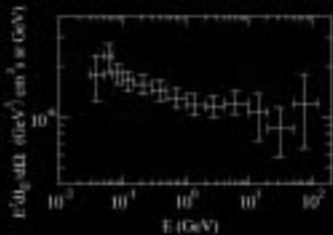


The Guaranteed Gamma-Ray Background

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The Observed Gamma Ray Background



EGRET observations of γ -ray background

(Sreekumar et al. 1998):

- *isotropic (extragalactic)*
- *single power law of index - 2.1 ± 0.03*
- *all-sky avg. $\sim 1.45 \times 10^{-5} \text{ cm}^{-2} \text{ s}^{-1} \text{ sr}^{-1}$*

EGRET observations of extragalactic γ -ray sources

(Hartman et al. 1999):

identified {
✓ AGN (blazars)
✓ normal galaxies

Unresolved sources of same class
contribute to diffuse background



The Guaranteed Gamma-Ray Background

guaranteed γ -ray background \equiv sum of γ -ray emission
from all unresolved identified sources
 $=$ unresolved blazars +
unresolved normal galaxies

Other proposed γ -ray sources

e.g.: annihilating dark matter (Bergstroem talk)

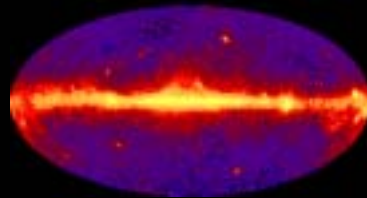
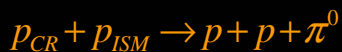
structure formation cosmic rays (Miniati talk),

constrained by :

(observed background) – (guaranteed background)

Gamma rays from Normal Galaxies

γ -rays in normal galaxies
produced through:



EGRET γ -ray sky

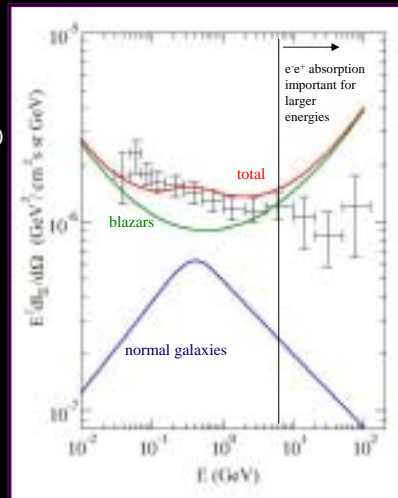
γ -ray flux of typical galaxy *higher* in the past because:

1. *Star formation rate higher*
 \Rightarrow more supernovae
 \Rightarrow larger cosmic ray flux
2. *More targets available*

- use *cosmic star formation rate* to calculate both effects.
- normalize γ -ray luminosity and spectrum produced per star formation rate unit to Milky Way

Gamma Ray Background – the Minimal Model

- ✓ Minimal 2-component model for γ -ray background :
normal galaxy contribution + blazar contribution (Stecker & Salamon 1996)
- ✓ Blazar spectrum: *concave*
Normal galaxy spectrum: *convex*
Summed spectrum: *flatter than either*
 \Rightarrow *better fit to observations*
- ✓ Relative normal galaxy contribution: highest at ~ 1 GeV
(about 1/3 of summed spectrum)

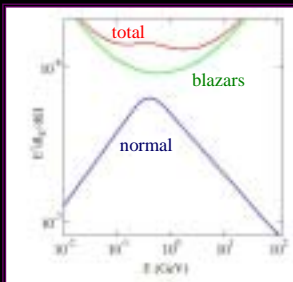


Observational Tests

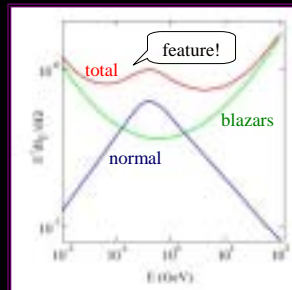
GLAST will:

1. *test the minimal model:*

will resolve many more blazars but at most 3 new normal galaxies
 \Rightarrow relative blazar contribution reduced
 \Rightarrow will detect normal galaxy peak at ~ 1 GeV



EGRET



GLAST

2. *improve observational inputs for both models*

Conclusions

1. Unresolved normal galaxies & blazars:
guaranteed, identified sources of diffuse γ -ray background.
Need to understand those to constrain any other source
and relevant physics.
2. Normal galaxy spectrum convex, blazar spectrum concave.
Sum flatter than either.
3. **GLAST** observations *will test* relative contribution of two
components to the gamma-ray background.