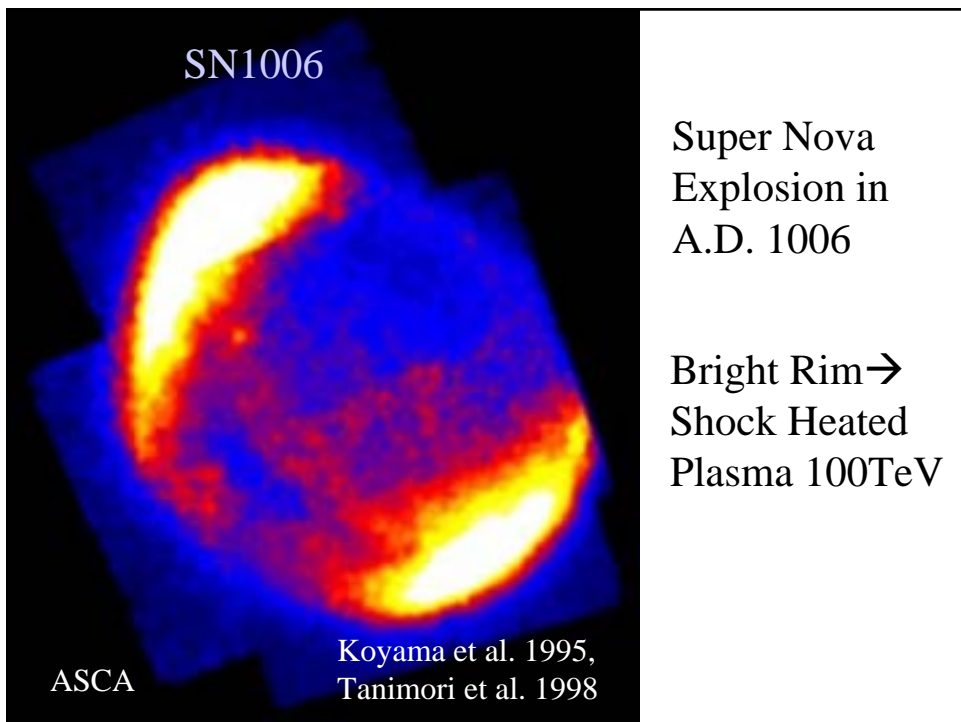


Plasma Physics View of Particle Acceleration

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University of Tokyo



SN1006

NE-Rim

Shock Fronts

shock heated plasma in narrow layers

$0.01\text{pc} \sim 10^{16}\text{cm} \sim$
electron gyro-radius
of 10 TeV

Fermi acceleration model

Why Shock Front is Laminar ???

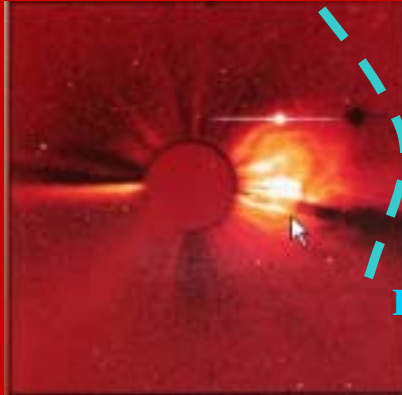
$\eta = D/D_B \sim 1$ (Bohm limit)

Chandra

Bamba et al. (2002)

Interplanetary Shock:
Diffusive Shock Acceleration
in Heliosphere

Strong Shocks --- Interplanetary Shock (IPS)
(Solar Flare-Initiated Shock)

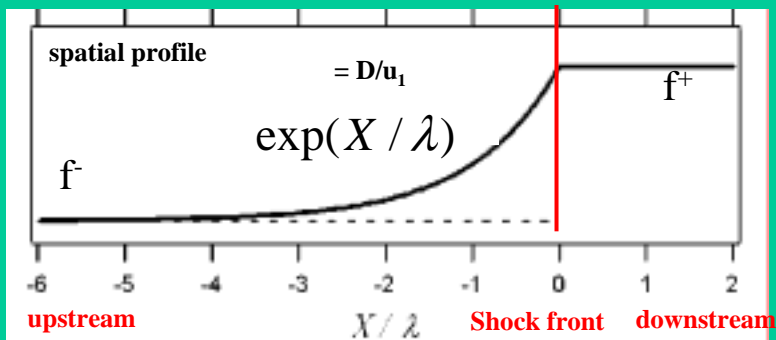


1000 ~ 4000 km/s

IPS

Coronal Mass Ejection (CME)

Diffusive Shock Acceleration (DSA)



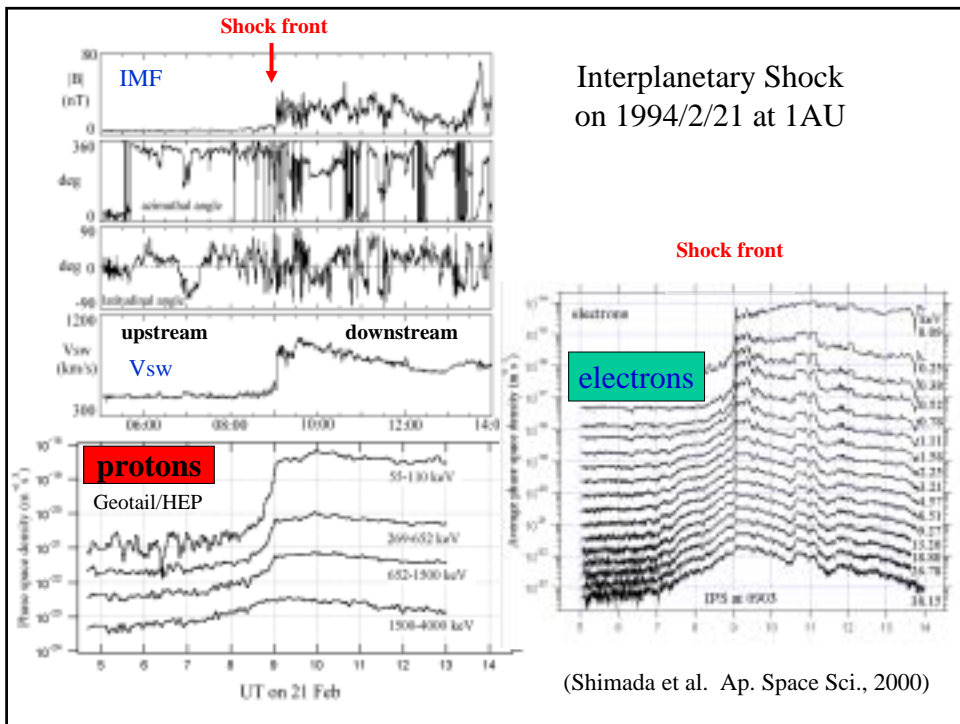
differential momentum spectrum

$$N(p) \propto f(p) \cdot p^2 \cdot p^{-q}$$

with $q = (r+2)/(r-1)$,

r : compression ratio

Nonthermal Energy Spectrum E^{-2}



Diffusion Coefficient: $D=(1/3) l w$
(l : mean free path; w :particle velocity)

“ l ” depends on level of turbulence

strongest limit :

Bohm limit: $l \approx \rho_c$, $D_B=(1/3) \rho_c w$

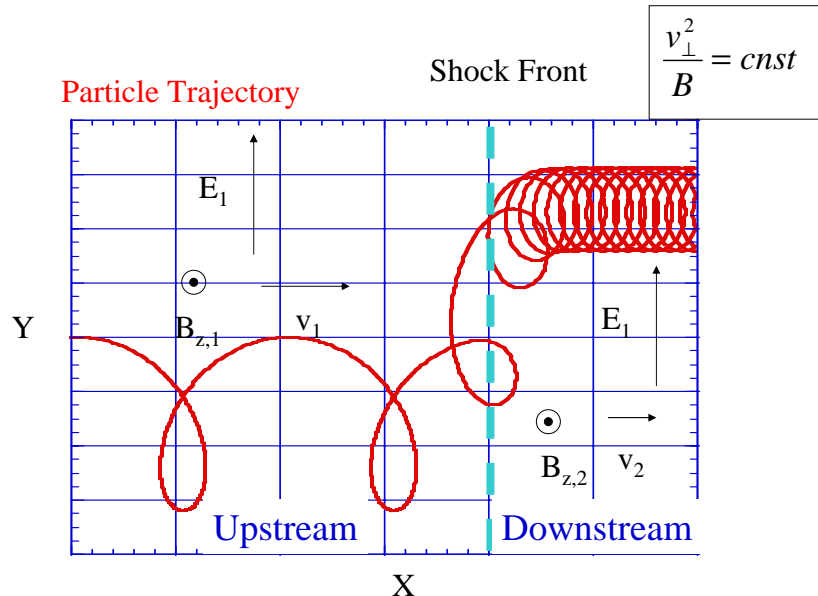
Heliosphere : Interplanetary Shocks

$\eta=D/D_B$ with $\eta \sim 10 - 100$ for proton
 $\eta \sim 500-10000$ for electron

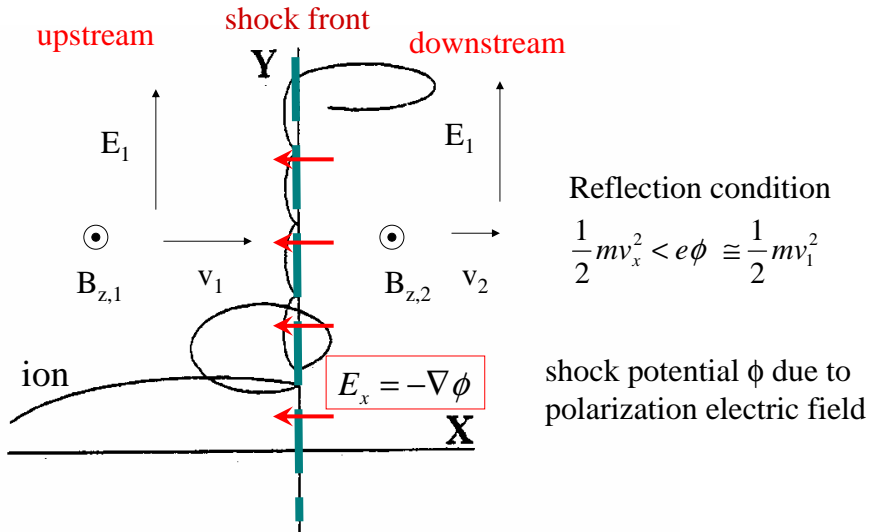
Other Acceleration Processes

Shock Drift Acceleration
Shock Surfing Acceleration

Shock Drift Acceleration



Ion "Shock Surfing" Acceleration



Sagdeev (1966), Sagdeev and Shapiro (1973), Katsouleas and Dawson (1983), Lembege et al. (1983), Ohsawa (1990),...Lee et al. (1996), Zank et al. (1996),....

Energy Spectrum in Shock Surfing Acceleration

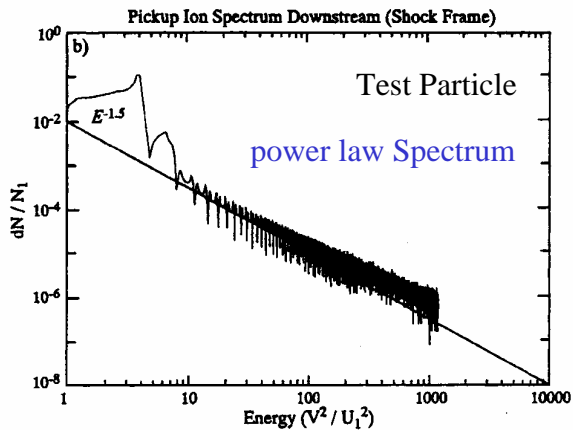


Figure 8. The transmitted PI spectrum at a perpendicular shock in the (a) fluid and (b) shock frames. An $E^{-1.5}$ power law has been plotted over the computed spectrum for the purposes of illustration. See text for further details. $V_{spec}/u_1 = 0.25$.

(Zank et al. JGR, 1996)

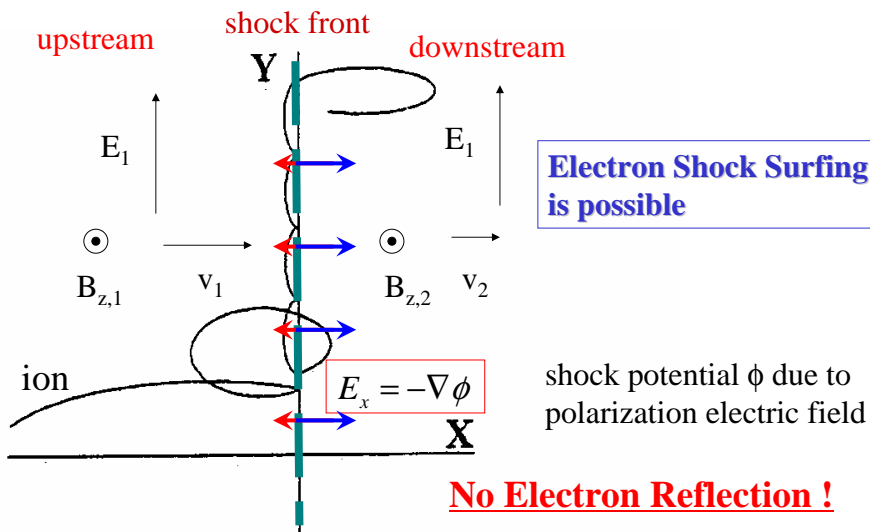
Shock Surfing Acceleration (SSA)

- Localized Region along Shock Front
- Fine Structure

But ...

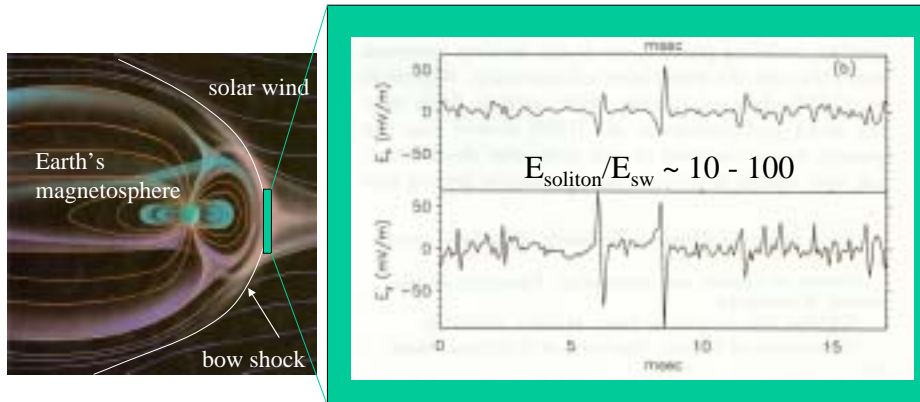
electron acceleration cannot be realized by
the standard shock surfing acceleration

Ion "Shock Surfing" Acceleration



“Solitary waves” in the Earth’s bow shock

quasi-perpendicular, $M_s \sim 9.5$



Scale of Solitary wave \sim Electron gyro-radius

(Bale et al. GRL 1998, Matsumoto et al. ASR 1997)

Plasma Particle-in-Cell Simulation

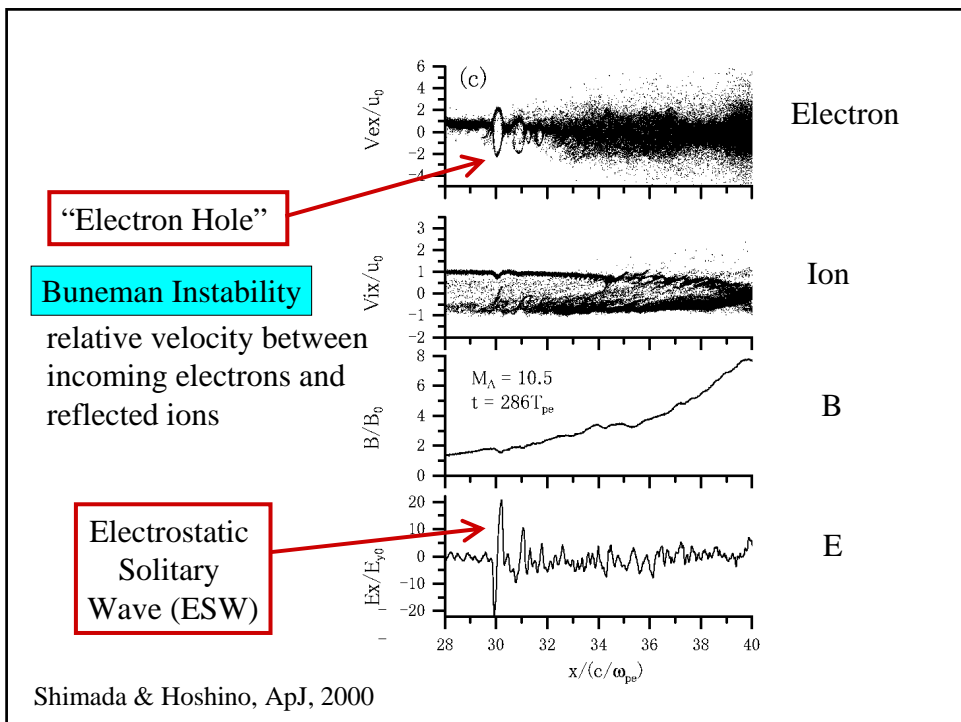
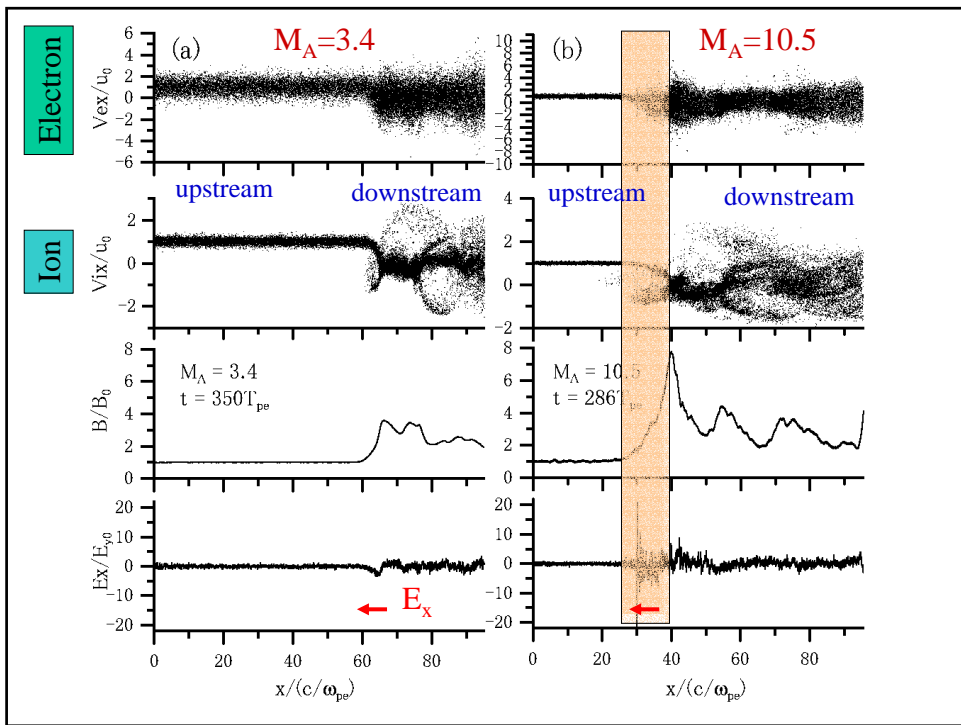
A set of Maxwell's Equations
+ relativistic Lorentz Equation

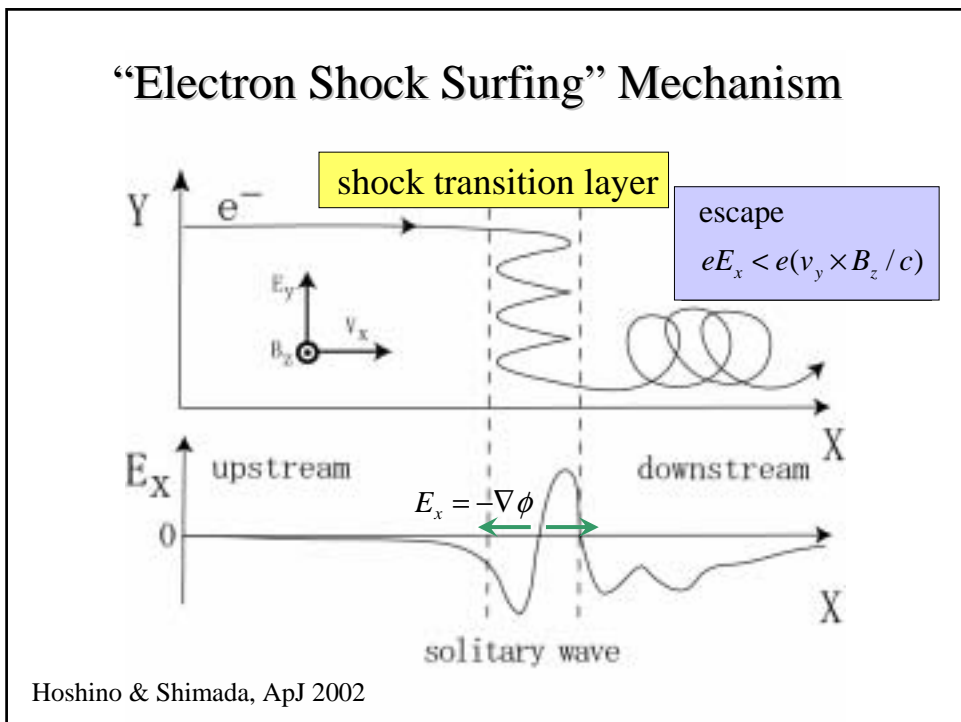
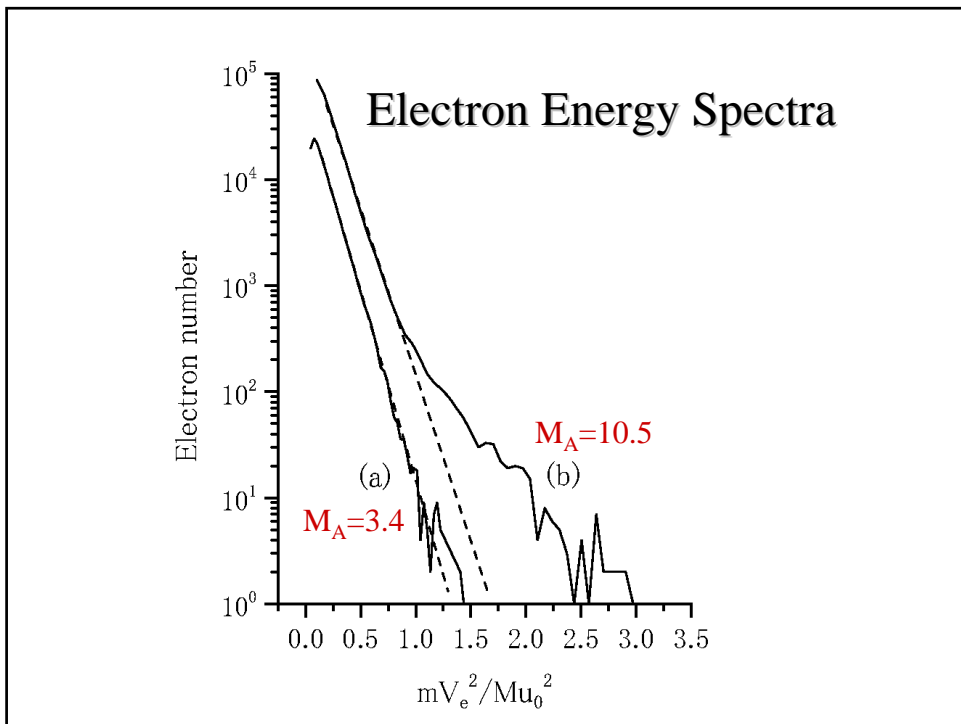
Mach number : 3 ~ 20

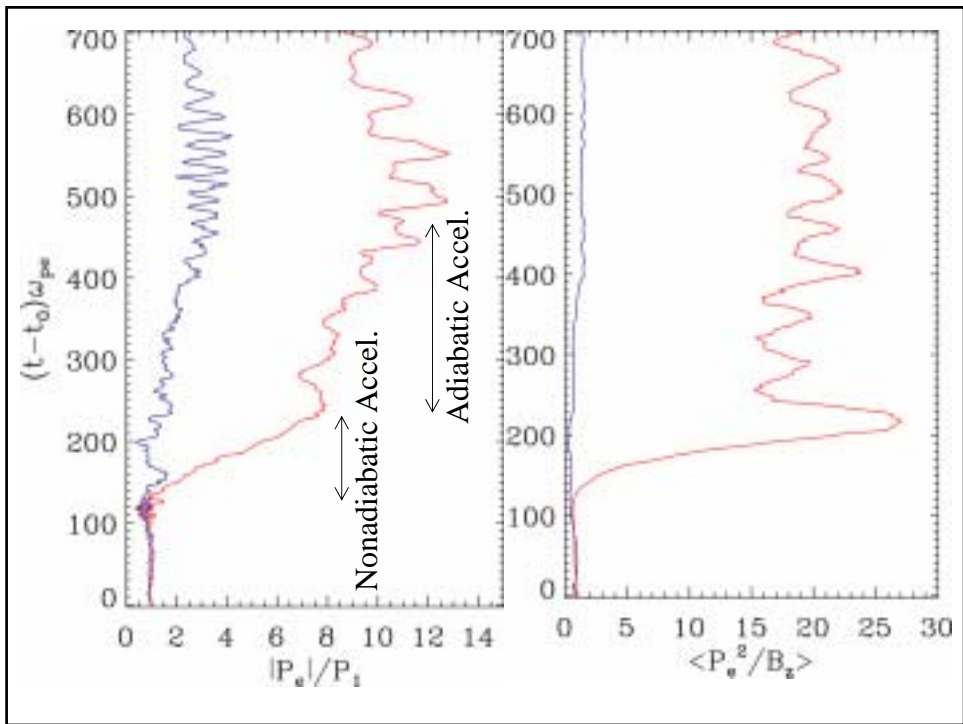
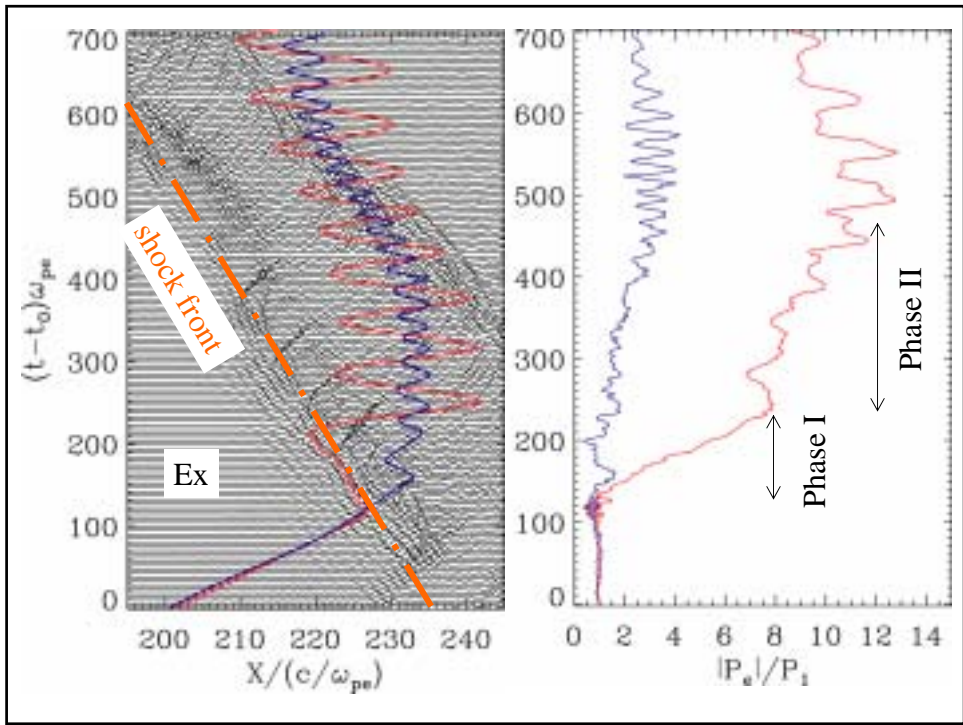
$M_{\text{ion}}/M_{\text{ele}} = 20$

100 grids for electron inertia length

Simulation system size: 100 ~ 300 electron inertia







Maximum Electron Energy

$$\frac{dp_x}{dt} = eE_x + e \frac{v_y \times B_z}{c}$$

$$F_E > F_B \rightarrow \text{trapped}$$

$$F_E < F_B \rightarrow \text{de trapped}$$

$$F_E \approx eE_{ESW}$$

$$F_B \approx \frac{e}{c} v_y B$$

Nonlinear Buneman Instability

$$E_{esw}^2 / 8\pi \approx \alpha m_e n V_d^2 / 2, \quad \alpha = 1/4 \sim (m_e / m_i)^{1/3}$$

Ishihara et al, 1981; McClements et al. 2001

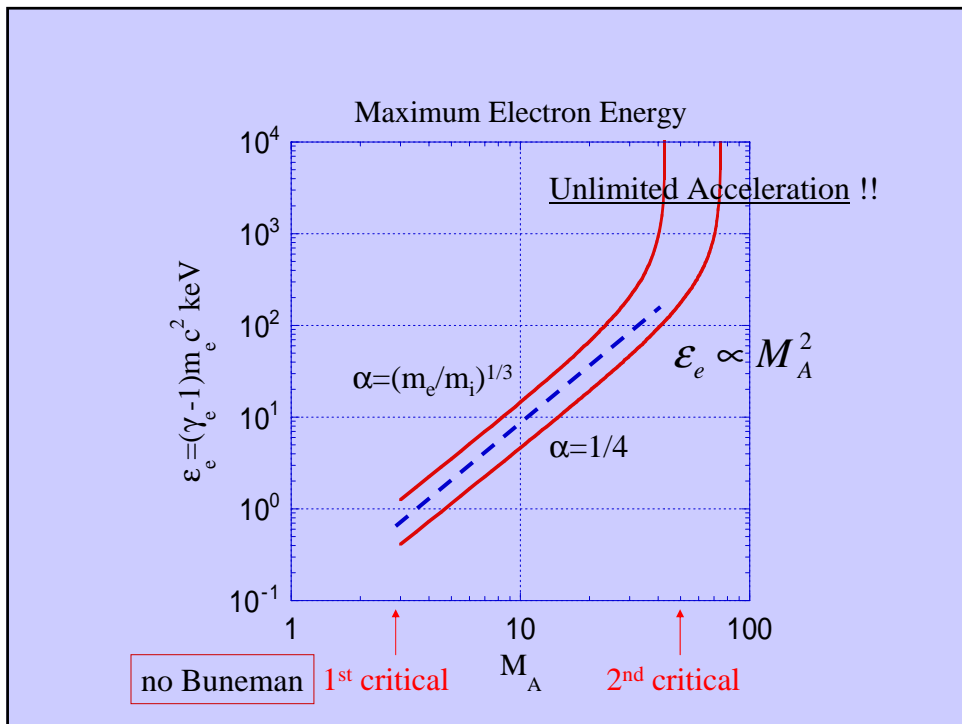
Maximum Electron Energy

$$F_E = F_B \Rightarrow$$

$$v_{y,\max} \sim \begin{cases} 2cM_A \sqrt{\alpha m_e / m_i} & \text{for } 2M_A \sqrt{\alpha m_e / m_i} < 1 \\ c & \text{for } 2M_A \sqrt{\alpha m_e / m_i} > 1 \end{cases}$$

unlimited electron acceleration condition

$$M_A > (1/2) \sqrt{m_i / (m_e \alpha)} \approx 43 \sim 75$$



Summary

- “Electron” Shock Surfing Acceleration
fine structure along shock front
- Unlimited Acceleration if $M_A > \text{several } 10$
- Acceleration Time Scale

$$\tau_{\text{SSA}} \Omega_c \sim c/v_s \quad (\text{cf. } \tau_{\text{DSA}} \Omega_c \sim \eta(c/v_s)^2)$$

possible to explain up to knee energy