

Observational Cosmology Group

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Outline

- 1) Science goals
- 2) Instruments and status
- 3) Scientific results
- 4) Size of Group, manpower, and budget
- 5) Relation with other universities/community
- 6) Education and mentoring young scientists
- 7) Future prospect

Goals and Projects

- Understanding the origin and evolution of the universe including
 - Energy density distribution,
 - Expansion history,
 - Structure/Galaxy formation, and
 - Cosmic reionization

Instruments and Status

- Projects

1. Sloan Digital Sky Survey (**SDSS**)

- Instrumentation until ~2000.
- Survey observations completed in 2008.
- Analysis and paper preparation up to 2012

2. Subaru Hyper Suprime-Cam (**HSC**)

- +complementary/preparatory observations w HST/ALMA
- ICRR started HSC-NB since 2010 (arrival of M. Ouchi)
- HSC instrument first-light (2012)
- HSC survey planned to start from FY 2013

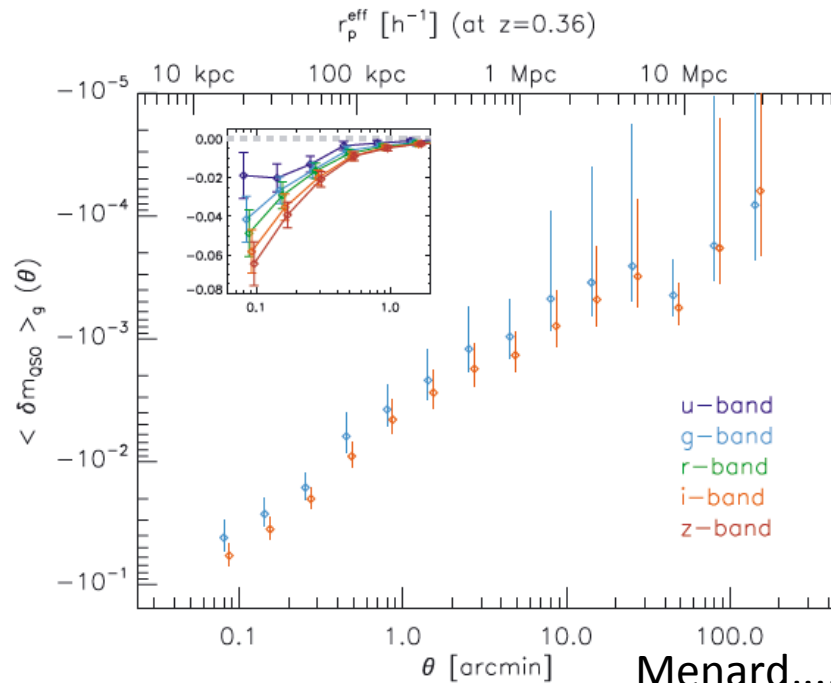
Scientific Results (Refereed Papers)

- The project papers
 - A total of 66 refereed papers
 - A total of 8622 citations
- ICRR-led papers
 - SDSS 8 papers, 234 citations (2006-2012)
 - HSC 4 papers, 237 citations (2010-2012)
 - SDSS: Fukugita et al. 2007, “A Catalog of Morphologically Classified Galaxies from the Sloan Digital Sky Survey: North Equatorial Region” (64 citations)
 - HSC: Ouchi et al. 2010, “Statistics of 207 Ly α Emitters at a Redshift Near 7: Constraints on Reionization and Galaxy Formation Models” (114 citations)

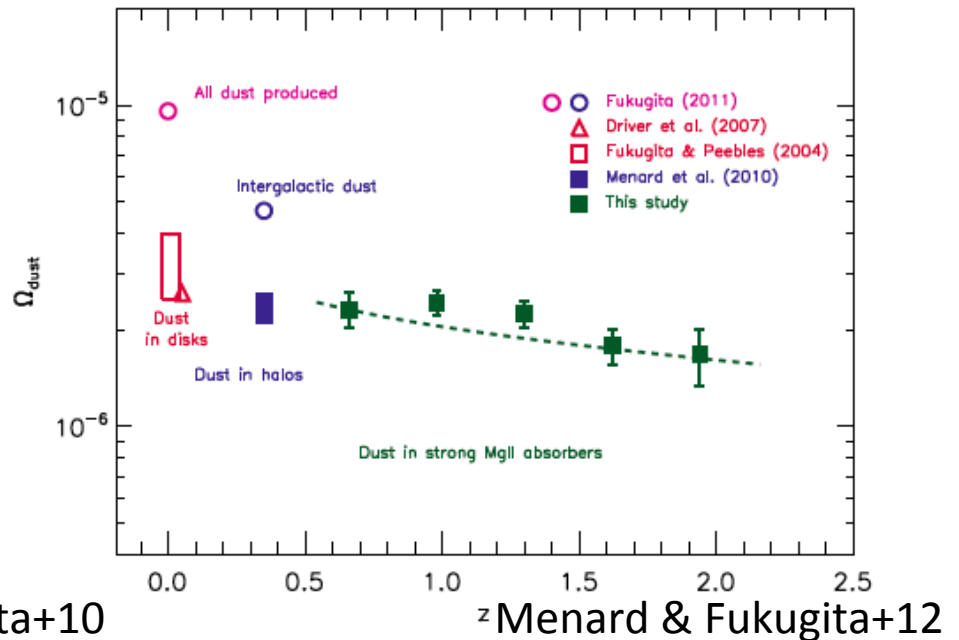
Citation counts as of Jan 2013

Highlights of Scientific Results (1)

First Identification of 'Cosmic' Dust



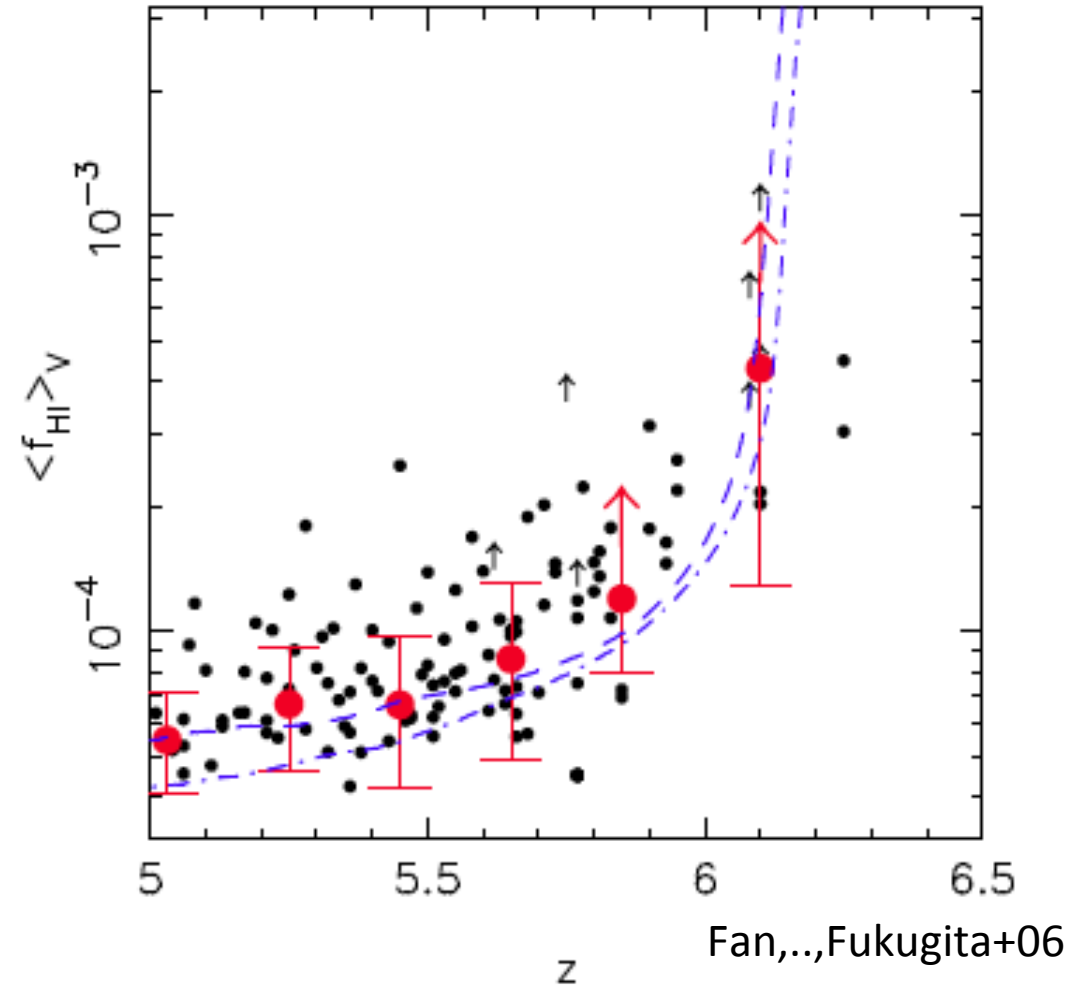
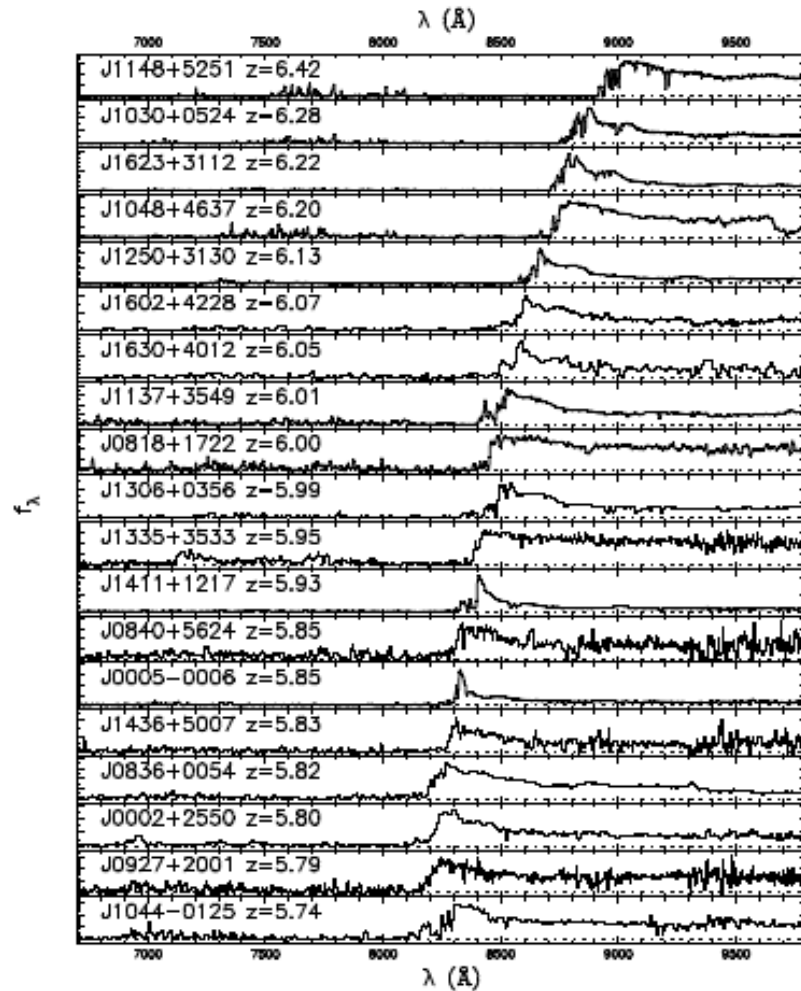
Menard,..,Fukugita+10



- Missing baryon problem: all the stars, gas, and dust within galaxies constitute at most 40% of the baryons predicted by the Big Bang. Where are the rest? → Baryon energy density dist?
- Statistics of QSO brightness (m) as a function of galaxy angular distance (θ). Color dependence → Dust in halo+IGM (Menard,..,Fukugita+10)
- Energy density of dust: $\Omega_{\text{dust}} \sim 5 \times 10^{-6}$. About twice large as those in galactic disks.

Highlights of Scientific Results (2)

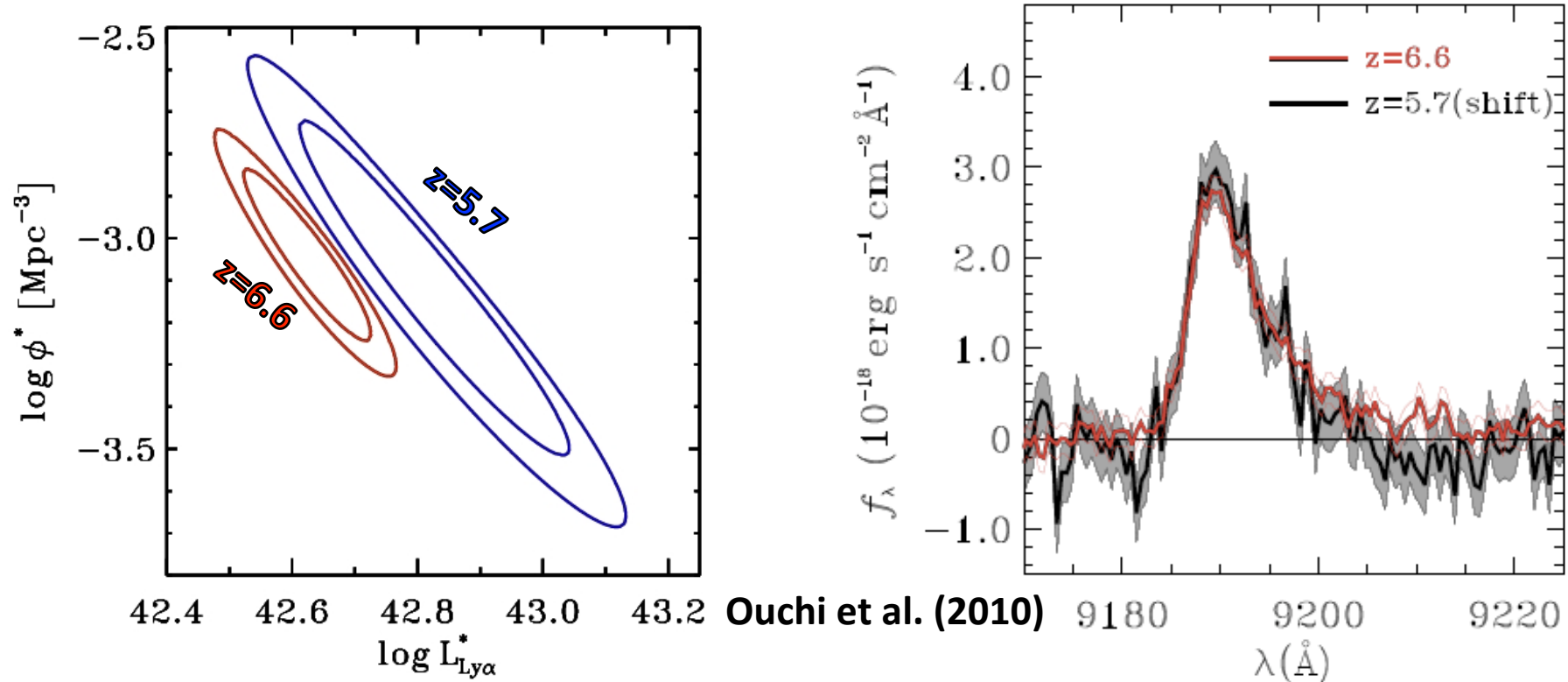
Cosmic Reionization Completion



- SDSS QSOs \rightarrow Ly-alpha optical depth (τ) of IGM (HI indicator)
- Rapid change of τ at $z \sim 6 \rightarrow$ completion of reionization

Highlights of Scientific Results (3)

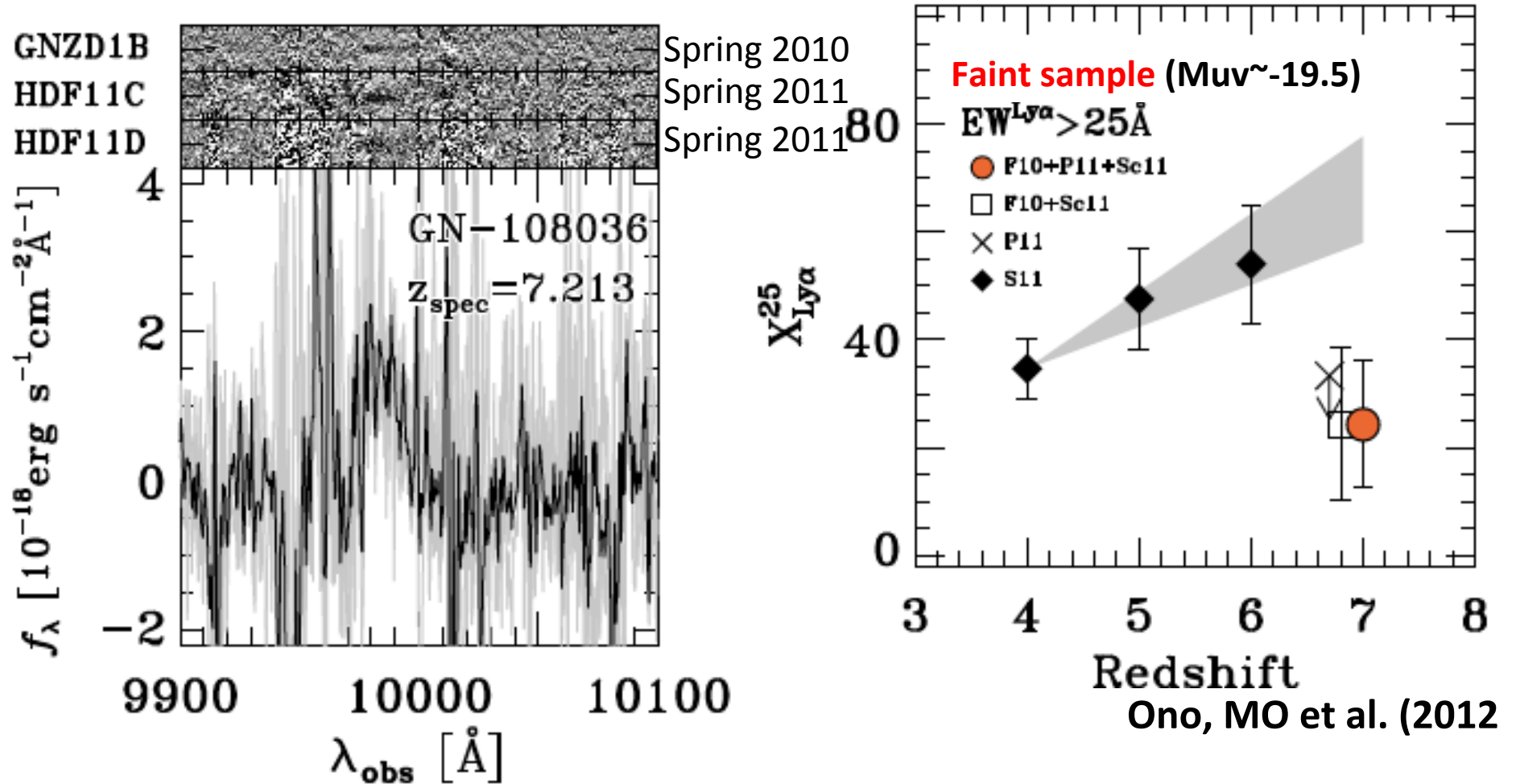
Cosmic Reionization Completion



- First precise measurement (incl. cosmic variance errors; cf. Malhotra&Rhoads+04, Kashikawa+06) of Ly α luminosity function at $z \sim 6$. Definitive identification of Ly α luminosity decrease (90% CL) from $z \sim 5.7$ to 6.6 that is likely made by IGM H I damping wing abs. (Ouchi et al. 2010). Moderately early cosmic reionization.
- $x_{\text{HI}} \sim 0.2 \pm 0.2$ at $z=6.6$. large statistical error \rightarrow next generation HSC survey with a size larger by x 30

Highlights of Scientific Results (4)

Cosmic Reionization Completion



- Identification of $z \sim 7$ galaxies including the most distant $z = 7.213$ galaxy.
- The fraction of Ly α emitting galaxies is low, $\sim 20\%$ \rightarrow Signature of reionization? The amplitude of drop is larger in faint galaxies than in bright galaxies \rightarrow **suggestive of inside-out reionization** (Ono, MO et al. 2011).

Size of Group, Manpower, and Budget

- Faculty
 - Masataka Fukugita, Professor, 1996 to 2012
 - Naoki Yasuda, Assoc. Professor, 2003 to 2009
 - Masami Ouchi, Assoc. Professor, 2010 to the present
 - Yoshiaki Ono, Research Associate, 2012 to the present
- Postdoctoral Fellows
 - Rieko Momose, 2012 to the present
 - Suraphong Yuma, 2012 to the present

Budget and Machine time

- SDSS
 - Grants-in-Aid for Scientific Research (A) for 2003-2006
- HSC

The basic design and tests for HSC narrow-band

 - Grants-in-Aid for Scientific Research (A) (Ouchi et al. 2011-2014) for
 - Grants-in-Aid for Scientific Research (start-up: Ono 2012-2014)
- Machine time (2010-2012)
 - Subaru: 14 nights (PI Ouchi)
 - Hubble Space Telescope:
 - Cycle 18: 18 orbits (PI Ouchi),
 - Cycle 19: 128 orbits (PI Ellis in collaboration)
 - Cycle 20: 16 orbits (PI Ono),
 - ALMA:
 - Cycle 0: 5 hours (PI Ouchi),
 - Cycle 1: 4+6 hours (PI Ouchi), 5 hours (PI Momose)
 - Keck: 5 nights (PI Ouchi) + VLT/Magellan etc.

Relation with other universities/community

- SDSS
 - Large international collaboration (US, Japan, German)
- Subaru Hyper Suprime-Cam (HSC) program
 - HSC survey consortium (Opt-NIR society in Japan, Princeton, Taiwanese institutes)
 - Narrow-band development w K. Shimasaku et al. from Tokyo, NAOJ, Ehime, Osaka-sangyo, Tsukuba etc.
 - ALMA Collaborations with NAOJ/Tokyo researchers (e.g. Nakanishi, K. Kohno). Accepting post-docs from ALMA community (e.g. R. Momose)

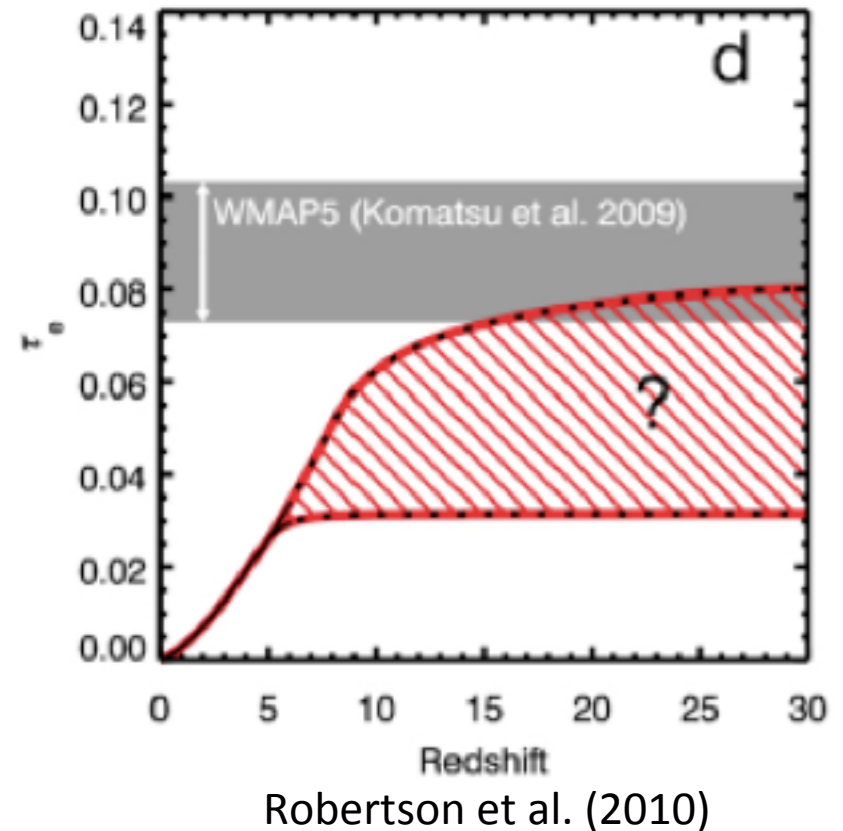
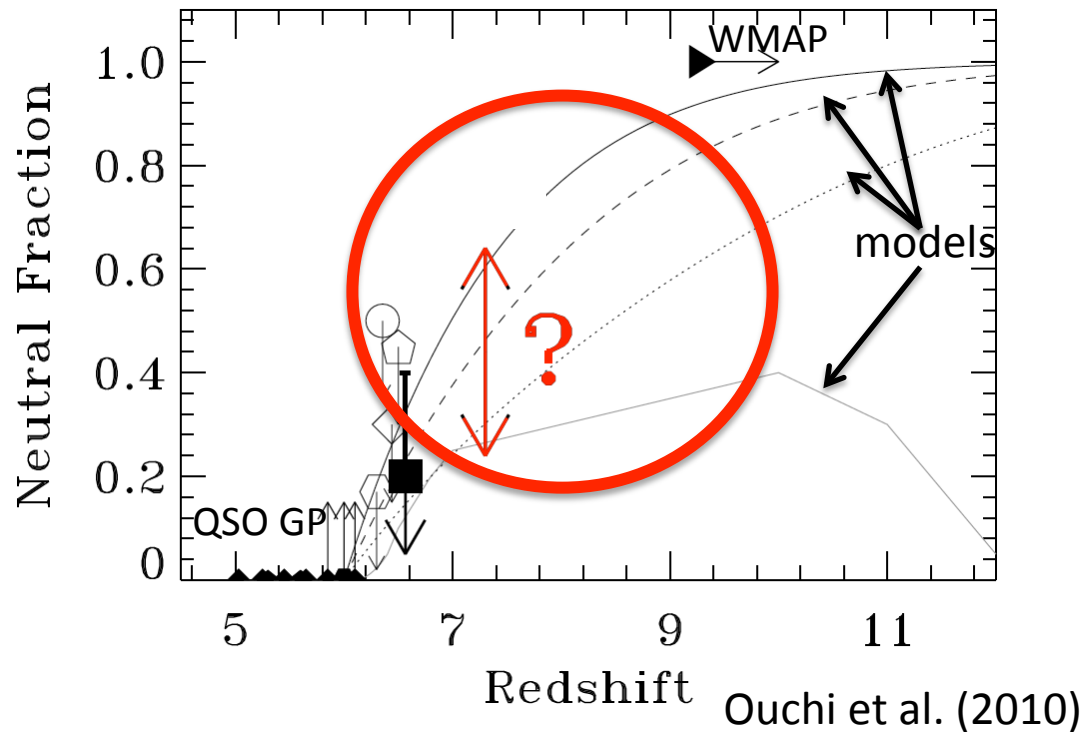
Education and mentoring young scientists

- A master and Ph.D degree for one student under N. Yasuda (-2009)
- No more students were accepted. No post-docs.
- Resuming education activities after the arrival of Ouchi
 - 2 students (2012) + 2 more students (2013)
from Physics/Astronomy departments
 - Spring school in 2012In ICRR, one of the most popular research group for which students apply.
- Resuming accepting post-docs
 - 2 post-docs (2012) + 1-2 post-docs (2013)
 - The 2 post-docs (2012) have submitted new papers in the past few months. One out of 2 won a JSPS fellowship.
- Young post-docs and students are gathering. (0→5 in 2012)
- Strong demands on this study subject from young people.

Future prospects

- HSC survey from FY 2013 (2013-2017)
- Synergistic observations with Hubble, ALMA etc.
- Preparation activities for 2020s programs
 - Subaru Prime Focus Spectrograph (PFS; ~2018-)
 - Thirty Meter Telescope; Infrared Imaging Spectrograph (IRIS; ~2020) TBA

Addressing Three Major Questions of Early Universe

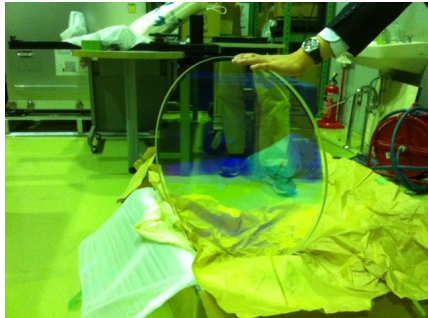


1. Uncovering the cosmic reionization history (determination of reionization scenario)
 2. Missing ionizing photons: Inconsistency of ionizing photon budget and τ
 3. First-generation galaxy formation strongly related to cosmic reionization
- interesting coincidence with the one of the priority sciences selected by US decadal survey Astro2010 (2012-2021)

Hyper Suprime-Cam Survey (2013-2017)

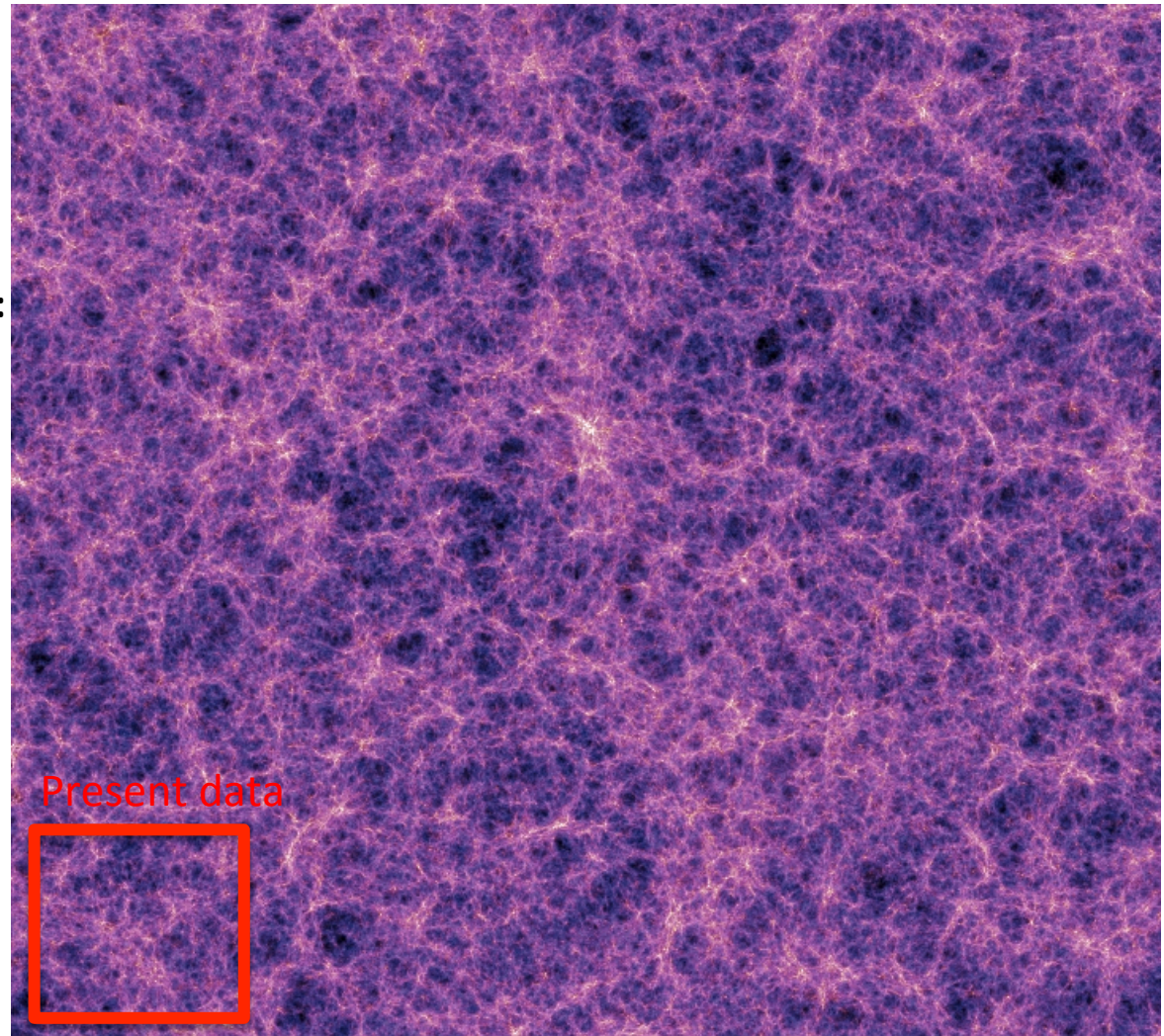


Hyper Suprime-Cam(HSC; 2012-) PFU
HSC NB filter(test piece: 2012 Mar.)



- Developing HSC narrow-band filters in ICRR
- HSC survey plans to be conducted in 2013-2017

Size of HSC survey for reionization ($z \sim 6-7$)



Summary

- **Science goals**
 - Origin/Evolution of the universe (Ω , galaxy formation, and reionization etc.)
- **Instruments and status**
 - SDSS (-2012). HSC+related studies (2010-)
- **Scientific results**
 - 66 refereed papers, 12 out of which are led by ICRR
- **Size of Group, manpower, and budget**
 - 2 faculty and 2 post-doc (2012-). JSPS grant, Subaru/HST/ALMA time
- **Relation with other universities/community**
 - Belonging to large international collaborations, SDSS and HSC survey.
- **Education and mentoring young scientists**
 - Education activities resumed in 2011→2 students, 2 PDs (more in 2013)
- **Future prospect**
 - Addressing 3 questions of early universe with HSC etc.