

STATUS OF THE MAGIC TELESCOPE

Eckart Lorenz (MPI-Munich), for the MAGIC collaboration

OVERVIEW

- * SOME BASIC MAGIC PARAMETERS
- * THE PHYSICS GOALS
- * TECHNICAL ELEMENTS
- * THE CURRENT STATUS OF THE CONSTRUCTION
- * SUMMARY

The MAGIC Collaboration (14 INSTITUTIONS, 74 PHYSICISTS)

Institut de Física d'Altes Energies, Barcelona:

Oscar Blanch, Juan Cortina, Eva Domingo, Enrique Fernández, Josef Flix, Markus Gaug, Javi López, Manel Martínez

Universitat Autònoma de Barcelona:

Carmen Baixeras, Carles Domingo, Lluís Font, Raul Orduna, Alejandro Sánchez, Andreu Torres

Crimean Astrophysical Observatory:

Arnold Stepanian

University of California, Davis:

Daniel Ferenc, Alvin Laille

Division of Experimental Physics, University of Lodz:

Maria Giller, Piotr Jacon, Dorota Sobczynska, Tadeusz Wibig

Universidad Complutense, Madrid:

Luis José Contreras, Victoria Fonseca, Marcos López, Emma Oña, Raquel Reyes

Max-Planck-Institut für Physik, München:

Rudolf K. Bock, José Antonio Coarasa, Markus Garzarczyk, Jurgen Gebauer, Florian Goebel, Eckart Lorenz, Keiichi Maze, Razmick Mirzoyan, David Paneque, Nadia Tonello, Vincenzo Vitale, Robert Wagner, Wolfgang Wittek

Dipartimento di Fisica, Università di Padova:

Laura Alciati, Denis Bastieri, Ciro Bigongiari, Nicola Galante, Mosè Mariotti, Abelardo Moralejo, Donatella Pascoli, Luigi Peruzzo, Antonio Saggion, Villi Scalzotto

Space Research Unit, Potchefstroom University:

Okkie C. de Jager

Fachbereich Physik, Universität-GH Siegen:

Thomas Hengstebeck, Nikolaj Pavel, Ralf Stiehler, Serguei Volkov

Dipartimento di Fisica, Università di Siena:

Mario Meucci, Riccardo Paoletti, R. Pegna, A. Piccioli, Antonio Stamera, Nicola Turini

Tuorla Observatory, Pikkio:

Aimo Sillanpää, Leo Takalo

Universität Würzburg:

Thomas Bretz, Eduardo Colombo, Tanja Kneiske, Karl Mannheim, Martin Merck

Yerevan Physics Institute, Cosmic Ray Division, Yerevan:

Ashot Chilingarian

THE MAGIC TELESCOPE

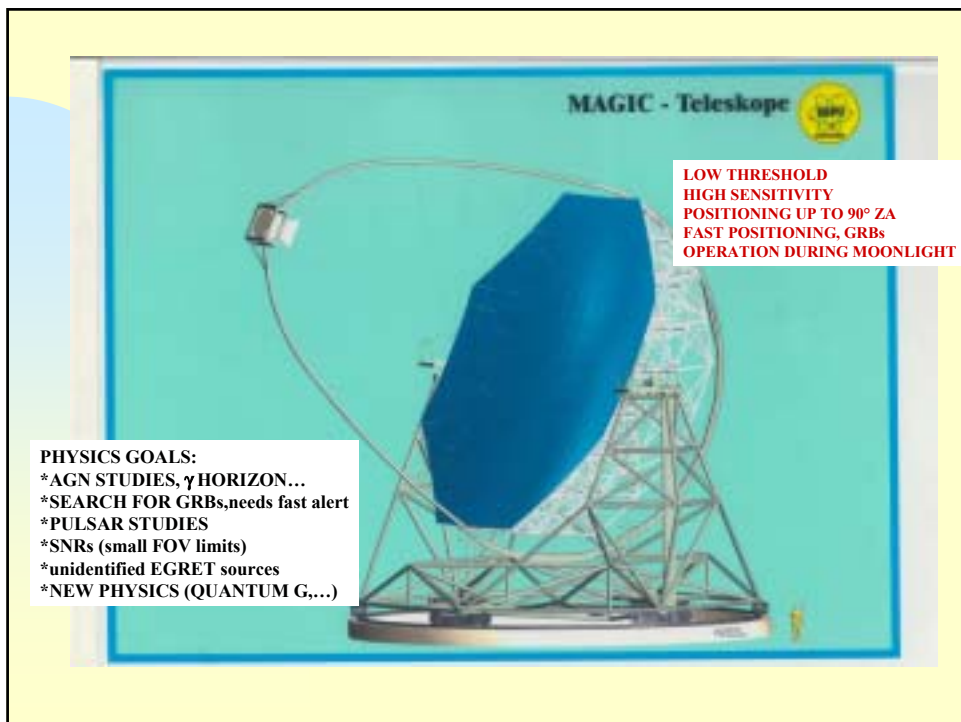
(Major Atmospheric Gamma Imaging Cherenkov Telescope)

A NEW TECHNOLOGY GROUND-BASED 17 m DIAMETER IMAGING TELESCOPE

LOCATED ON LA PALMA 28.8° N, 17.8° W, 2225 m asl (at old HEGRA site)

**LOW THRESHOLD: 30 GeV (phase I, classical pmts)
12-15 GeV (phase II, high QE pmt)**

First of a multitelescope system (see talk R. Mirzoyan: MAGIC II)



Brief history of the project

- Summer 1994: 1st feasibility discussions
- 1995: 1st presentations at conferences
- 1995-98: major R&D, at the end proposal
- 1996-98: several int. review committees
- 1997-99: formation of a collaboration
- FALL 2000: funding
- Early 2001: start production
- 2002: construction, assembly & 1st tests

NOVEL COMPONENTS

- * LOW WEIGHT CF SPACE FRAME CONSTRUCTION, 5 TONS OK
- * LIGHT WEIGHT ALL-ALU MIRRORS, DIAMOND TURNED, INTERNAL HEATING OK
- * ACTIVE MIRROR CONTROL, FULL TEST STILL MISSING (OK)
- * HEMISPHERICAL PMTS WITH 6 STAGES, IMPROVED QE OK
- * FAST PMT SIGNAL TRANSMISSION IN ANALOG MODE BY OPTICAL FIBERS WORKS, BUT MORE DIFFICULT THAN ANTICIPATED
DYN. RANGE 60 dB
BW: 230 MHZ
- * 300 MHZ F-ADC WITH HIGH, LOW RANGE SWITCHING, IN PRODUCTION
- * 3 LEVEL TRIGGER: STILL SOME DEVELOPMENT OF THE 2. LEVEL TRIGGER
- * \approx 1 KHZ DATA RATE, PROOF STILL MISSING

THE MIRROR

17 m DIAMETER, 234 m**2

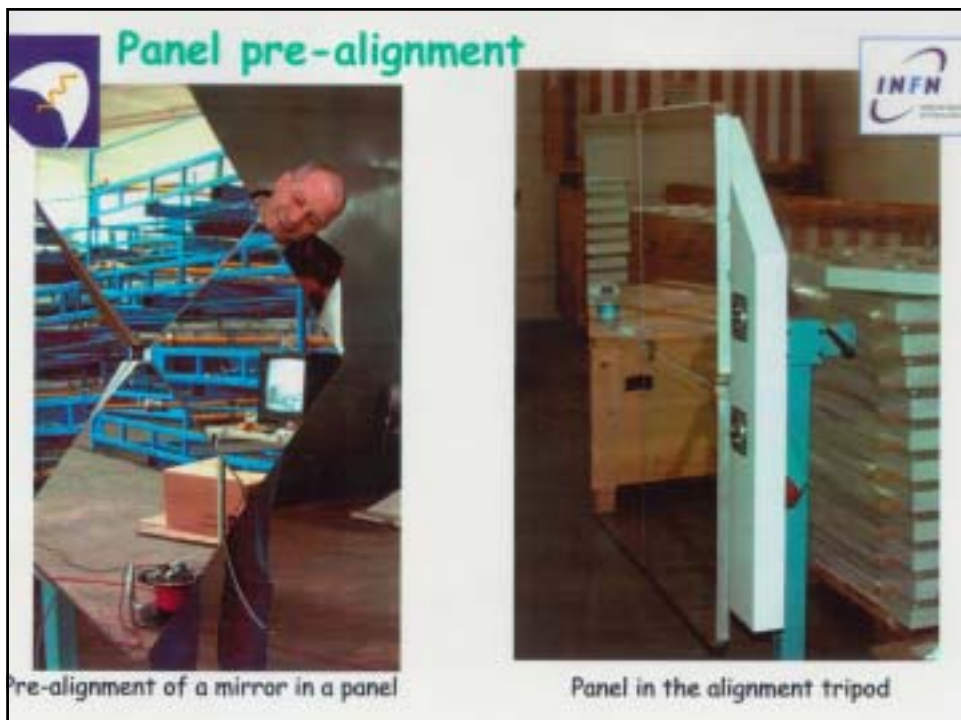
TESSALETED, PARABOLIC PROFILE

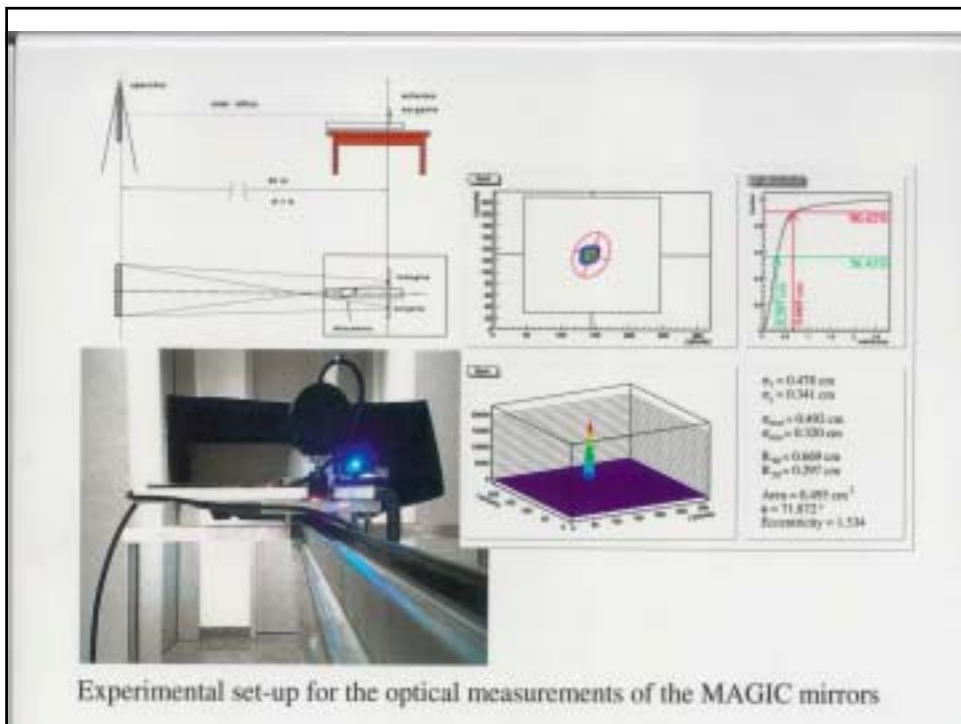
934 SMALL MIRRORS 49.5x 49.5 cm,

All raw blanks made,

700 mirrrors diamond turned

45 m** 2 installed -> 100 m**2 in October





Experimental set-up for the optical measurements of the MAGIC mirrors

MIRRORS

Zur Anzeige wird der QuickTime™
Dekompressor "Foto - JPEG"
benötigt.

TEST SURFACE ROUGHNESS OF DIAMOND TURNED MIRRORS

Zur Anzeige wird der QuickTime™
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CAMERA IMPROVEMENTS

**(4° DIAMETER :394 PIXELS 0.1°, 180 PIXELS 0.2°
6 DYNODE , HEMISPHERICAL CATHODE PMTS,
GAIN≈ 2X10**4 + FAST PREAMP -> OPERATION DURING
MOONSHINE POSSIBLE)**

**SPECIAL GEOMETRY WINSTON CONES IN CENTRAL REGION:
OFTEN PHOTON TRAJECTORY PASSES 2x CATHODE**

**SPECIAL TREATMENT OF CATHODE WINDOW WITH
LAQUER LOADED WITH WLS, FROSTED SURFACE
-> QE INCREASE BY 20-30 %**

**NEXT: LINING WINSTON CONES WITH 98% REFLECTIVITY
FOIL
-> GAIN ≈ 13-20%**

CAMERA TO TELESCOPE IN OCT/NOV

CAMERA

**A METHOD TO INCREASE
THE QE:COAT WINDOW
WITH A LAQUER LOADED
WITH WLS AND USING A
FAST EVAPORATING
SOLVENT ->
FORMS FROSTED WINDOW
SURFACE LAYER**

Zur Anzeige wird der QuickTime™
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CAMERA

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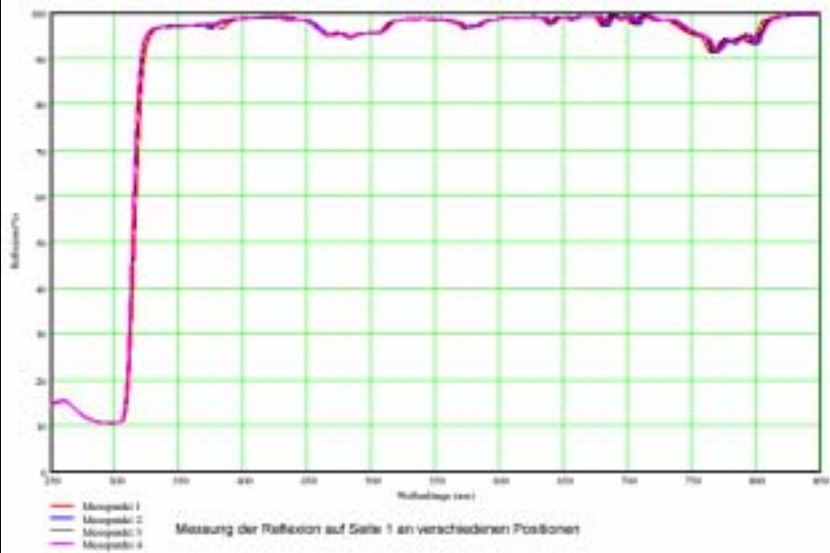
CAMERA

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benötigt.

CAMERA

**A NEW LINING MATERIAL FOR THE WINSTON CONES: DIELECTRIC
MIRROR FOIL FROM 3M, STILL EXPERIMENTAL PRODUCTION**

Reflexionsmessung hochreflektierender Folien der Firma 3M - Folie Nummer 2 - Seite 1



MAGIC USES OPTICAL FIBERS FOR THE ANALOG SIGNAL TRANSFER FROM CAMERA PMTS TO THE 100 M AWAY COUNTING HOUSE

**PROS: LOW WEIGHT OF CAMERA (OSCILLATIONS REDUCED)
REDUCED ELECTRONICS IN CAMERA -> LESS DOWNTIME DUE TO DEFECTS
LOW POWER CONSUMPTION IN CAMERA
EASIER TO MODIFY ELECTRONICS (TRIGGER...)
EASIER TO IMPLEMENT TRIGGER FOR MULTITELESCOPE SYSTEM**

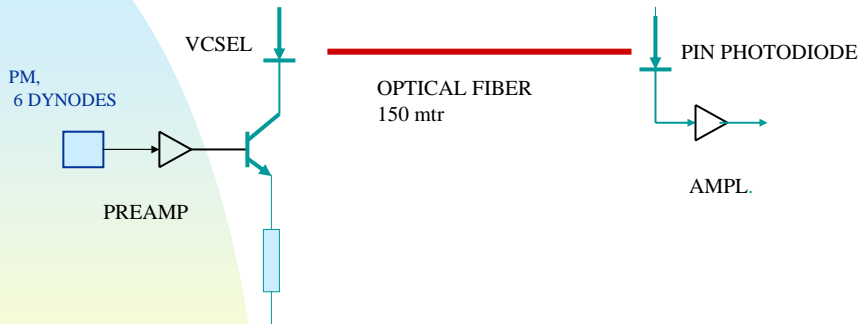
**CONS: NEW PRINCIPLE, NO EXPERIENCE
EXTRA COSTS**

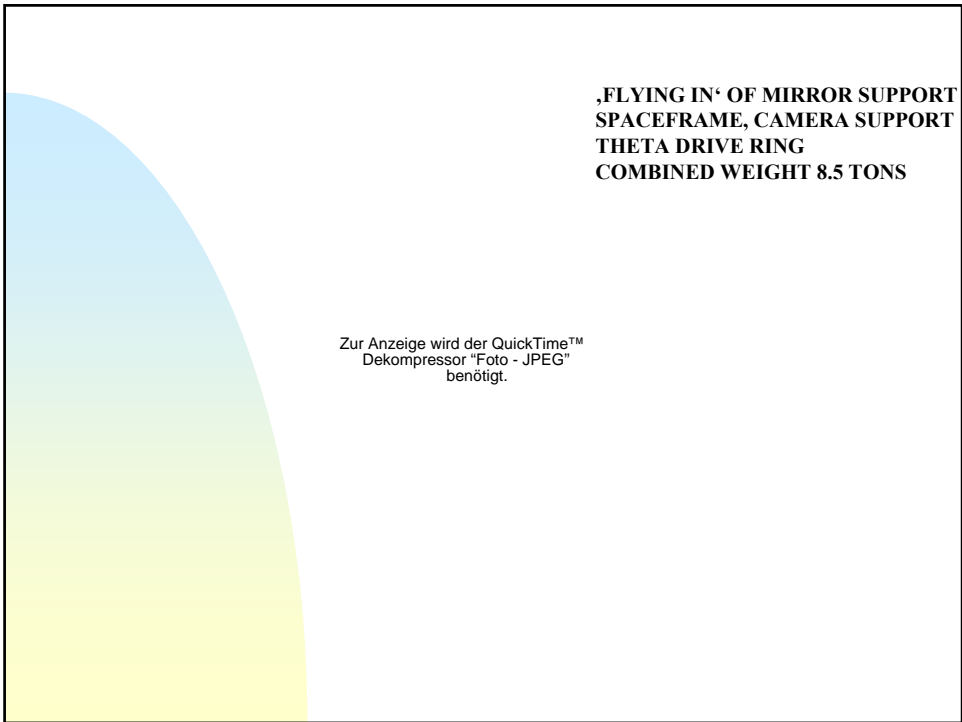
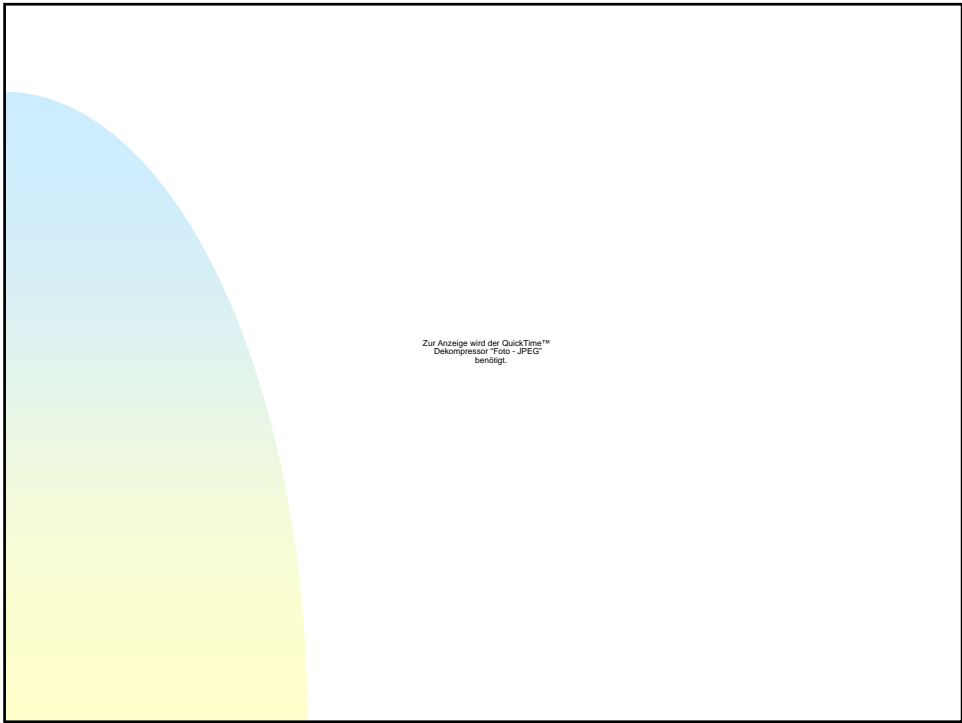
USE OF VCSELS AS LIGHT EMITTER (CAN DRIVE TO 200 MA PULSE CURRENT FOR LOW DUTY CYCLE.
HIGH BANDWIDTH (> 1 GHZ POSSIBLE, NOW 230 MHZ BW)
NO ATTENUATION OVER 165 M LENGTH

**PROBLEM AT THE BEGGINNING: WE WERE TOO ABITIOUS ABOUT DYN RANGE:
72 dB TRIED BUT OSCILLATIONS AND INSTABILITY FOR SMALL SIGNALS**

**SOLUTION: SELECT VCSELS
BW 230 Mhz
INCREASE INPUT GAIN x4 -> DYNAMIC GAIN \approx 1000**

FAST PM SIGNAL TRANSMISSION BY OPTICAL FIBER SYSTEM WORKING IN ANALOG MODE





THE TELESCOPE FRAME BEFORE THE MIRROR INSTALLATION



INSTALLATION OF MIRROR PANELS



THETA DRIVE INSTALLATION



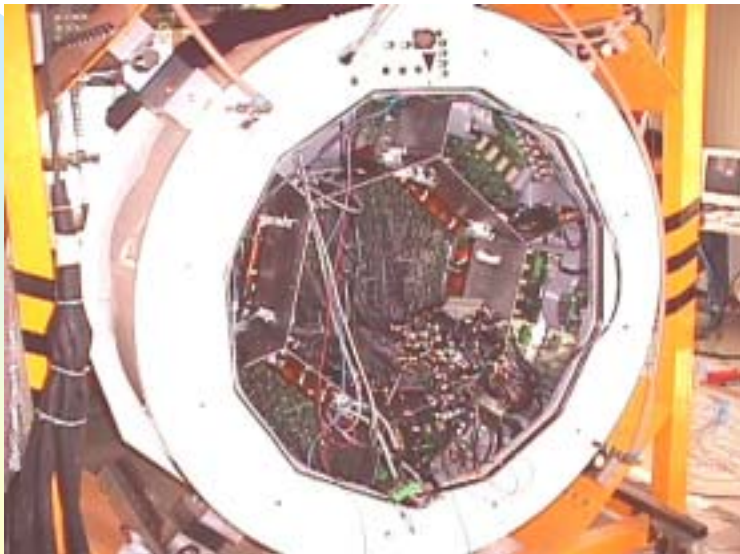
PHOTOGRAPH OF CAMERA, PMT SIDE, WITHOUT WINSTON CONE PLATE



Camera in temporary frame, front side



Camera in temporary frame, backside



FIRST TEST RUNS WITH 100 m2 MIRROR
STILL THIS YEAR**

COMPLETION FULL MIRROR: SPRING 2003

PHYSICS RUNS BEGIN IN 2003

COUNTING HOUSE LATE, BACKUP OLD HEGRA CONT.

TELESCOPE WEIGHT \approx 60 TONS

COSTS: 3.5 m \$ (ONLY TELESCOPE)

CONCLUSIONS

- * **MAGIC INSTALLATION PROCEEDS WELL , TO BE COMPLETED EARLY 2003
1 YEAR BEHIND ORIGINAL PLANS
BUT FUNDING WAS DELAYED BY \approx 2 YEARS**
- * **ALL ' NEW TECHNOLOGY ' ELEMENTS WORK AS EXPECTED
HARD FIELD TESTS STILL MISSING**
- * **MAGIC FULLY OPERATIONAL FOR PHYSICS IN 2003
(IF NO LATE SURPRISES)**
- * **NO UPDATE ON SENSITIVITY OR PERFORMANCE
BECAUSE SOON EXPERIMENTAL VERIFICATION
30 GEV THRESHOLD STILL STANDING (SOME IMPROVEMENTS
IN LIGHT COLLECTION)**
- * **NOW PLANNING OF UPGRADE PATH TO MULTIPLE/STEREO TELESCOPES:
MAGICII**