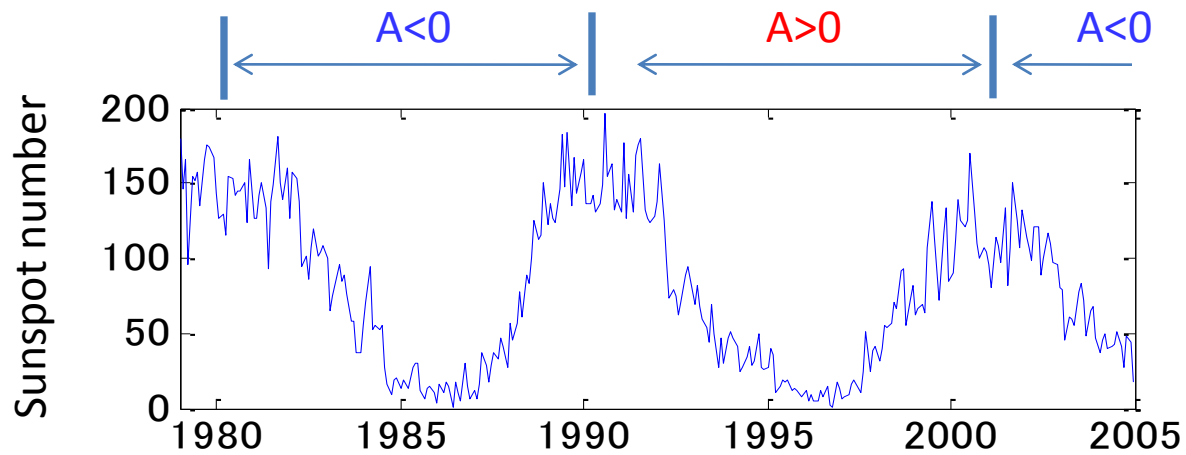
10~50日程度の周期性
(平均は27日周期)

Forbush Decrease

AD 2000

(Solar Max)





← short-term variability

Larger at 11-year **max**

(※ 未発表分は削除いたしました)

宇宙線に由来する何がどう効いているのか
の物理は今後の課題

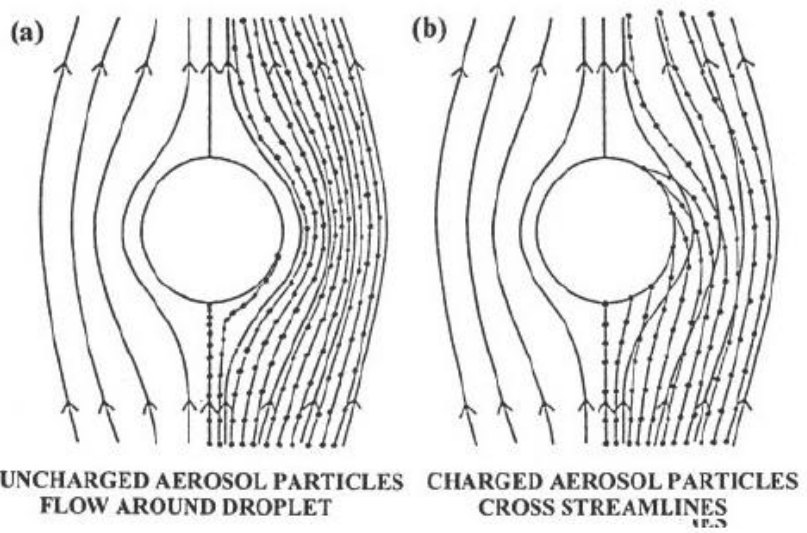


Figure 5.3. (a) Schematic of aerosol flow around a falling droplet in the absence of electrical forces. (b) Schematic of effect of electrical forces in moving aerosol particles across streamlines.

Tinsley & Yu 2006

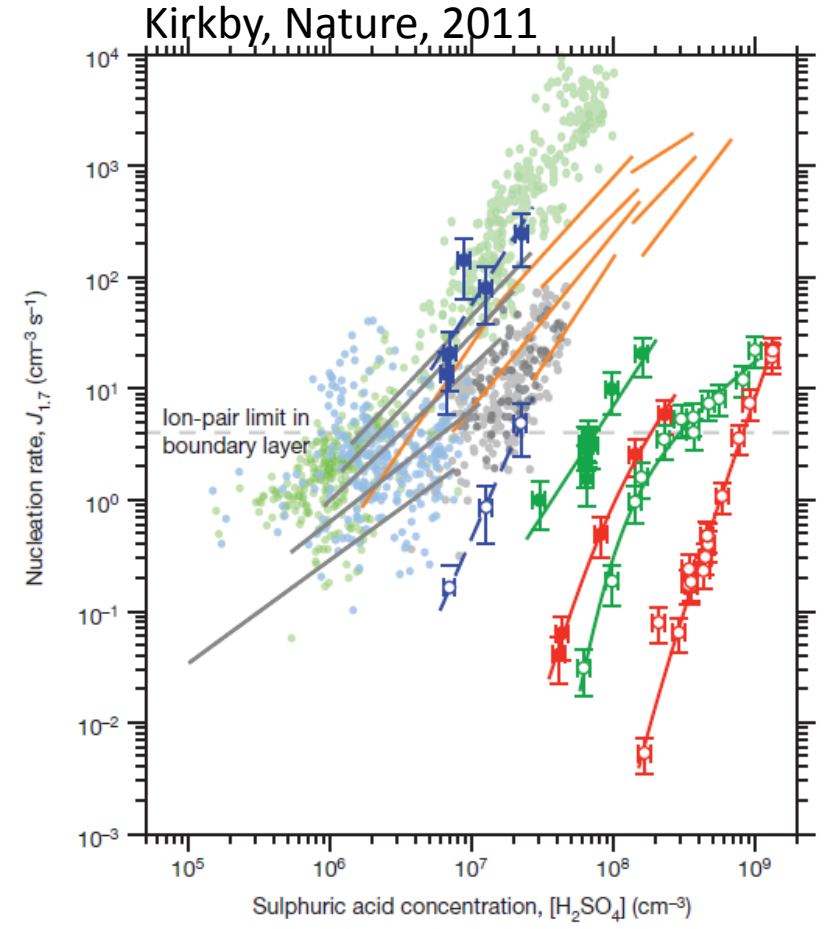
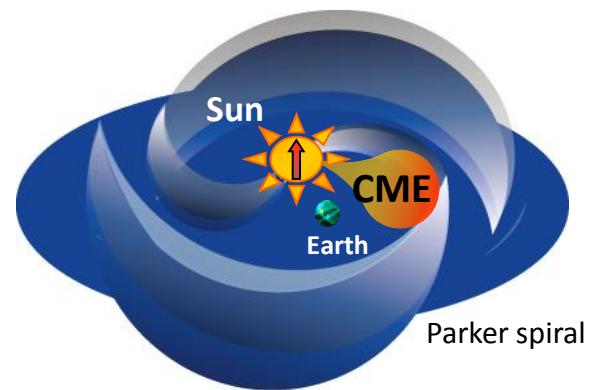


Figure 5 | Nucleation rate comparison. Comparison of CLOUD data with measurements of the nucleation rate of new particles as a function of $[H_2SO_4]$ in the atmospheric boundary layer (pale filled circles^{8,33} and pale open circles³²) and with recent laboratory experiments at room temperature (grey¹⁹ and orange²⁹ lines). The CLOUD data (large, darker symbols and lines) show the galactic cosmic ray nucleation rates, J_{gcr} , measured at 248 K (blue), 278 K (green) and 292 K (red) and at NH_3 mixing ratios of <35 p.p.t.v. (open green and red circles), <50 p.p.t.v. (open blue circles), 150 p.p.t.v. (filled blue and green circles) and 190 p.p.t.v. (filled red circles). The bars indicate 1σ total errors, although the overall factor 2 scale uncertainty on $[H_2SO_4]$ is not shown. The measurements at 278 and 292 K bracket the typical range of boundary-layer temperatures, whereas those at 248 K reflect exceptionally cold conditions. Ion-induced nucleation in the boundary layer is limited by the ion-pair production rate to a maximum of about $4\text{ cm}^{-3}\text{ s}^{-1}$.

Kirkby, Nature, 2011

まとめ

- ・太陽圏システムとして地球気候、気象を捉えなおす必要がある



(※ 未発表分は削除いたしました)