



CTAで迫る物理

CTA Physics Drivers

井岡 邦仁

Kunihito IOKA

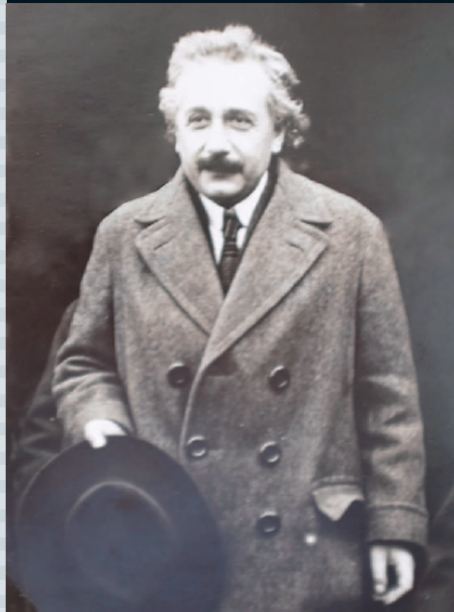
(KEK, SOKENDAI)



National University
SOKENDAI
The Graduate University for Advanced Studies



B U T S U R I
日本物理学会誌



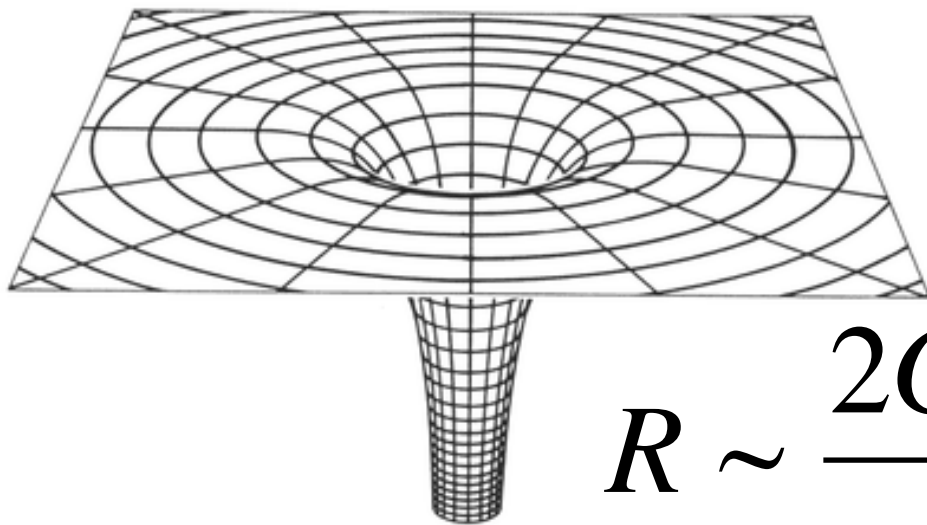
General Relativity



Amazing Predictions

Black Hole

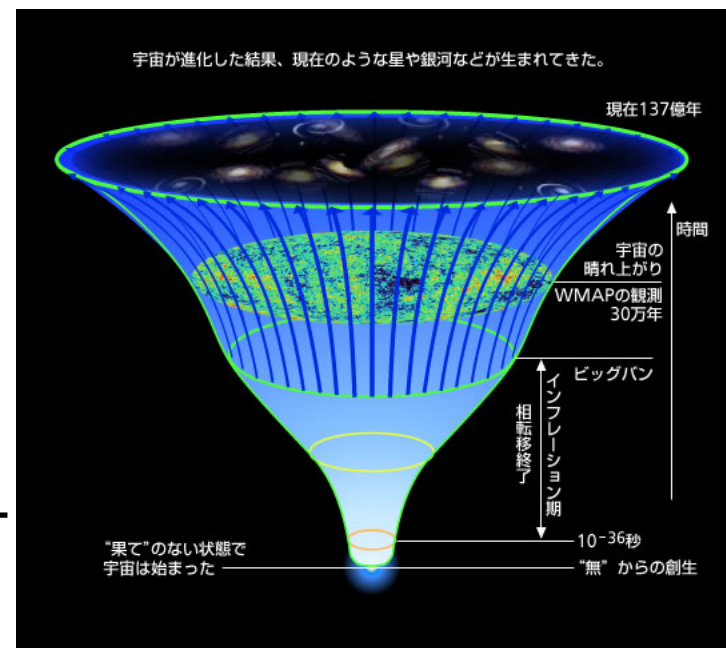
Expanding Universe



$$R \sim \frac{2GM}{c^2}$$

Final fate of stellar evolution

Astrophysics



Big Bang

Cosmology

Evolution to Black Hole

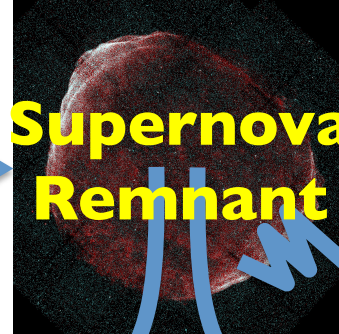
Gravity
Quantum



Star

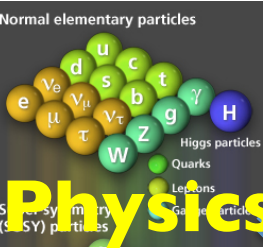


Supernova GRB



Supernova Remnant

W. A. Fowler

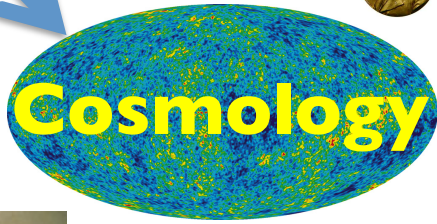


Physics

**Chandrasekhar
Bethe**

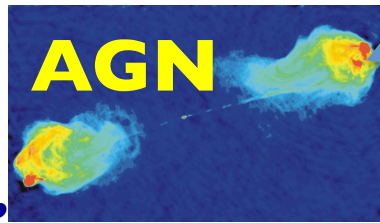


**Davis
Koshiba**



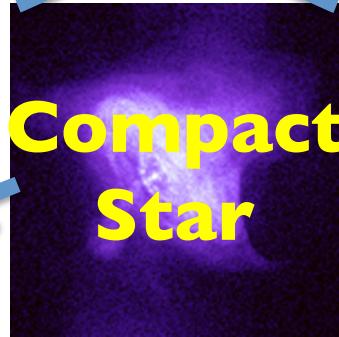
Cosmology

**Penzias, Wilson
Perlmutter, Riess,
Schmidt**



AGN

Active galactic nuclei
Massive black hole
at the center of galaxy



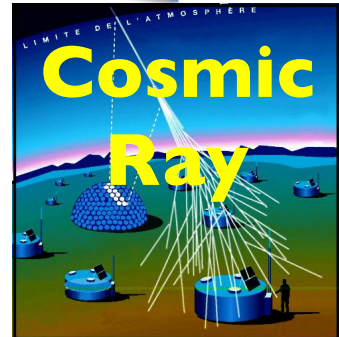
Compact Star

Black Hole
Neutron Star
White Dwarf

**Giacconi
Hewish**

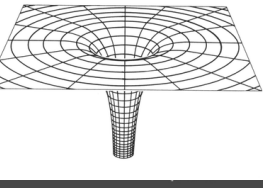


Hulse, Taylor



Cosmic Ray

**Hess
Anderson
Yukawa
Powell**



Computer

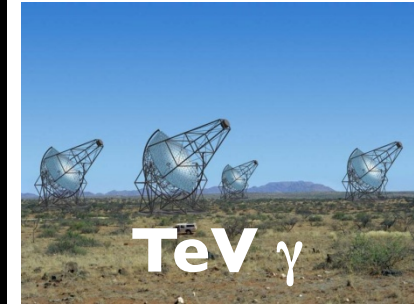
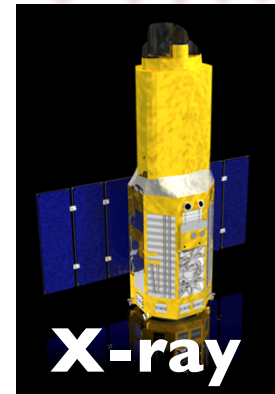
High Energy Universe

Multi-Wavelength

Wavelength
(meters)

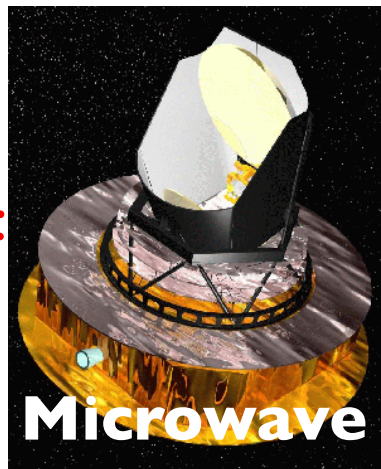


About the size of...



TeV = 10^{12} eV
 GeV = 10^9 eV

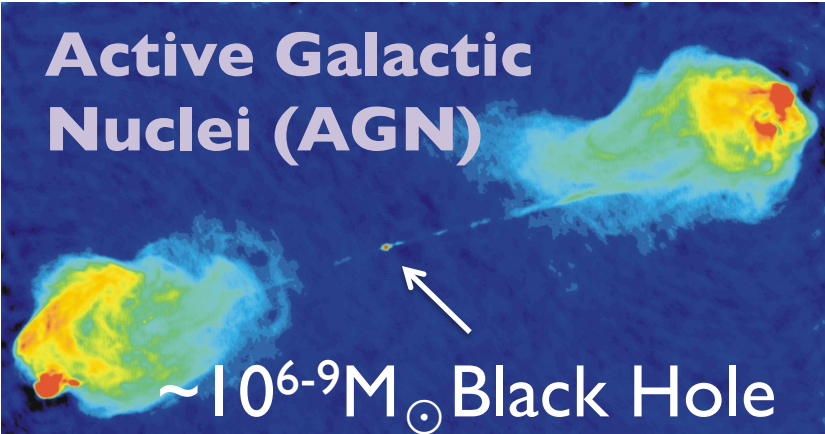
20th century observations:
 In particular,
 X- γ -VHE γ



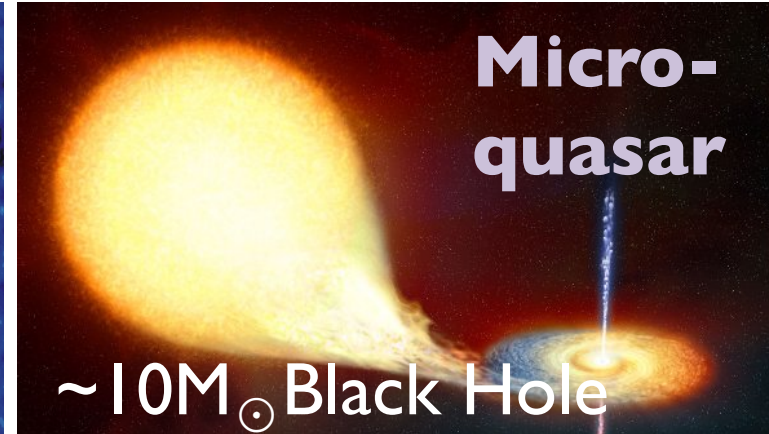
Cosmic Accelerators

High Energy Phenomena pervade the Universe

**Active Galactic
Nuclei (AGN)**



**Micro-
quasar**



**Dark
Matter
?**

Gravity, EM, Weak, Strong + New?

**Pulsar
PWN**

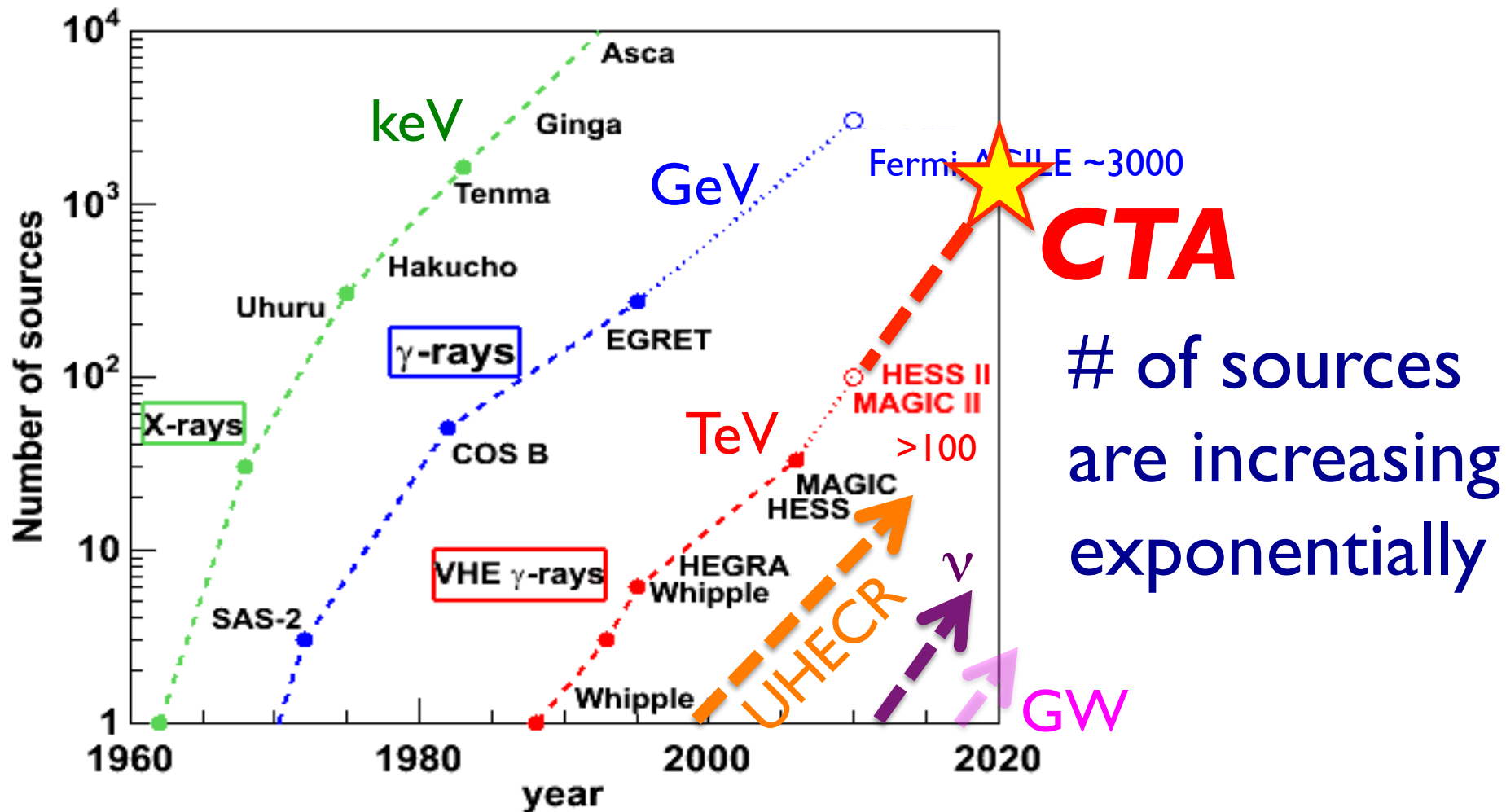
Neutron Star
w/ $10^8\text{-}14 G$

**Supernova
Remnant**

**Gamma-Ray
Burst (GRB)**

Explosive Development

Kifune Plot

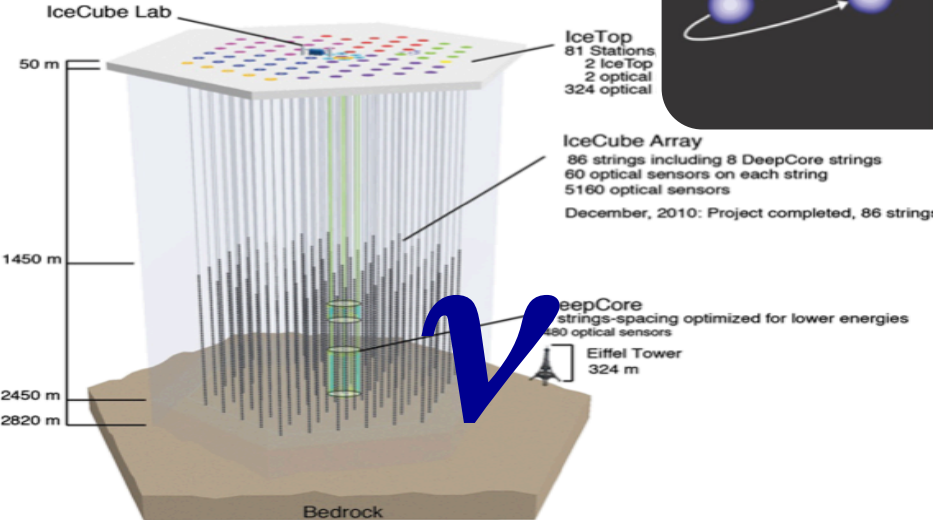
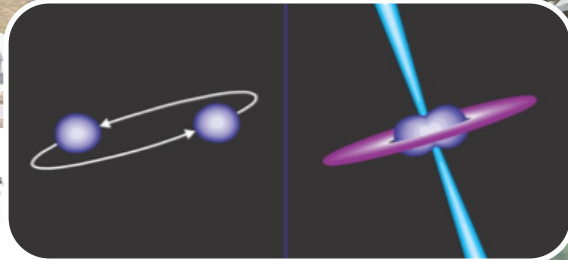
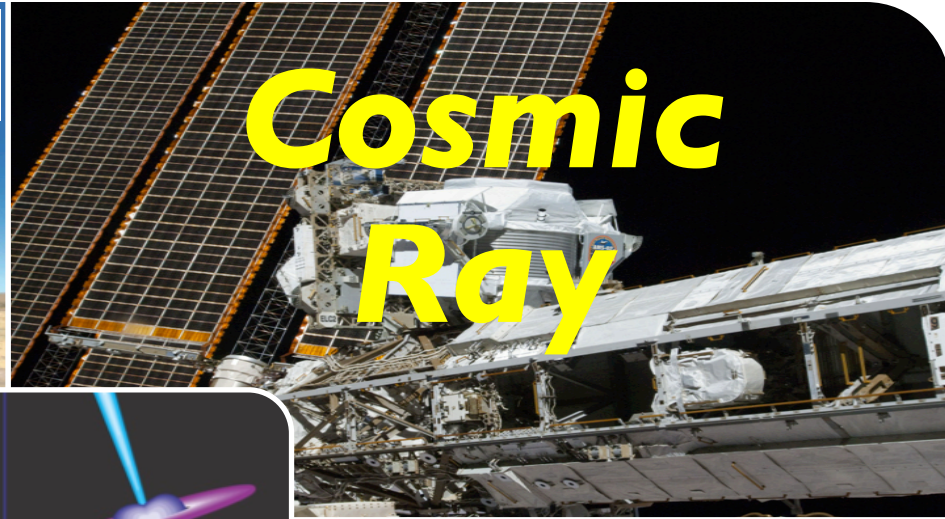


Multi-Messenger Era

Photon



Cosmic Ray

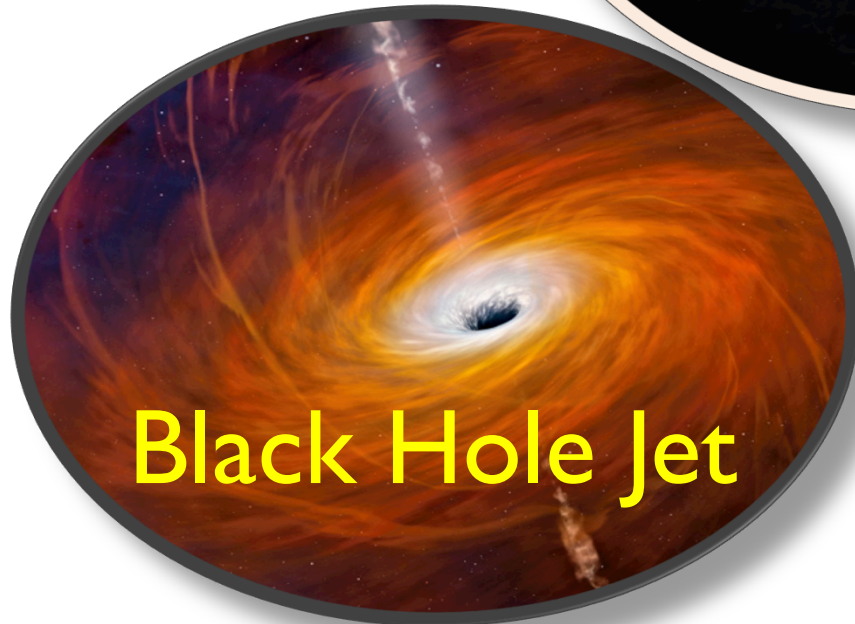


Gravitational Wave

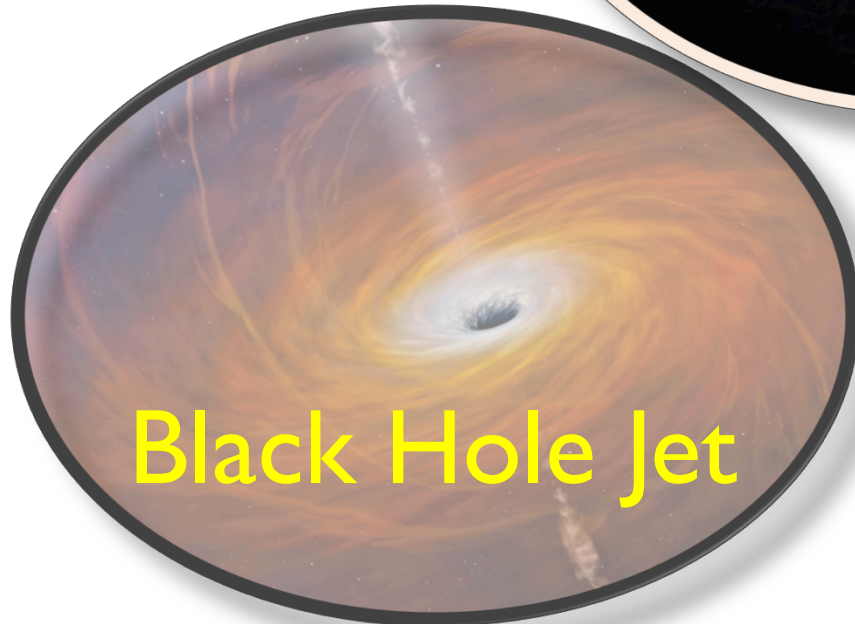
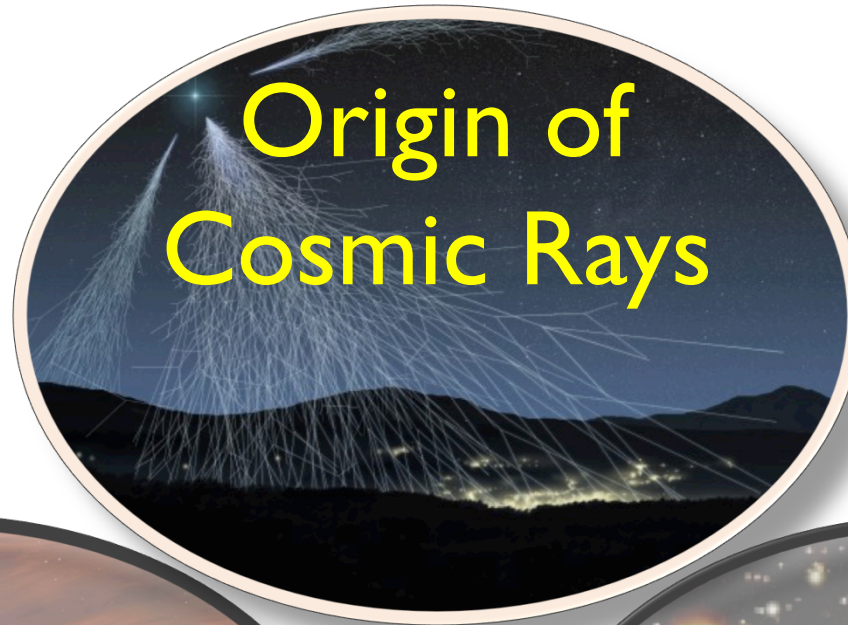


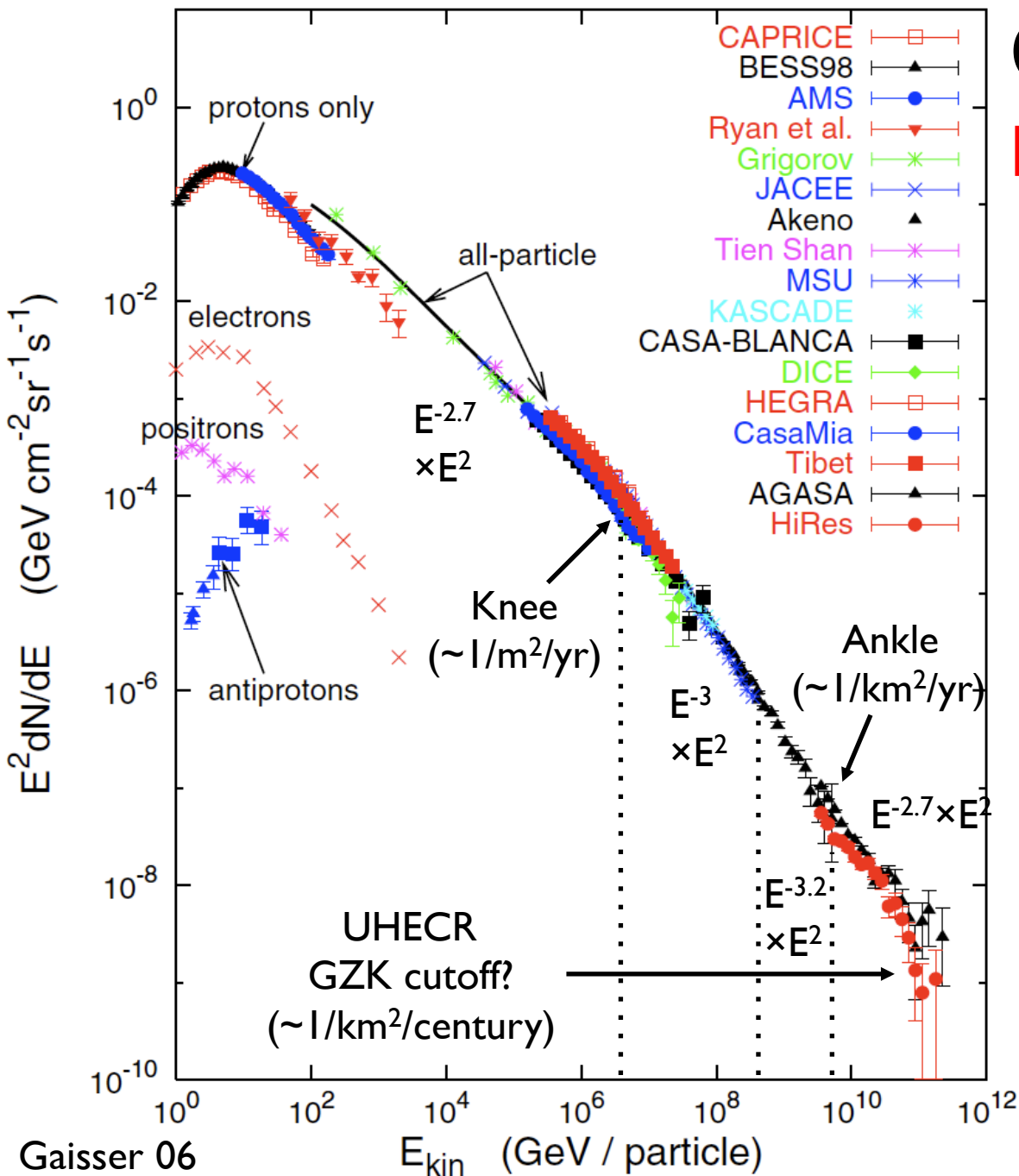
21st Century: Multi-Messenger Era

Mysteries in 21th Century



Mysteries in 21th Century





Cosmic Ray

$E < 10^{15-16} \text{ eV}$ (Knee)

$F \propto E^{-2.7}$

Supernova remnant(?)

$L_{\text{CR}} \sim 10^{41} \text{ erg/s}$
 $\sim 0.1 E_{\text{SN}}/t_{\text{SN}}$

$10^{15-16} < E$

$< 10^{18} \text{ eV}$ (Ankle)

$F \propto E^{-3-3.2}$

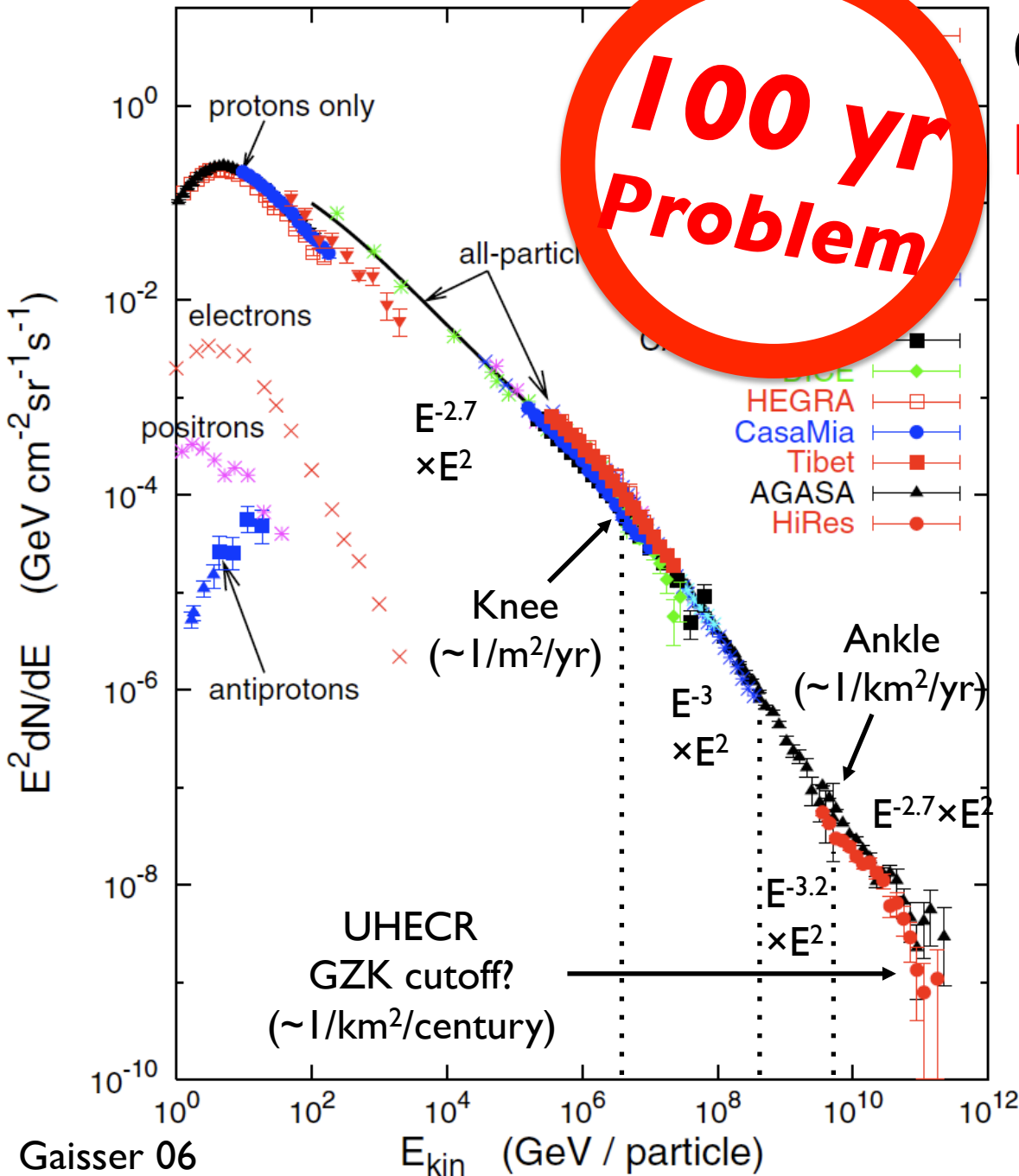
Galactic origin?

$< 10^{14-15} \text{ eV}$ by SNR

$10^{18} \text{ eV} < E$

$F \propto E^{-2.7}$

Extra-Gal. AGN? GRB?



100 yr Problem

Cosmic Ray

$E < 10^{15-16} \text{ eV}$ (Knee)

$F \propto E^{-2.7}$

Supernova remnant(?)

$L_{\text{CR}} \sim 10^{41} \text{ erg/s}$
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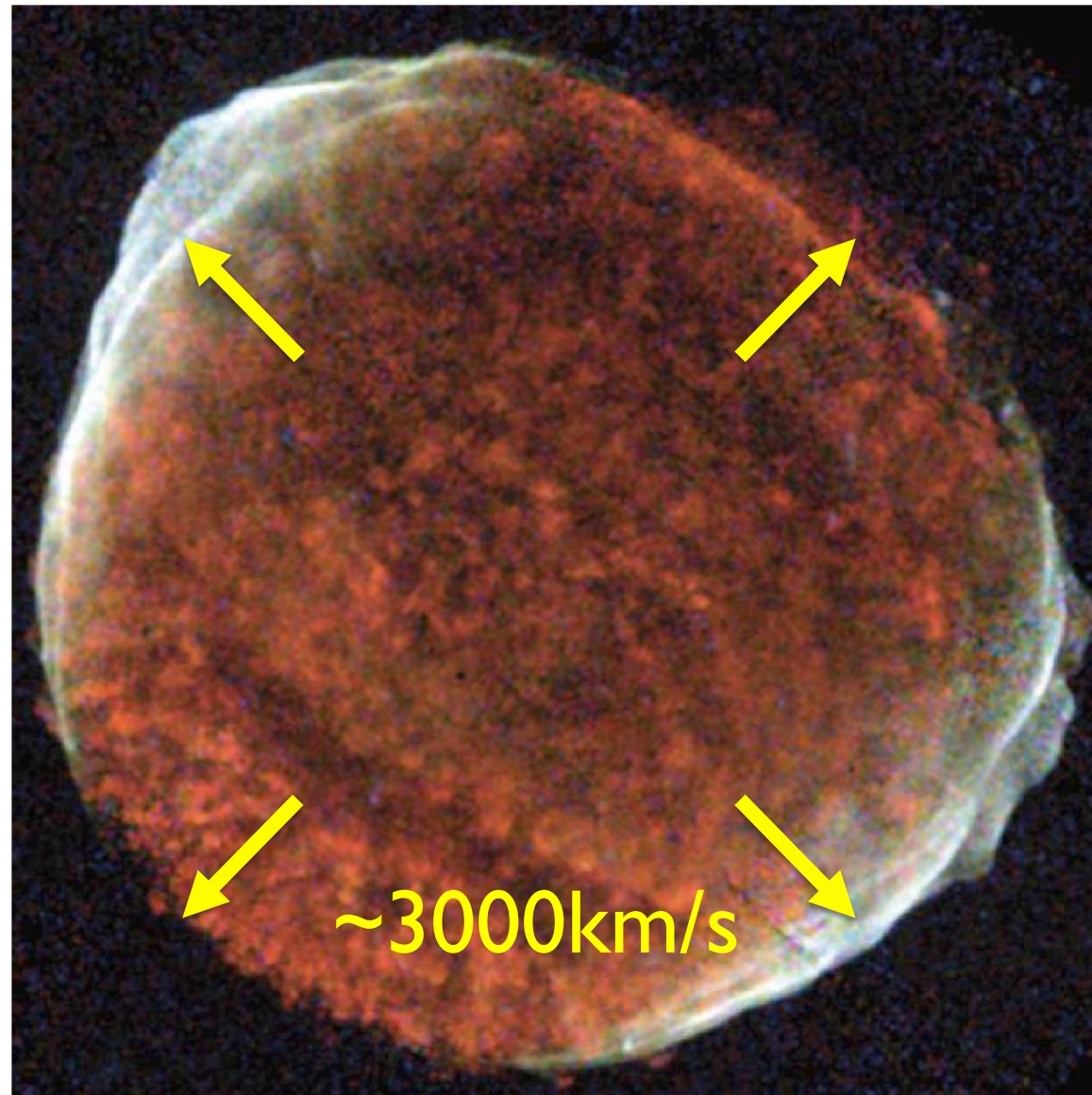
$< 10^{14-15} \text{ eV}$ by SNR

$10^{18} \text{ eV} < E$

$F \propto E^{-2.7}$

Extra-Gal. AGN? GRB?

Supernova Remnant



Supernova
explosion

(gravity, nuclear)

⇒ Shock

⇒ Heating

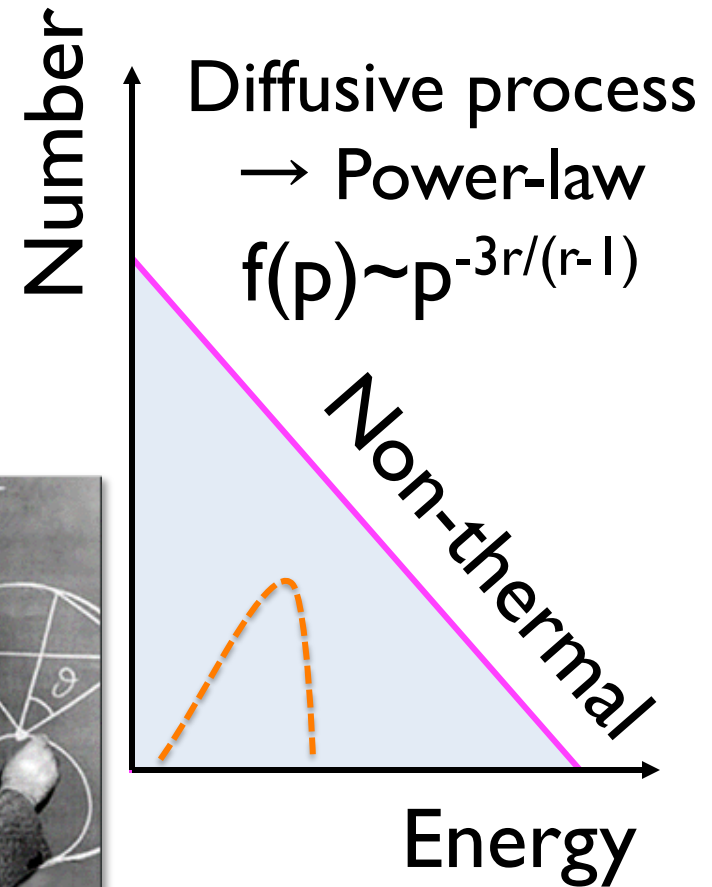
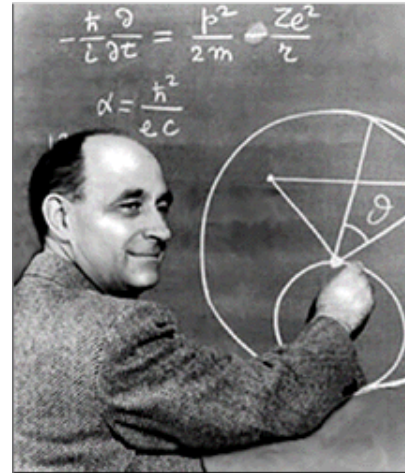
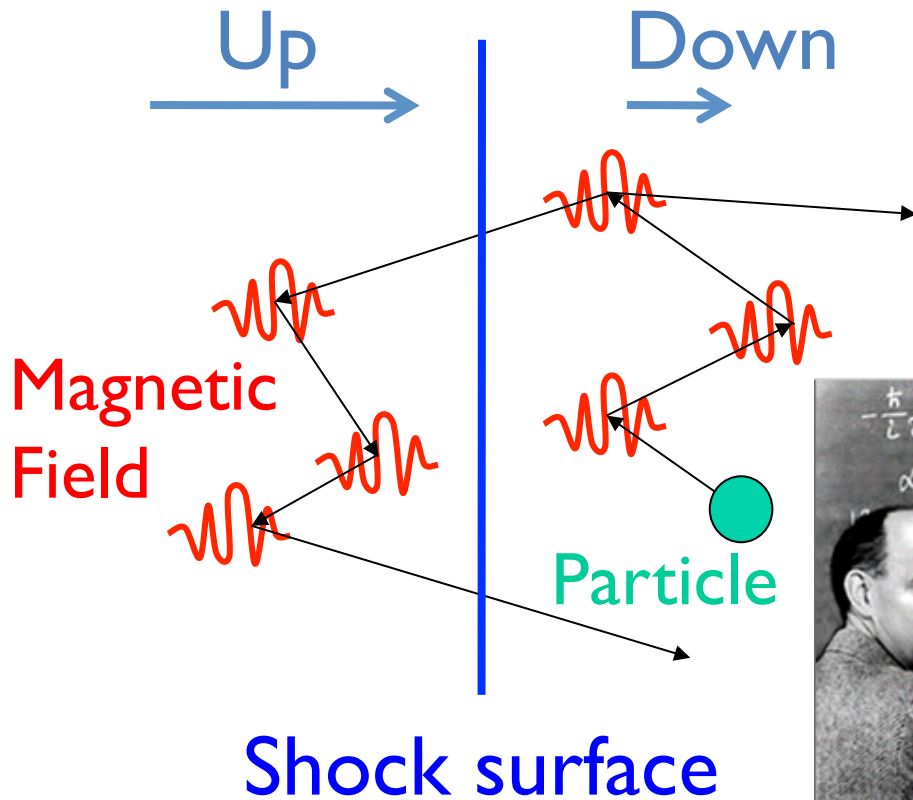
~ keV ~ X-ray

Interstellar Medium

$c_s \sim 10 \text{ km/s}$

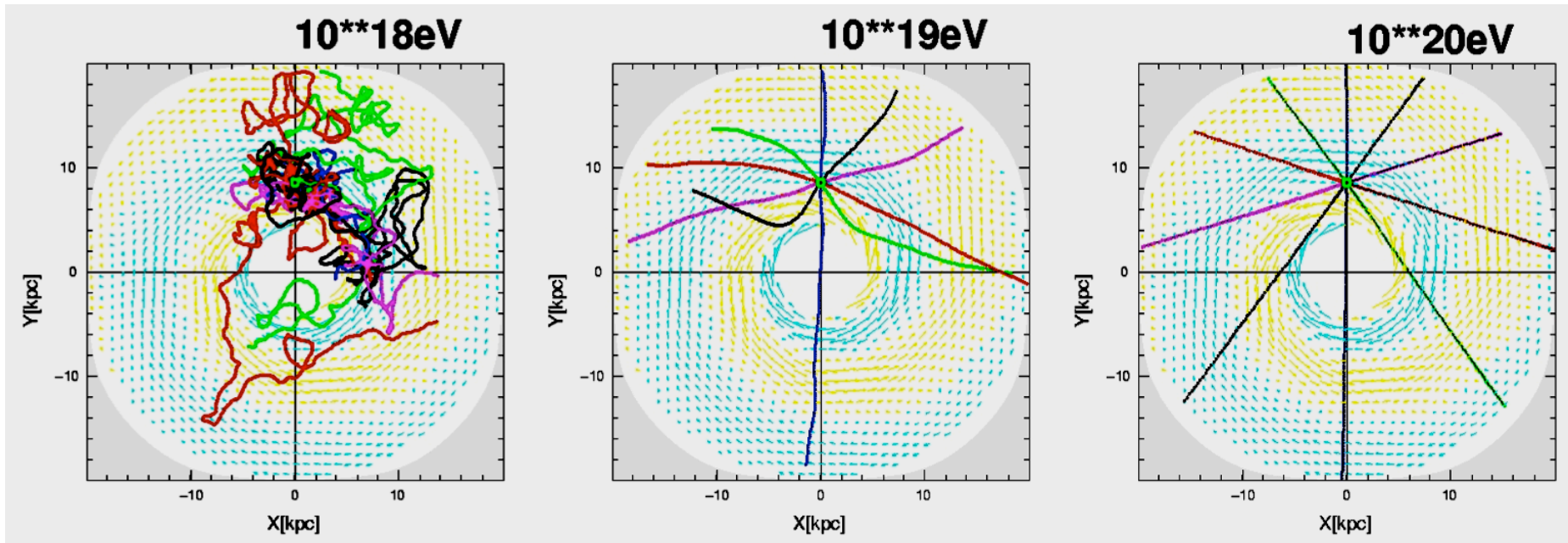
Particle Acceleration

Collisionless Shock



Diffusive Shock (Fermi) Acceleration

Galactic Magnetic Fields

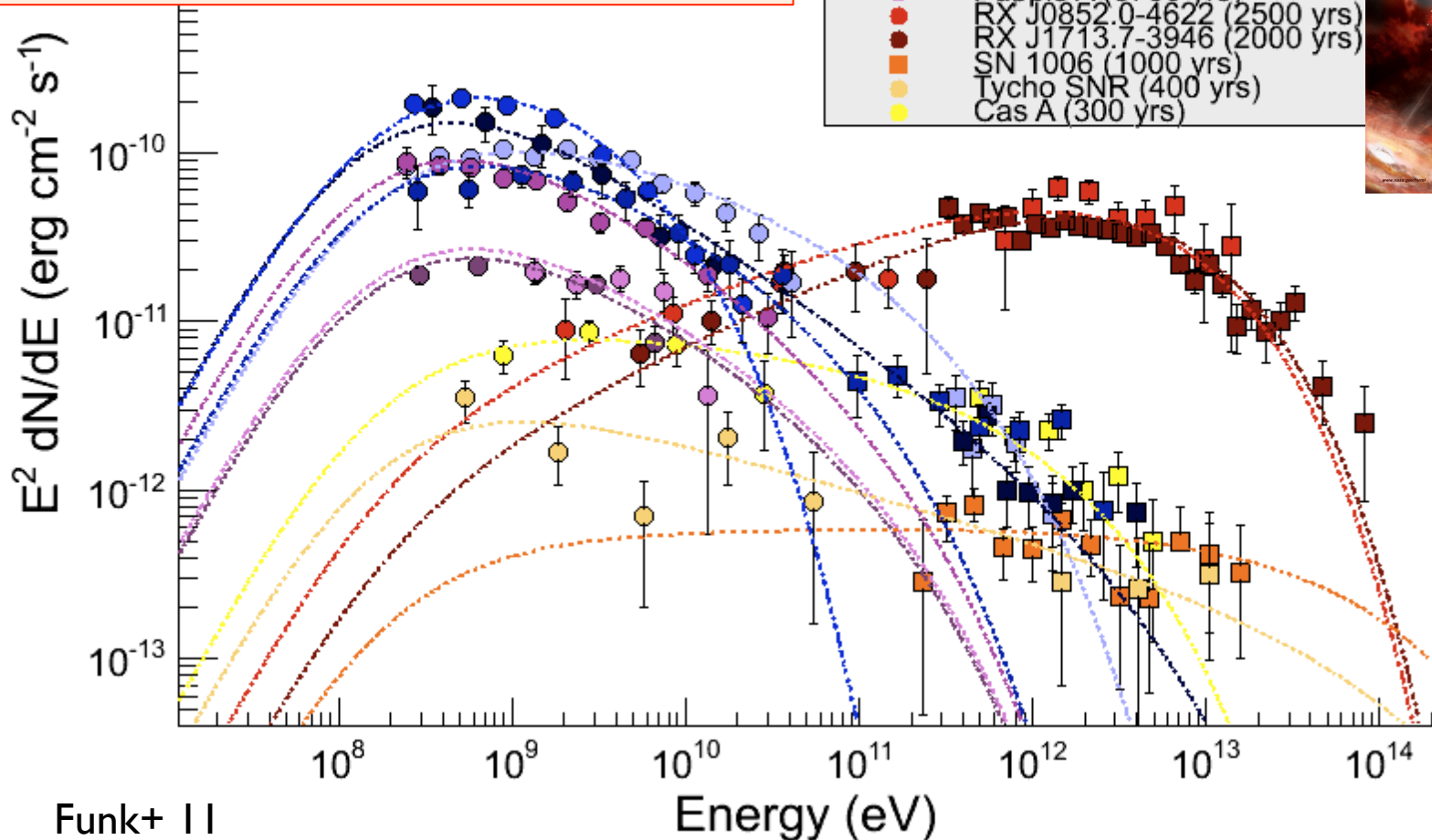


Cosmic Rays below $\sim 10^{18}$ eV are deflected
by Galactic Magnetic Fields $\sim \mu\text{G}$

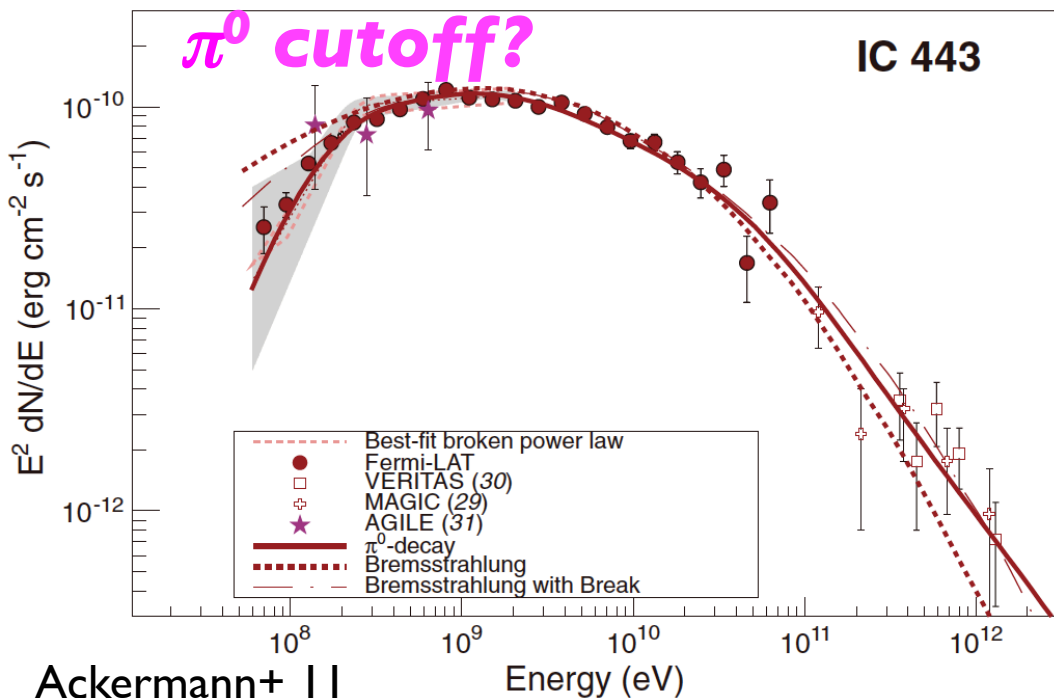
Cosmic-Ray Origin?

Supernova Remnant \Rightarrow CR?

$pp \rightarrow \pi^0 \rightarrow \gamma\gamma$ Detection?



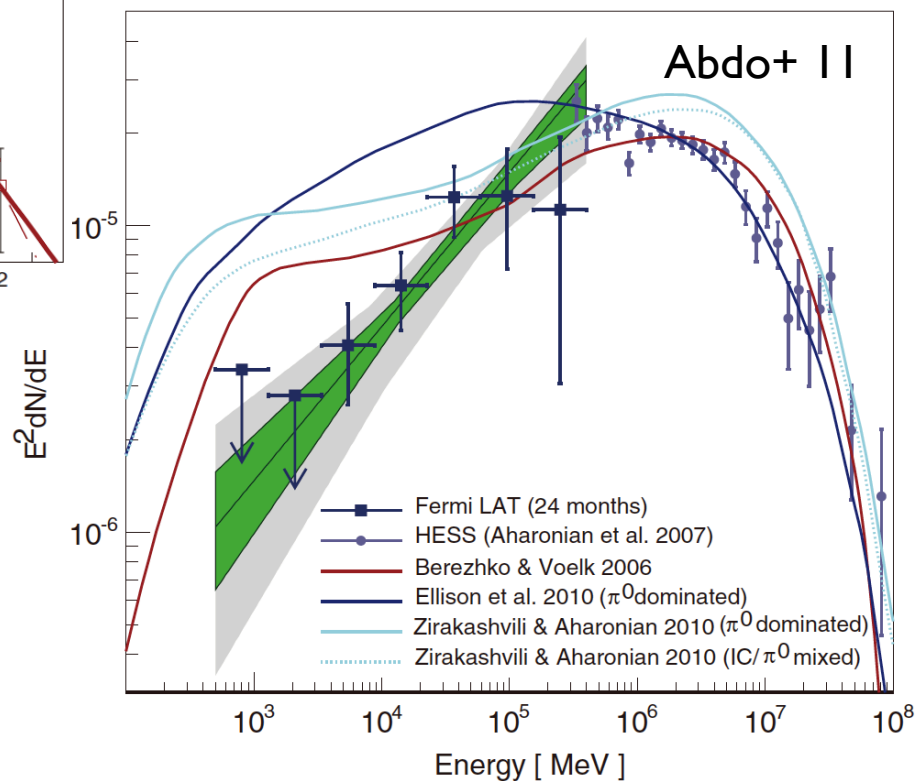
Hadronic or Leptonic?



$$pp \rightarrow \pi^0 \rightarrow \gamma\gamma$$

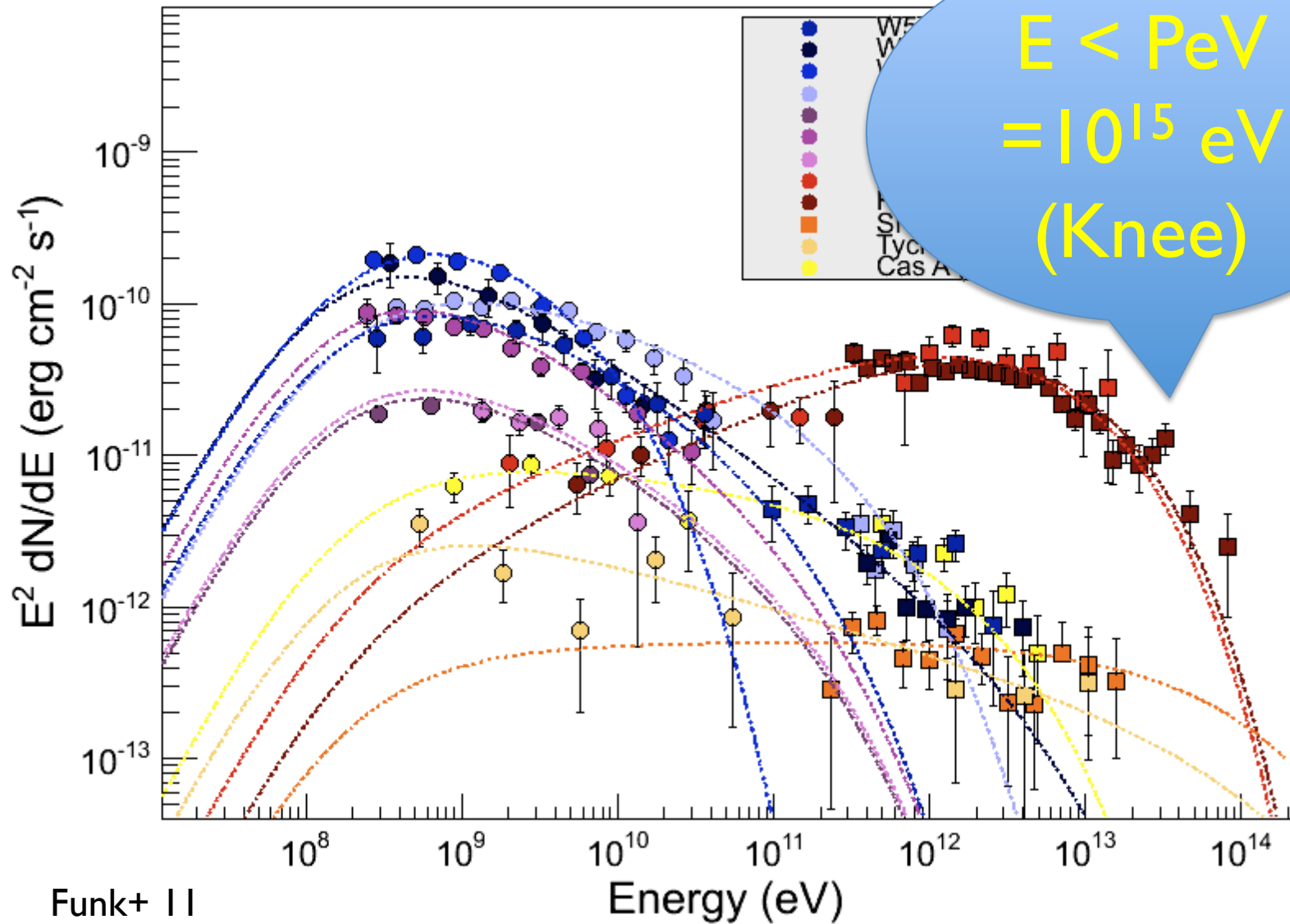
$$e\gamma \rightarrow e\gamma$$

inverse Compton

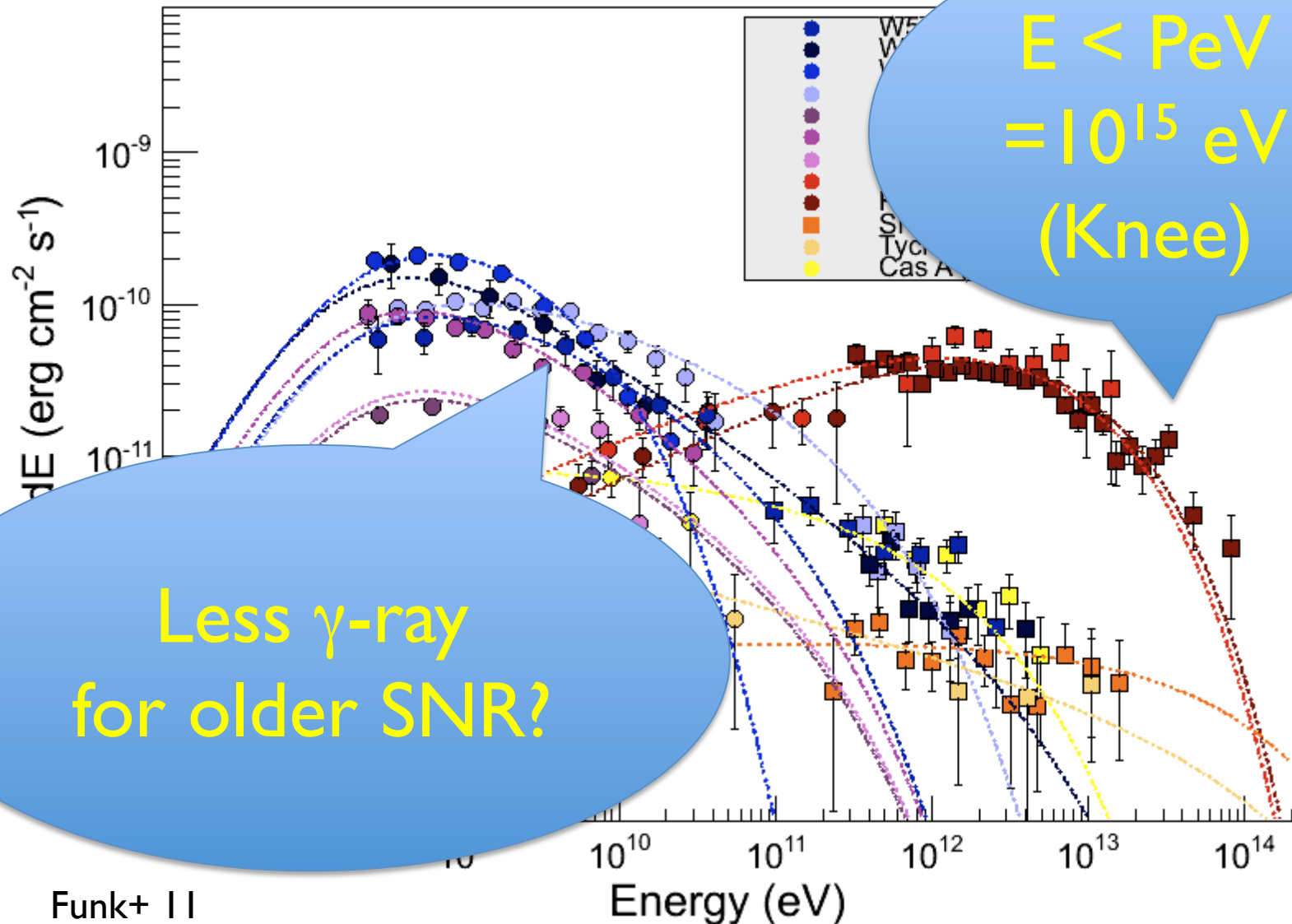


Evidence of CR Origin
Still under discussions

Where is PeVatron?



Where is PeVatron?



66th Fujihara Seminar

X-raying the Gamma-Ray Universe

-- CTA-X-ray LINK Meeting --

4 (Mon) - 6 (Wed) Nov. 2013

Yumoto Fujiya Hotel, Hakone-Yumoto, Kanagawa-ken, Japan

[Top](#)

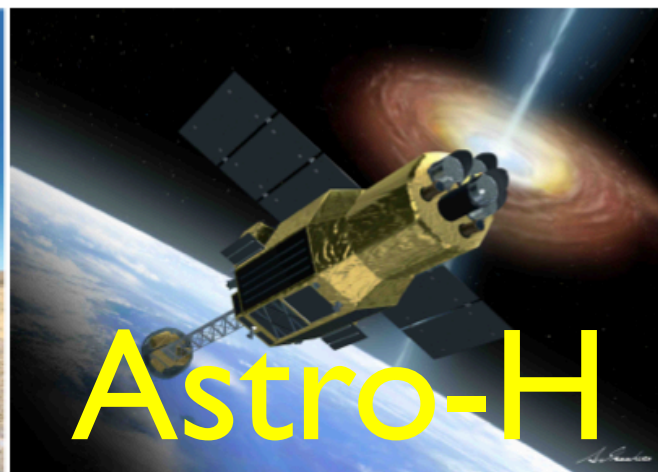
[Overview](#)

[Registration](#)

[Program](#)

[Participants](#)

[Venue](#)



Supported by

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European Union, within the Framework Programme 7: Preparatory Phase of the Cerenkov Telescope Array,

INFRA-2010-2.2.10 / 262053

JSPS Grant-in-Aid for Specially Promoted Research 24000004

KEK

Chair of the organizer: Kunihiro Ioka (KEK)

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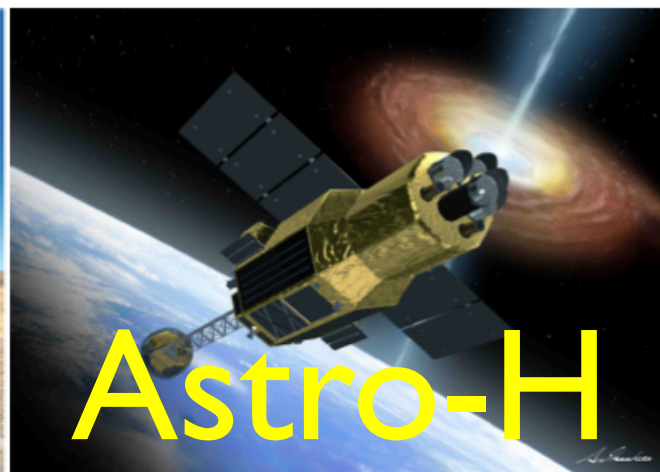
Overview

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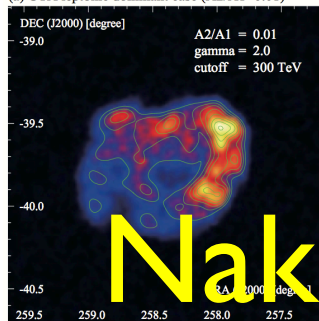


+ NANTEN
Submm radio

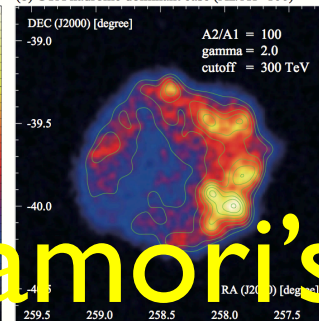
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KEK
Chair of the organizer: Kunihiro Ioka (KEK)

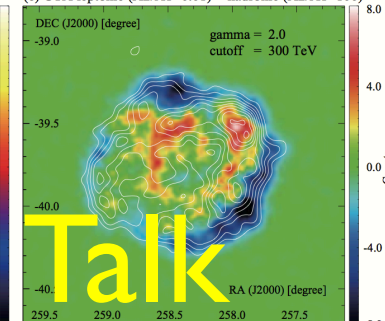
(a) CTA leptonic dominant case ($A2/A1=0.01$)



(b) CTA hadronic dominant case ($A2/A1=100$)



(c) CTA leptonic ($A2/A1=0.01$) - hadronic ($A2/A1=100$)



Nakamori's Talk

The Fujihara Foundation of Science

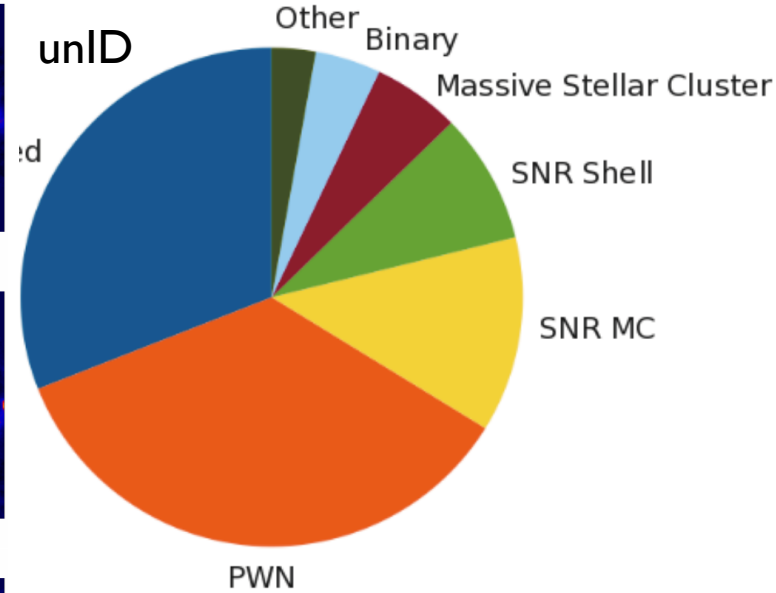
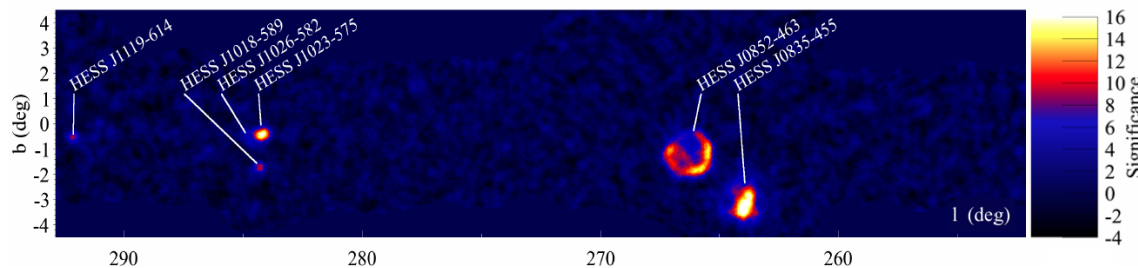
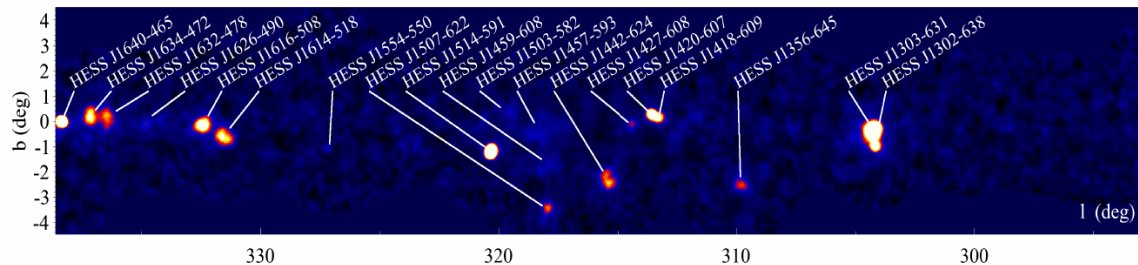
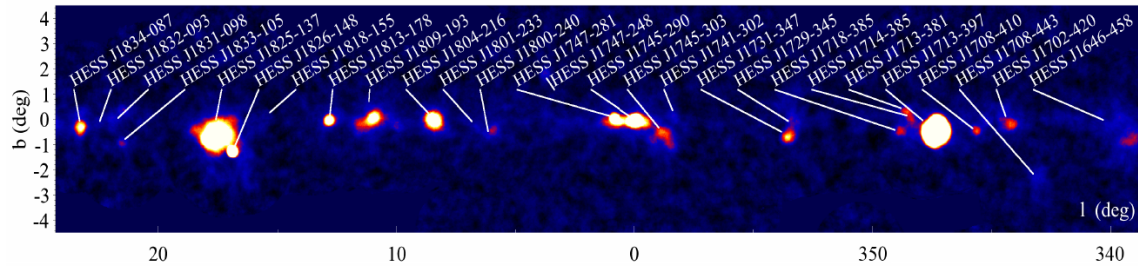
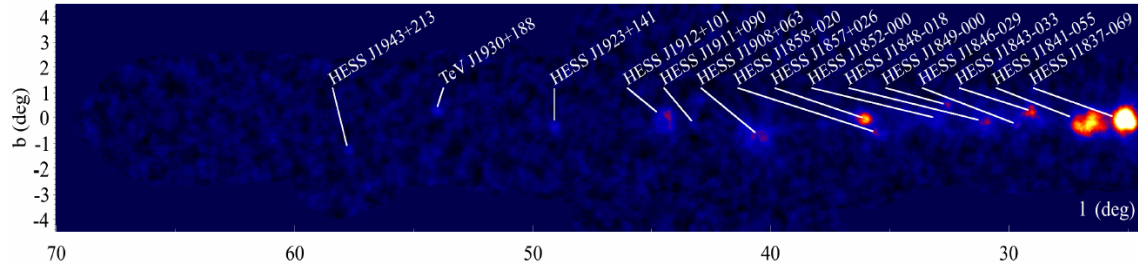
藤原科学財団



HIGH ENERGY ACCELERATOR RESEARCH ORGANIZATION

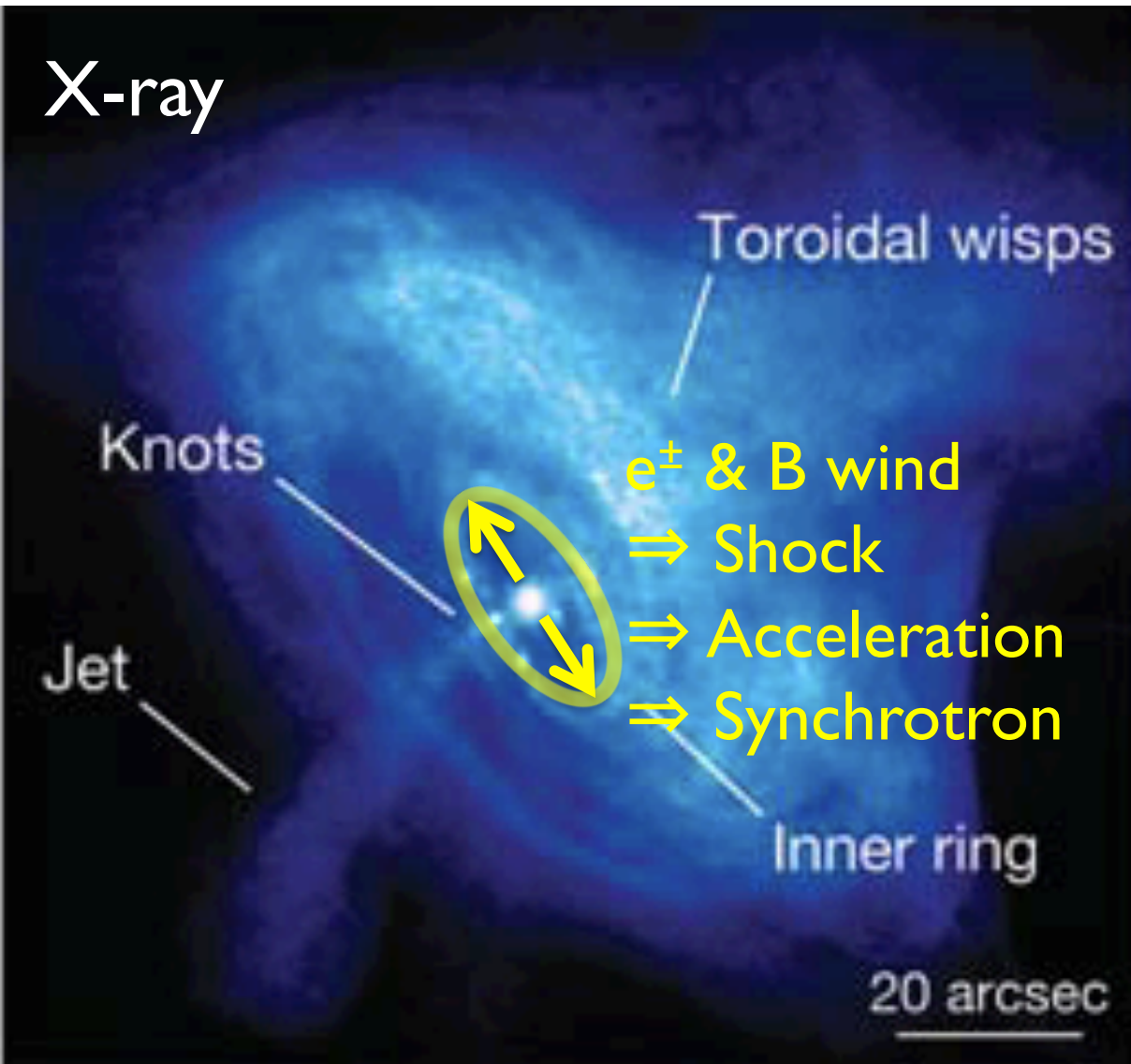
TeV Gamma-Ray Sky

I 307.4690



**Pulsar
Wind Nebula
dominates
TeV γ -ray sky**

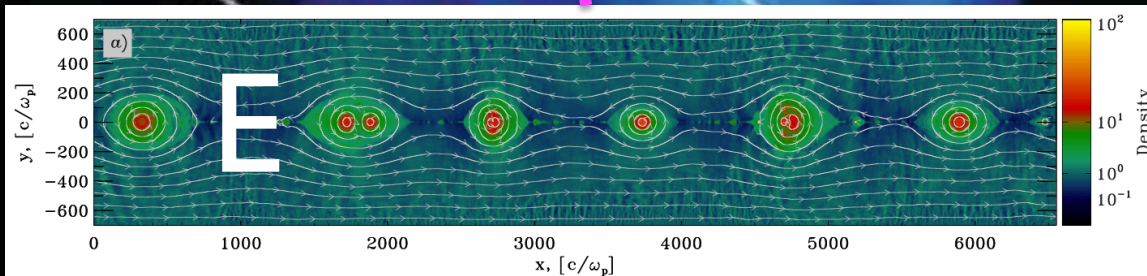
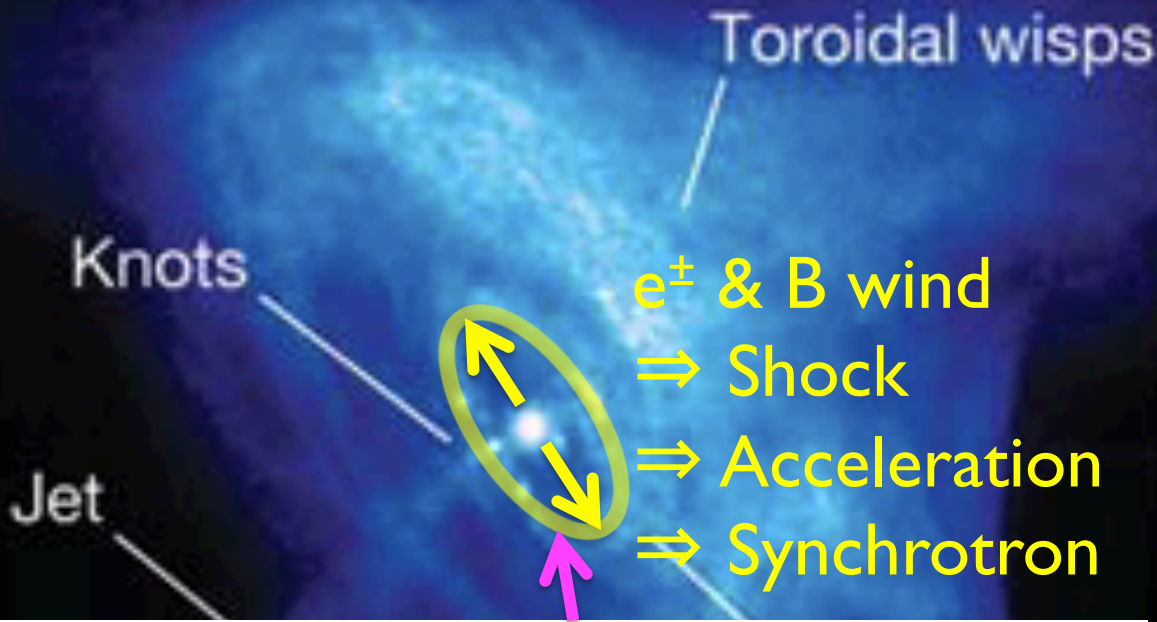
Pulsar Wind Nebula



- Neutron star
Rotation + B
- Crab nebula
- $r_{\text{shock}} \sim 0.1 \text{ pc}$
- $r_{\text{knot}} \sim 0.01 \text{ pc}$
- **~ 4 days flares**
 \Rightarrow **PeV electron**

Pulsar Wind Nebula

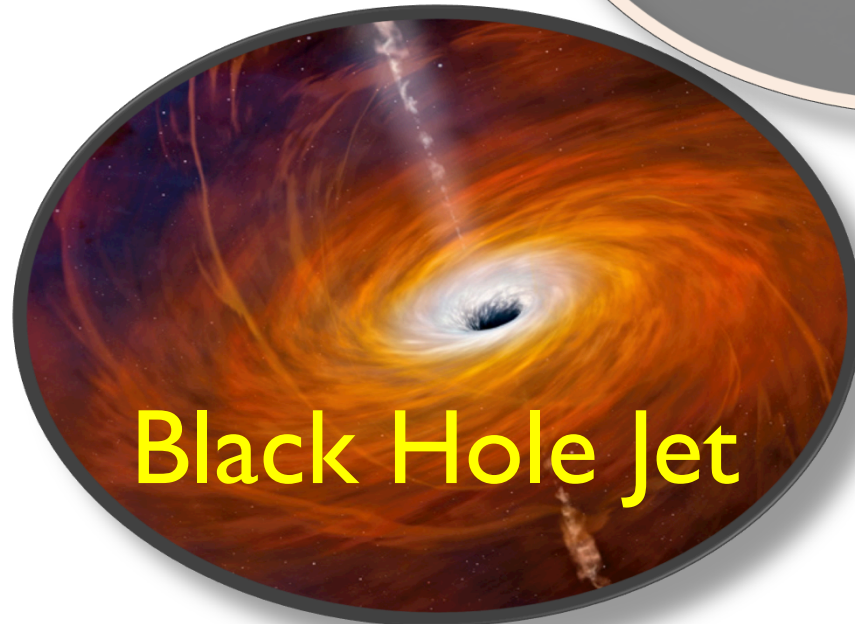
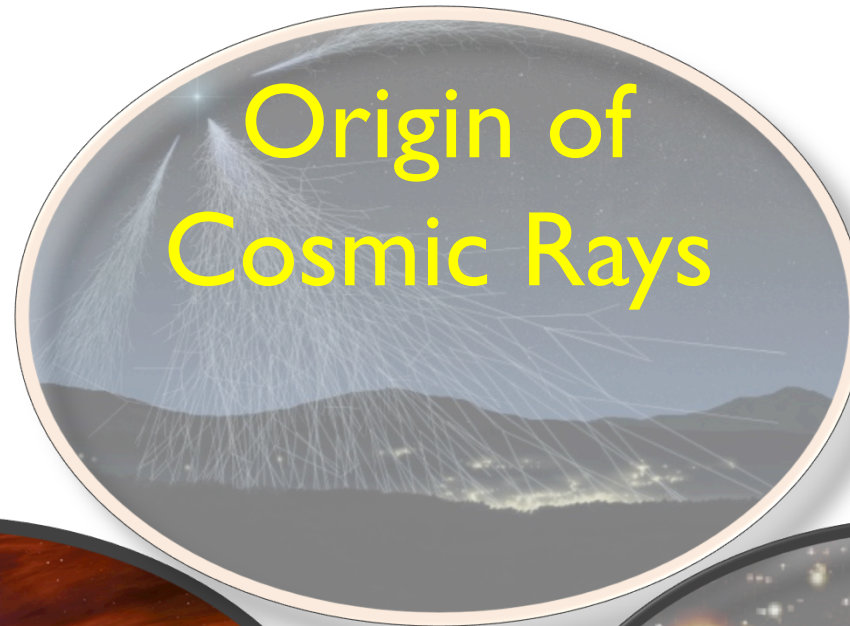
X-ray



- Neutron star Rotation + B
- Crab nebula
- $r_{\text{shock}} \sim 0.1 \text{ pc}$
- $r_{\text{knot}} \sim 0.01 \text{ pc}$
- **$\sim 4 \text{ days flares}$**
- ⇒ **PeV electron**

Linear accelerator?

Mysteries in 21th Century

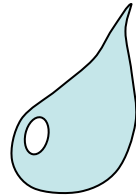


Gamma-Ray Burst

$E=mc^2$ (by Einstein)

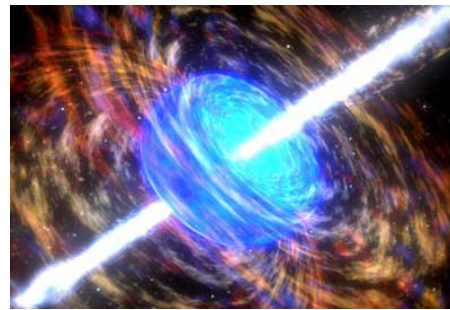


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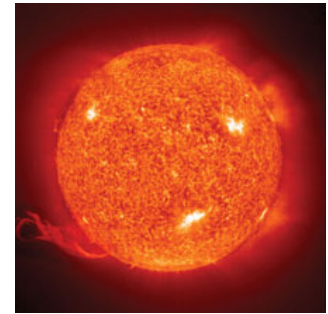


Atomic bomb

$\sim 1\text{kg}$



=



GRB

$\sim 10^{52}\text{erg}$

Sun

$\sim 10^{33}\text{g}$

In \sim sec, GRB release energy Sun emit over lifetime

GRB is the most luminous object

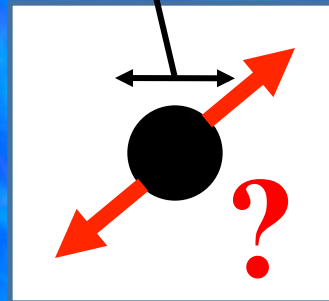
GRB \approx Black Hole Formation?

GRB jet breaks out of massive stellar envelope

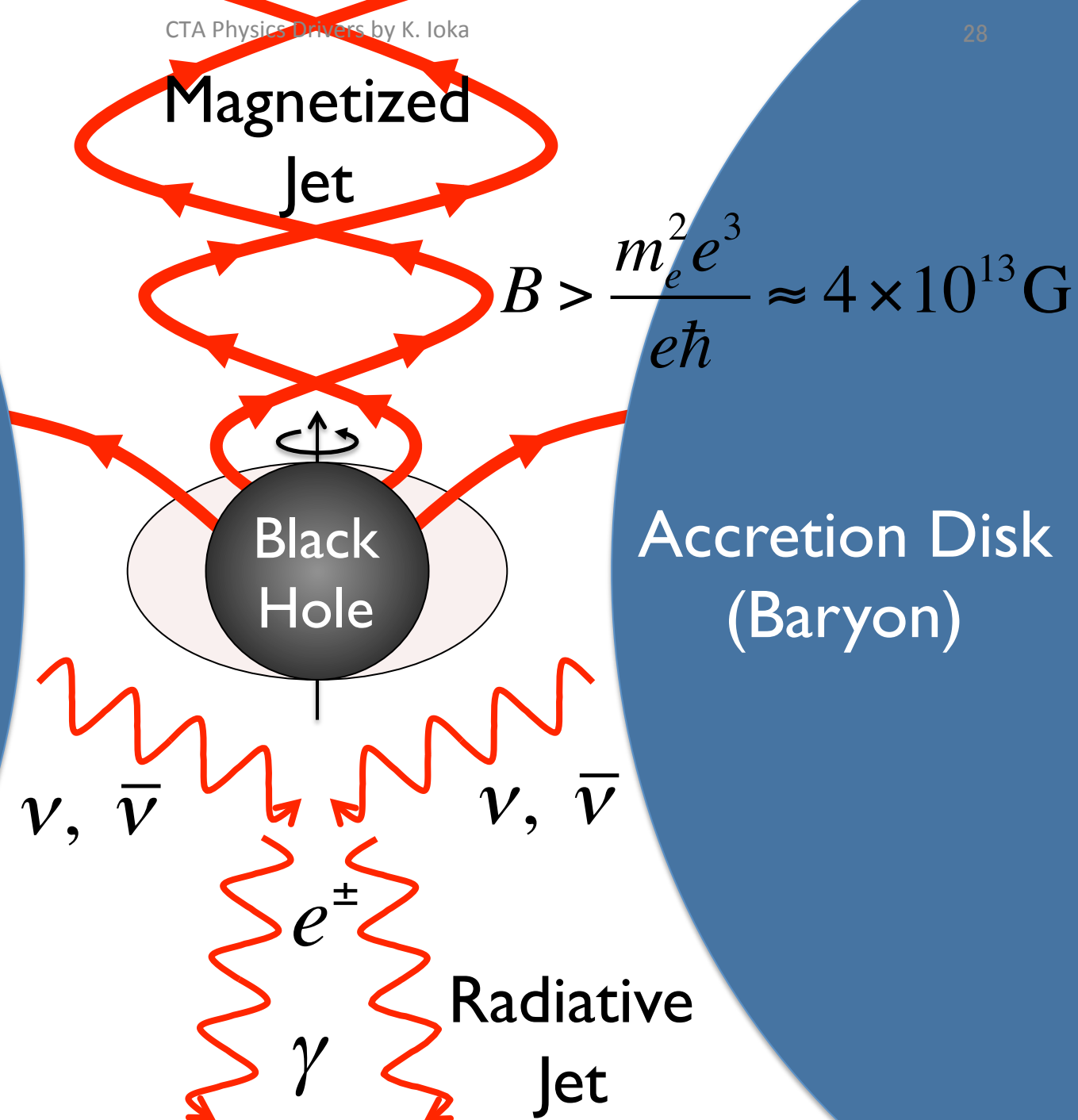
msec \times $c \sim 300\text{km} \sim$ BH radius

Collapsar

$v > 0.9999 \times c$

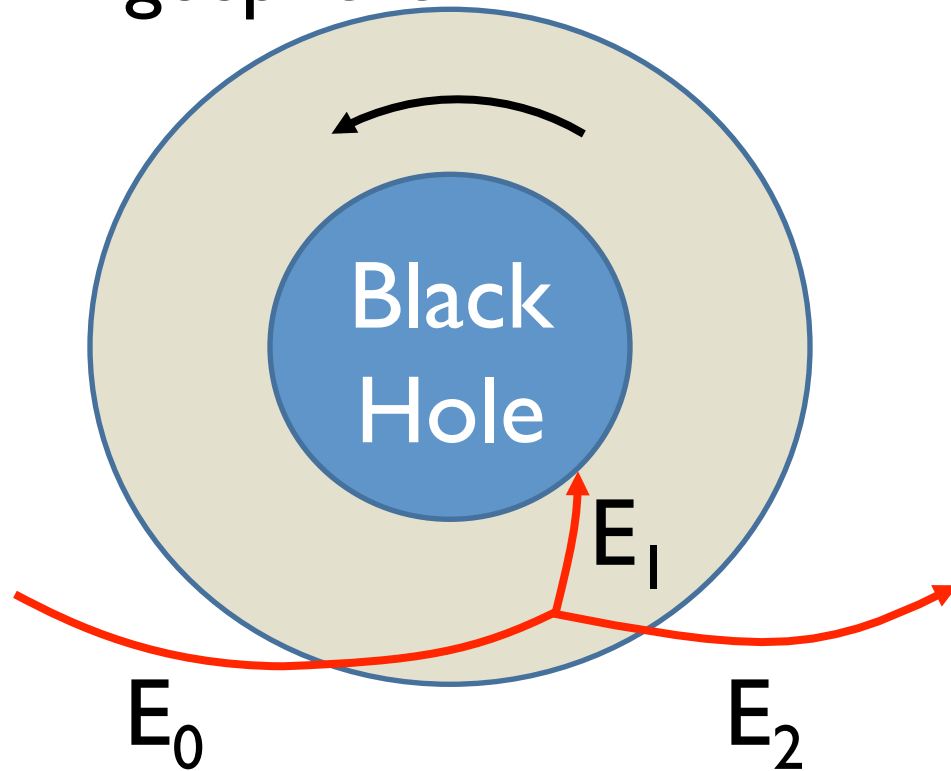


Why BH launches relativistic jet?



Penrose Process

Ergosphere



$E_1 < 0$ is possible

since $g_{tt} < 0$

in Ergosphere

$$E_0 = E_1 + E_2$$

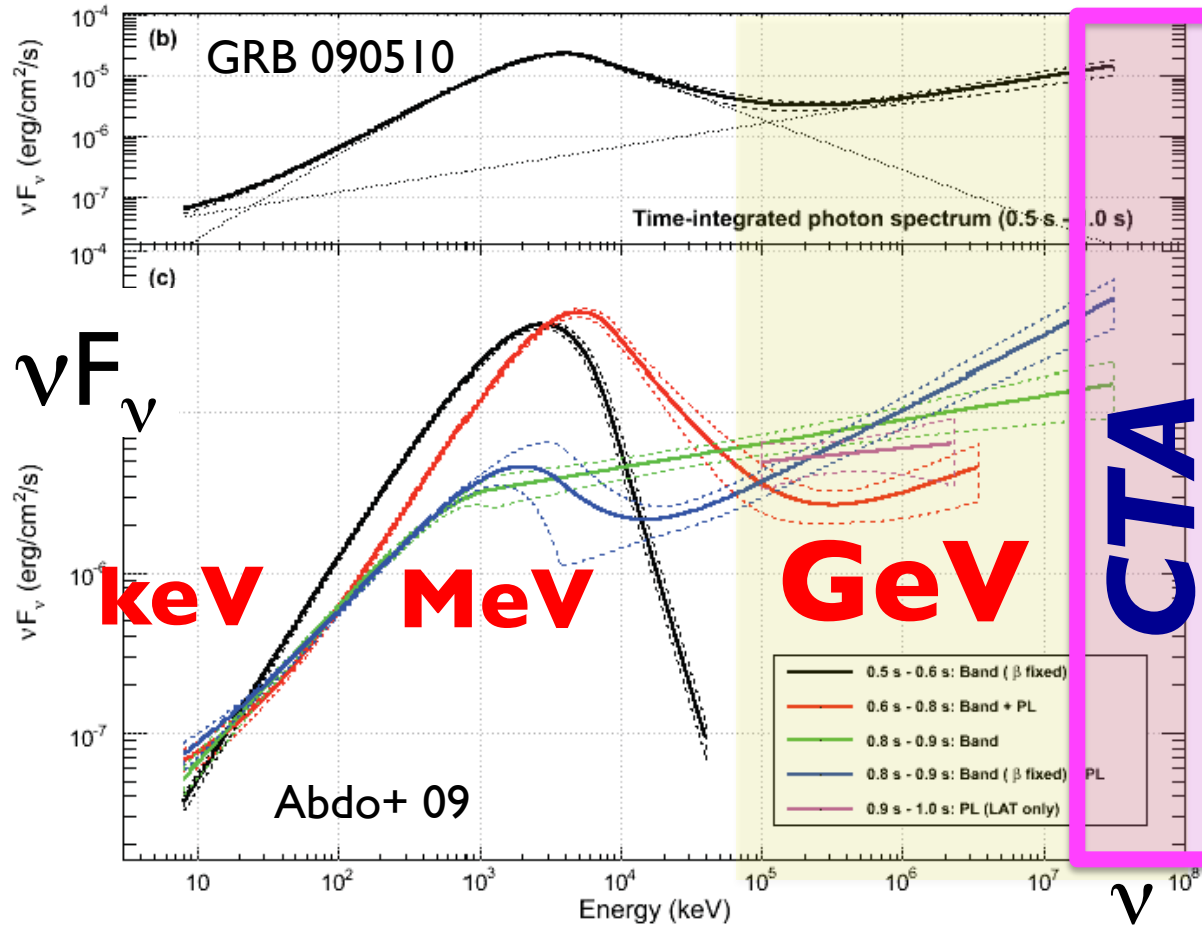
$$\Rightarrow E_2 > E_0$$

Penrose 69; Blandford & Znajek 77

We can extract energy from BH!!

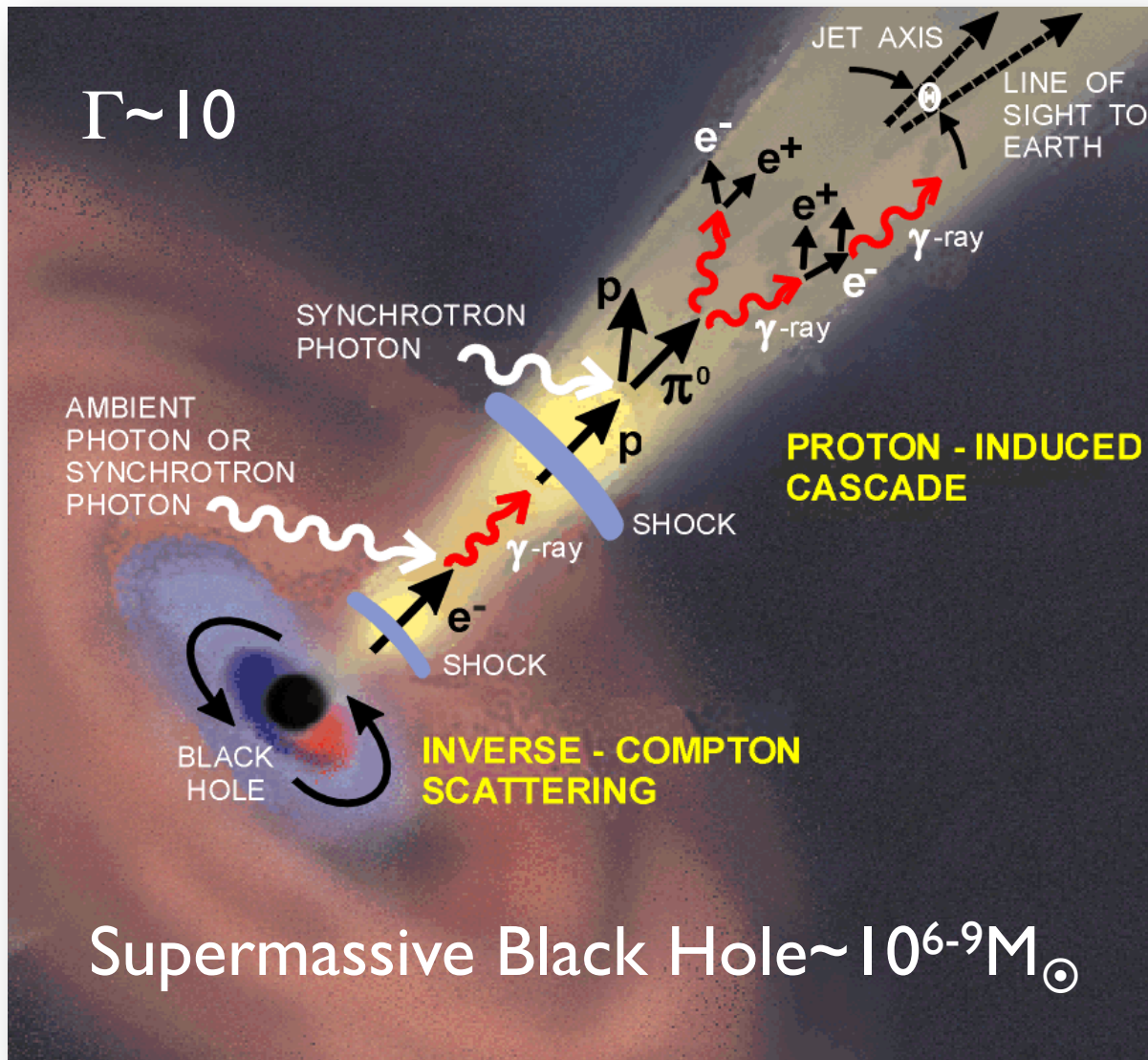
Max $\sim 29\% Mc^2$ (if $a=M$)

> GeV γ -ray from GRBs



- $\Gamma > 1000$ to avoid $\gamma\gamma \rightarrow e^+e^-$
- $> 10^{17} \text{ eV}$ $p + \gamma \rightarrow \pi \rightarrow \gamma, e?$

Active Galactic Nuclei

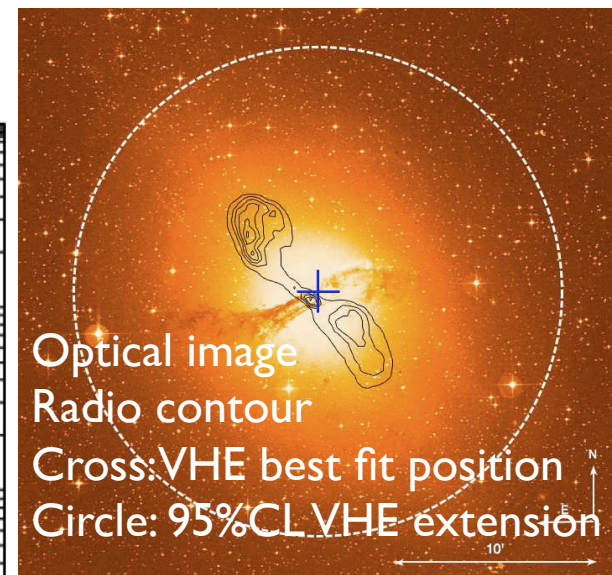
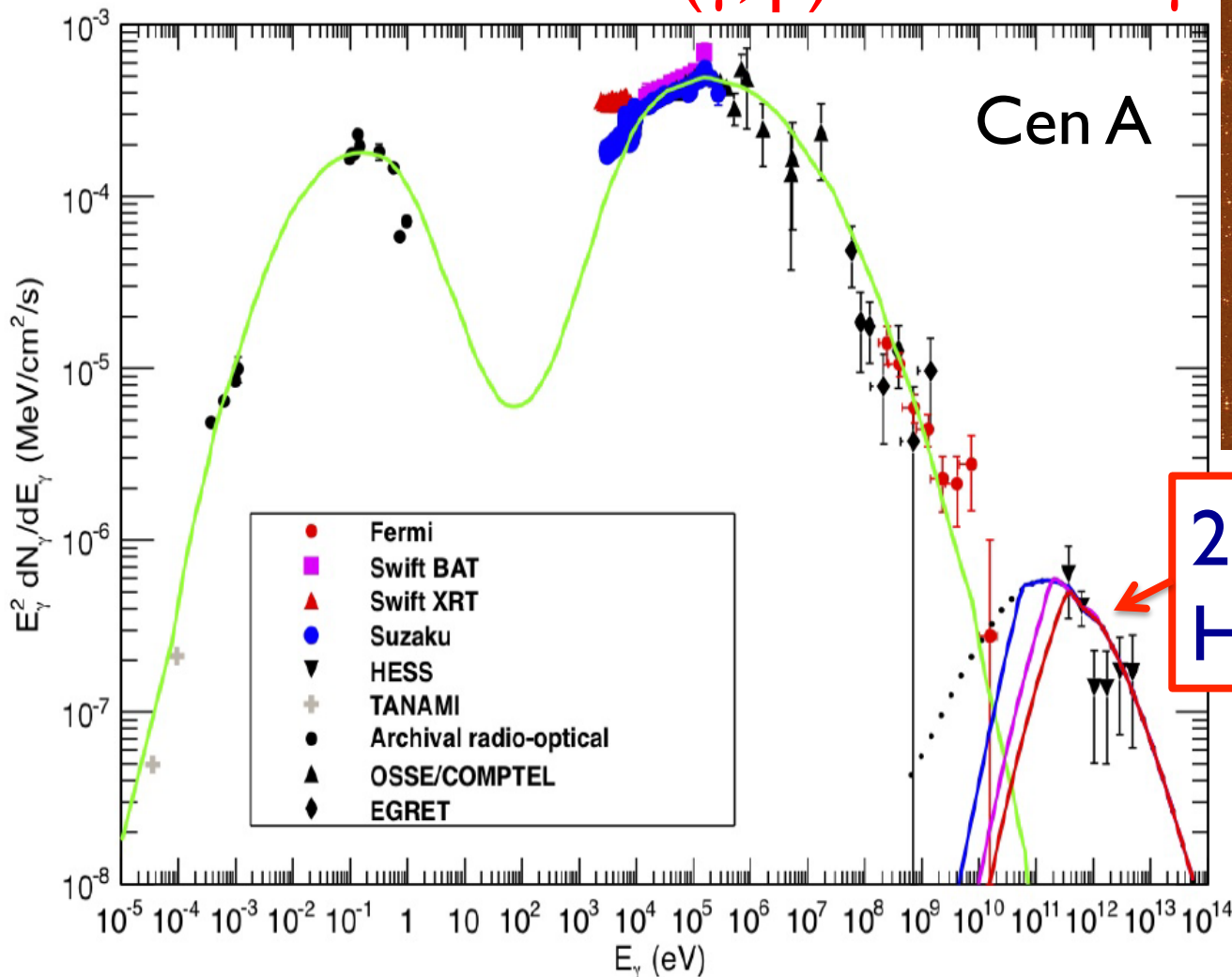


Blazar

Radio galaxy
(BLR, NLR)
(FR I, II)

Off-Axis AGN

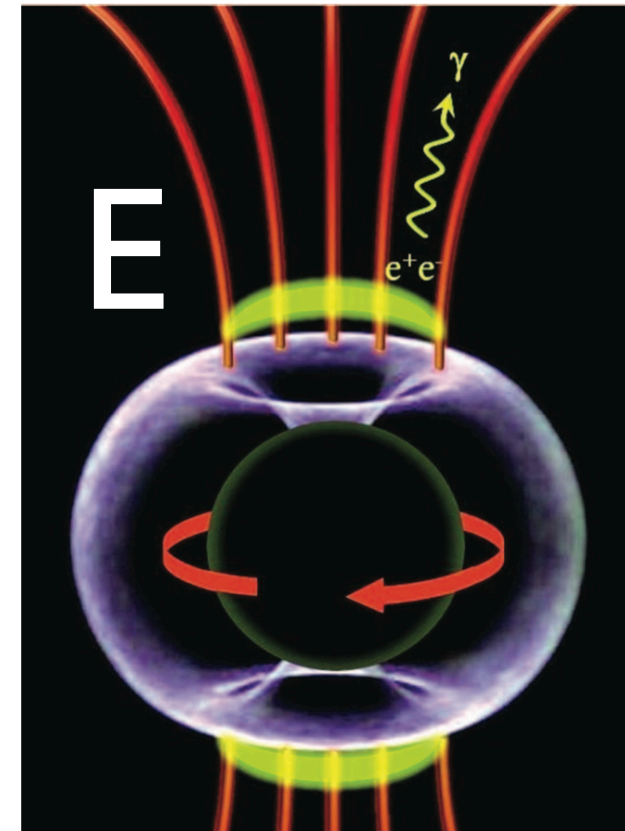
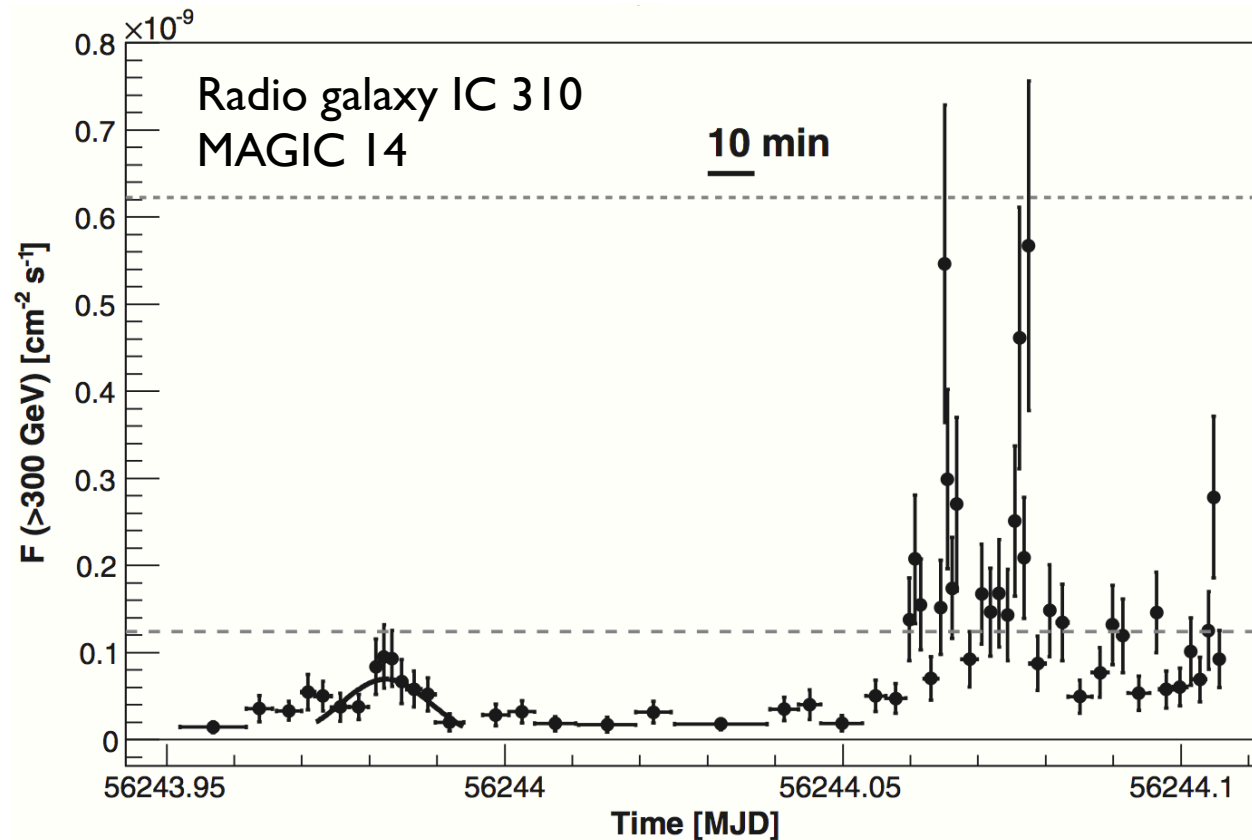
$$\text{UHECR} + (\gamma, p) \rightarrow \pi \rightarrow \text{HE } \gamma$$



**2nd Component
Hadronic origin?**

Aharonian+ 09
Abdo+ 10
Kachelrieß+ 09
Sahu+ 12

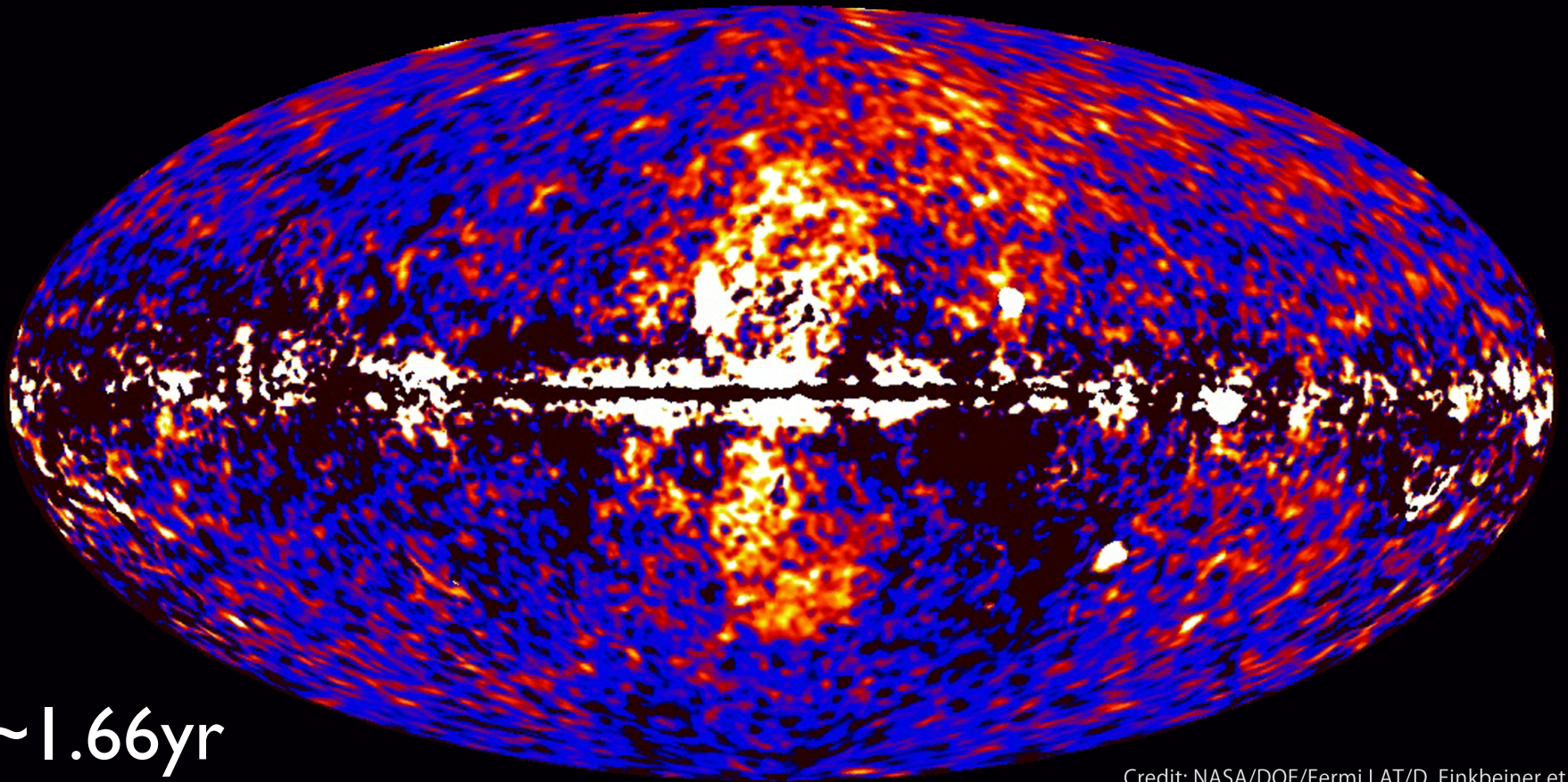
Fast Variability



$\Delta t \sim 4.8 \text{ min} \sim 20\% r_g \text{ (Sub-horizon)}/c?$

γ -ray Bubble Sky Image

Fermi data reveal giant gamma-ray bubbles

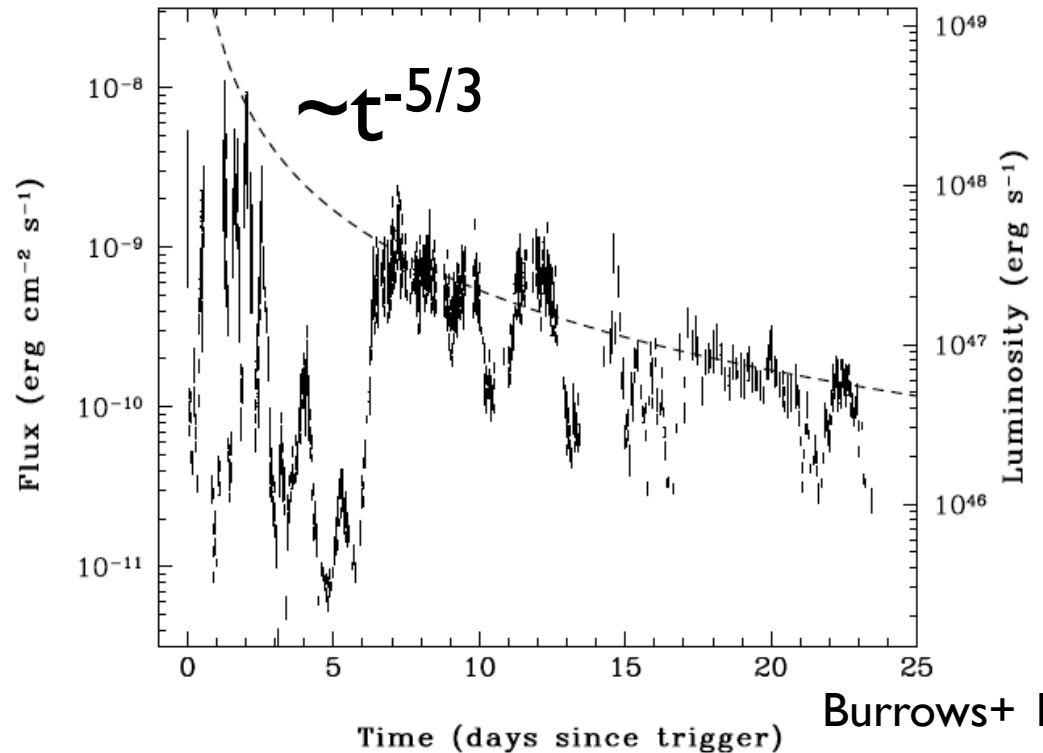
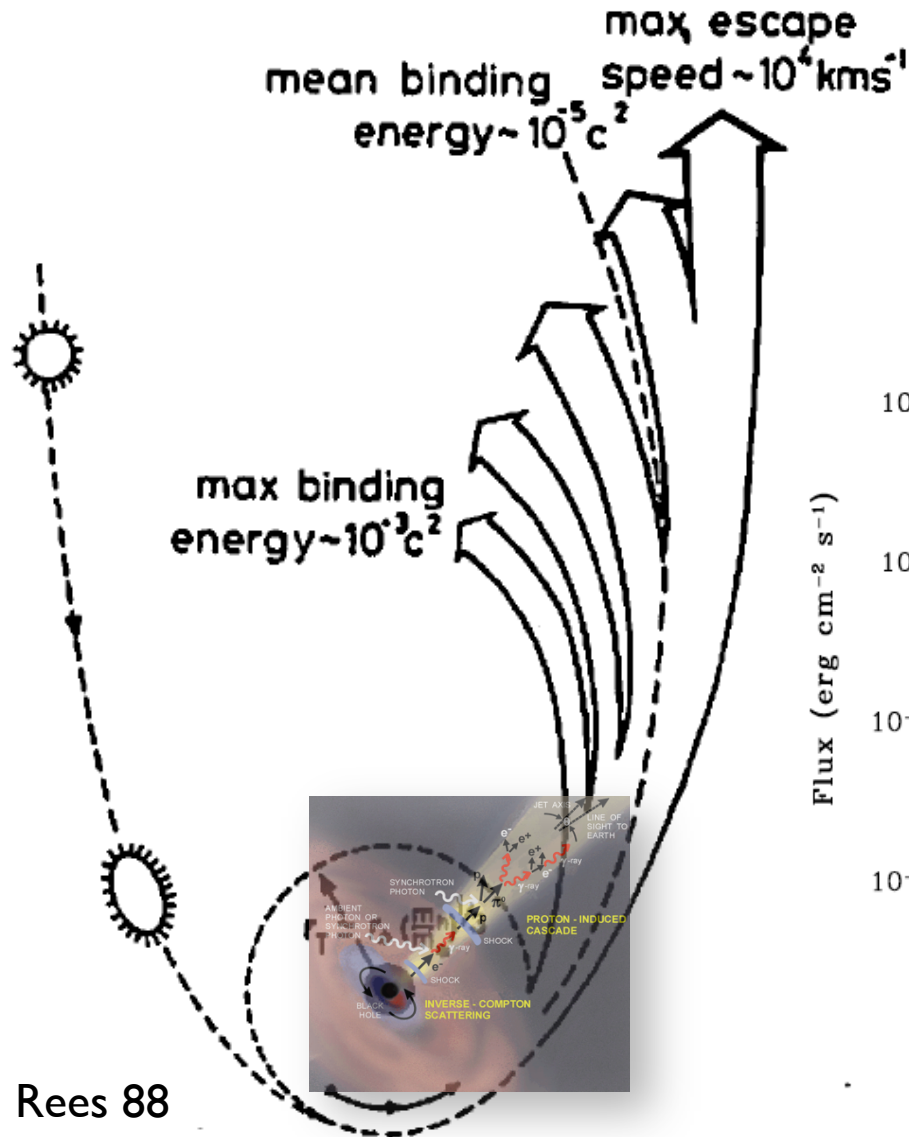


Credit: NASA/DOE/Fermi LAT/D. Finkbeiner et al.

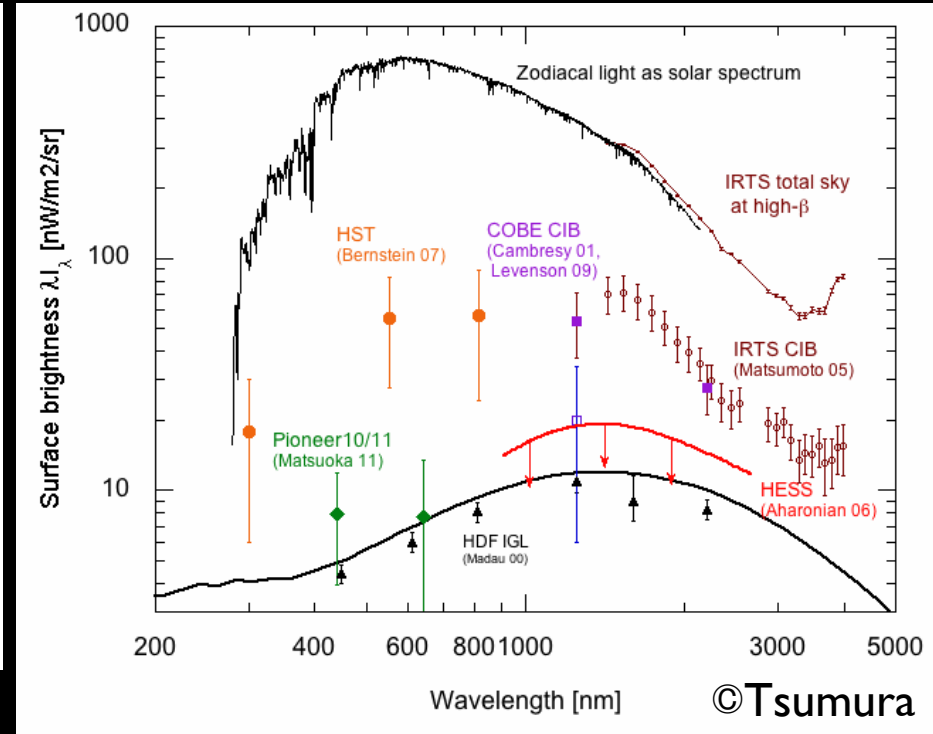
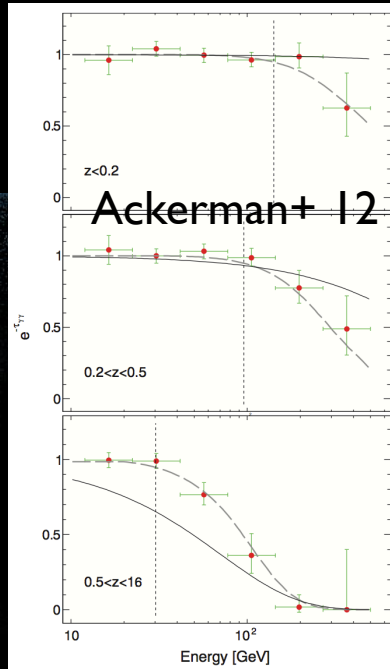
Data subtraction reveals the gamma-ray bubbles

Tidal Disruption Jet

Supermassive BH
disrupts a star \Rightarrow Jet

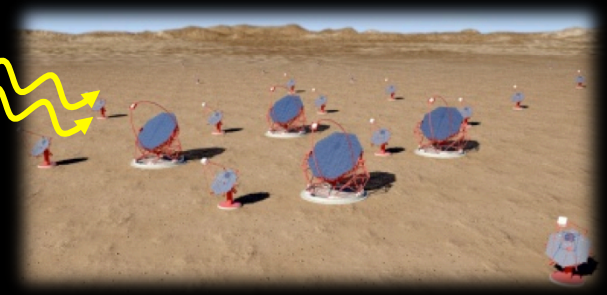
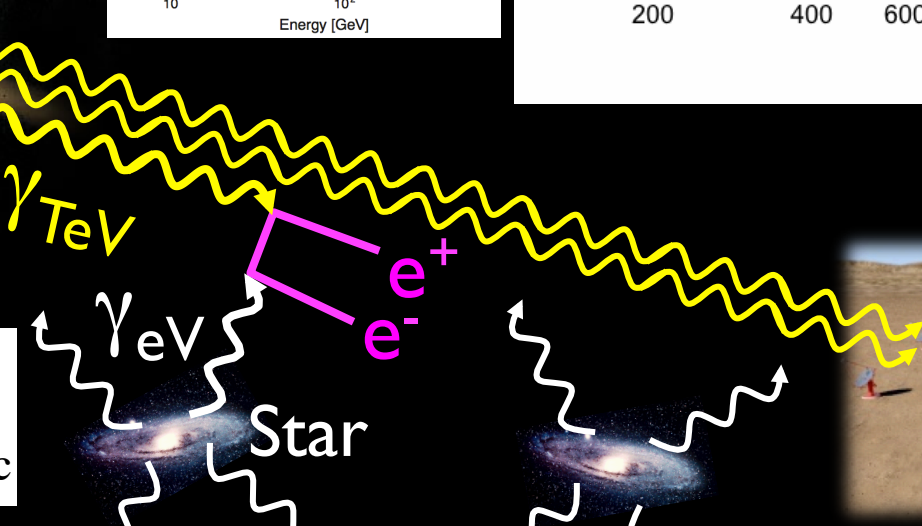


Cosmic Infrared Background



$$eV \times TeV \sim (m_e c^2)^2$$

$$l_{\gamma\gamma} \sim 1/\sigma_T n_{IR} \sim 10 n_{IR}^{-1} \text{Mpc}$$

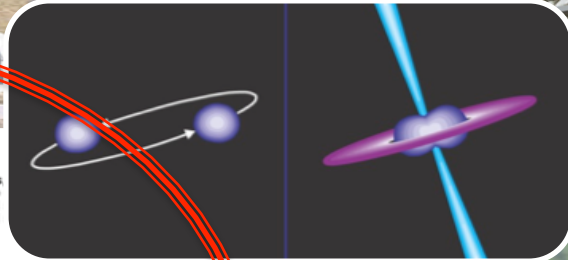
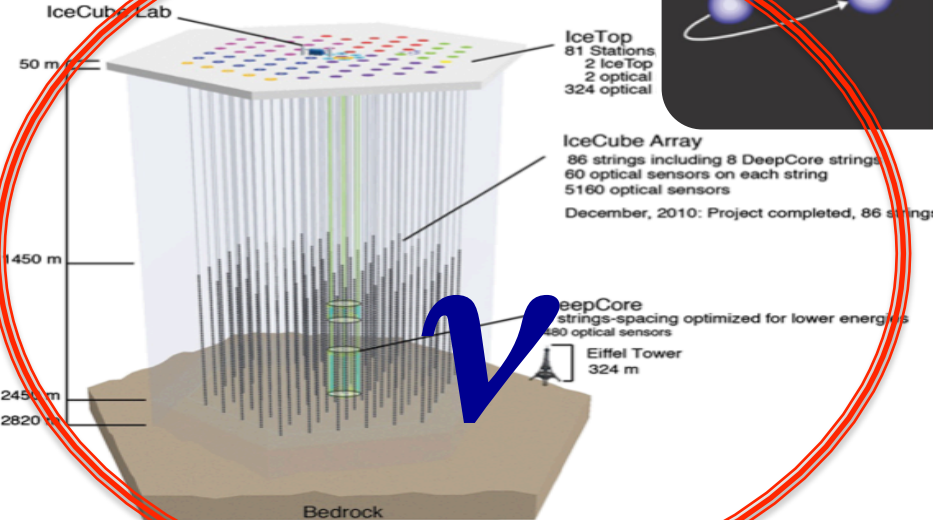
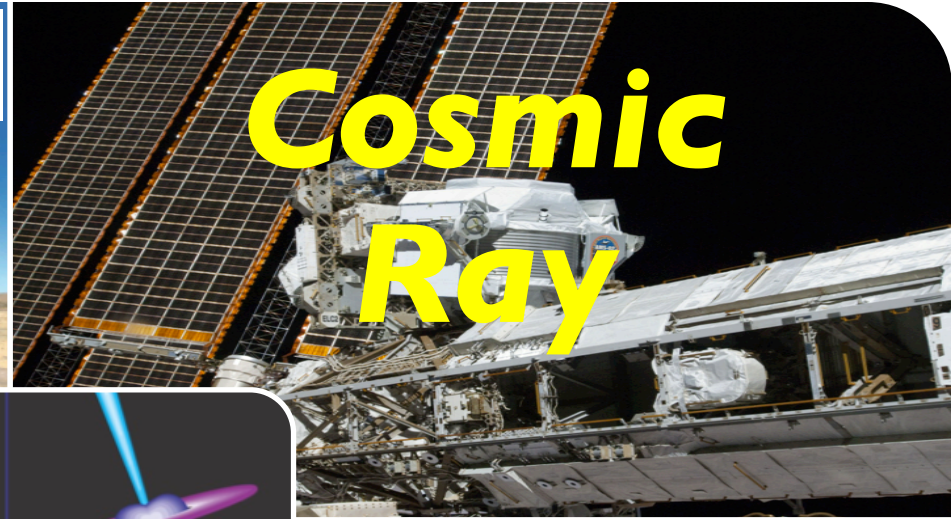


Multi-Messenger

Photon



Cosmic Ray

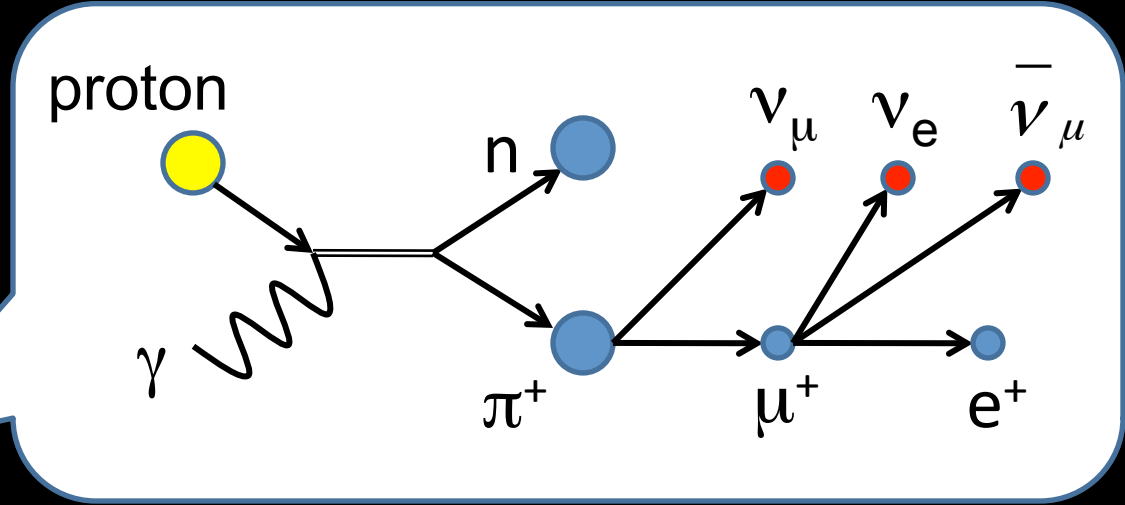
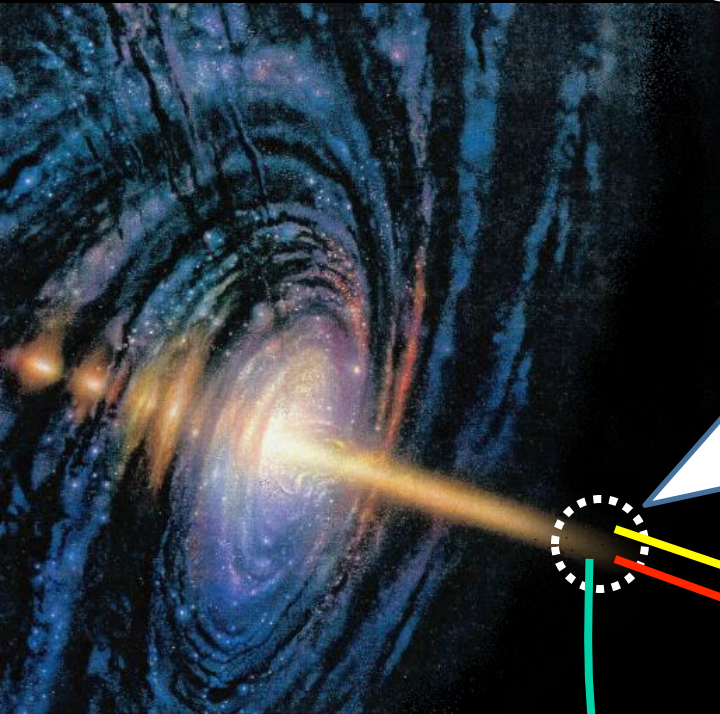


Gravitational Wave



21st Century: Multi-Messenger Era

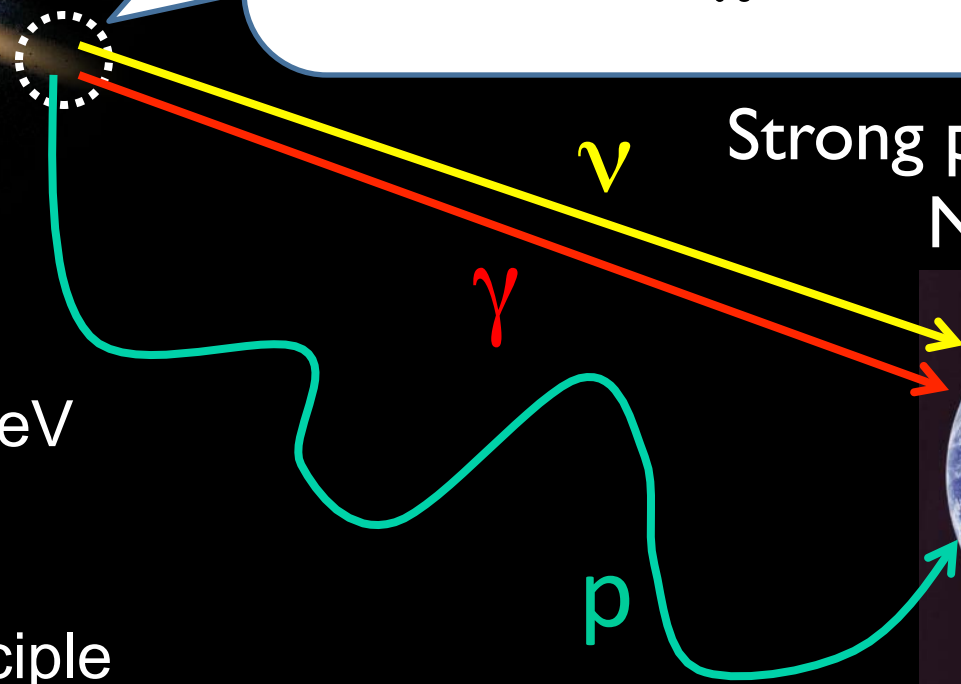
Cosmic ν



Strong penetration
Not delayed



- CR origin
- ν interaction @ $> \text{TeV}$
- τ appearance
- Limiting ν speed
- Equivalence principle



First 2 PeV ν s

PeV = 10^{15} eV

"Bert"

8 Aug 2011

3 Jan 2012

"Ernie"

**Breakthrough of
the year 2013**

1.04 ± 0.16 PeV



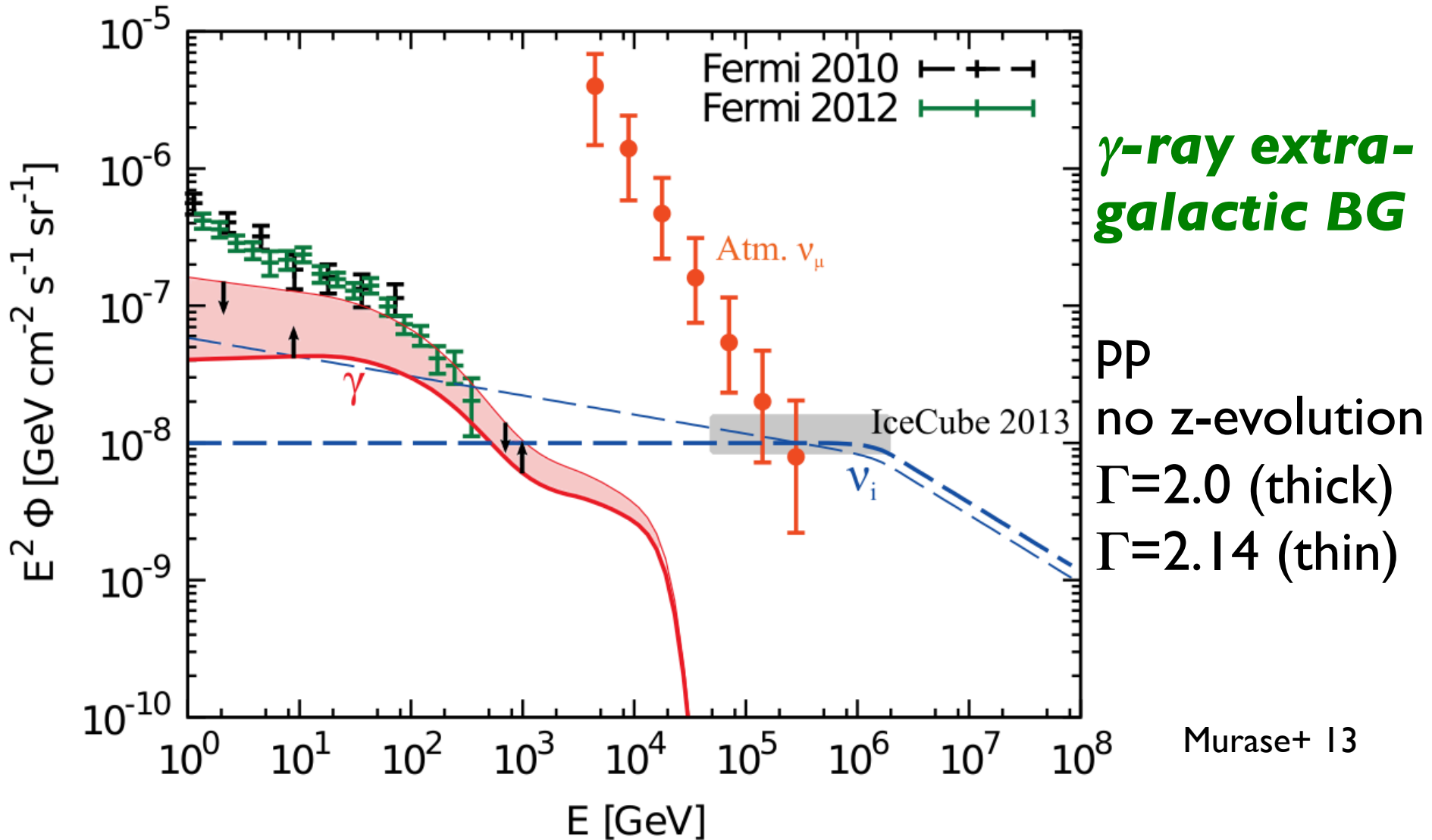
1.14 ± 0.17 PeV

Reported in Kyoto ν 2012

Aartsen+(IceCube), arXiv:1304.5356

Dawn of High-Energy ν Astronomy!!!

Strong γ -Ray Limits

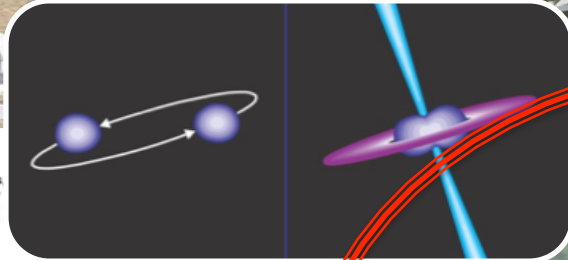
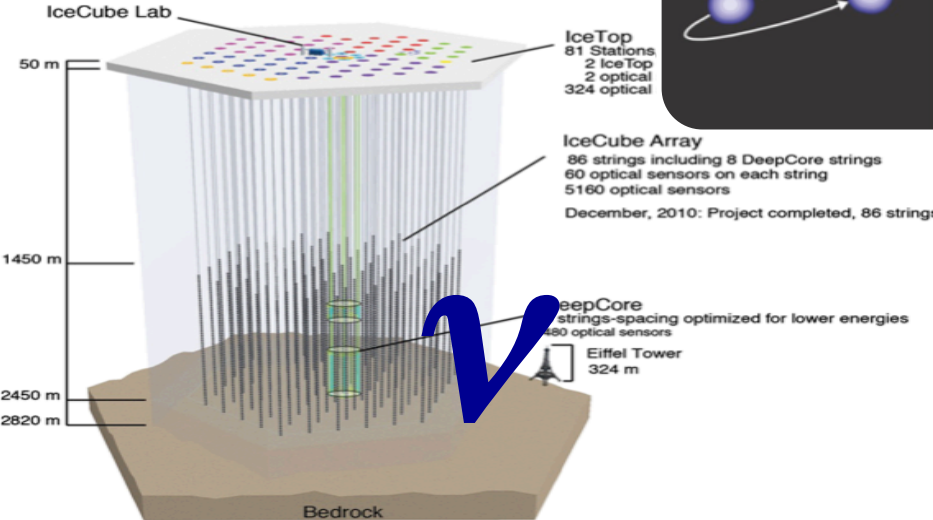
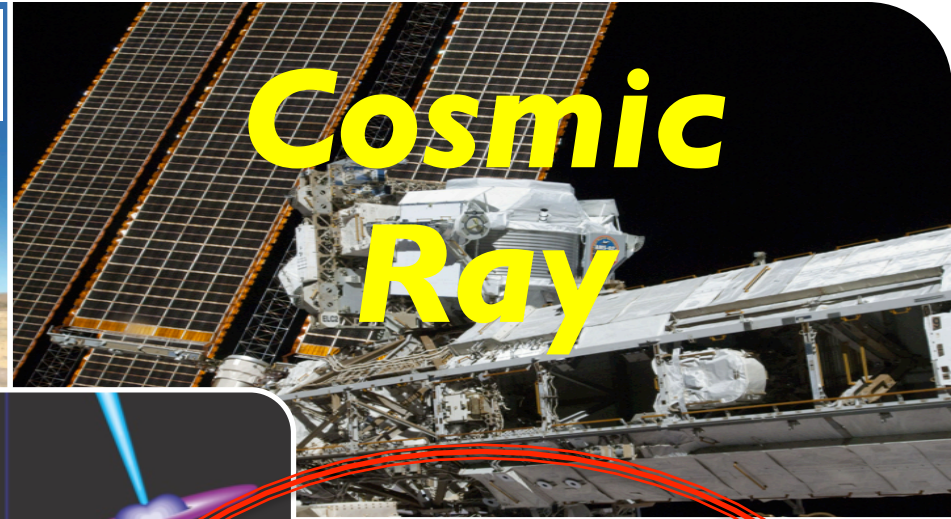


Multi-Messenger

Photon



Cosmic Ray



Gravitational Wave



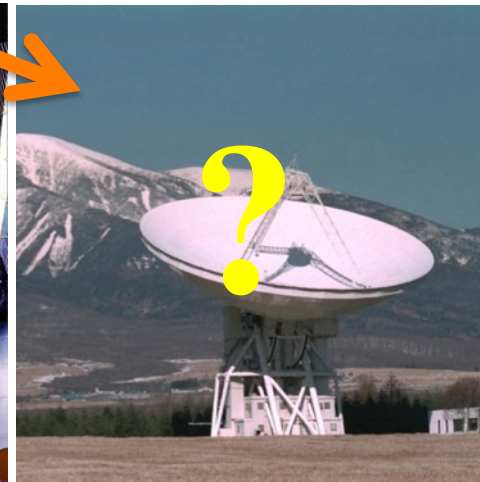
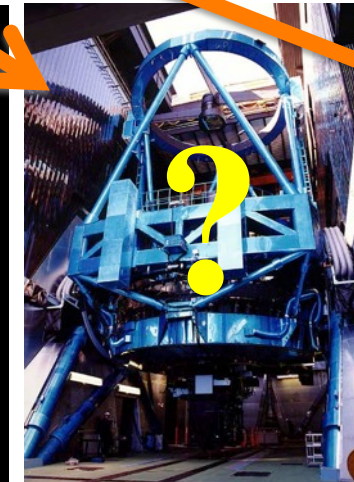
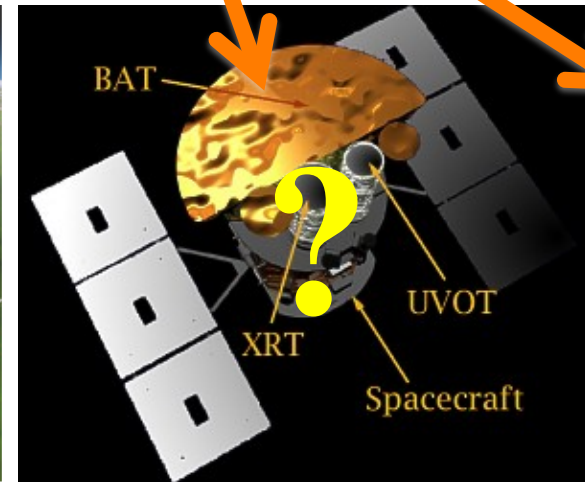
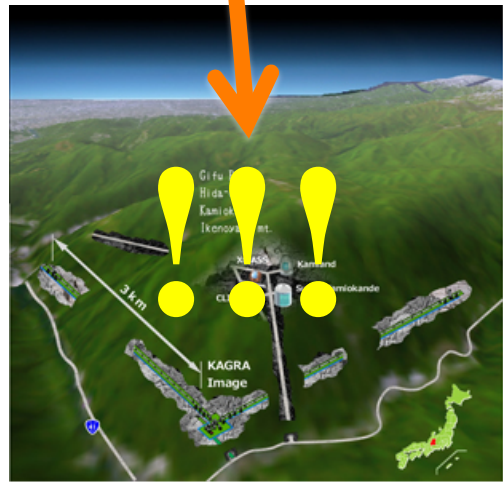
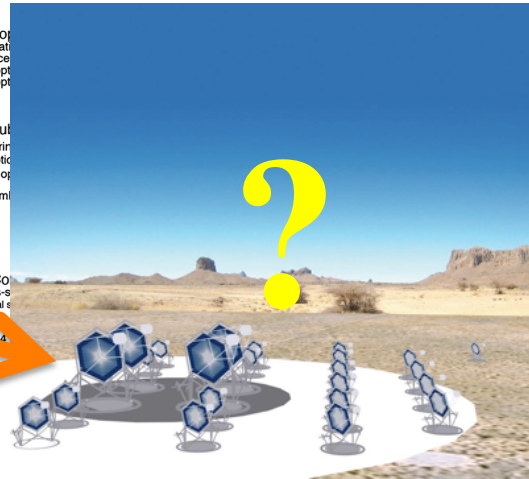
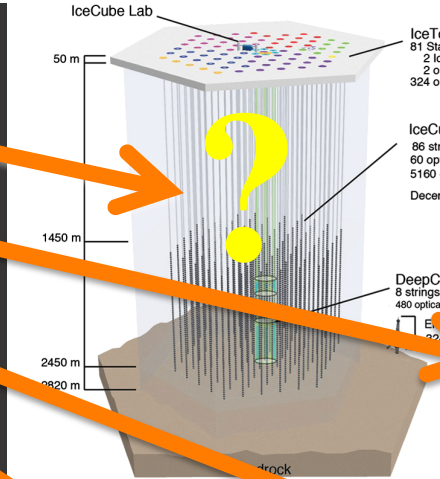
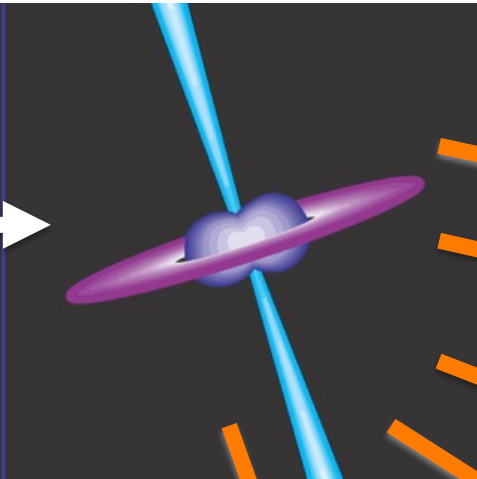
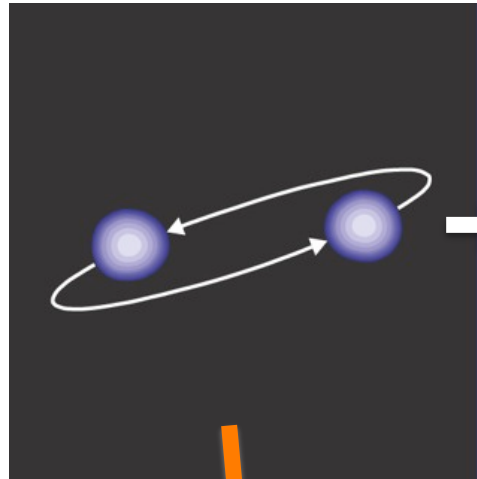
21st Century: Multi-Messenger Era

Counterparts to GW

Gravitational Wave Sources

Neutrino

Gamma-ray



Gravitational wave

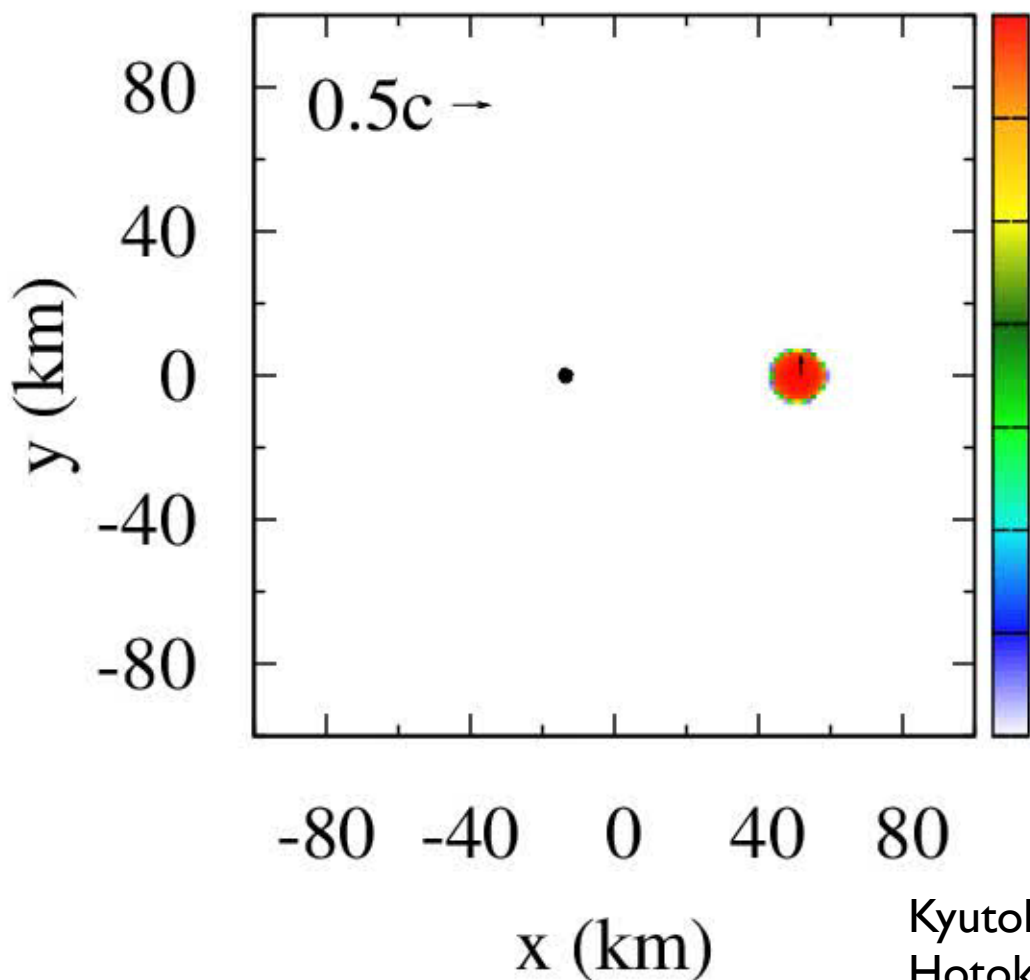
X-ray

IR-Opt

Radio

Full General Relativity

t=0 ms $\log \rho \text{ (g/cm}^3\text{)}$

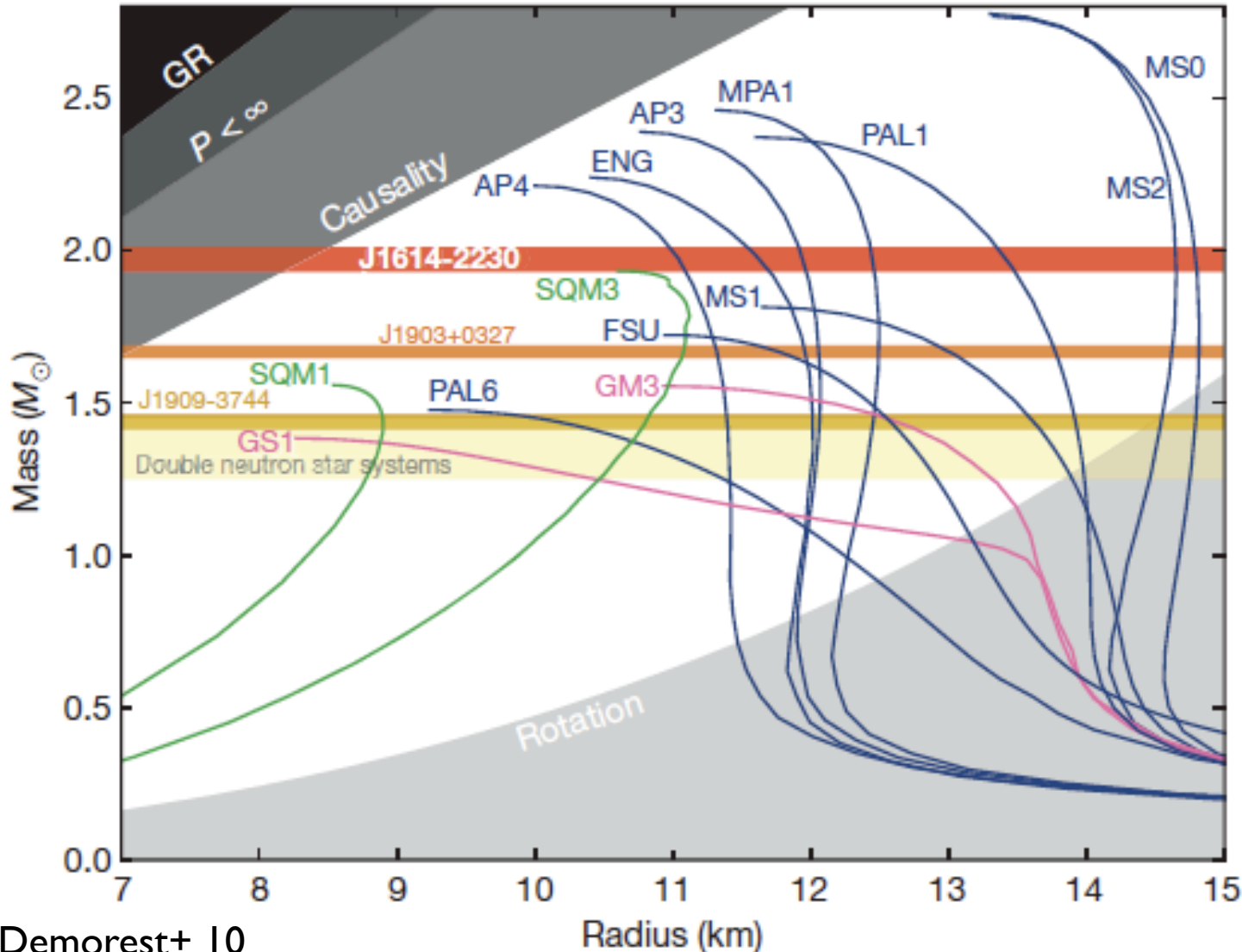


15 **Ejecta energy**
14 **~ Supernova**
13
12 **Ejecta velocity**
11 **~0.1-0.3 c**
10 **>> Supernova**
9
8 **⇒ TeV γ -ray?**

Takami, Kyutoku & KI 14

Kyutoku, KI, Okawa, Shibata & Taniguchi 15
Hotokezaka+ 13

Max Mass of Neutron Star



$2M_{\odot}$

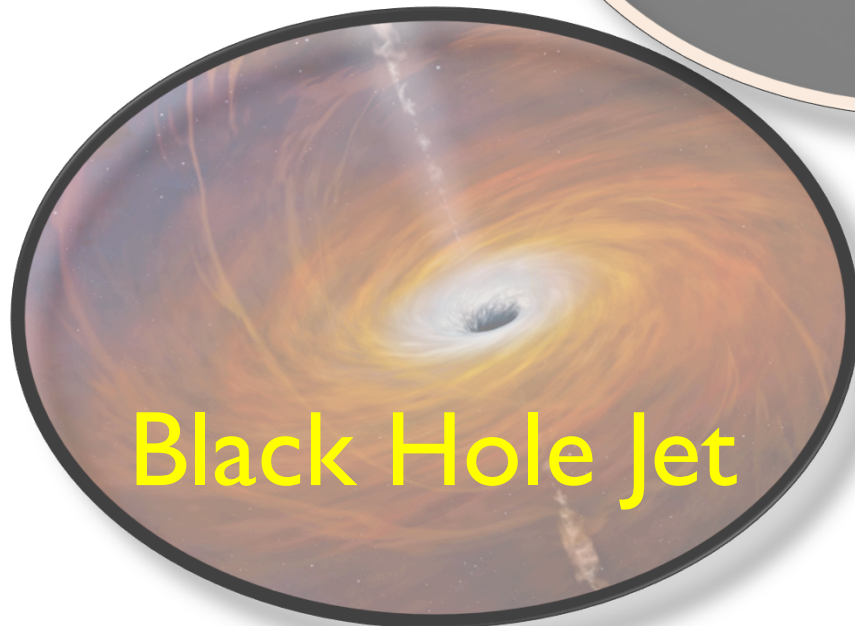
msec NS +
 $0.5M_{\odot}$ WD

Shapiro delay
 $\sim (1.97 \pm 0.04) M_{\odot}$

Stiff EOS

(No hyperon?
3-body?
vector force?)

Mysteries in 21th Century

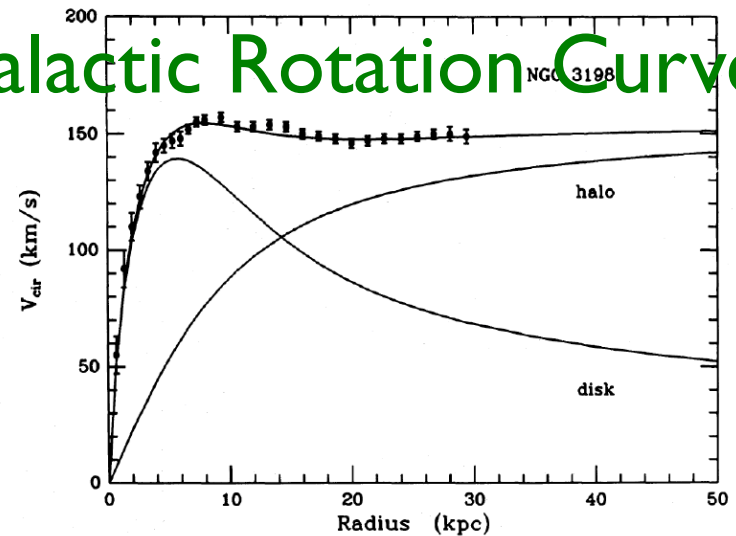


Evidence of Dark Matter

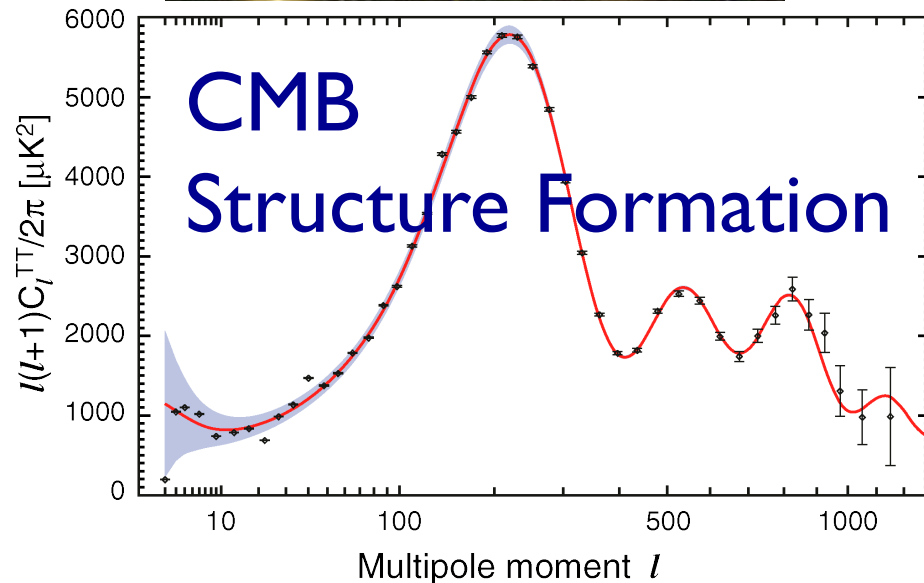
Gravitational Lensing



Galactic Rotation Curve



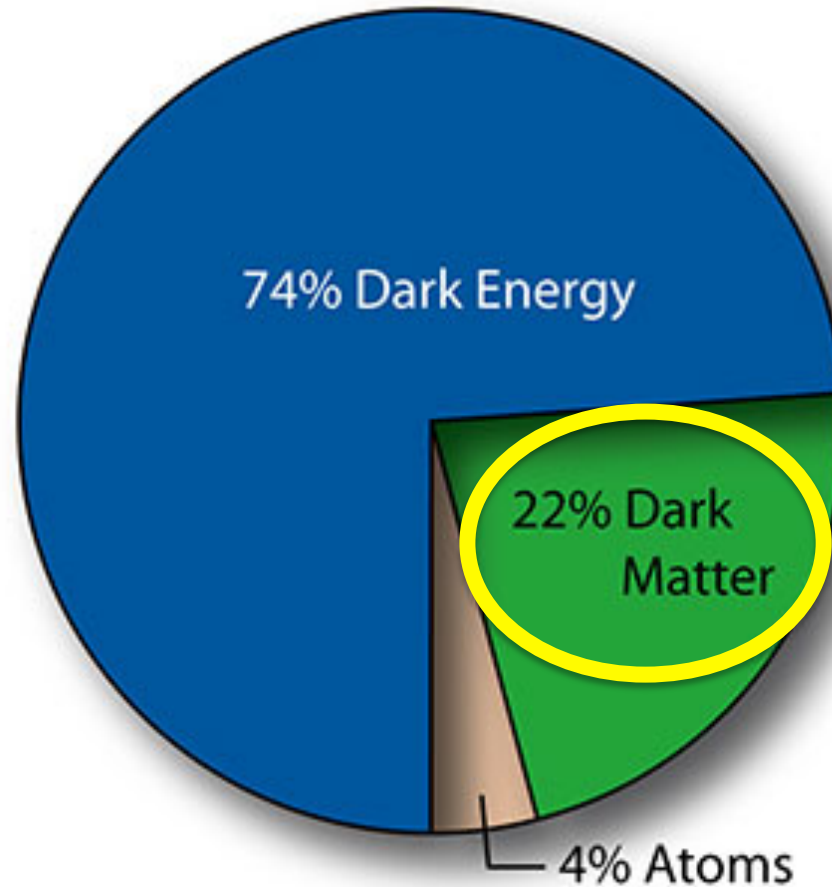
CMB
Structure Formation



Bullet Cluster

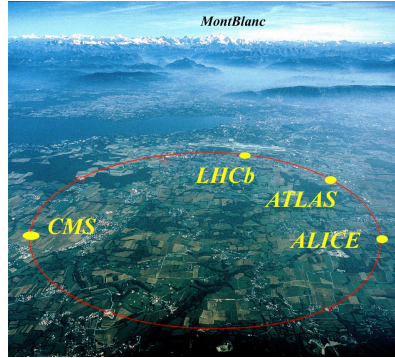


Content of the Universe



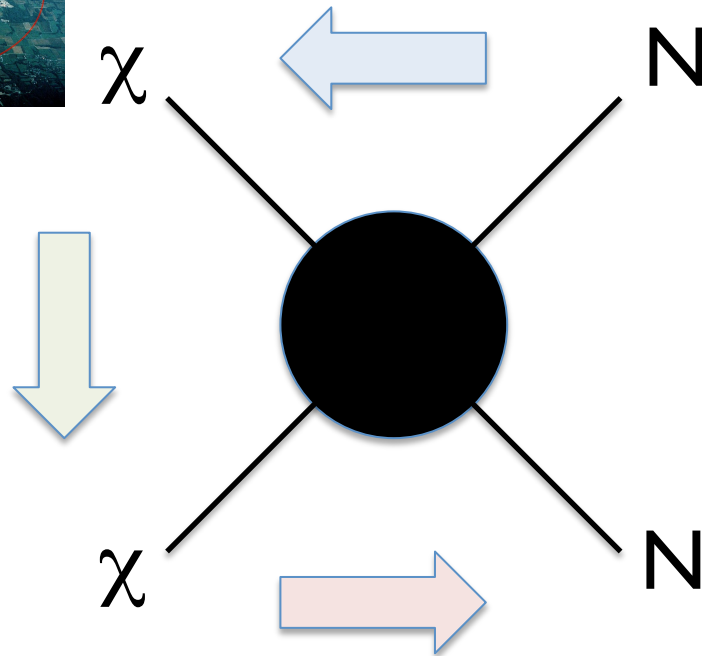
Know particles can NOT be Cold Dark Matter

Searches for Dark Matter

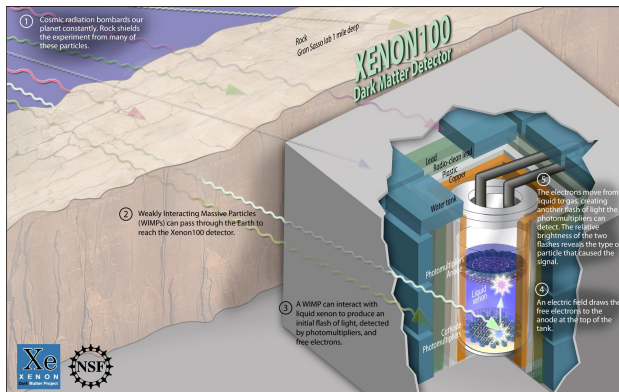


Accelerator

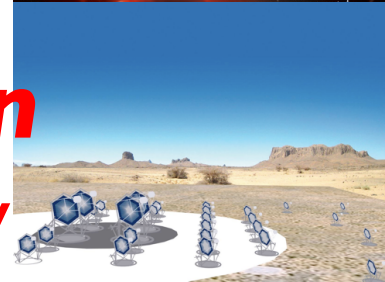
Missing p_T



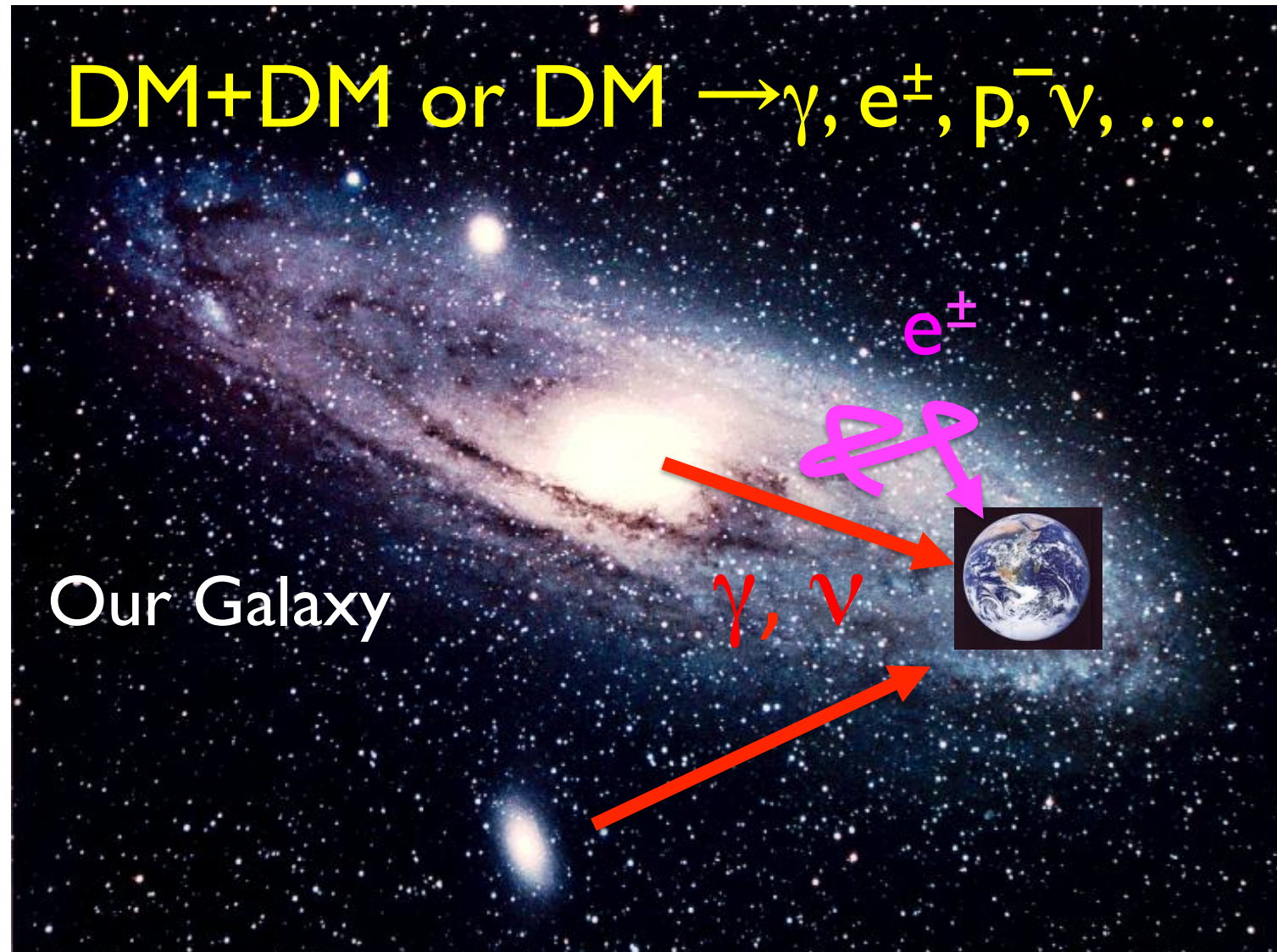
Direct Detection
DM scatters Nuclei



Indirect Detection
High-energy cosmic ray



Indirect Detection



ATLAS SUSY Searches* - 95% CL Lower Limits

Status: SUSY 2013

ATLAS Preliminary

$$\int \mathcal{L} dt = (4.6 - 22.9) \text{ fb}^{-1} \quad \sqrt{s} = 7, 8 \text{ TeV}$$

Model	e, μ, τ, γ Jets	E_T^{miss}	$\int \mathcal{L} dt [\text{fb}^{-1}]$	Mass limit	Reference			
Inclusive Searches	MSUGRA/CMSSM	0	2-6 jets	Yes	20.3	\tilde{q}, \tilde{g} 1.7 TeV	$m(\tilde{q})=m(\tilde{g})$	ATLAS-CONF-2013-047
	MSUGRA/CMSSM	1 e, μ	3-6 jets	Yes	20.3	\tilde{g} 1.2 TeV	any $m(\tilde{q})$	ATLAS-CONF-2013-062
	MSUGRA/CMSSM	0	7-10 jets	Yes	20.3	\tilde{g} 1.1 TeV	any $m(\tilde{q})$	1308.1841
	$\tilde{q}\tilde{q}, \tilde{q} \rightarrow q\tilde{l}_1^0$	0	2-6 jets	Yes	20.3	\tilde{q} 740 GeV	$m(\tilde{l}_1^0)=0 \text{ GeV}$	ATLAS-CONF-2013-047
	$\tilde{g}\tilde{g}, \tilde{g} \rightarrow q\tilde{q}\tilde{l}_1^0$	0	2-6 jets	Yes	20.3	\tilde{g} 1.3 TeV	$m(\tilde{l}_1^0)=0 \text{ GeV}$	ATLAS-CONF-2013-047
	$\tilde{g}\tilde{g}, \tilde{g} \rightarrow q\tilde{q}\tilde{l}_1^0 \rightarrow q\tilde{q}W^\pm\tilde{\chi}_1^0$	1 e, μ	3-6 jets	Yes	20.3	\tilde{g} 1.18 TeV	$m(\tilde{l}_1^0) < 200 \text{ GeV}, m(\tilde{\tau}^\pm)=0.5(m(\tilde{l}_1^0))+m(\tilde{g}))$	ATLAS-CONF-2013-062
	$\tilde{g}\tilde{g}, \tilde{g} \rightarrow q\tilde{q}(\ell\ell/\nu\nu/\nu\nu)\tilde{\chi}_1^0$	2 e, μ	0-3 jets	-	20.3	\tilde{g} 1.12 TeV	$m(\tilde{l}_1^0)=0 \text{ GeV}$	ATLAS-CONF-2013-089
	GMSB ($\tilde{\ell}$ NLSP)	2 e, μ	2-4 jets	Yes	4.7	\tilde{g} 1.24 TeV	$\tan\beta < 15$	1208.4688
	GMSB ($\tilde{\ell}$ NLSP)	1-2 τ	0-2 jets	Yes	20.7	\tilde{g} 1.4 TeV	$\tan\beta > 18$	ATLAS-CONF-2013-026
	GGM (bino NLSP)	2 γ	-	Yes	4.8	\tilde{g} 1.07 TeV	$m(\tilde{l}_1^0) > 50 \text{ GeV}$	1209.0753
	GGM (wino NLSP)	1 $e, \mu + \gamma$	-	Yes	4.8	\tilde{g} 619 GeV	$m(\tilde{l}_1^0) > 50 \text{ GeV}$	ATLAS-CONF-2012-144
	GGM (higgsino-bino NLSP)	γ	1 b	Yes	4.8	\tilde{g} 900 GeV	$m(\tilde{l}_1^0) > 220 \text{ GeV}$	1211.1167
GGM (higgsino NLSP)	2 $e, \mu (Z)$	0-3 jets	Yes	5.8	\tilde{g} 690 GeV	$m(\tilde{H}) > 200 \text{ GeV}$	ATLAS-CONF-2012-152	
Gravitino LSP	0	mono-jet	Yes	10.5	$F^{1/2}$ scale 645 GeV	$m(\tilde{g}) > 10^{-4} \text{ eV}$	ATLAS-CONF-2012-147	
3rd gen. \tilde{g} med.	$\tilde{g} \rightarrow b\tilde{b}_1^0$	0	3 b	Yes	20.1	\tilde{g} 1.2 TeV	$m(\tilde{l}_1^0) < 600 \text{ GeV}$	ATLAS-CONF-2013-061
	$\tilde{g} \rightarrow t\tilde{t}\tilde{\chi}_1^0$	0	7-10 jets	Yes	20.3	\tilde{g} 1.1 TeV	$m(\tilde{l}_1^0) < 350 \text{ GeV}$	1308.1841
	$\tilde{g} \rightarrow t\tilde{t}\tilde{\chi}_1^0$	0-1 e, μ	3 b	Yes	20.1	\tilde{g} 1.34 TeV	$m(\tilde{l}_1^0) < 400 \text{ GeV}$	ATLAS-CONF-2013-061
	$\tilde{g} \rightarrow b\tilde{t}\tilde{\chi}_1^0$	0-1 e, μ	3 b	Yes	20.1	\tilde{g} 1.3 TeV	$m(\tilde{l}_1^0) < 300 \text{ GeV}$	ATLAS-CONF-2013-061
	3rd gen. squarks direct production	$\tilde{b}_1\tilde{b}_1, \tilde{b}_1 \rightarrow b\tilde{l}_1^0$	0	2 b	Yes	20.1	\tilde{b}_1 100-620 GeV	$m(\tilde{l}_1^0) < 90 \text{ GeV}$
$\tilde{b}_1\tilde{b}_1, \tilde{b}_1 \rightarrow t\tilde{l}_1^0$		2 e, μ (SS)	0-3 b	Yes	20.7	\tilde{b}_1 275-430 GeV	$m(\tilde{l}_1^0) = 2 m(\tilde{l}_1^0)$	ATLAS-CONF-2013-007
$\tilde{t}_1\tilde{t}_1$ (light), $\tilde{t}_1 \rightarrow b\tilde{l}_1^0$		1-2 e, μ	1-2 b	Yes	4.7	\tilde{t}_1 110-167 GeV	$m(\tilde{l}_1^0) = 55 \text{ GeV}$	1208.4305, 1209.2102
$\tilde{t}_1\tilde{t}_1$ (light), $\tilde{t}_1 \rightarrow Wb\tilde{l}_1^0$		2 e, μ	0-2 jets	Yes	20.3	\tilde{t}_1 130-220 GeV	$m(\tilde{l}_1^0) = m(\tilde{t}_1) - m(W) - 50 \text{ GeV}, m(\tilde{t}_1) < m(\tilde{t}_1^+)$	ATLAS-CONF-2013-048
$\tilde{t}_1\tilde{t}_1$ (medium), $\tilde{t}_1 \rightarrow t\tilde{l}_1^0$		2 e, μ	2 jets	Yes	20.3	\tilde{t}_1 225-525 GeV	$m(\tilde{l}_1^0) = 0 \text{ GeV}$	ATLAS-CONF-2013-065
$\tilde{t}_1\tilde{t}_1$ (medium), $\tilde{t}_1 \rightarrow b\tilde{l}_1^0$		0	2 b	Yes	20.1	\tilde{t}_1 150-580 GeV	$m(\tilde{l}_1^0) < 200 \text{ GeV}, m(\tilde{t}_1^+) - m(\tilde{t}_1^0) = 5 \text{ GeV}$	1308.2631
$\tilde{t}_1\tilde{t}_1$ (heavy), $\tilde{t}_1 \rightarrow t\tilde{l}_1^0$		1 e, μ	1 b	Yes	20.7	\tilde{t}_1 200-610 GeV	$m(\tilde{l}_1^0) = 0 \text{ GeV}$	ATLAS-CONF-2013-037
$\tilde{t}_1\tilde{t}_1$ (heavy), $\tilde{t}_1 \rightarrow t\tilde{l}_1^0$		0	2 b	Yes	20.5	\tilde{t}_1 320-660 GeV	$m(\tilde{l}_1^0) = 0 \text{ GeV}$	ATLAS-CONF-2013-024
$\tilde{t}_1\tilde{t}_1, \tilde{t}_1 \rightarrow c\tilde{l}_1^0$		0	mono-jet/c-tag	Yes	20.3	\tilde{t}_1 90-200 GeV	$m(\tilde{l}_1^0) = 0 \text{ GeV}$	ATLAS-CONF-2013-068
$\tilde{t}_1\tilde{t}_1$ (natural GMSB)		2 $e, \mu (Z)$	1 b	Yes	20.7	\tilde{t}_1 500 GeV	$m(\tilde{l}_1^0) - m(\tilde{l}_1^0) < 85 \text{ GeV}$	ATLAS-CONF-2013-025
$\tilde{t}_2\tilde{t}_2, \tilde{t}_2 \rightarrow \tilde{t}_1 + Z$		3 $e, \mu (Z)$	1 b	Yes	20.7	\tilde{t}_2 271-520 GeV	$m(\tilde{t}_1) = m(\tilde{t}_1^+) + 180 \text{ GeV}$	ATLAS-CONF-2013-025
EW direct		$\tilde{t}_1\tilde{t}_1, \tilde{t}_1 \rightarrow \tilde{t}_1^0\tilde{\chi}_1^0$	2 e, μ	0	Yes	20.3	\tilde{t}_1 85-315 GeV	$m(\tilde{l}_1^0) = 0 \text{ GeV}$
	$\tilde{t}_1^+\tilde{t}_1^-, \tilde{t}_1^+ \rightarrow \tilde{t}_1^0\tilde{\chi}_1^0$	2 e, μ	0	Yes	20.3	\tilde{t}_1^* 125-450 GeV	$m(\tilde{l}_1^0) = 0 \text{ GeV}, m(\tilde{\tau}, \tilde{\nu}) = 0.5(m(\tilde{t}_1^+) + m(\tilde{l}_1^0))$	ATLAS-CONF-2013-049
	$\tilde{t}_1^+\tilde{t}_1^-, \tilde{t}_1^+ \rightarrow \tilde{t}_1^0\tilde{\chi}_1^0$	2 τ	-	Yes	20.7	\tilde{t}_1^* 180-330 GeV	$m(\tilde{l}_1^0) = 0 \text{ GeV}, m(\tilde{\tau}, \tilde{\nu}) = 0.5(m(\tilde{t}_1^+) + m(\tilde{l}_1^0))$	ATLAS-CONF-2013-028
	$\tilde{t}_1^+\tilde{t}_1^-, \tilde{t}_1^+ \rightarrow \tilde{t}_1^0\tilde{\chi}_1^0$	3 e, μ	0	Yes	20.7	\tilde{t}_1^* 600 GeV	$m(\tilde{l}_1^0) = m(\tilde{t}_1^0), m(\tilde{l}_1^0) = 0, m(\tilde{\tau}, \tilde{\nu}) = 0.5(m(\tilde{t}_1^+) + m(\tilde{l}_1^0))$	ATLAS-CONF-2013-035
	$\tilde{t}_1^+\tilde{t}_1^0 \rightarrow W\tilde{t}_1^+\tilde{\chi}_1^0$	3 e, μ	0	Yes	20.7	\tilde{t}_1^* 315 GeV	$m(\tilde{l}_1^0) = m(\tilde{t}_1^0), m(\tilde{l}_1^0) = 0, m(\tilde{\tau}, \tilde{\nu}) = 0.5(m(\tilde{t}_1^+) + m(\tilde{l}_1^0))$	ATLAS-CONF-2013-035
	$\tilde{t}_1^+\tilde{t}_1^0 \rightarrow W\tilde{t}_1^+\tilde{\chi}_1^0$	1 e, μ	2 b	Yes	20.3	\tilde{t}_1^* 285 GeV	$m(\tilde{l}_1^0) = m(\tilde{t}_1^0), m(\tilde{l}_1^0) = 0, m(\tilde{\tau}, \tilde{\nu}) = 0, m(\tilde{l}_1^0) = 0, m(\tilde{l}_1^0) = 0$	ATLAS-CONF-2013-093
	Long-lived particles	Direct $\tilde{t}_1^+\tilde{t}_1^0$ prod., long-lived \tilde{t}_1^+	Disapp. trk	1 jet	Yes	20.3	\tilde{t}_1^+ 270 GeV	$m(\tilde{t}_1^+) - m(\tilde{t}_1^0) = 160 \text{ MeV}, \tau(\tilde{t}_1^+) = 0.2 \text{ ns}$
Stable, stopped \tilde{g} R-hadron		0	1-5 jets	Yes	22.9	\tilde{g} 832 GeV	$m(\tilde{t}_1^+) = 100 \text{ GeV}, 10 \mu\text{s} < \tau(\tilde{g}) < 1000 \text{ s}$	ATLAS-CONF-2013-057
GMSB, stable $\tilde{\tau}, \tilde{\chi}_1^0 \rightarrow \tilde{\tau}(\tilde{g}, \tilde{\mu}) + \tau(e, \mu)$		1-2 μ	-	-	15.9	$\tilde{\tau}^0$ 475 GeV	$10 < \tan\beta < 50$	ATLAS-CONF-2013-058
GMSB, $\tilde{\chi}_1^0 \rightarrow \tilde{\chi}_1^0\tilde{G}$, long-lived $\tilde{\chi}_1^0$		2 γ	-	Yes	4.7	$\tilde{\chi}_1^0$ 230 GeV	$0.4 < \tau(\tilde{\chi}_1^0) < 2 \text{ ns}$	1304.6310
$\tilde{q}\tilde{q}, \tilde{\chi}_1^0 \rightarrow q\tilde{q}$ (RPV)		1 μ , displ. vtx	-	-	20.3	\tilde{q} 1.0 TeV	$1.5 < c\tau < 156 \text{ mm}, \text{BR}(\mu) = 1, m(\tilde{l}_1^0) = 108 \text{ GeV}$	ATLAS-CONF-2013-092
RPV	LFV $pp \rightarrow \tilde{\nu}_\tau + X, \tilde{\nu}_\tau \rightarrow e + \mu$	2 e, μ	-	-	4.6	$\tilde{\nu}_\tau$ 1.61 TeV	$\lambda_{113}^c = 0.10, \lambda_{132} = 0.05$	1212.1272
	LFV $pp \rightarrow \tilde{\nu}_\tau + X, \tilde{\nu}_\tau \rightarrow e(\mu) + \tau$	1 $e, \mu + \tau$	-	-	4.6	$\tilde{\nu}_\tau$ 1.1 TeV	$\lambda_{113}^c = 0.10, \lambda_{1(2)33} = 0.05$	1212.1272
	Bilinear RPV CMSSM	1 e, μ	7 jets	Yes	4.7	\tilde{q}, \tilde{g} 1.2 TeV	$m(\tilde{q}) = m(\tilde{g}), c\tau_{LSP} < 1 \text{ mm}$	ATLAS-CONF-2012-140
	$\tilde{\chi}_1^+\tilde{\chi}_1^-, \tilde{\chi}_1^+ \rightarrow W\tilde{\chi}_1^0\tilde{\chi}_1^0 \rightarrow ee\tilde{\nu}_\mu, e\mu\tilde{\nu}_e$	4 e, μ	-	Yes	20.7	$\tilde{\chi}_1^*$ 760 GeV	$m(\tilde{l}_1^0) > 300 \text{ GeV}, \lambda_{121} > 0$	ATLAS-CONF-2013-036
	$\tilde{\chi}_1^+\tilde{\chi}_1^-, \tilde{\chi}_1^+ \rightarrow W\tilde{\chi}_1^0\tilde{\chi}_1^0 \rightarrow \tau\tilde{\nu}_e, e\tau\tilde{\nu}_\tau$	3 $e, \mu + \tau$	-	Yes	20.7	$\tilde{\chi}_1^*$ 350 GeV	$m(\tilde{l}_1^0) > 80 \text{ GeV}, \lambda_{113} > 0$	ATLAS-CONF-2013-036
	$\tilde{g} \rightarrow q\tilde{q}$	0	6-7 jets	-	20.3	\tilde{g} 916 GeV	$\text{BR}(t) - \text{BR}(b) - \text{BR}(c) = 0\%$	ATLAS-CONF-2013-091
$\tilde{g} \rightarrow \tilde{t}_1 t, \tilde{t}_1 \rightarrow b\tilde{s}$	2 e, μ (SS)	0-3 b	Yes	20.7	\tilde{g} 880 GeV		ATLAS-CONF-2013-007	
Other	Scalar gluon pair, $\text{sgluon} \rightarrow q\tilde{q}$	0	4 jets	-	4.6	sgluon 100-287 GeV	incl. limit from 1110.2693	1210.4826
	Scalar gluon pair, $\text{sgluon} \rightarrow t\tilde{t}$	2 e, μ (SS)	1 b	Yes	14.3	sgluon 800 GeV		ATLAS-CONF-2013-051
	WIMP interaction (D5, Dirac χ)	0	mono-jet	Yes	10.5	M^* scale 704 GeV	$m(\chi) < 80 \text{ GeV}$, limit of $< 687 \text{ GeV}$ for D8	ATLAS-CONF-2012-147

$\sqrt{s} = 7 \text{ TeV}$ full data
 $\sqrt{s} = 8 \text{ TeV}$ partial data
 $\sqrt{s} = 8 \text{ TeV}$ full data

10⁻¹ 1 Mass scale [TeV]

*Only a selection of the available mass limits on new states or phenomena is shown. All limits quoted are observed minus 1 σ theoretical signal cross section uncertainty.

γ -ray Limits on DM

GeV

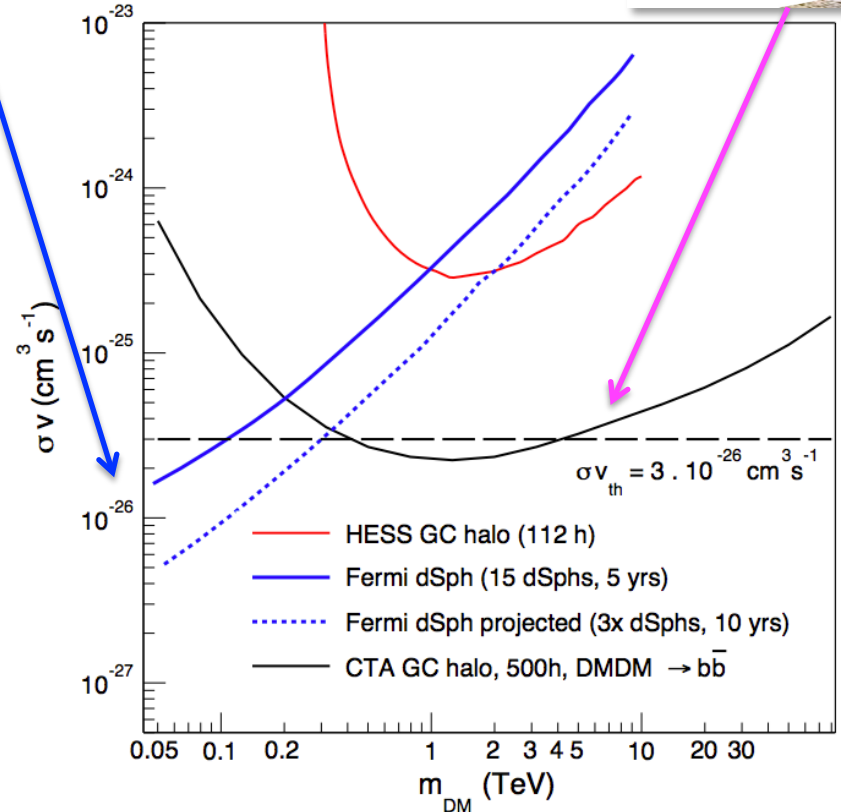
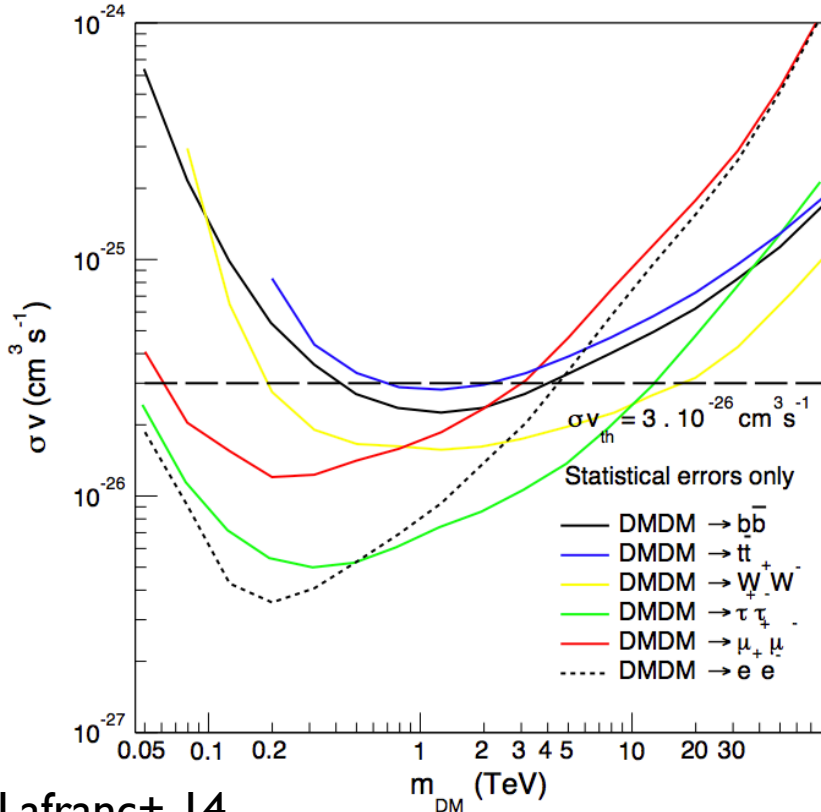


TeV



Five Rols, Einasto profile, 500 h

Weak scale, Higgs

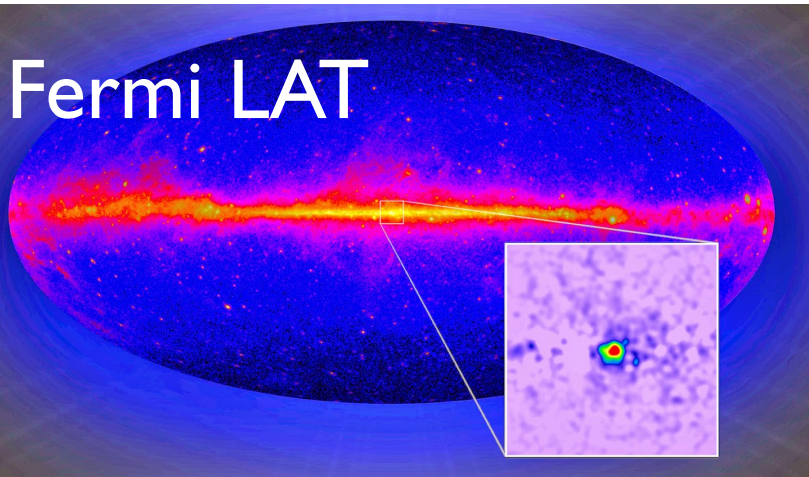


Lafranc+ 14

Complimentary to LHC & Direct detections

Galactic Center GeV Excess

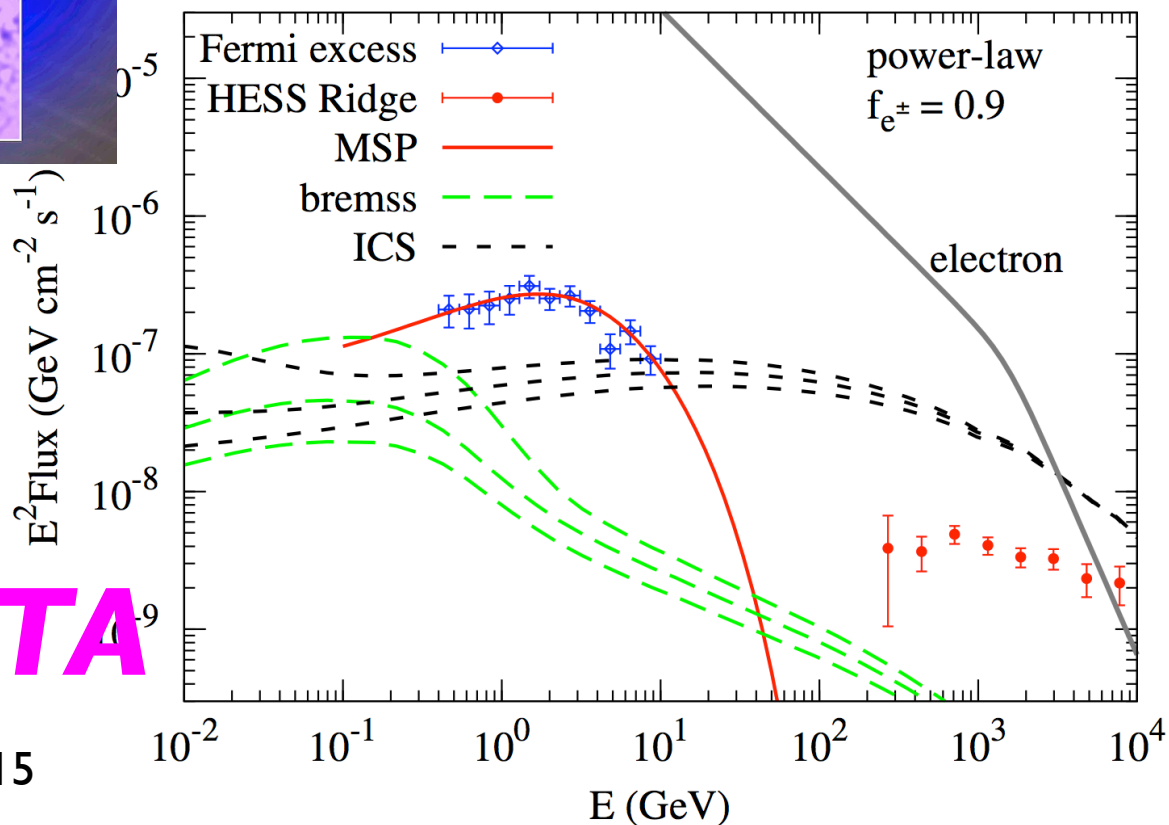
Fermi LAT



Circular, Extended $\propto \theta^{-2.4}$
Bump-like spectrum

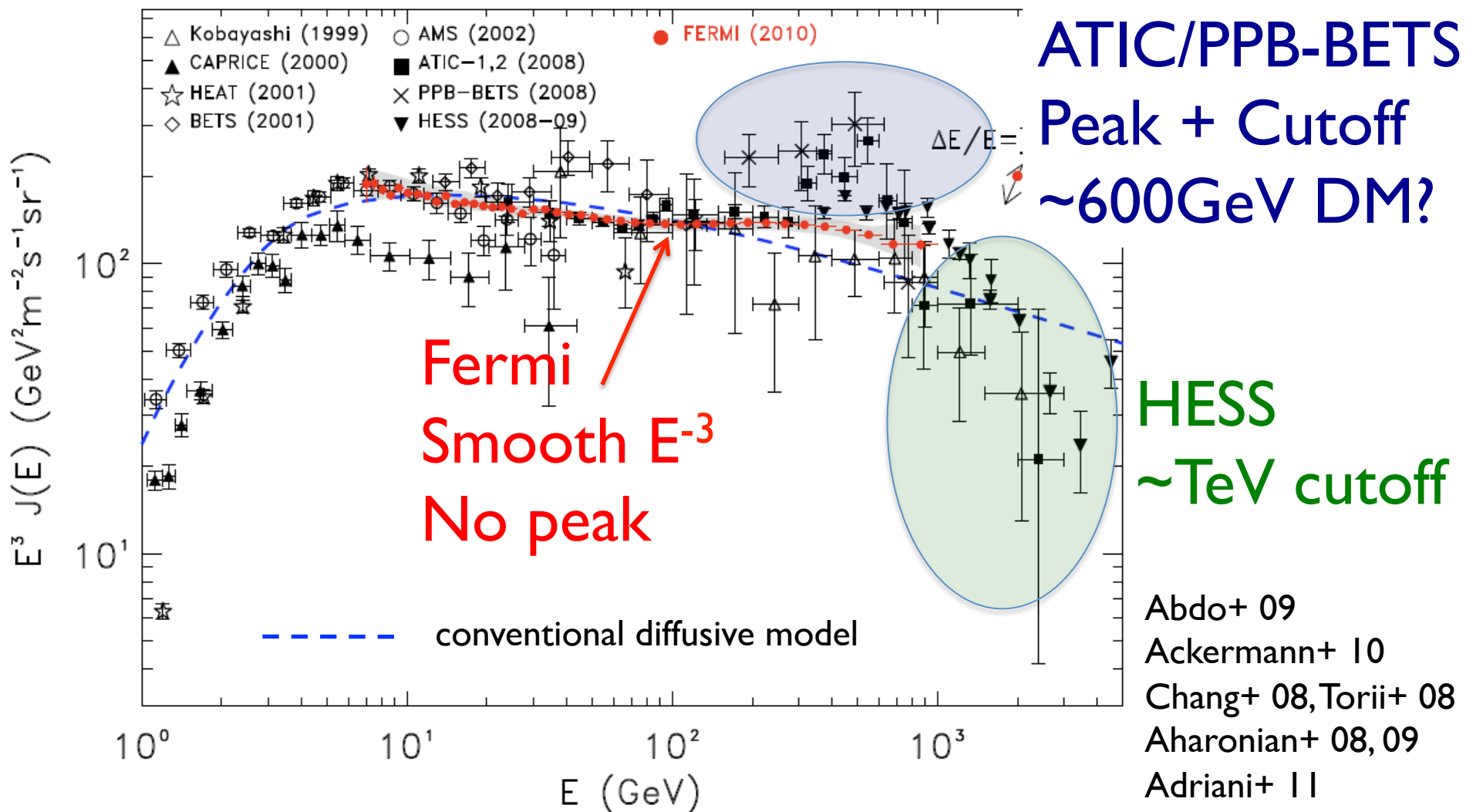
GeV dark matter?
(so-called Hooperon)
msec Pulsars?

\Rightarrow **TeV by CTA**



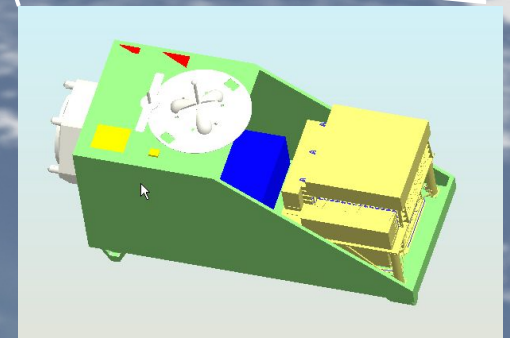
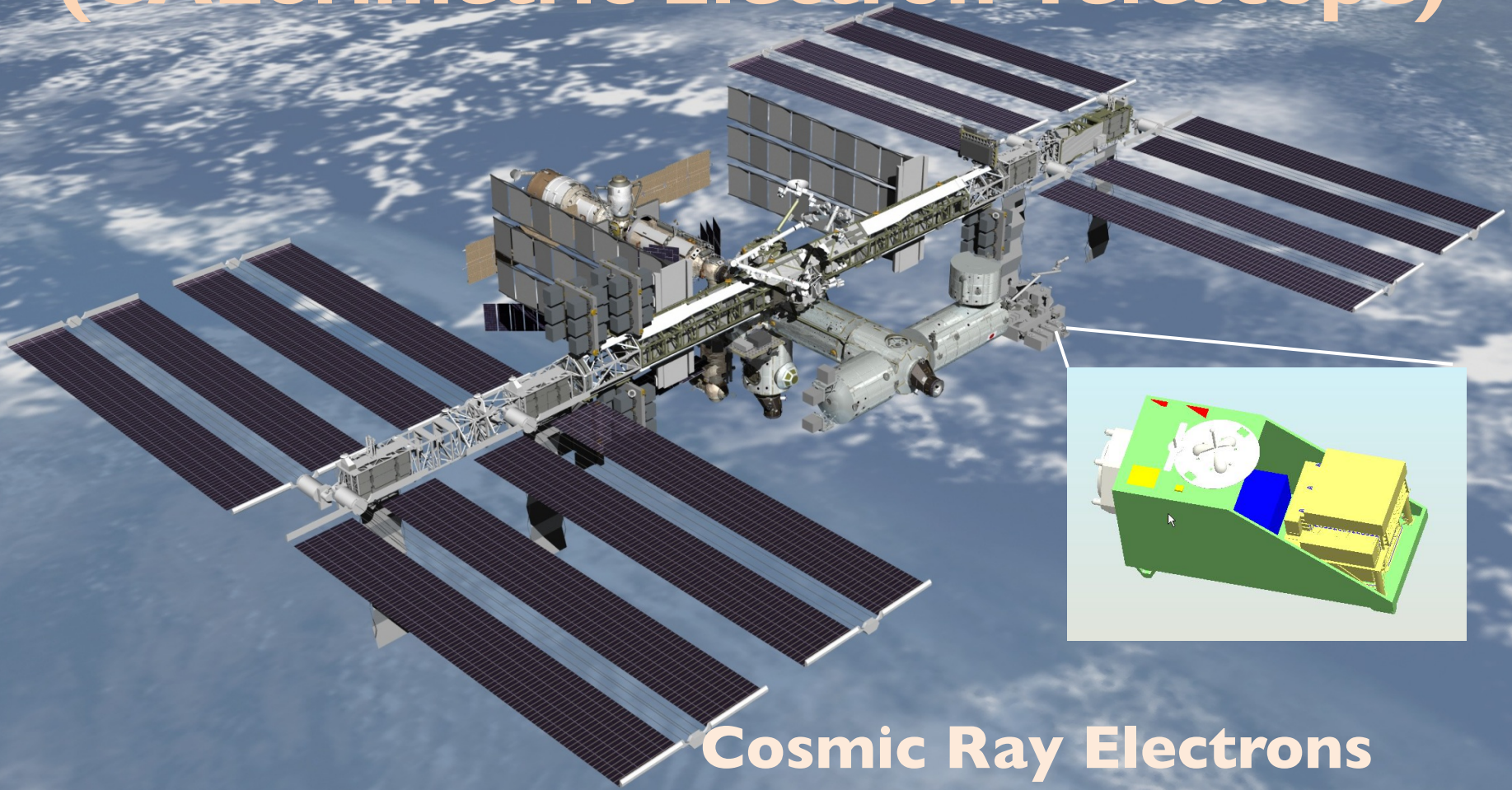
Cosmic-Ray Electron

An Excess also in $(e^+ + e^-)$ Spectrum



CALET

(CALorimetric Electron Telescope)

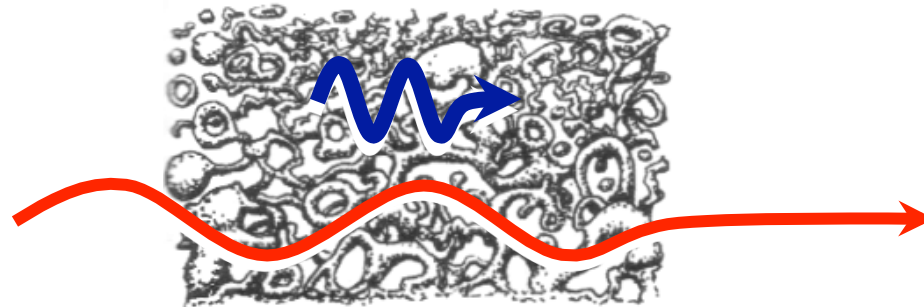


**Cosmic Ray Electrons
up to $\sim 10\text{TeV}$
w/ $\Delta E \sim$ a few % ($> 100\text{GeV}$)**

Lorentz Invariance Violation

- Some Quantum Gravity Theories violate LI

⇒ Energy dependent dispersion & speed of light



$$s_n = \pm 1$$

$$\xi \sim O(1)$$

$$\frac{p_\gamma^2 c^2}{E_\gamma^2} - 1 = s_n \left(\frac{E_\gamma}{\xi M_{\text{Planck}} c^2} \right)^n \quad M_{\text{Planck}} = \left(\frac{\hbar c}{G} \right)^{1/2} \approx 1.22 \times 10^{19} \text{ GeV}/c^2$$

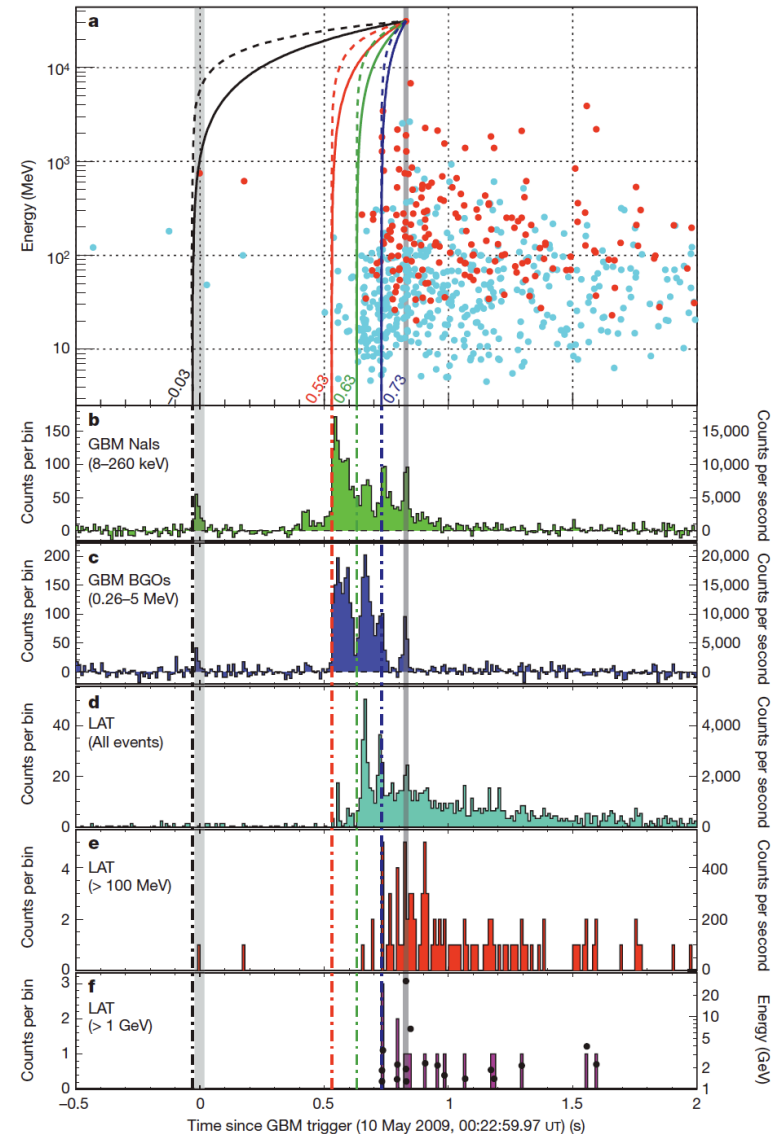
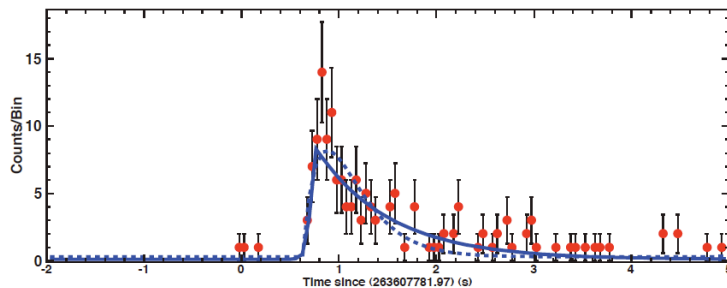
$$v_\gamma = \frac{\partial E_\gamma}{\partial p_\gamma} \approx c \left[1 - s_n \frac{n+1}{2} \left(\frac{E_\gamma}{\xi M_{\text{Planck}} c^2} \right)^n \right]$$

LIV with γ -ray

GRB090510 ($z=0.9$, $E_h=28\text{GeV}$)

$$\xi = M_{\text{QG},1} / M_{\text{Planck}} > 1.2$$

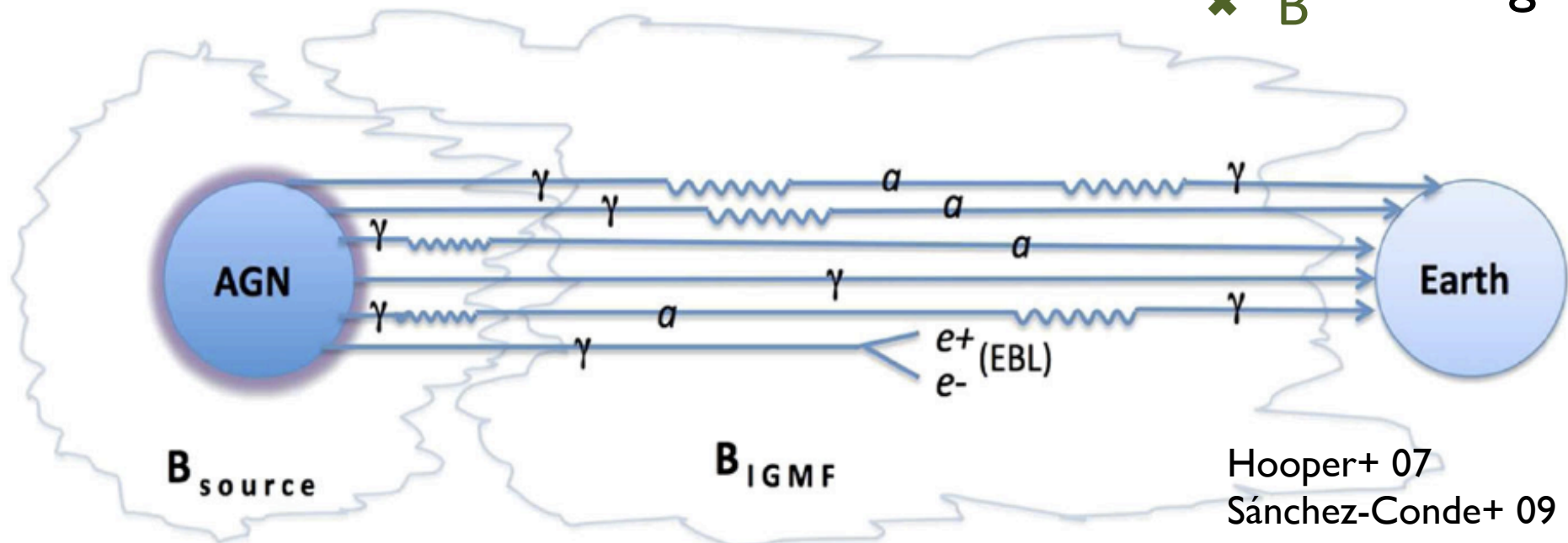
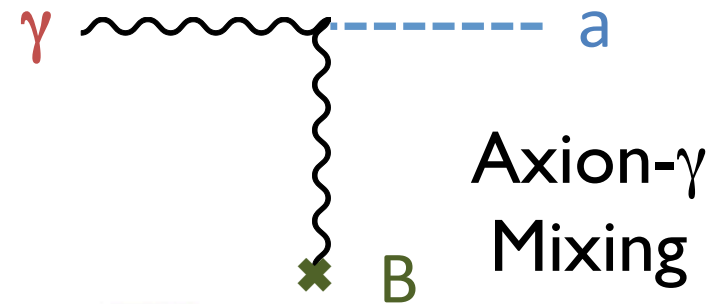
lower limit on $M_{\text{QG},1}/M_{\text{Planck}}$	limit on $M_{\text{QG},2}$ in $10^{10} \text{ GeV}/c^2$
> 1.19	> 2.99
> 3.42	> 5.06
> 5.12	> 6.20
> 10.0	> 8.79
> 102	> 27.7
> 1.33	> 0.54
> 1.22	—



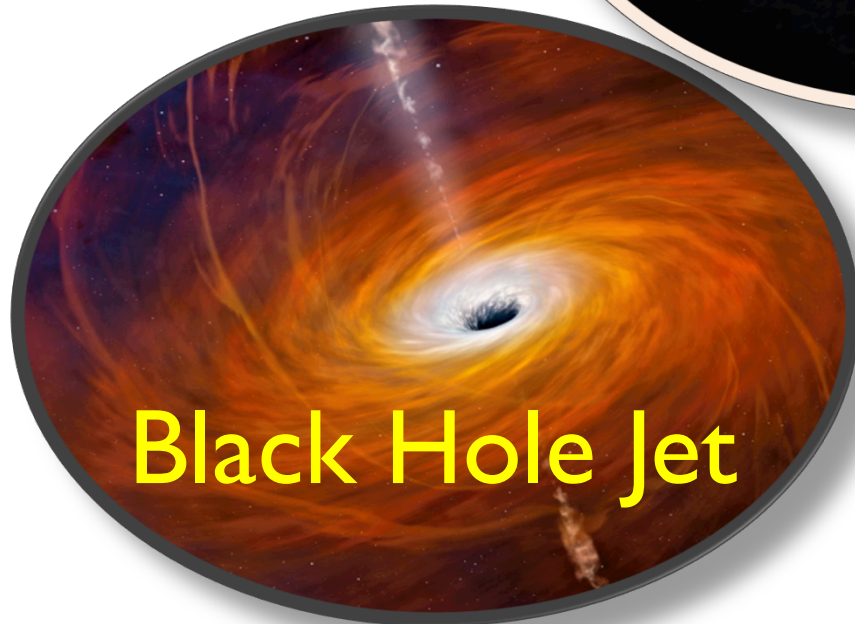
Axion-Like Particle

- $$a\gamma\gamma \quad \mathcal{L}_{a\gamma\gamma} = -\frac{1}{4}g_{a\gamma\gamma}F_{\mu\nu}\tilde{F}^{\mu\nu}a = g_{a\gamma\gamma}\mathbf{E} \cdot \mathbf{B}a,$$

$\gamma(\text{TeV } \gamma) + \gamma(\text{Background})$
 $\rightarrow e^+e^-$ is avoided



Mysteries in 21th Century



Mysteries in 21th Century



Thank

You