

加速器データを用いたニュートリノ相互作用シミュレーションの研究

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申請額 旅費 20万円 交付額 5万円

- ニュートリノ相互作用シミュレーション
 - QE interaction
 - $\Delta \rightarrow N\gamma$ decay
 - (atmospheric neutrino flux)
- SciBooNE 実験

ν 相互作用シミュレーションの研究

実験データ

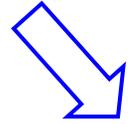
e 散乱等

ν 相互作用事象

理論

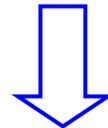
ν 相互作用モデル

π 、K等の核内相互作用



ν 相互作用の理解

ν interaction simulation (NEUT)の改良
核内相互作用simulationの改良



ニュートリノ振動

大気ニュートリノ

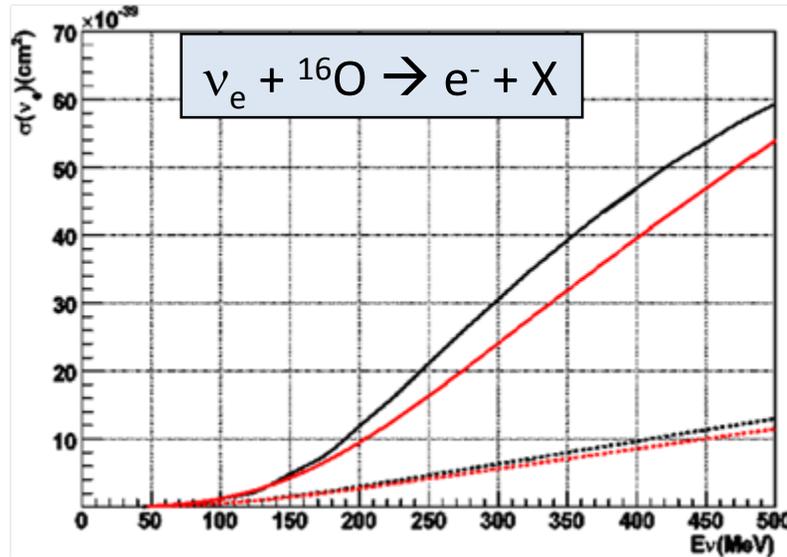
長基線ニュートリノ実験

核子崩壊の探索

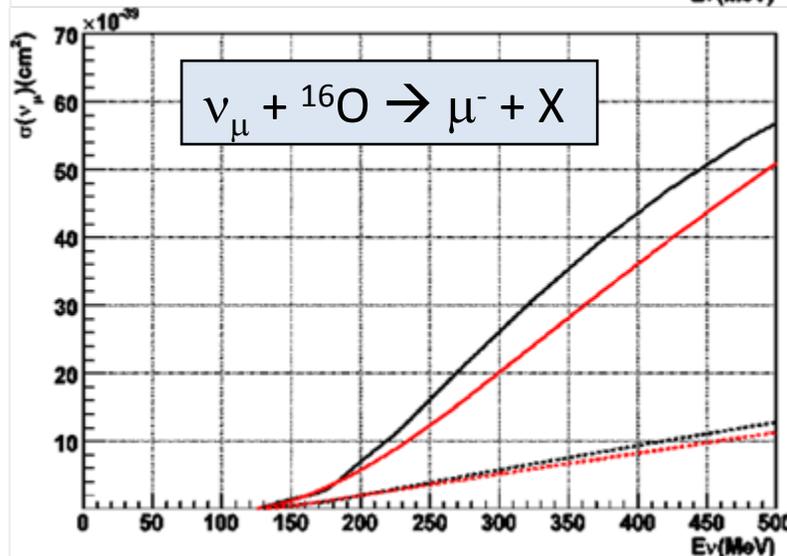
Quasi-elastic interaction

- Smith-Moniz model
 - nuclear effects are too simple
- New model (J.Nieves et al)
 - Random phase approximation (RPA) and Final State interaction (FSI) are applied. (good agreement with electron scattering data)
- This work is done by Mitsuka and M.Varverde (Univ. of Granada, Spain)

Quasi-elastic interaction



NEUT(Smith-Moniz)
Nieves et al.
solid: ν , dashed: $\bar{\nu}$



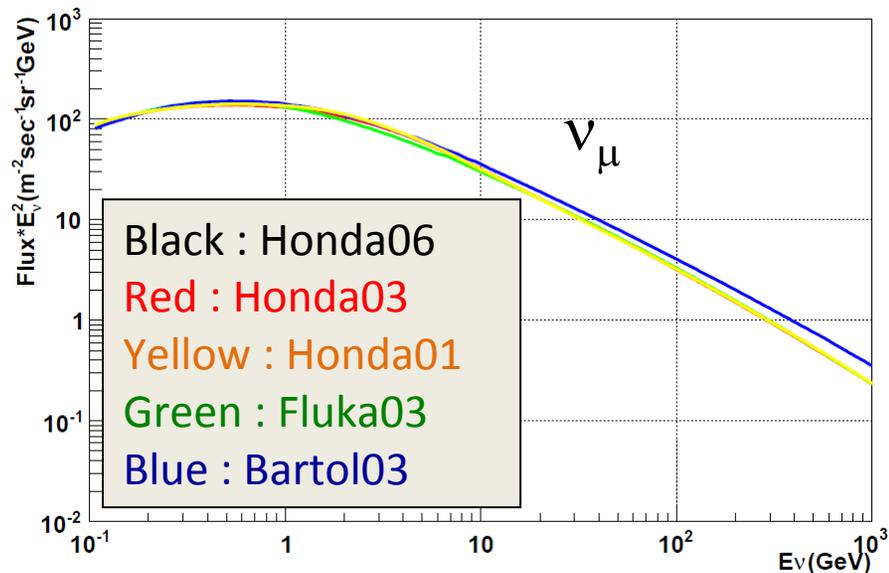
systematic error の評価に使用

$\Delta \rightarrow N\gamma$ decay

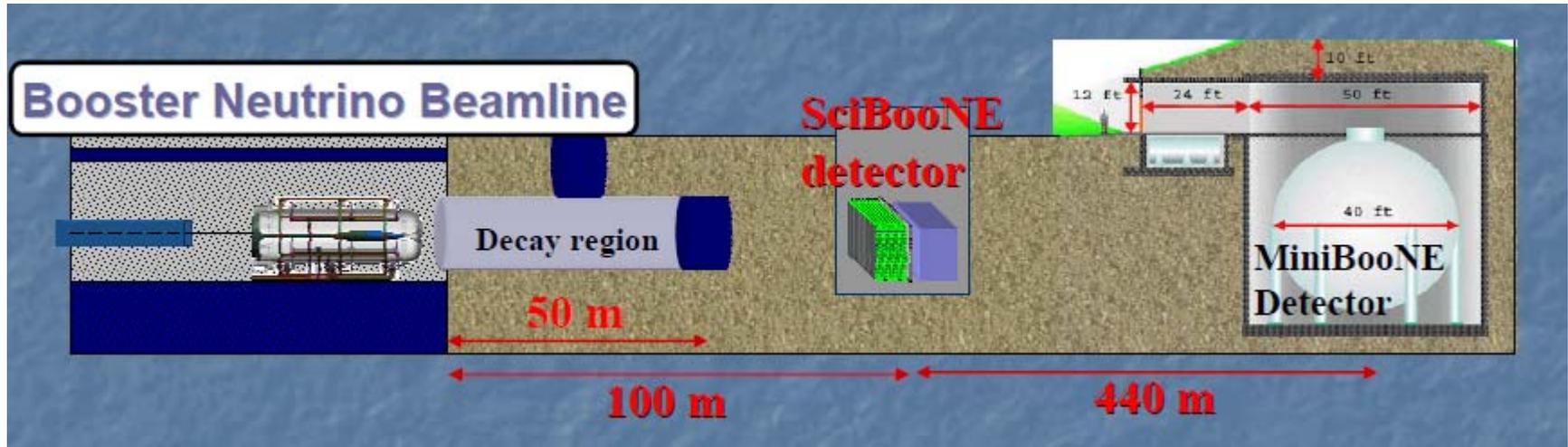
- $\text{Br} \sim < 0.5\%$
- But **NC $\Delta \rightarrow N\gamma$ can be background events in $\nu_\mu \rightarrow \nu_e$ appearance search**
- Following modes are considered
 - CC : $\nu + n \rightarrow l + p + \gamma$
 - NC : $\nu + n \rightarrow \nu + n + \gamma$
 - NC : $\nu + p \rightarrow \nu + p + \gamma$

atmospheric neutrino flux update

- Honda flux for atmospheric ν is updated (**Honda06**)
 - It take the most reasonable primary flux model
 - Hadron interaction model is modified
It reproduces the observed μ fluxes
 - Calculation for solar modulation is updated with the BESS observation



SciBooNE実験

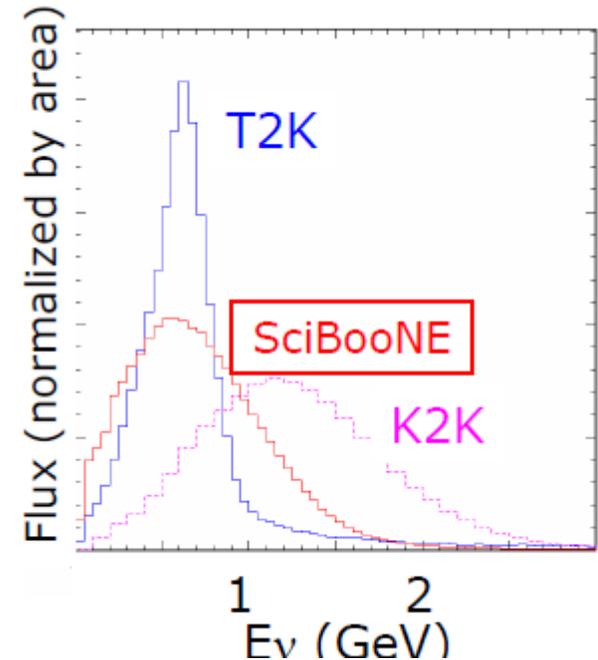


SciBooNE (FNAL E954)

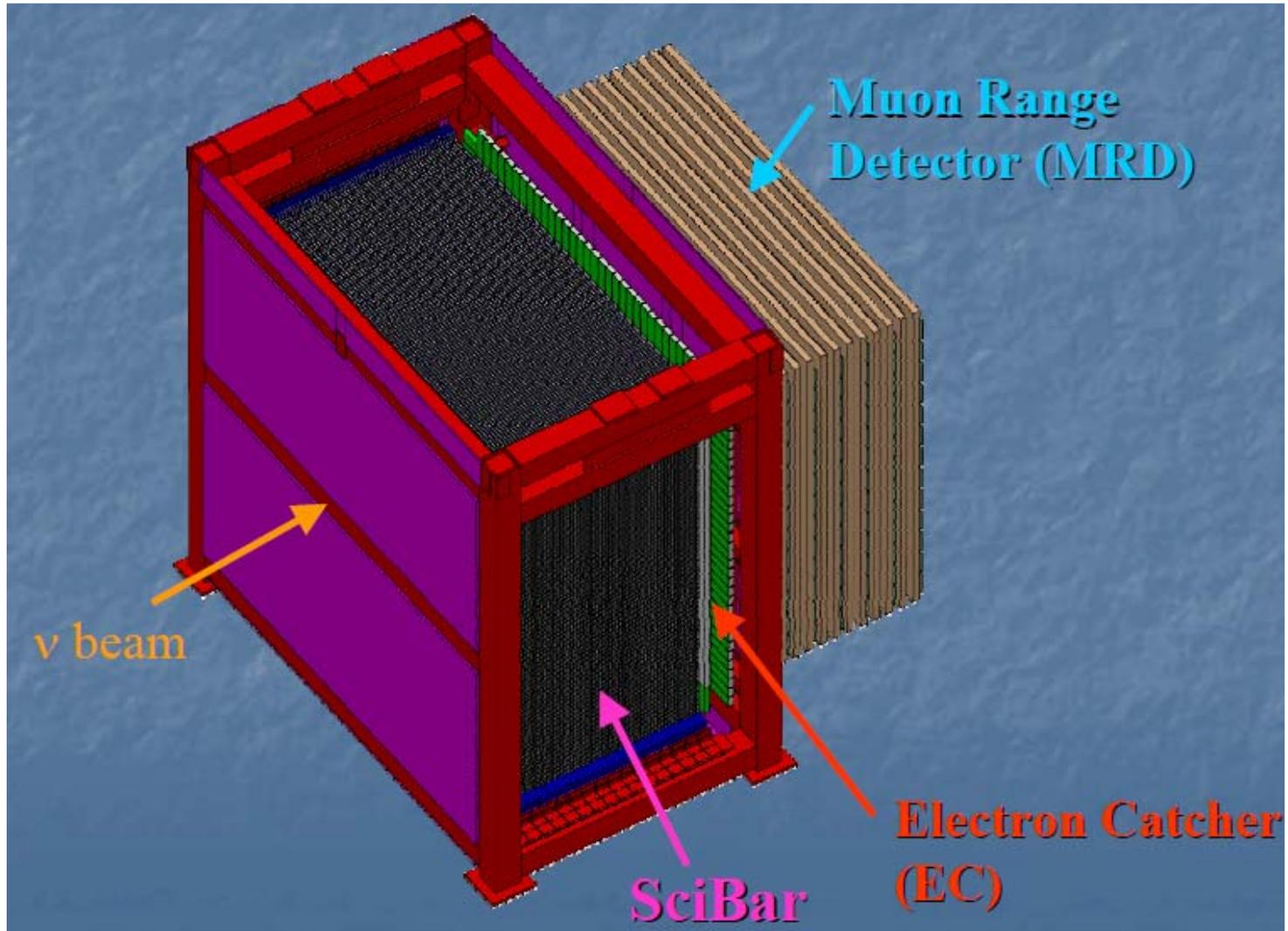
- Precise study of neutrino-nucleus cross section around 1 GeV
- Combination of Booster Neutrino Beam at FNAL and SciBar Detector used in K2K

Physics motivation

- Precision study of neutrino cross section for T2K
 - Anti-neutrino run
 - MiniBooNE near detector
-
- 1×10^{20} POT for neutrino
 - 1×10^{20} POT for anti-neutrino



SciBooNE 検出器



SciBooNE

- Sep/2006 Ground breaking
- Apr/2007 Detector installation
- May/2007 Commissioning
- Jun/2007 anti-neutrino data run
- Oct/2007 neutrino data run

Summary

- neutrino interaction simulation is updated
 - QE interaction model (systematic study)
 - $\Delta \rightarrow N\gamma$
 - atmospheric neutrino flux
 - etc
- SciBooNE experiment
 - neutrino/anti-neutrino beam around 1 GeV
 - precise information will be obtained
- T2K will start from Apr/2009

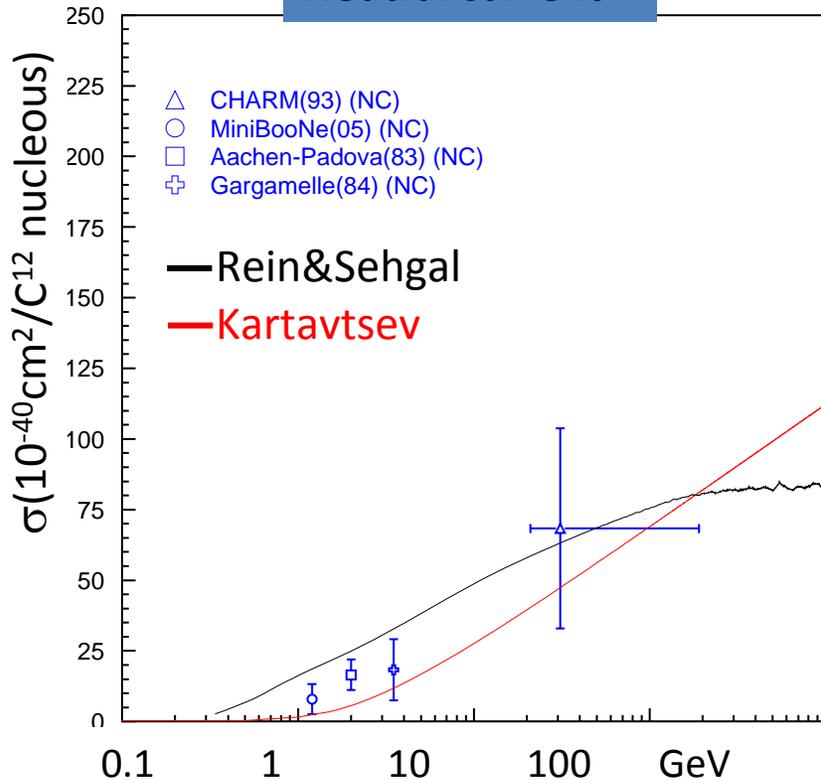
backup

Coherent pion production

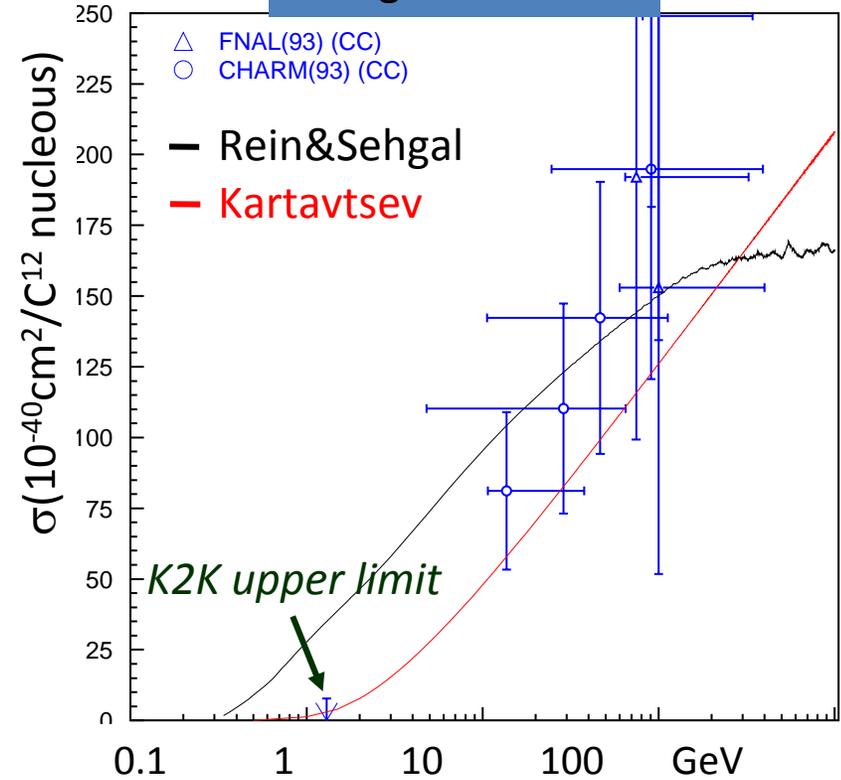
Comparison of Rein & Sehgal and Kartavtsev, Paschos & Gounars model

by Mitsuka, Hayato

Neutral current



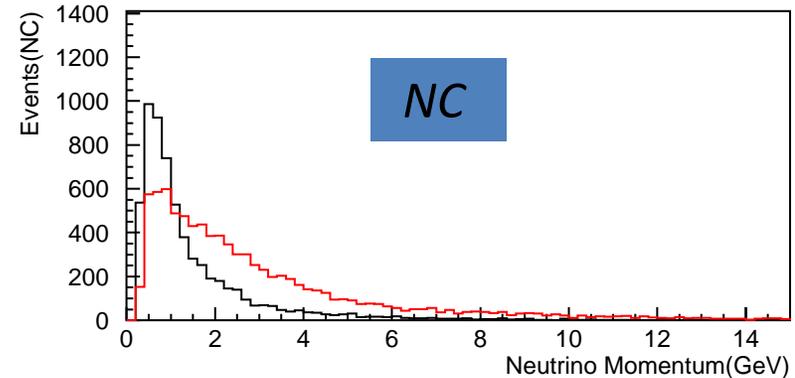
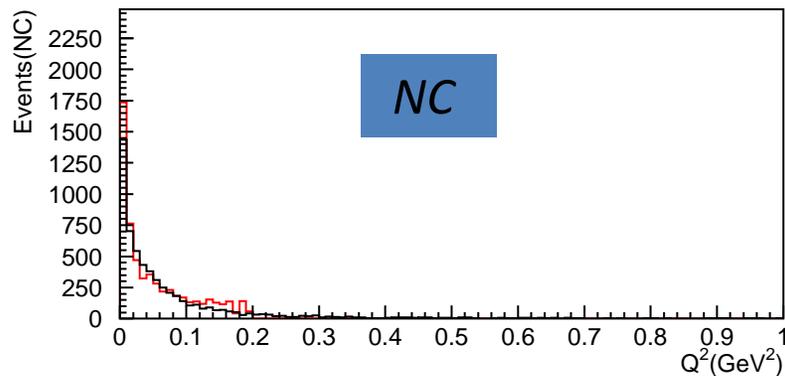
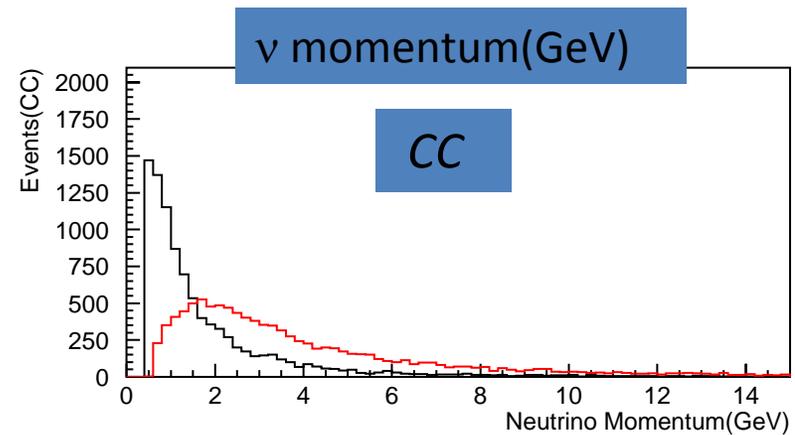
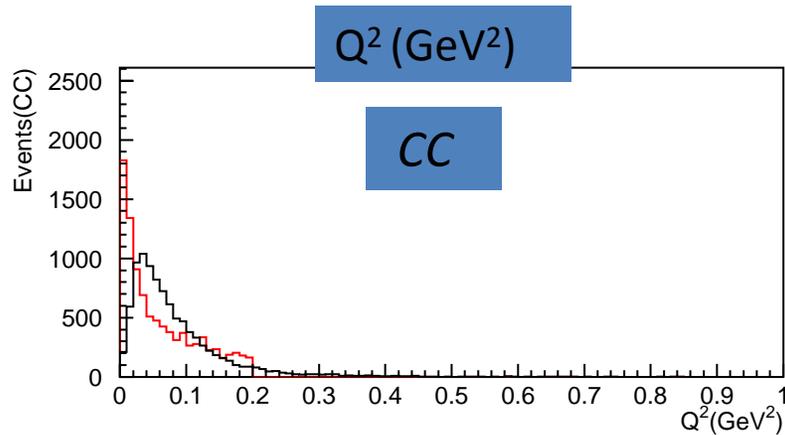
Charged current



- Both Rein&Sehgal and Kartavtsev model agree experimental data in view of the large experimental errors, except for K2K result in CC
- *Kartavtsev model agrees K2K result*

Coherent pion production

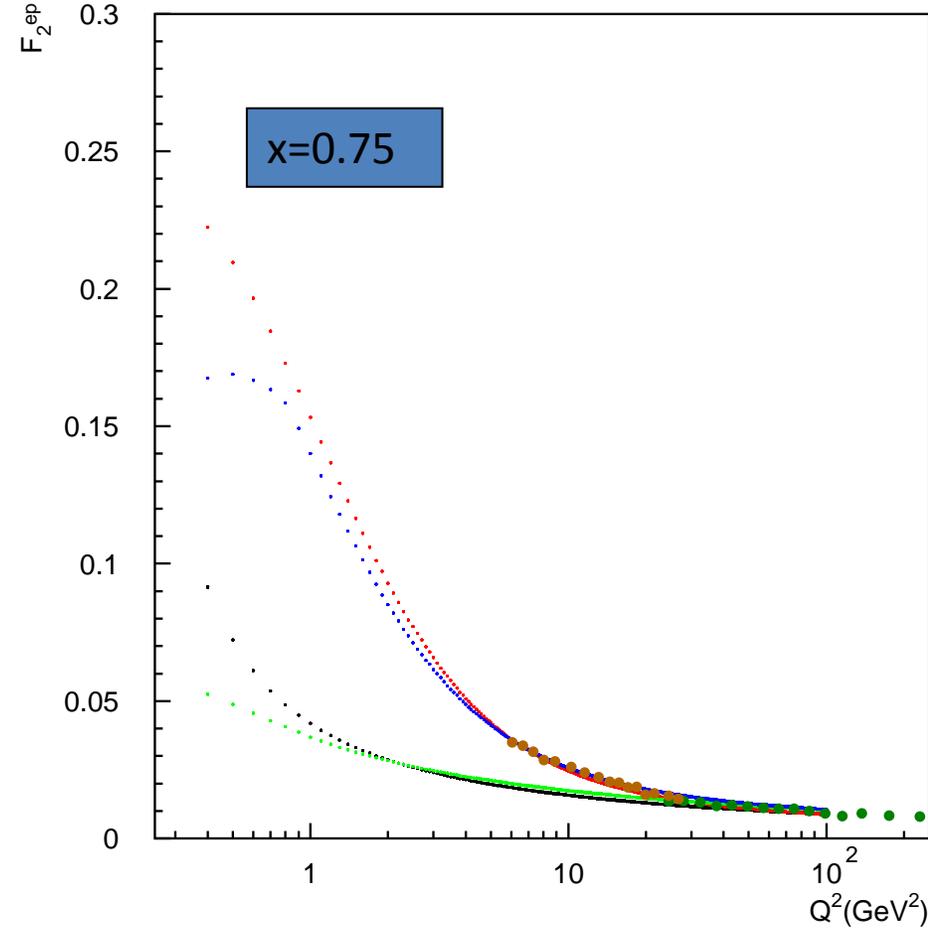
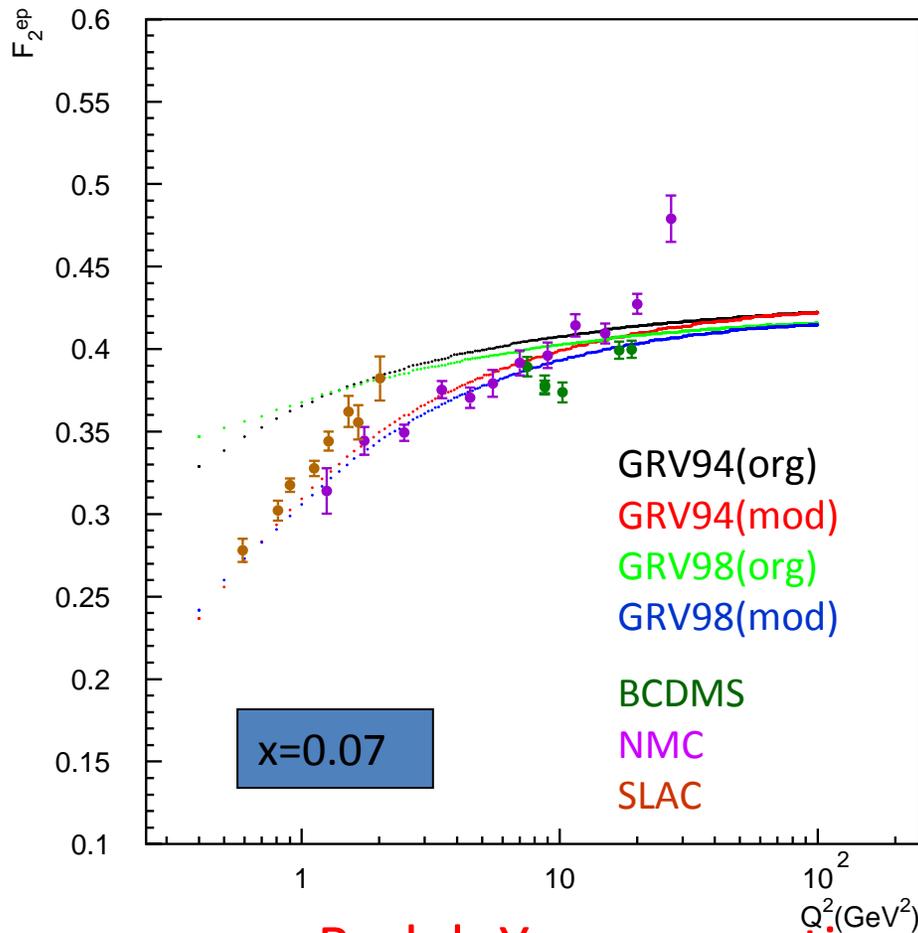
Black : Rein&Sehgal, Red : Kartavtsev (normalized by number of events)



We need more data and study of the effect for neutrino oscillation study.

Deep inelastic scattering

Comparison in $F_2(e^- \rightarrow p)$

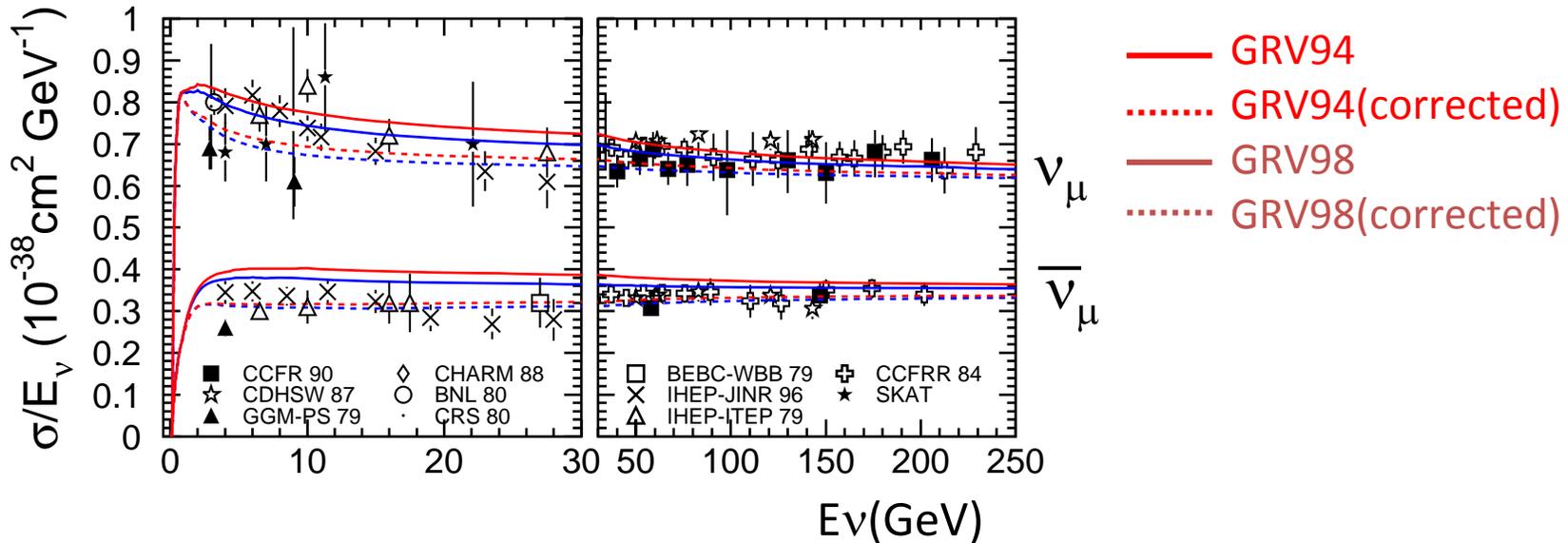


Bodek-Yang correction

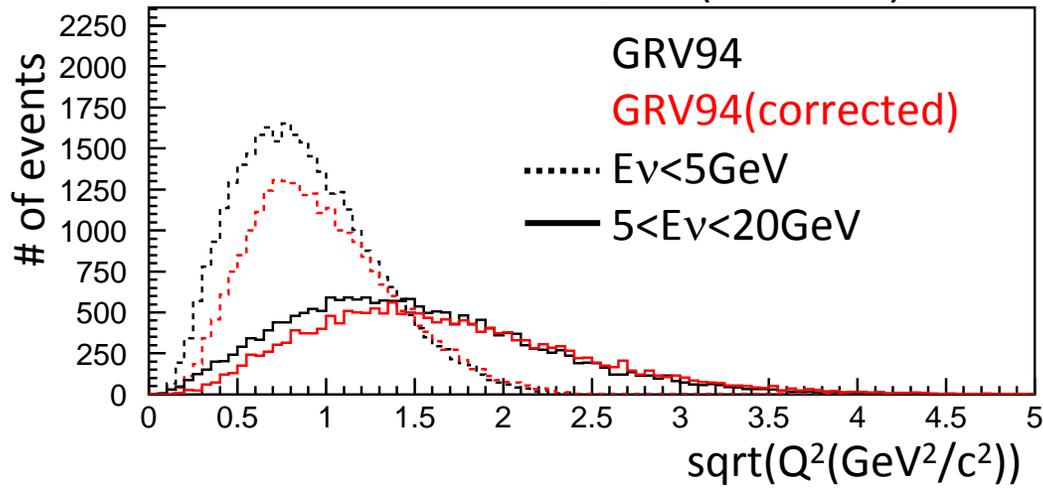
Corrected PDFs show good agreement with experimental data in both x region

Deep inelastic scattering

CC muon neutrino cross section



CC muon neutrino Q^2 dist(Honda03)



Deep inelastic scattering

Summary of Bodek-Yang correction

1. Bjorken scaling $x \rightarrow x_w$

$$x_w = x \frac{Q^2 + B}{Q^2 + Ax}$$

*A : target mass effect
higher twist effect
B : nonzero F_2 as $Q^2 \rightarrow 0$*

2. Structure function F_2

$$F_2(x) \rightarrow \frac{Q^2}{Q^2 + C} F_2(x_w)$$

*In order to fit both intermediate-x
and low-x*

3. d/u ratio

$$d_v \rightarrow d'_v(d_v, u_v)$$

$$u_v \rightarrow u'_v(d_v, u_v)$$

Correction for conversion $F_2^d \rightarrow F_2^n$

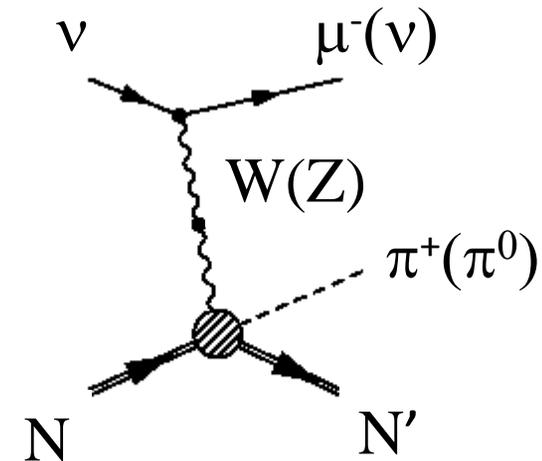
4. Longitudinal R

$$2xF_1 = F_2 \frac{1 + 4Mx^2/Q^2}{1 + R}$$

Correction for spin of target particle

Coherent pion production

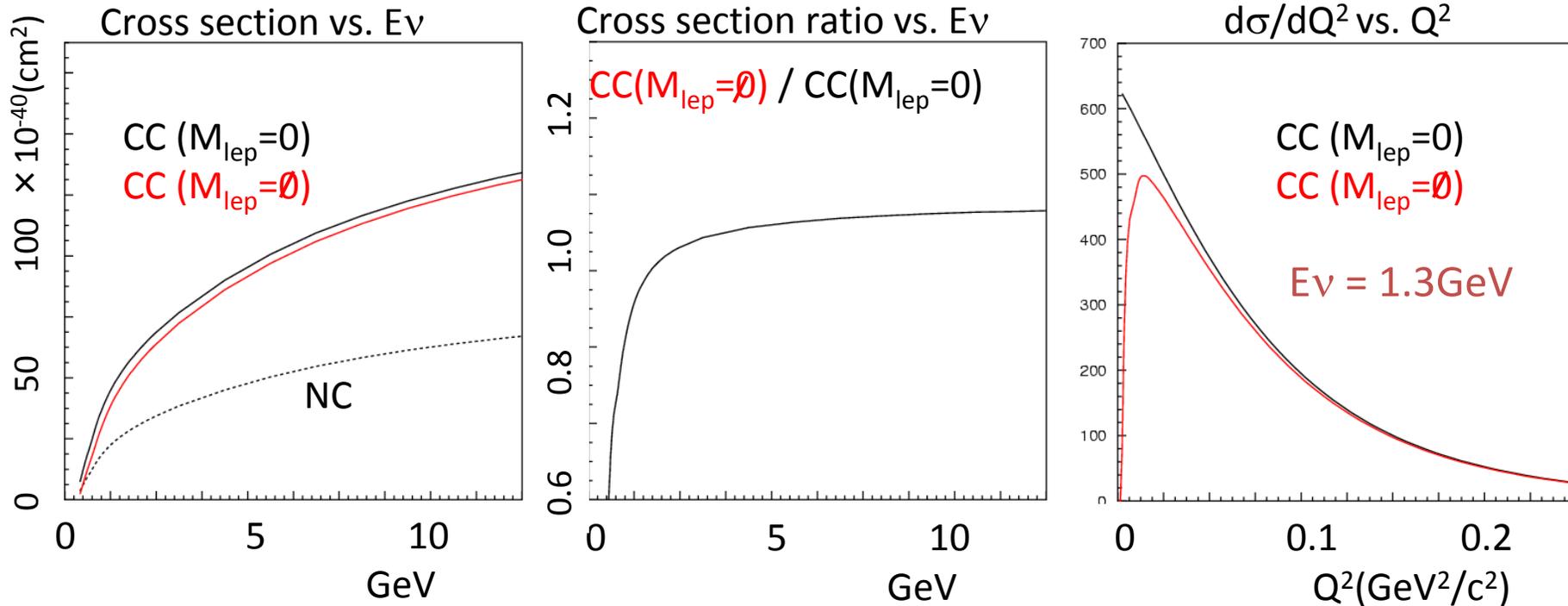
- K2K-SciBar group reports that CC coherent pion production is not observed at about 1GeV
- In CC, $\nu + {}^{16}\text{O} \rightarrow l^{(-)} + {}^{16}\text{O} + \pi^{+(-)}$ produces a charged lepton, **effect of lepton mass on cross section is not negligible at lower energy**



$$\frac{d\sigma^{\pi^+}_{\text{CC}}}{dx dy} = 2 \frac{d\sigma^{\pi^0}_{\text{NC}}}{dx dy} \times \left\{ \left(1 - \frac{1}{2} \frac{Q_{\text{min}}^2}{Q^2 + m_{\pi}^2} \right)^2 + \frac{1}{4} y \frac{Q_{\text{min}}^2 (Q^2 - Q_{\text{min}}^2)}{(Q^2 + m_{\pi}^2)^2} \right\}$$

x, y : Bjorken scaling where $Q_{\text{min}}^2 \propto m_{\text{lepton}}^2$, y $Q^2 \propto x, y$

Coherent pion production



- Lepton mass correction reduces cross section by 15% at 1GeV, 5% at 3GeV
- Effect is appeared in small Q^2 region ($< \sim 0.1 \text{GeV}^2/c^2$)
- Difference between Rein&Sehgal model ($M_{\text{lep}}=0$) and K2K results becomes smaller, but not compatible

Coherent pion production

- **Kartavtsev, Paschos & Gounaris model is tried**
(Phys. Rev. D **74**, 054007 (2006)) as one option
- Kartavtsev model describes kinematics and cross section **taking into account lepton mass**

Summary of Kartavtsev model

- Based on the partial conservation of axial current(PCAC), same as Rein & Sehgal
- Adler relation(formula for weak current cross section) is not used
- need π -nucleus scattering experimental data