

Sidereal daily variation of ~10TeV GCR intensity observed by the **Super-Kamiokande/Tibet AS array**

宗像、加藤、安江、伏下、溝口、森下、宮坂、稲葉(信州大理)、瀧田(ICRR)

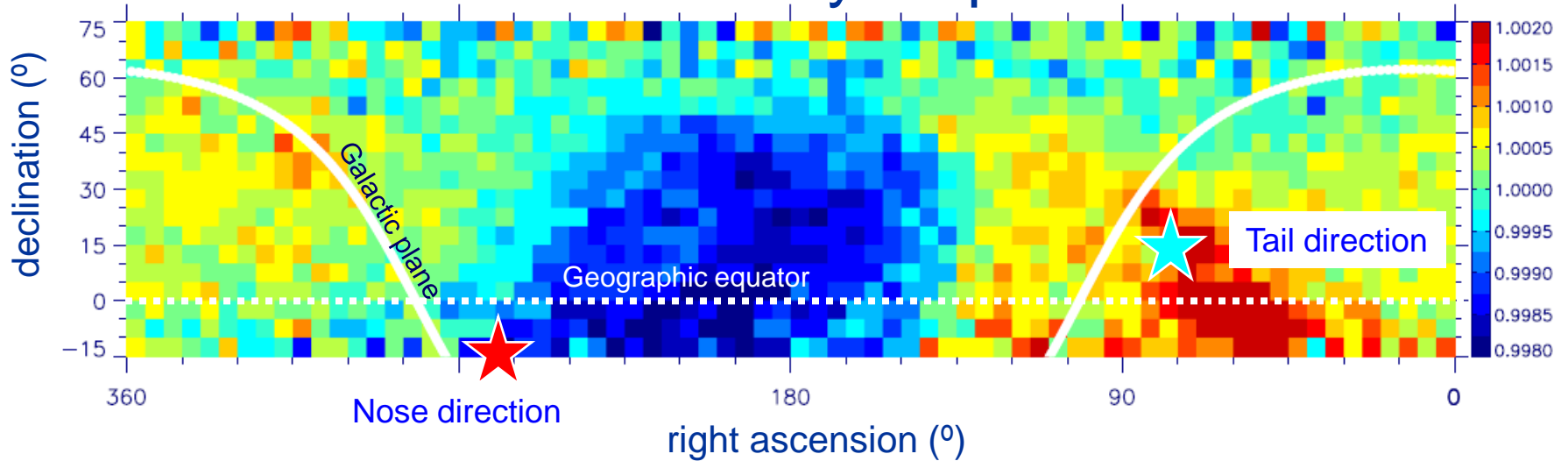
旅費(松本 \leftrightarrow 柏): 100千円(SK) / 100千円(Tibet)

- Modeling the **large-scale anisotropy** with **Global Anisotropy (GA)** and **Additional Excess (AE)**.
 \Rightarrow Amenomori et al., Proc. 31st ICRC, 2009.
- **Solar cycle dependence** of the diurnal anisotropy observed with the Matsushiro UG- μ detector.
 \Rightarrow Munakata et al., ApJ, submitted.

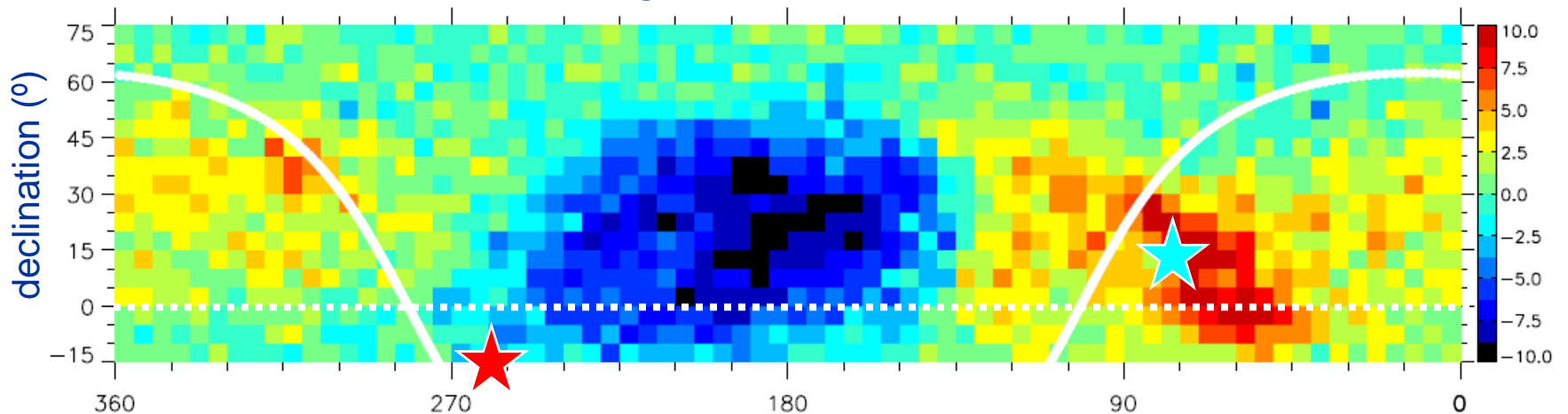
2D sky map of CR intensity by Tibet AS γ

6 years data in Nov.1999-Oct.2005 (Amenomori et al., *Science*, **314**, 2006)

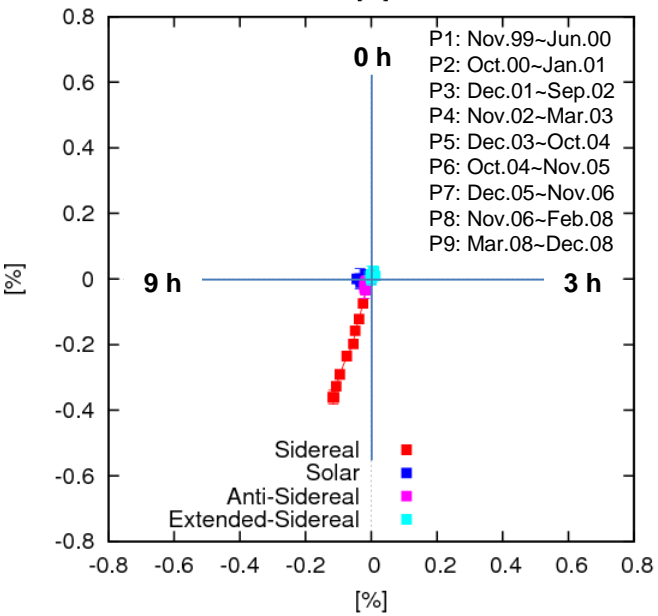
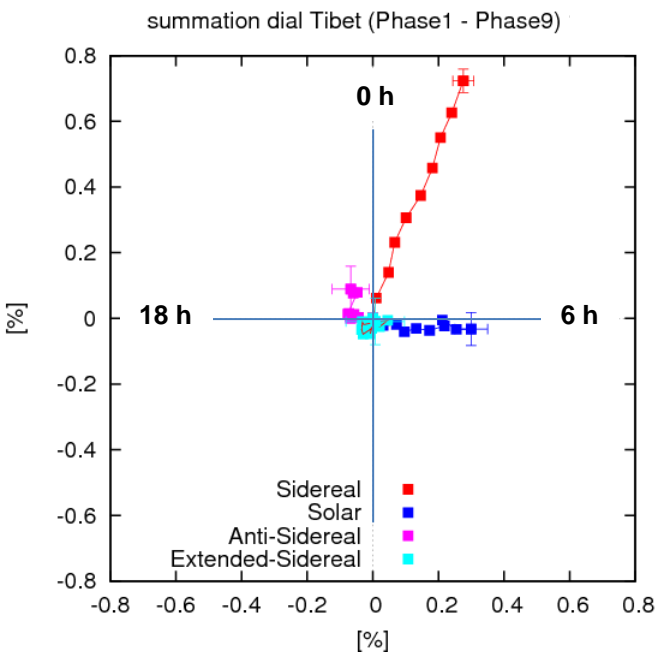
Intensity map (5°x5° pixels)



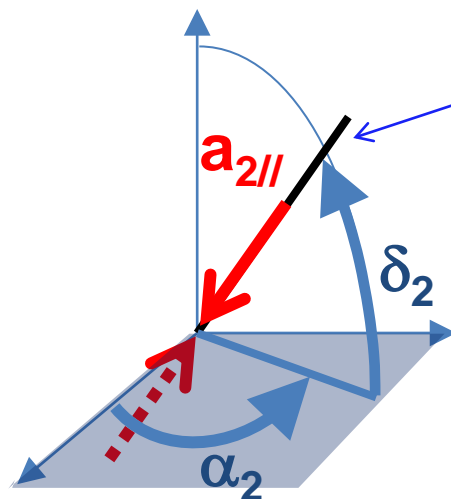
Significance map



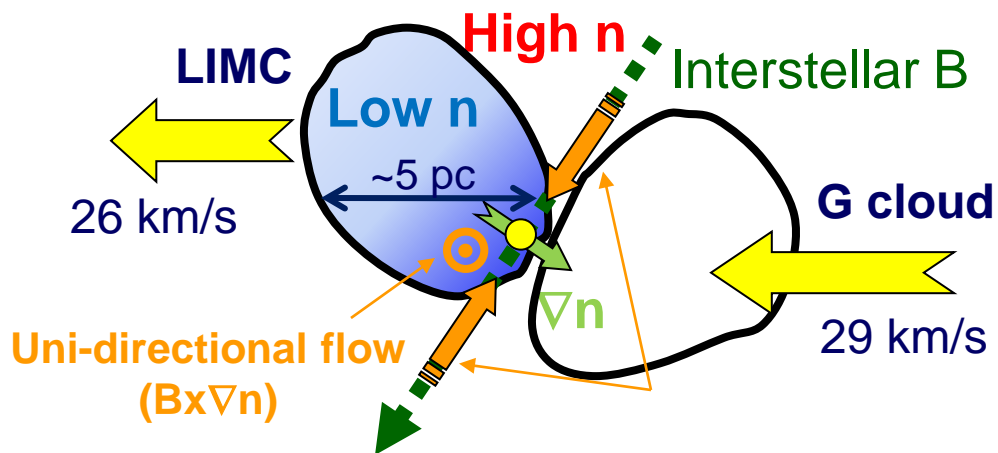
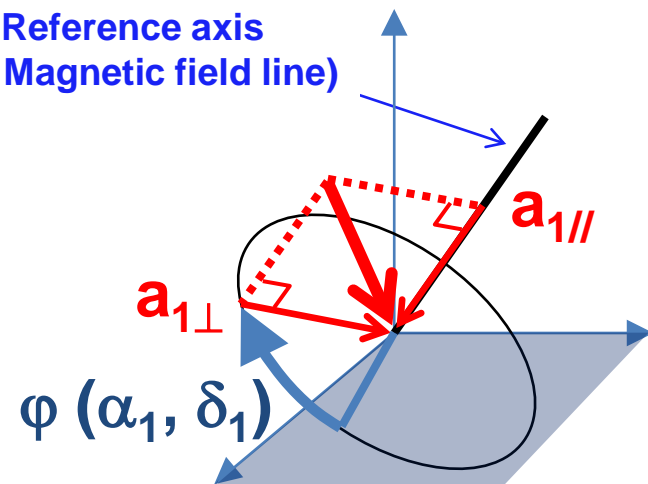
Global Anisotropy model (GA)

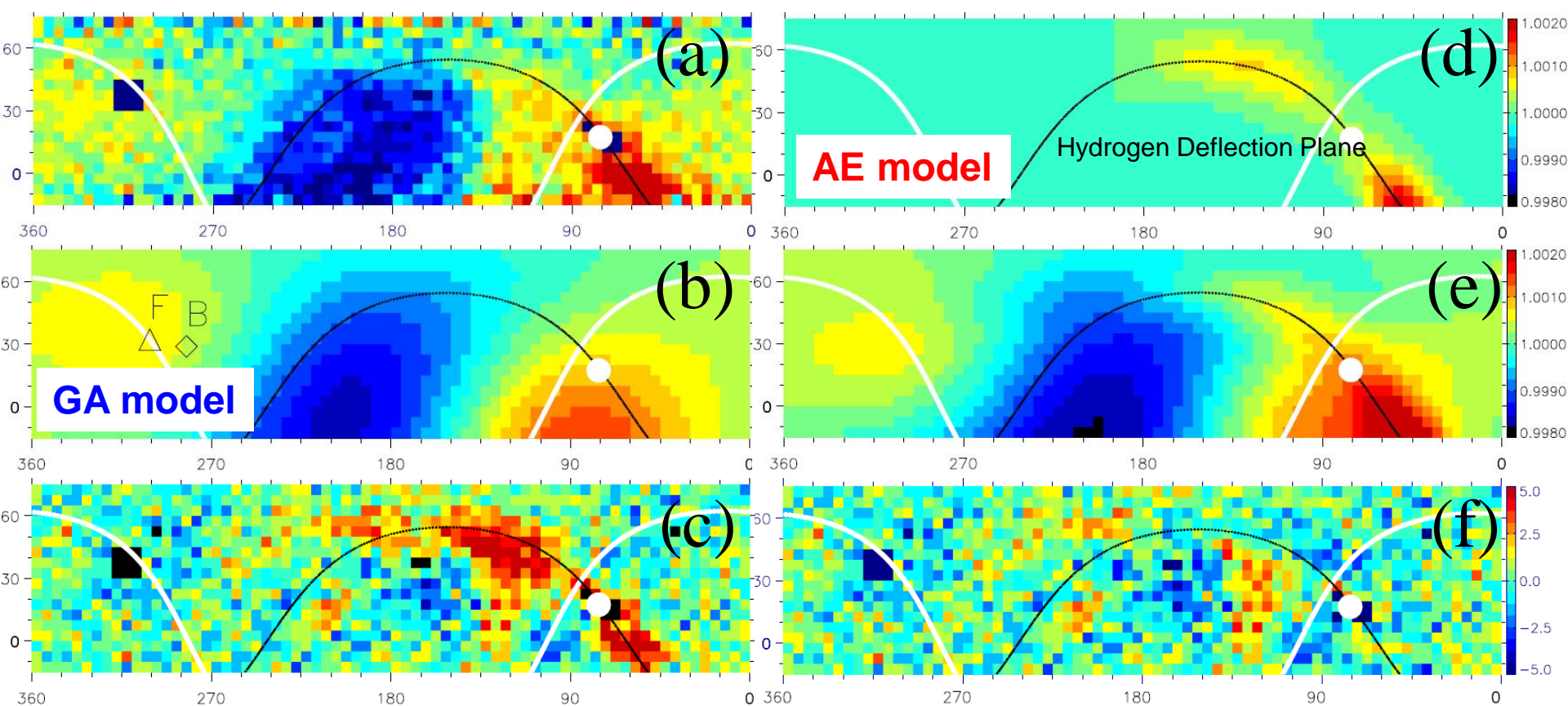


Bi-Directional Flow (BDF)



Uni-Directional Flow (UDF)





Best-fit parameters

GA:

$a_{1//}$ (%)	$a_{1\perp}$ (%)	$a_{2//}$ (%)	$\alpha_{1\perp}$ (°)	$\delta_{1\perp}$ (°)	$\alpha_{2//}$ (°)	$\delta_{2//}$ (°)
0.006	0.141	0.140	37.5	37.5	102.5 (120.9)	-28.9 (-32.2)

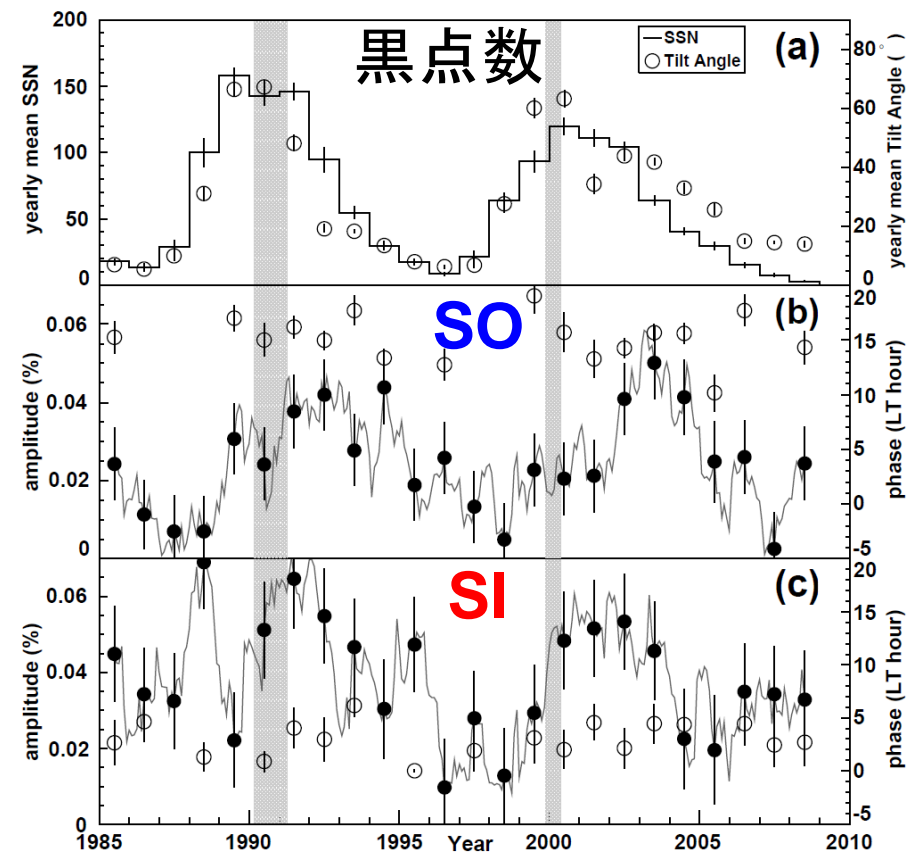
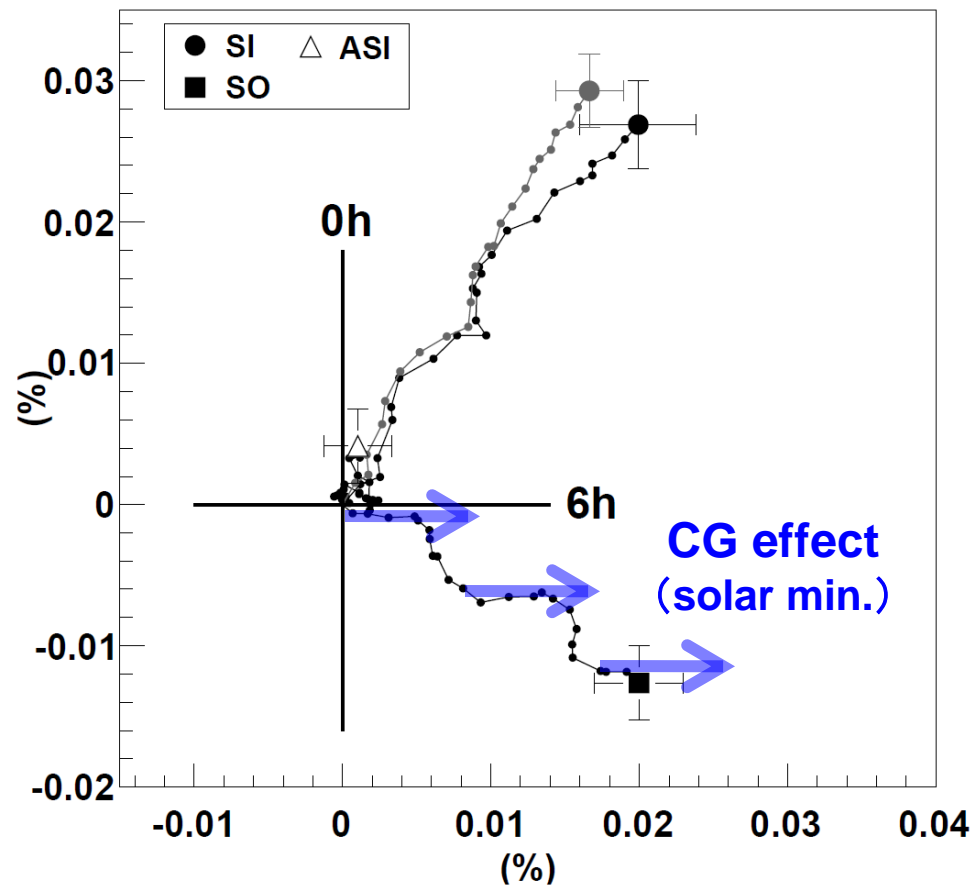
(Frisch's B)

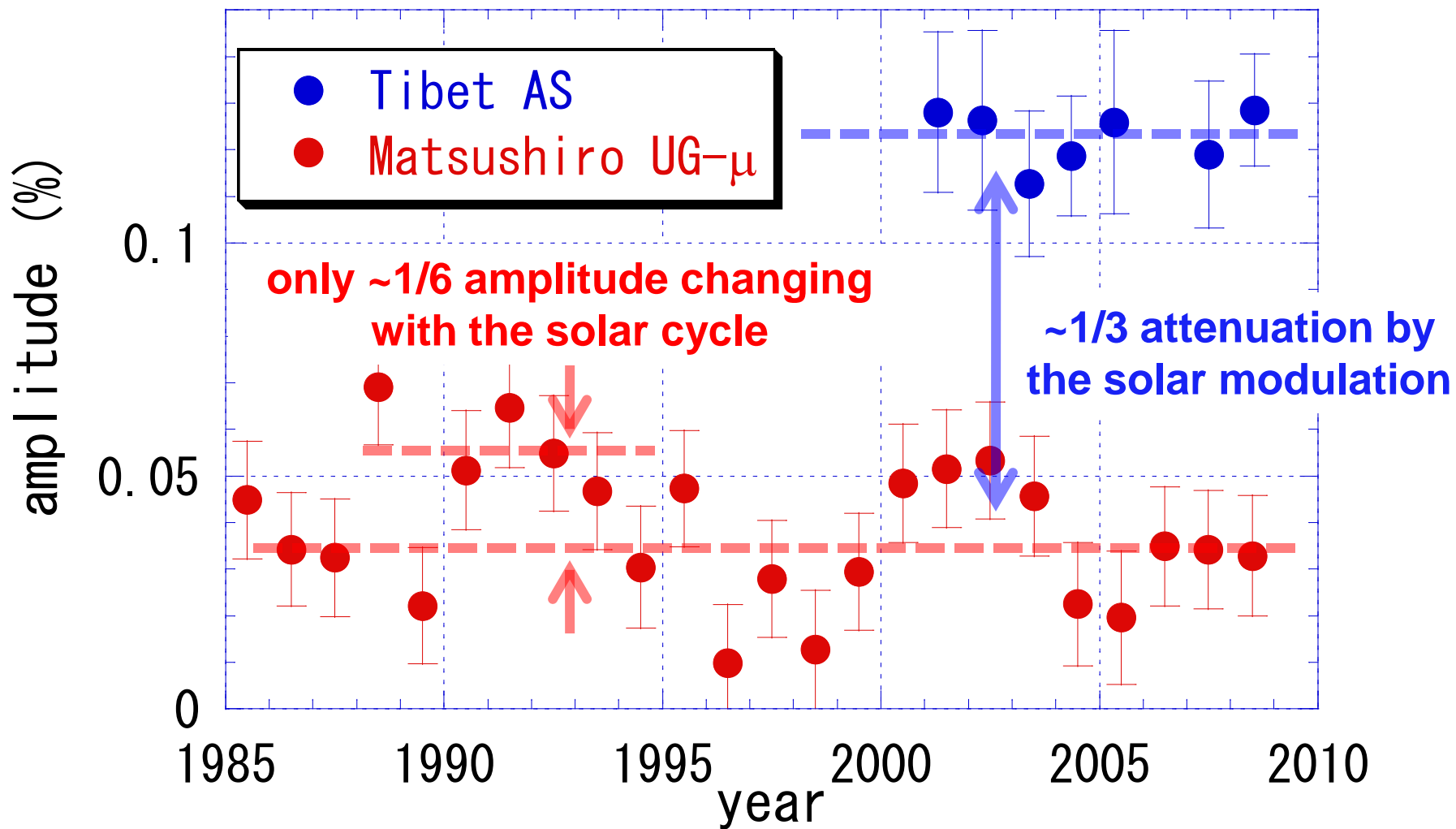
AE:

b_1 (%)	b_2 (%)	$\sigma_{//}$ (°)	σ_{\perp} (°)	Φ (°)
0.234	0.100	25.0	10.0	52.5

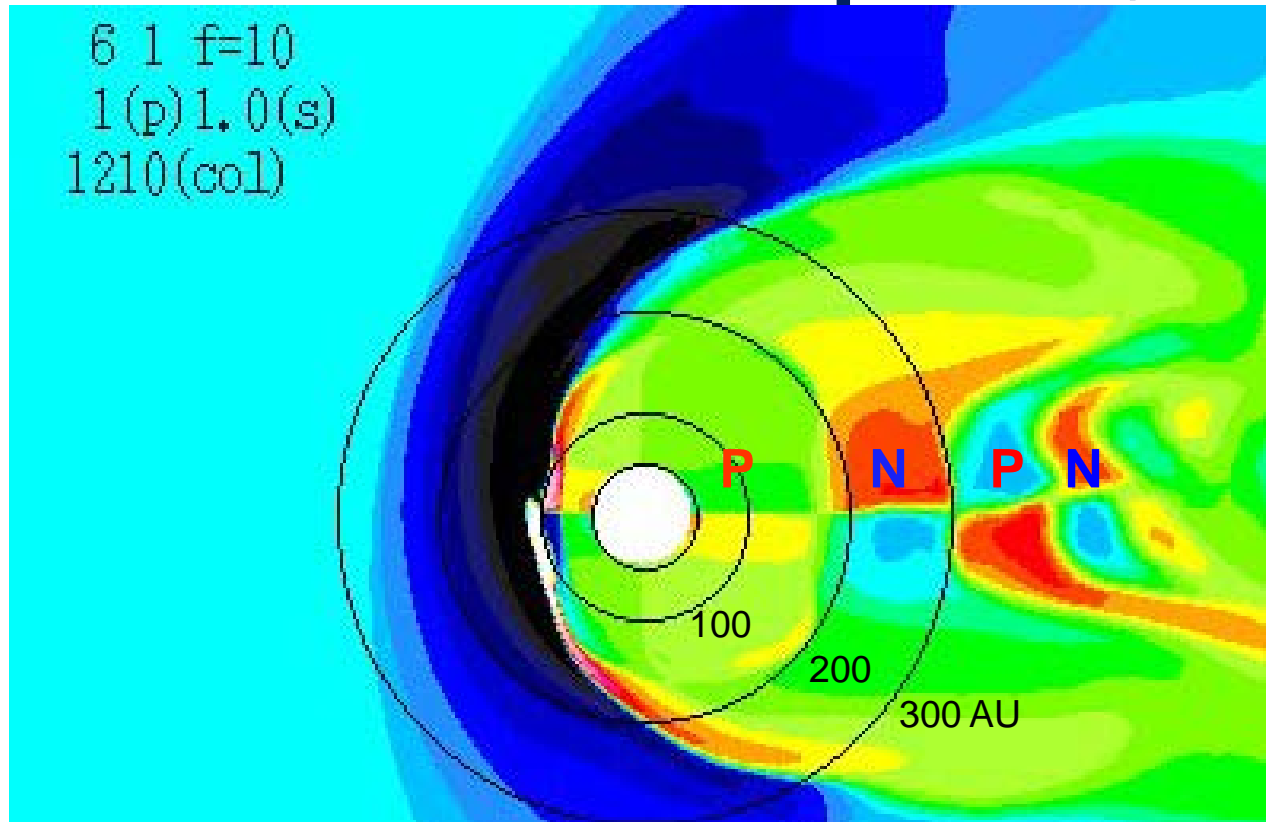
Solar cycle dependence of 0.6 TeV GCR anisotropy

(by Matsushiro UG- μ detector in 1985-2008)

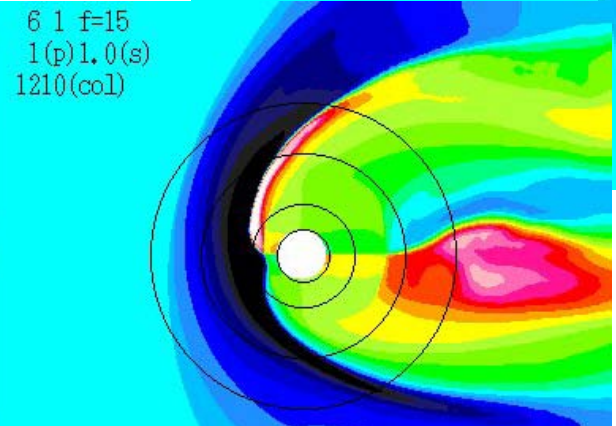
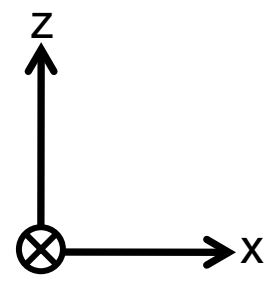




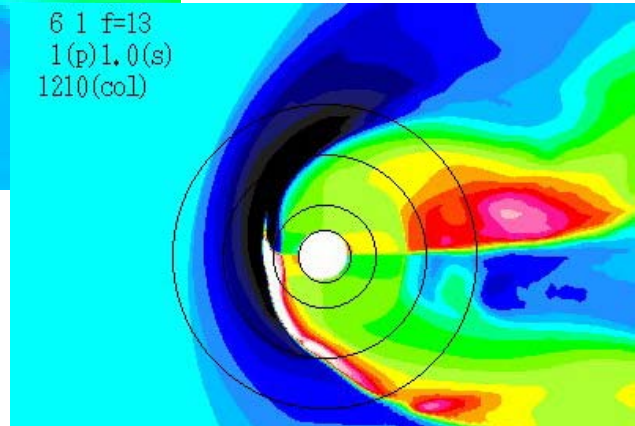
MHD model heliosphere (Washimi & Tanaka)



By



Positive



Negative

Summary

- GA+AE model developed for the 2D-map observed by Tibet III AS-array suggests...
 - GCR propagation in the **local structure** needed for a large density gradient.
 - **LISMF almost lays in the galactic plane.**
 - **Modulation in the heliotail (AE along HDP).**
- The average amplitude @0.6 TeV is roughly one third of the amplitude in multi-TeV region.
- **Only one fourth of the total attenuation varies in a correlation with the solar activity cycle and/or the solar magnetic cycle.**